

THE CHALLENGE OF GALAXY FORMATION

**Joe Silk, Oxford
August, 2010**

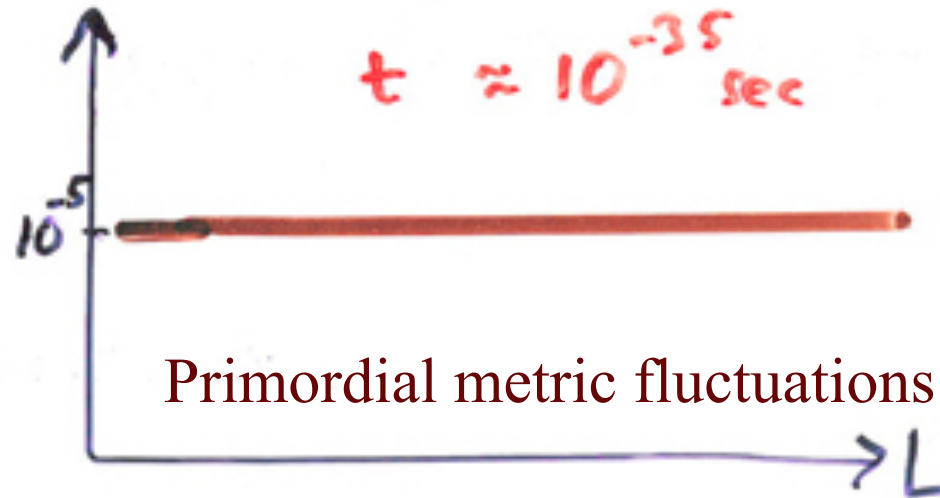
1. Structure formation
2. Disk galaxy formation
3. Spheroidal galaxy formation
4. Star formation

1. Structure formation

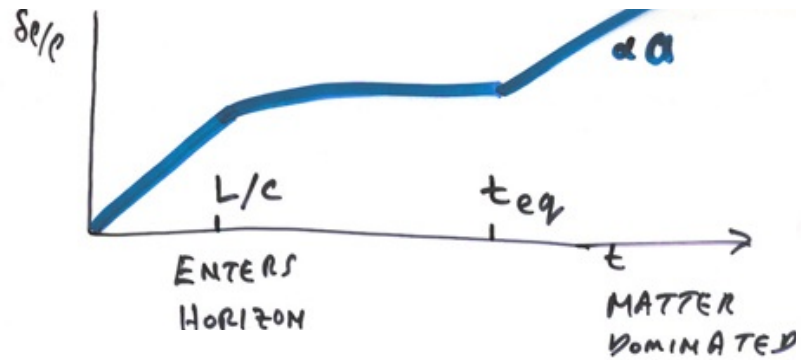
INFLATION



δK



Primordial metric fluctuations

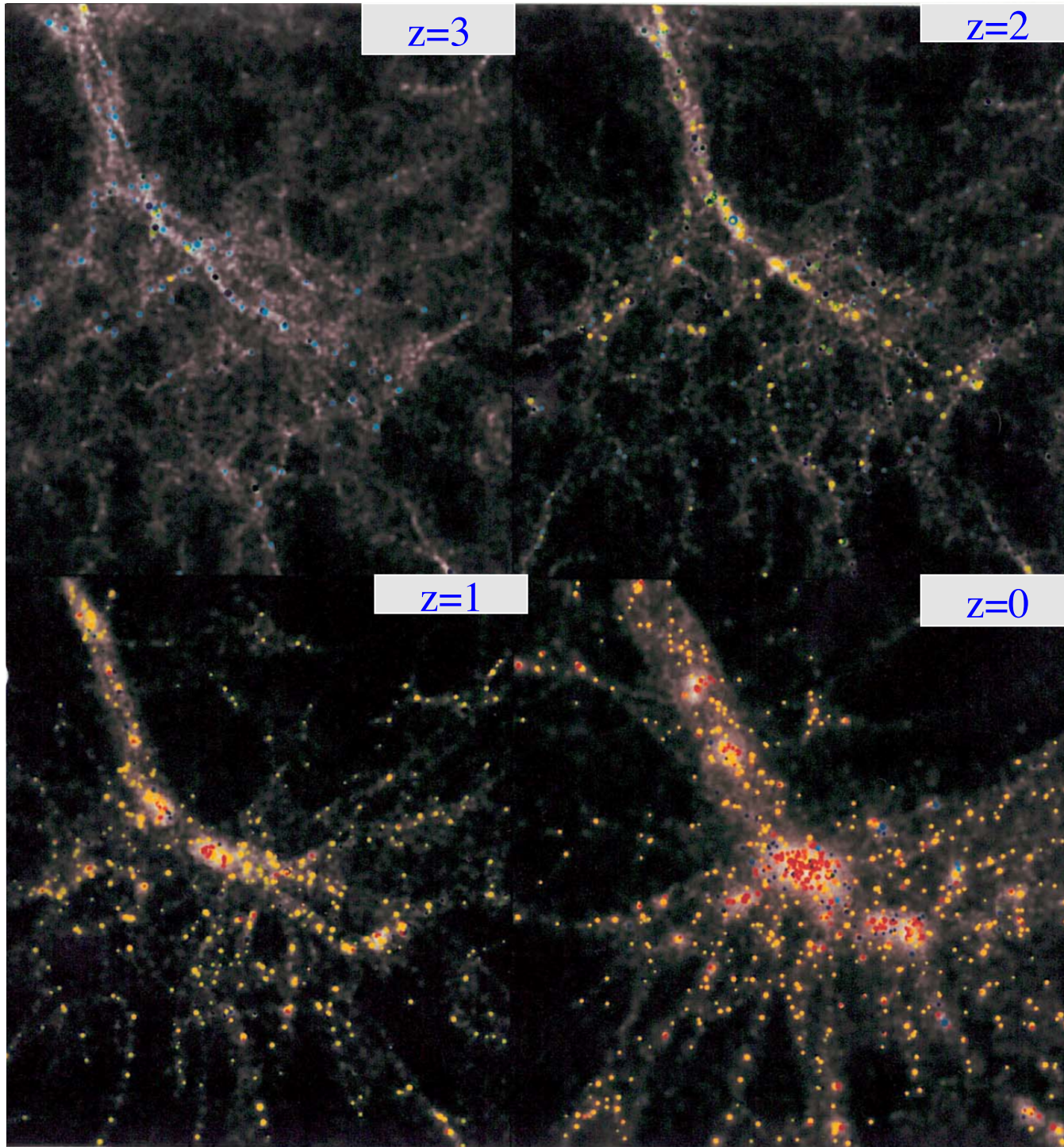


Bottom-up growth



$t \approx 10^{10}$ yr

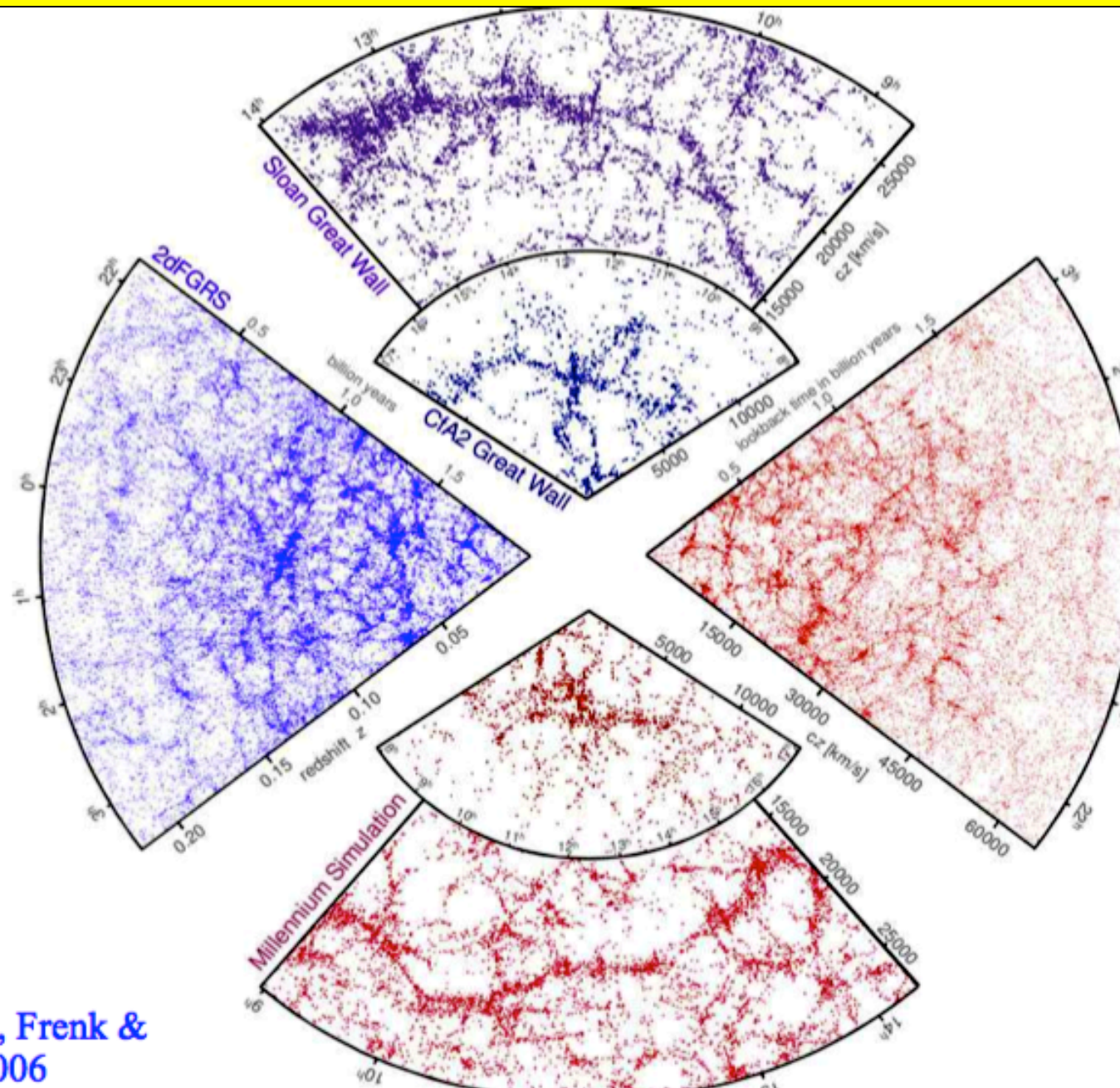




Galaxies form early in rare peaks

Virtual universe

Observed universe



Springel, Frenk &
White 2006



2. Disk galaxy formation

Some analytic estimates

Gas surface density

Star surface density

$$\Sigma_{\text{SFR}} = (\text{SFE}) \frac{\Sigma_{\text{gas}}}{t_{\text{dyn}}}$$

Star formation efficiency

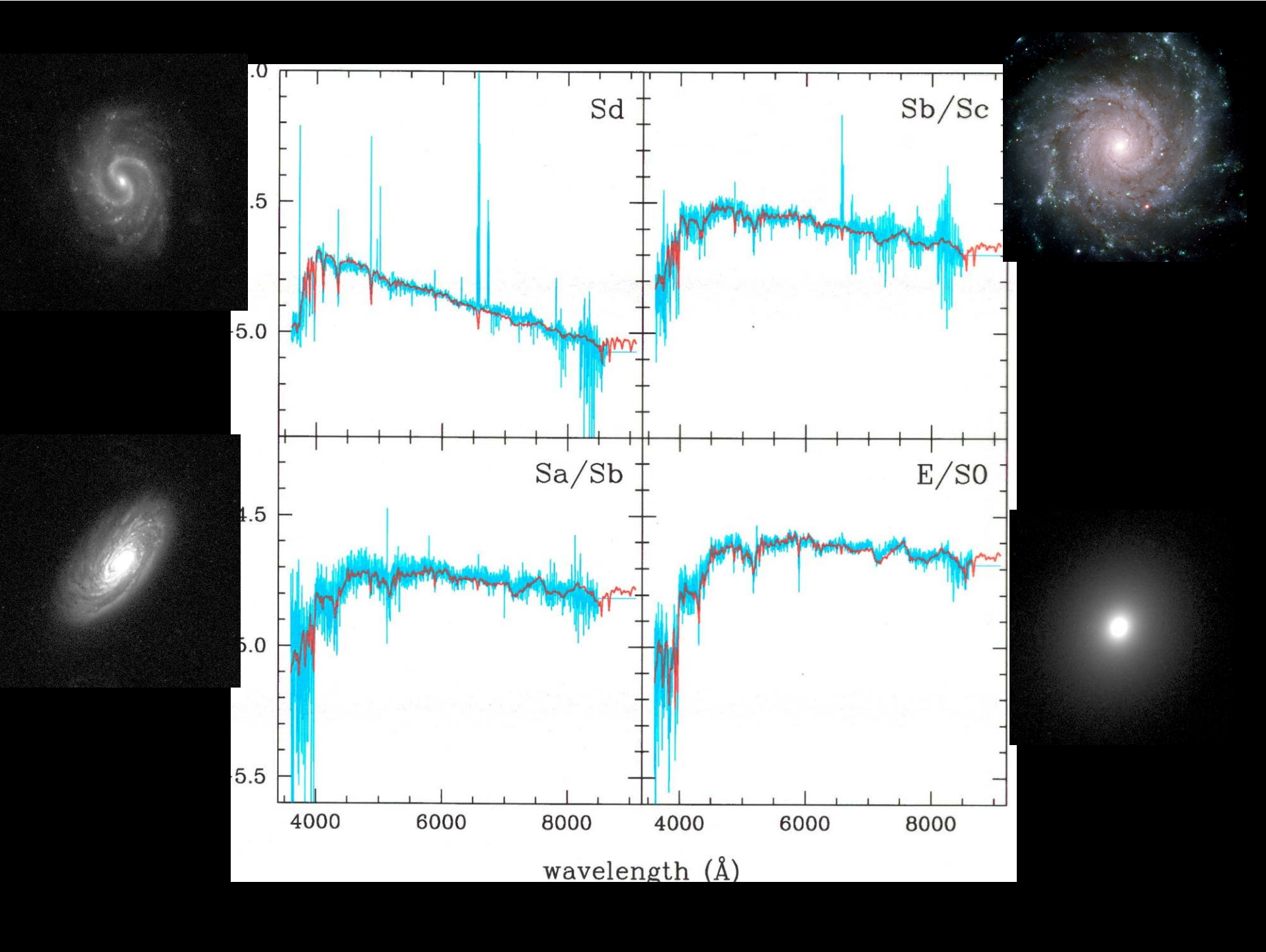
$$\text{SFE} = \frac{\sigma_{\text{gas}} V_{\text{cool}} m_{*,\text{SN}}}{E_{\text{SN}}^{\text{initial}}}$$

$$\approx 0.02$$

$$\propto \Sigma_{\text{gas}} \Omega$$
$$\approx \Sigma_{\text{gas}}^{3/2}$$

Infrared (24 microns)

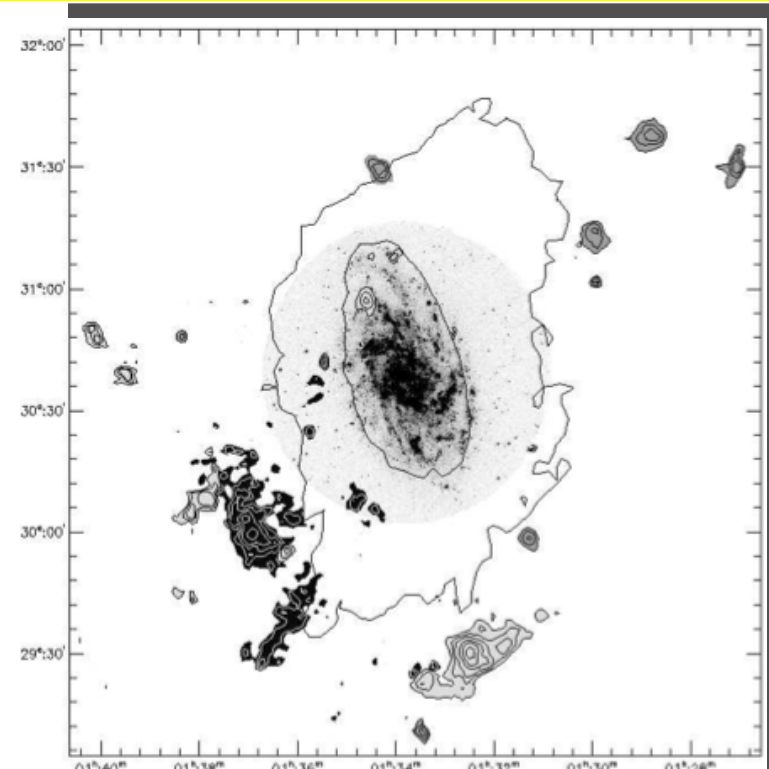
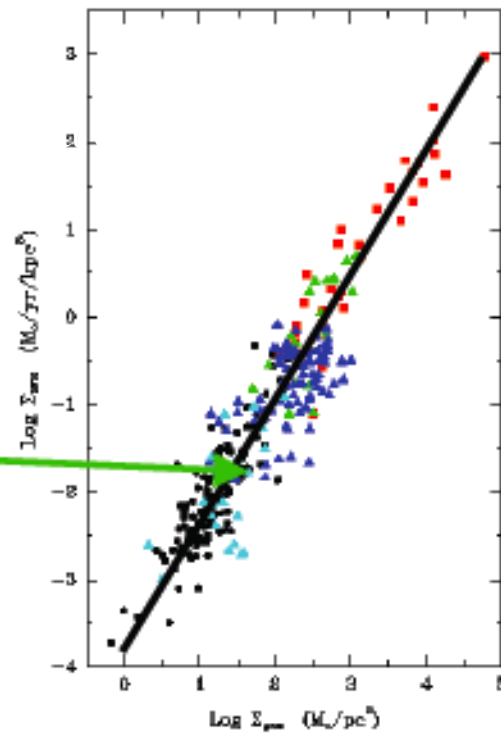




A GLOBAL STAR FORMATION LAW



K₀

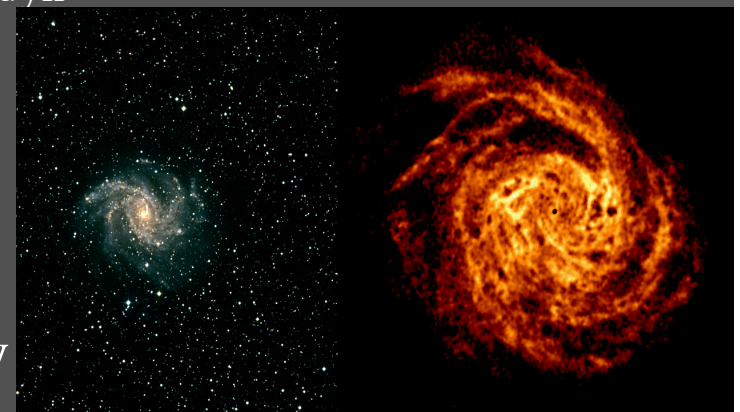


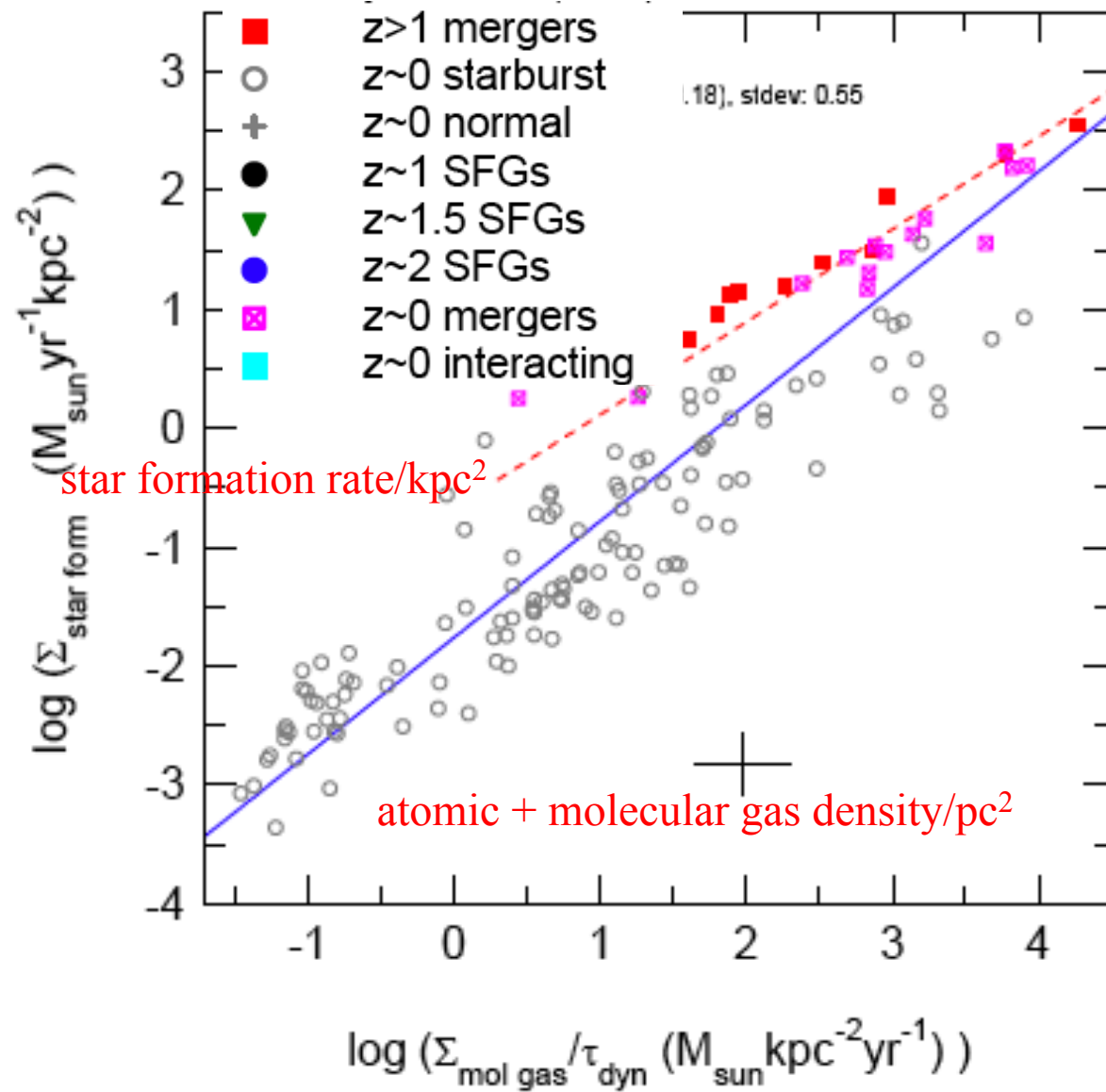
$$\text{SFR} = 0.02 (\text{GAS SURFACE DENSITY}) / t_{\text{dyn}}$$

fits disk galaxies (& M51 complexes)

low efficiency due to SN feedback
+ cold gas accretion/global disk instability

NGC 6946





Star Formation efficiency

=SFR/GAS MASS x ROTATION TIME

=2.5%

Genzel et al 2010

Molecular 'Elmegreen-Silk'-relation for mergers and SMGs. Symbols

Press-Schechter theory
(refined by Sheth & Tormen...)

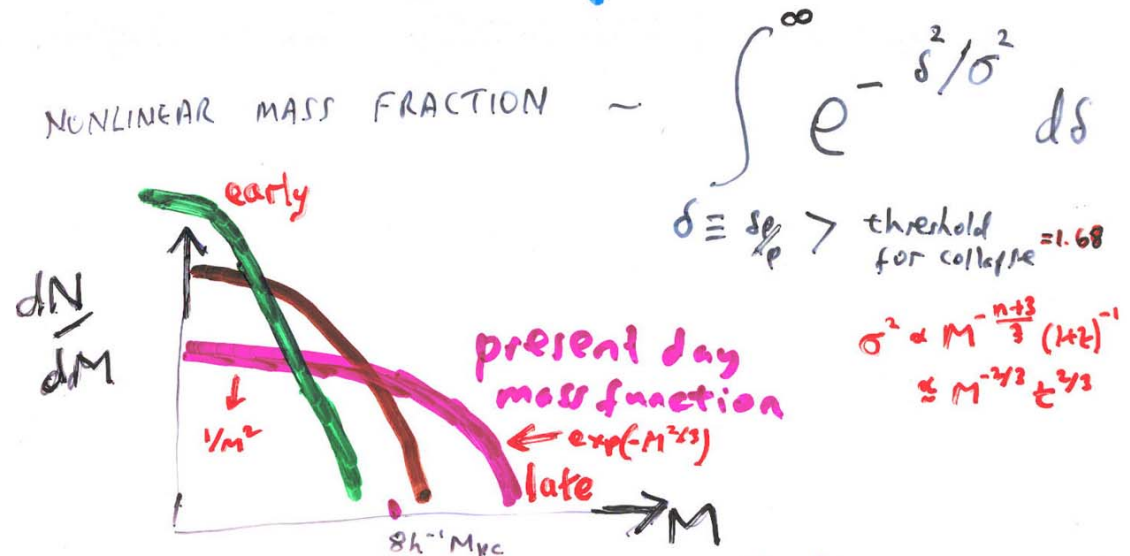
$$\left\langle \left(\frac{\delta\rho}{\rho} \right)^2 \right\rangle_{\text{mass}}^{y_2} \equiv \sigma(M) \propto t^{2/3} \quad (\text{if } \Omega=1)$$

$\left\langle \left(\frac{\delta\rho}{\rho} \right)^2 \right\rangle_{\text{light}}^{y_2}$ at $z=0$ ~ 1 averaged over $8h^{-1}\text{Mpc}$ spheres

growth suppressed at $1+z < \left(\frac{\Omega_m}{\Omega} \right)^{1/2}$ if $\Omega_m < 1 \Rightarrow$ need larger σ at specified M

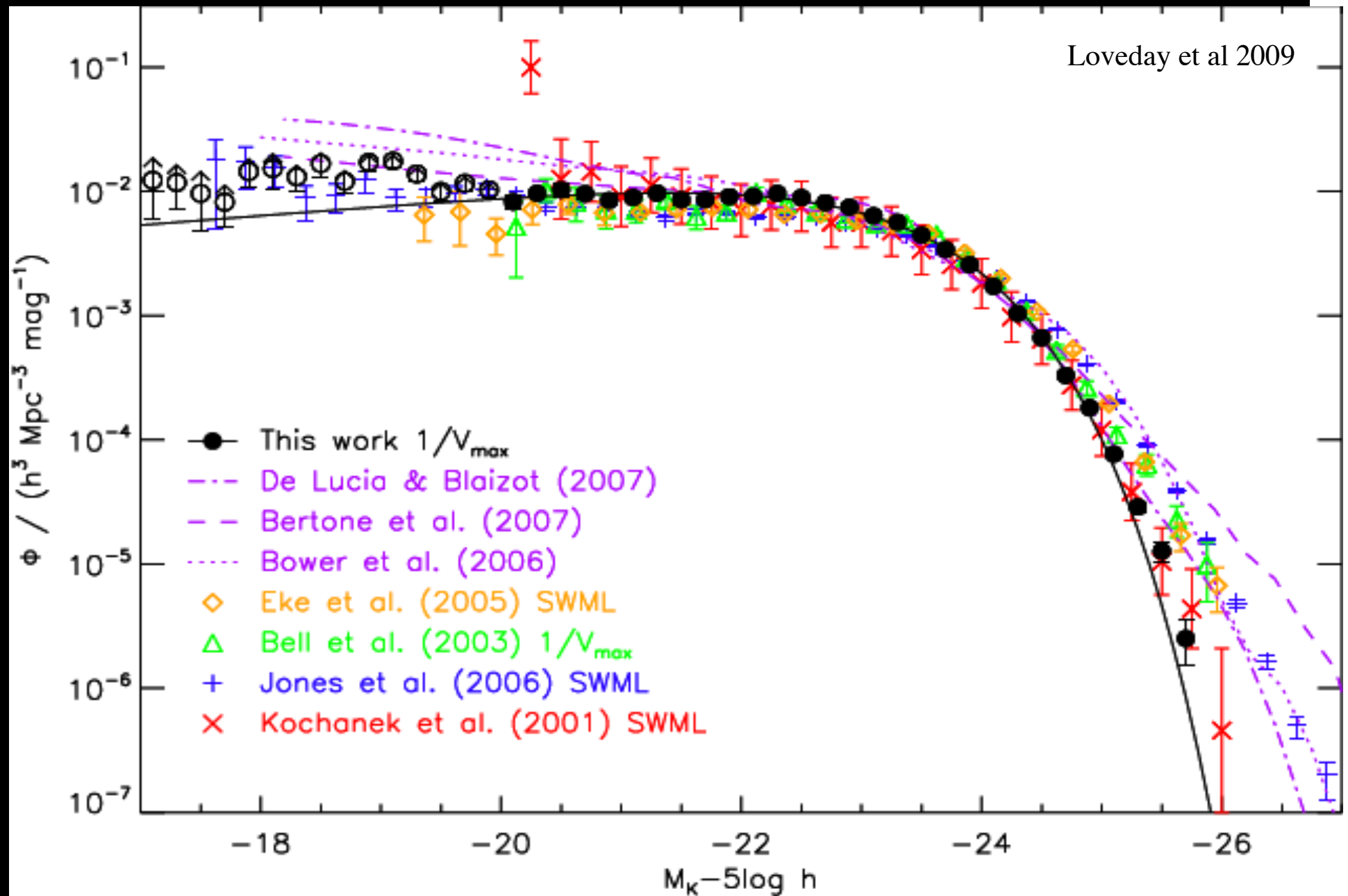
calculate number of newly non-linear objects

sensitive to gaussian tail i.e. rare peaks



Luminosity function of galaxies

Loveday et al 2009



FEEDBACK IS ESSENTIAL

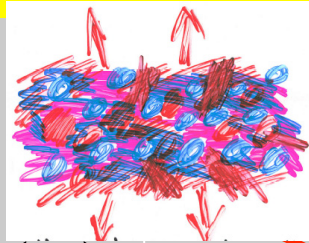
Reionization

Supernovae

Tidal stripping

Active galactic nuclei

Feedback is needed



$$M_{\text{cooled-baryons}} \sim \alpha_g^{-2} \alpha^3 \left(\frac{m_p}{m_e} \right) \left(\frac{t_{\text{cool}}}{t_{\text{dyn}}} \right) T^{1+2\beta}$$

theory (CDM-motivated)

$$\alpha_g = Gm_p^2/e^2$$

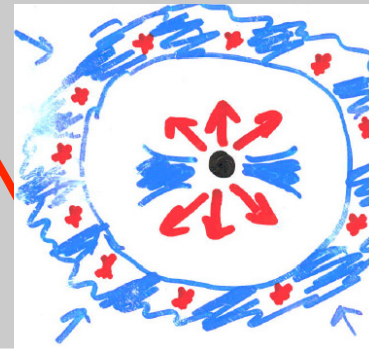
$\phi(L)$

$$L_* \sim 3 \times 10^{10} L_{\odot}$$

$$t_{\text{cool}} \sim \frac{nkT}{\Lambda(T)n^2}$$

$$t_{\text{dyn}} \sim \frac{1}{\sqrt{Gm_p n}}$$

observations

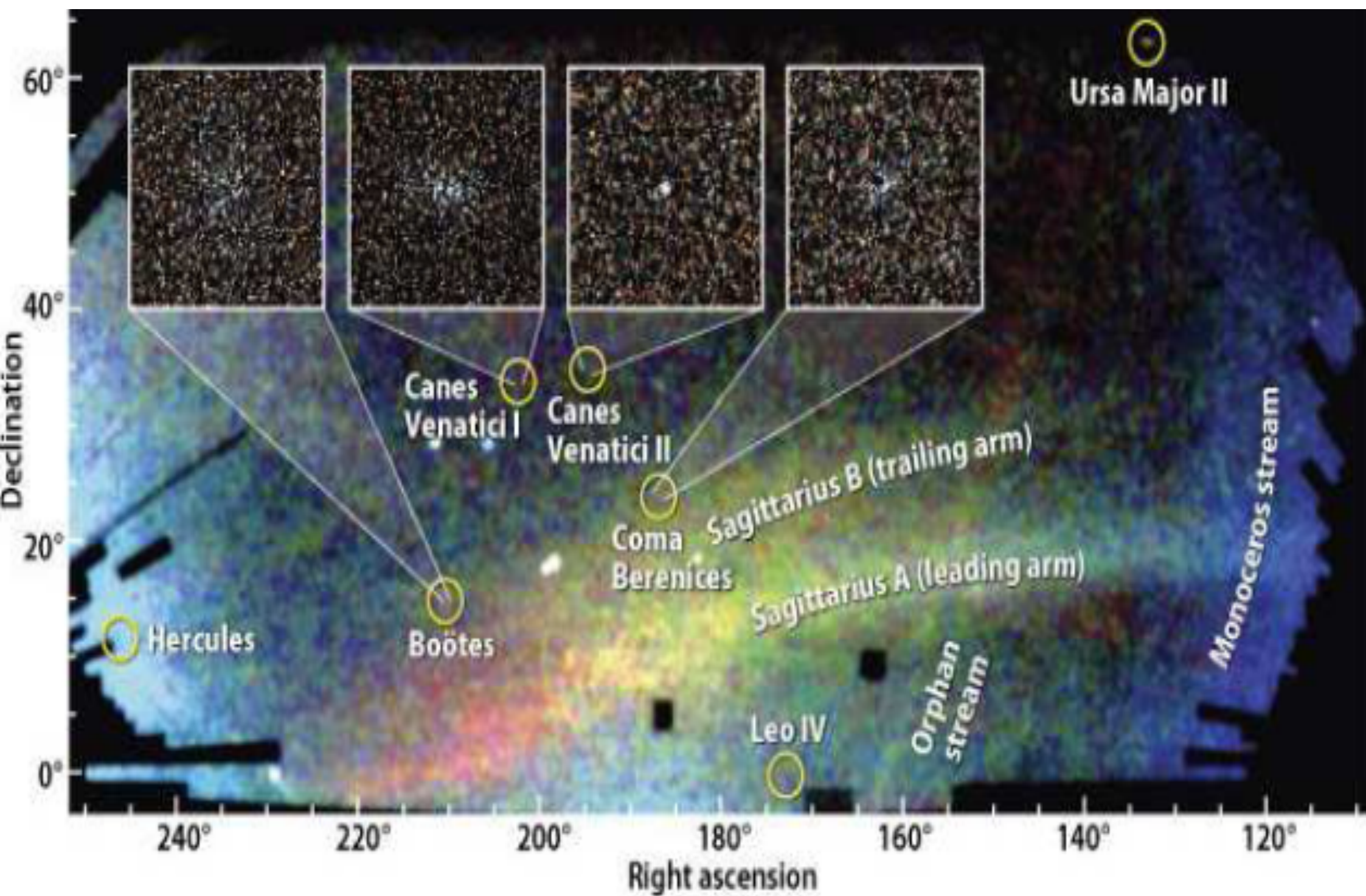


luminosity

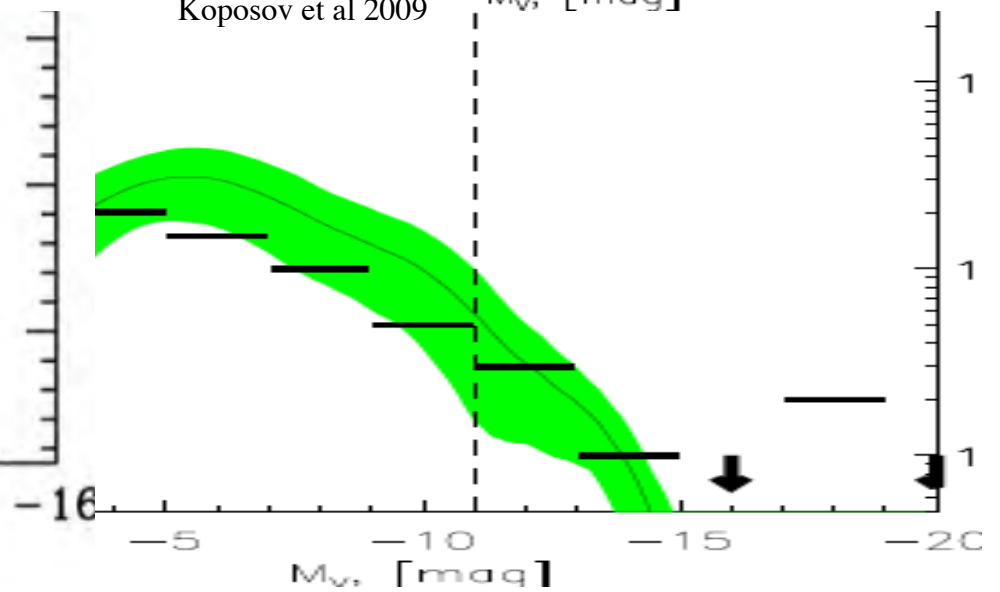
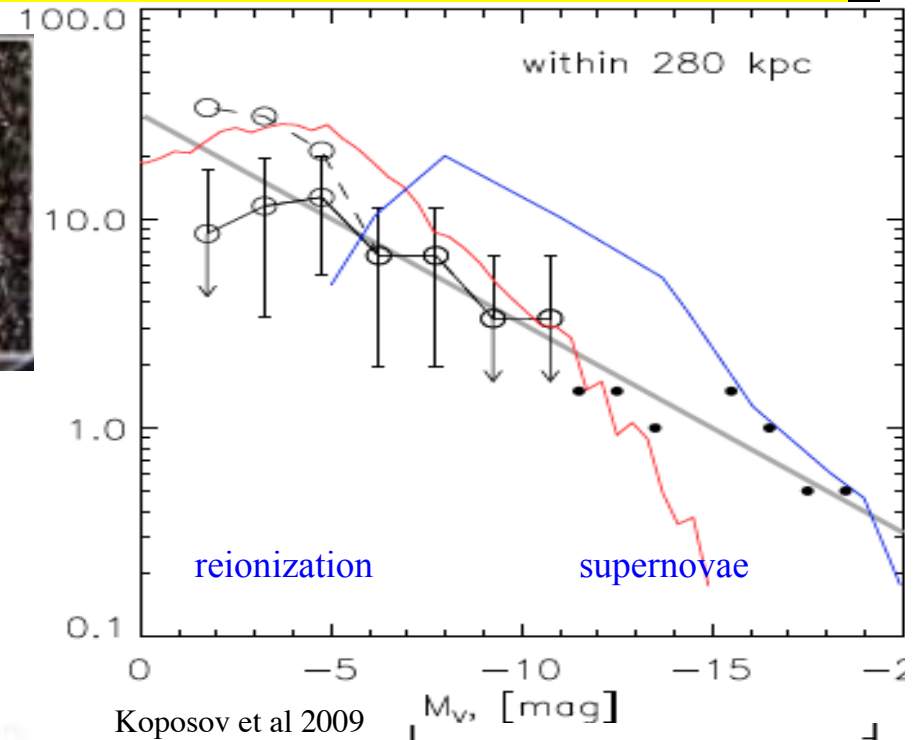
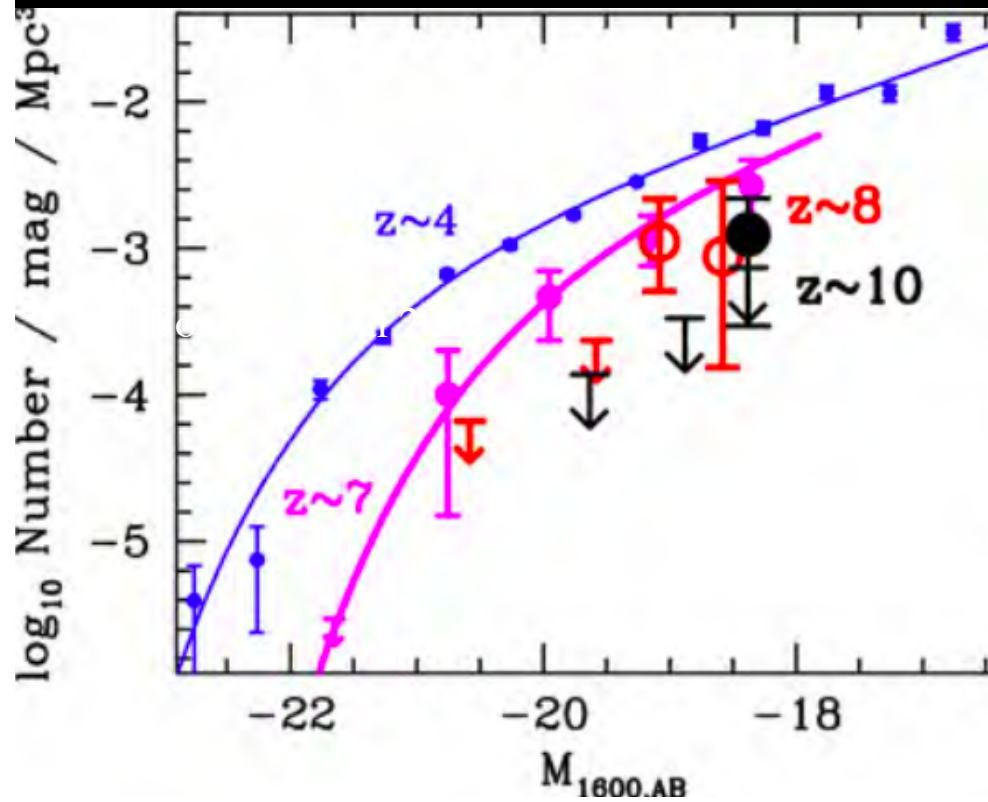
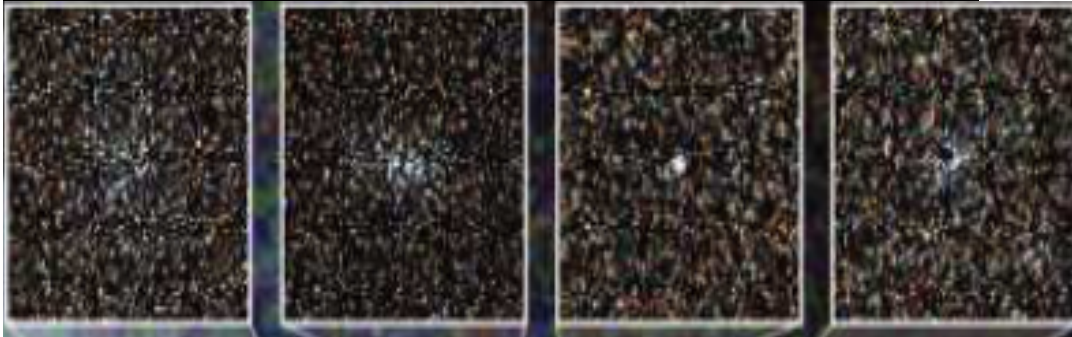
SN

AGN



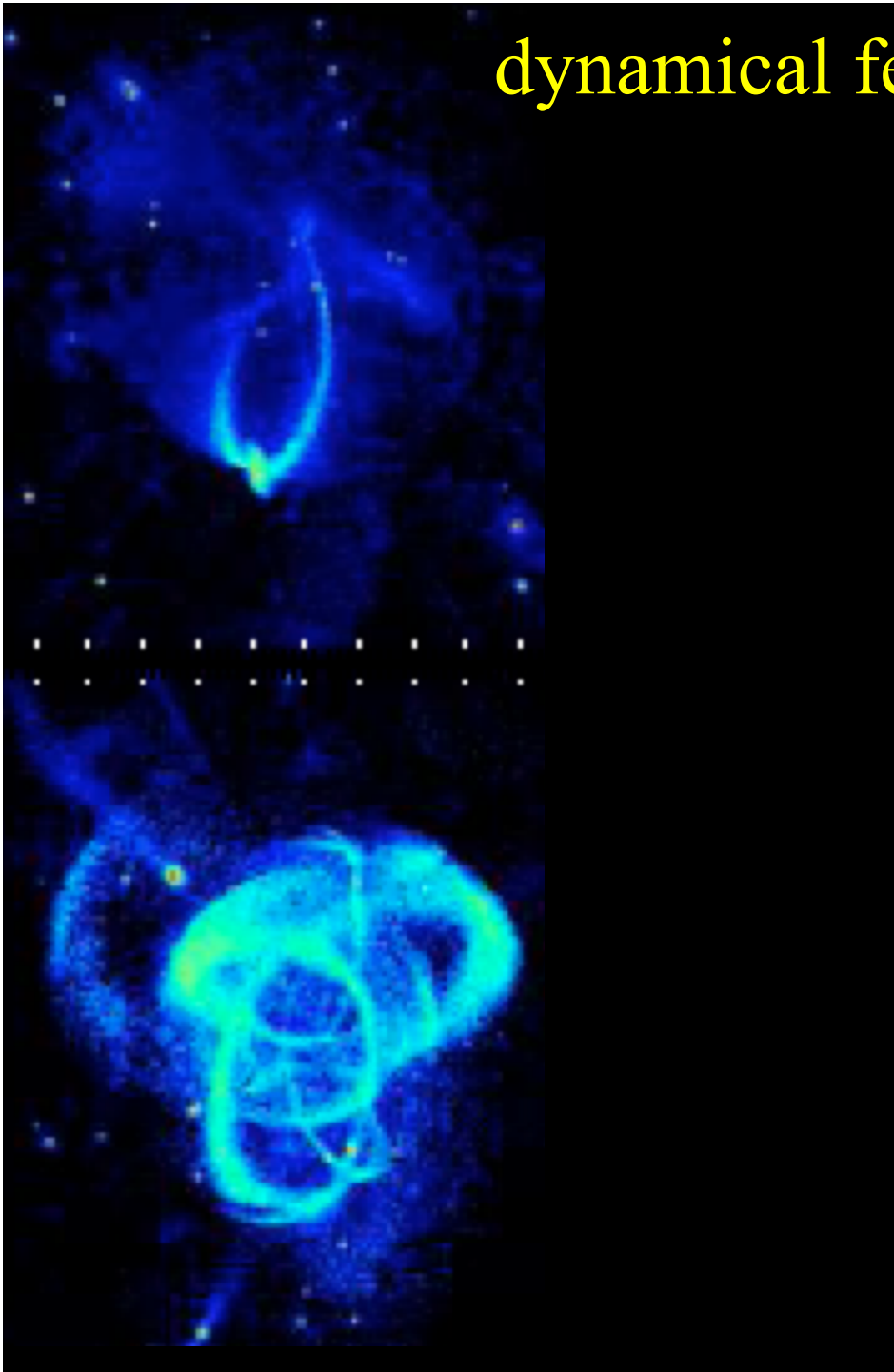


Feedback in low mass galaxies

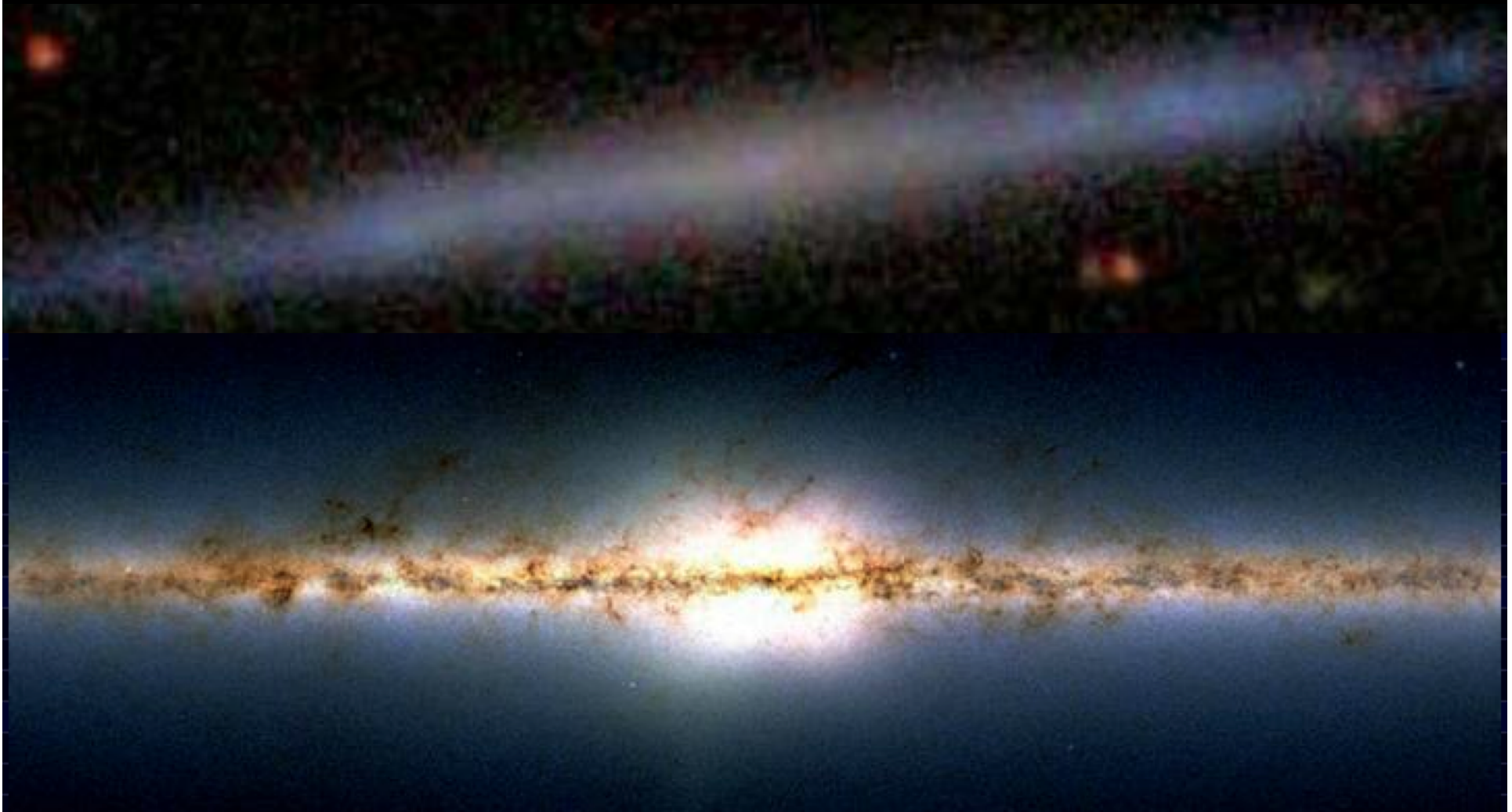


dynamical feedback

Martinez-Delgado et al 2008



**BUT WE CAN'T EXPLAIN
> 15% OF GALAXIES!**



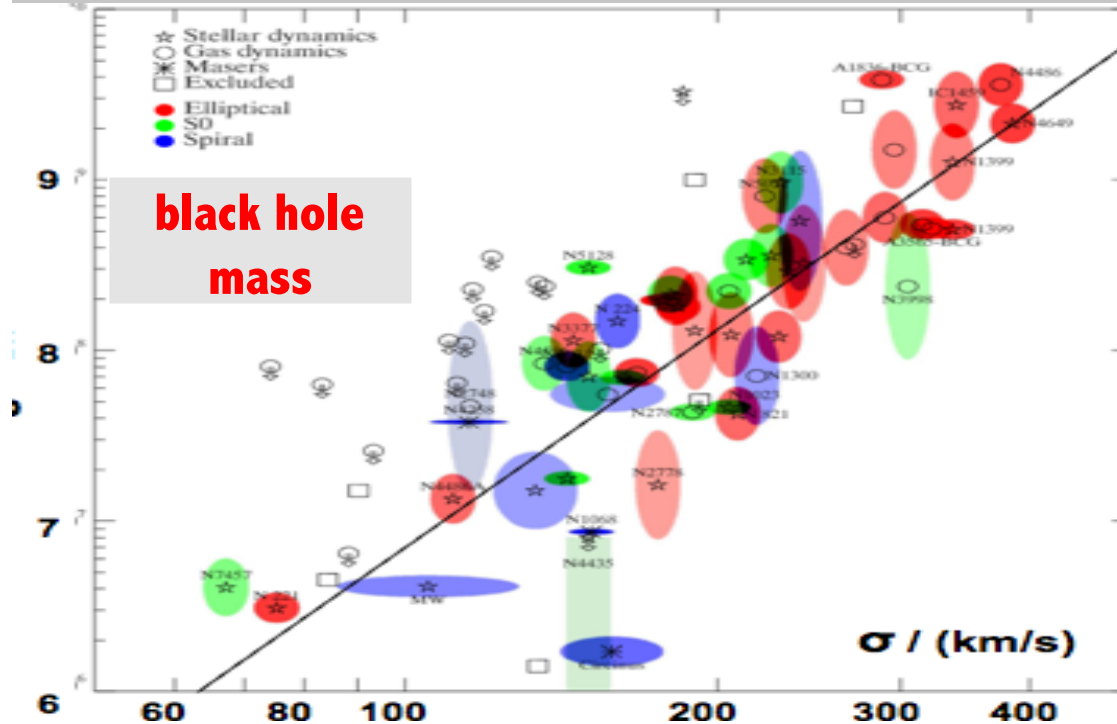
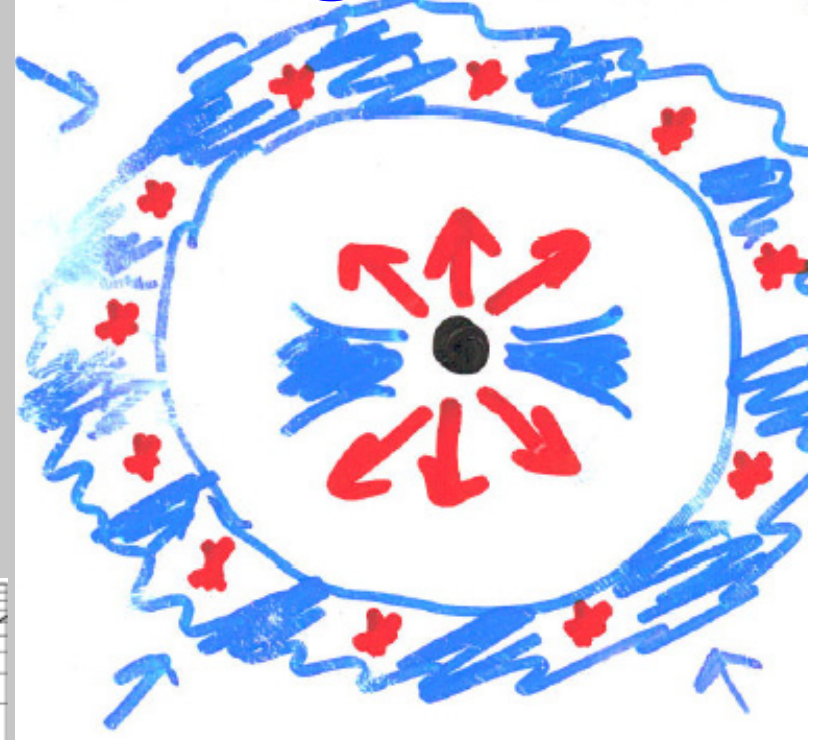
3. Spheroidal galaxy formation

Feedback by massive galaxies

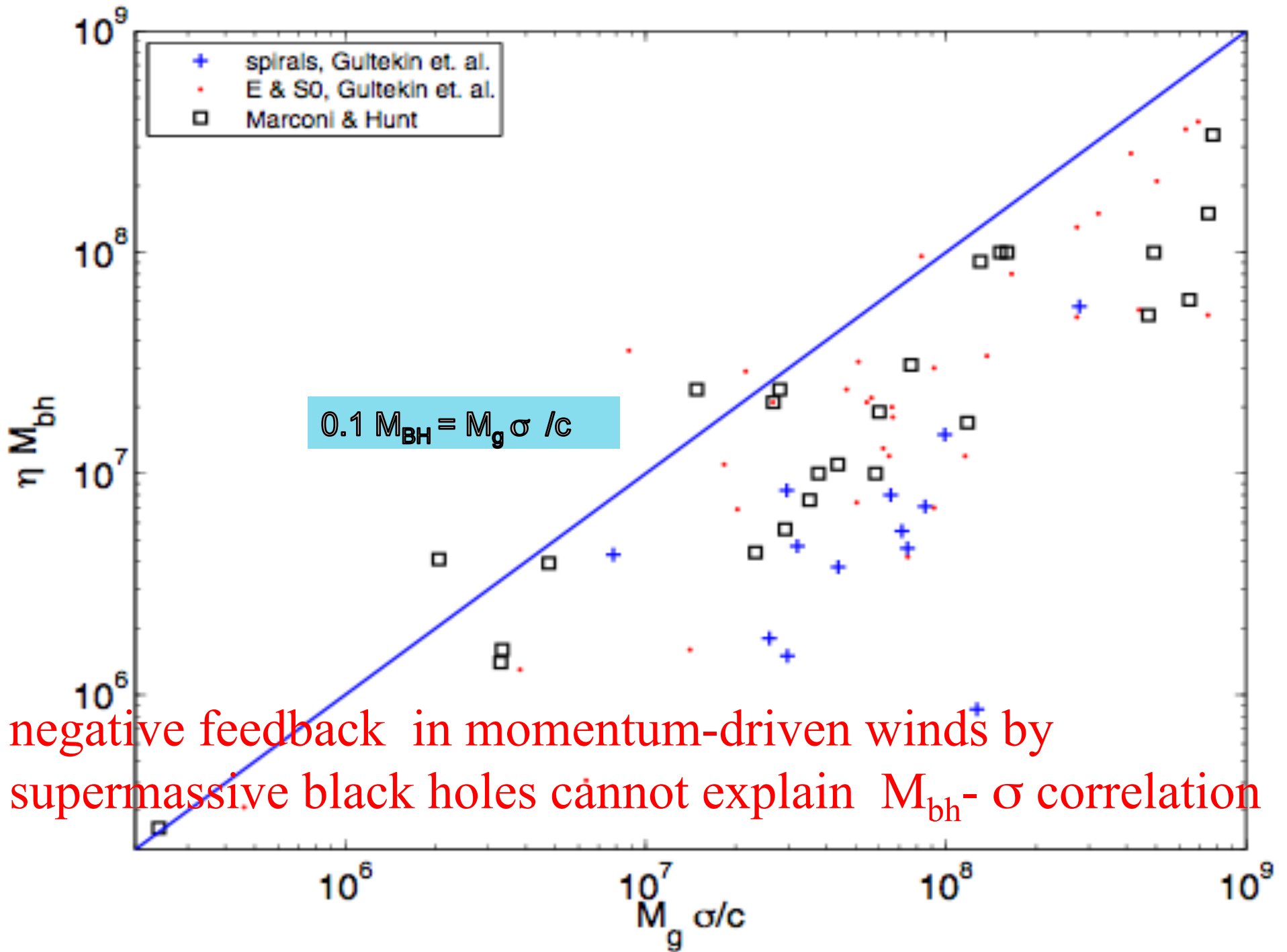
$$L_{\text{Edd}}/c = G M M_{\text{gas}}/r^2$$

$$M_{\bullet} = 3 \times 10^9 M_{\text{sun}} \left(\frac{\sigma}{300 \frac{\text{km}}{\text{s}}} \right)^4$$

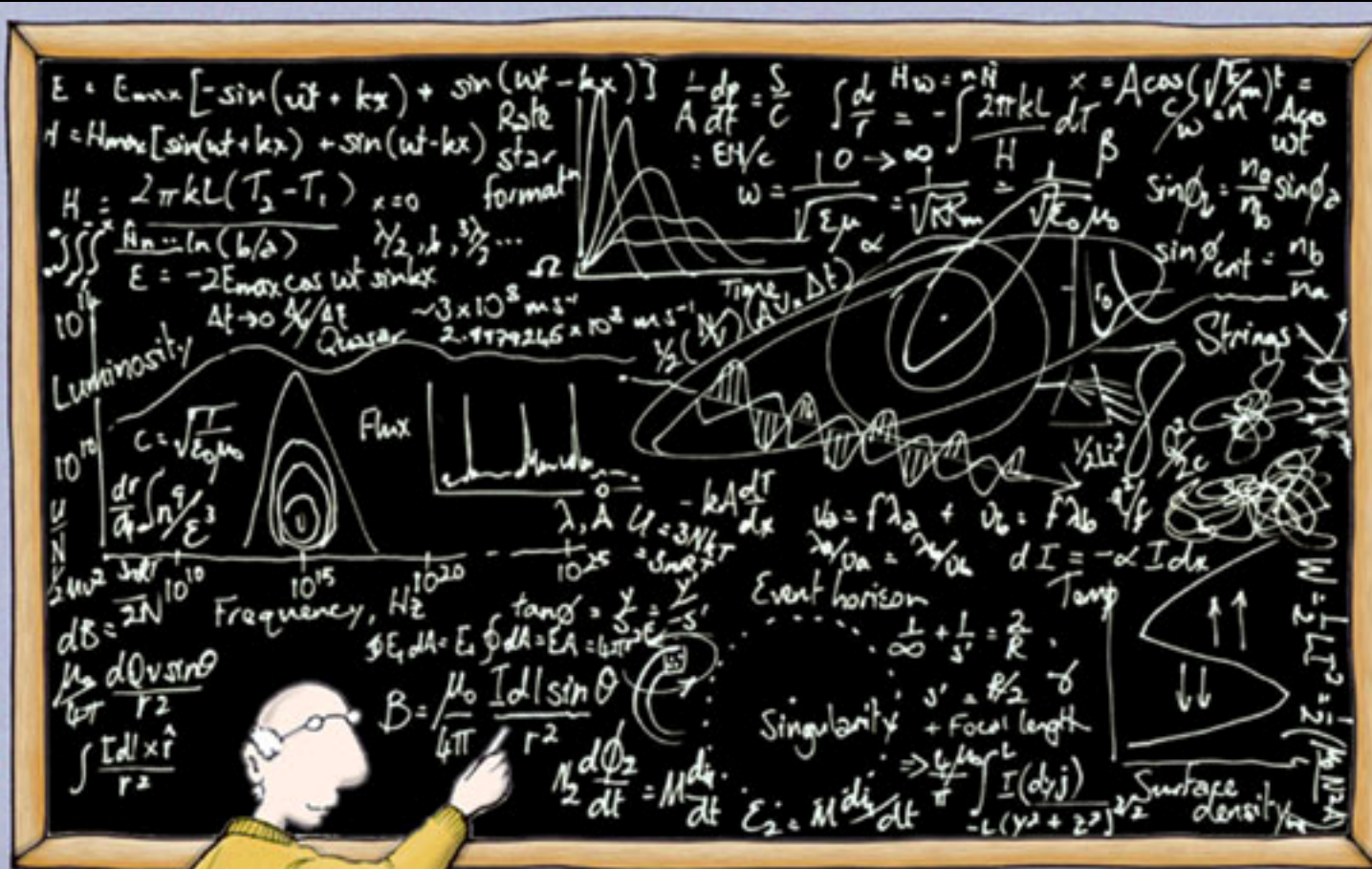
Blowout occurs/star formation terminates
when SMBH- σ relation saturates



$$L_{\text{Edd}} = 4\pi c G M_{\text{BH}} m_p / \sigma_T$$



negative feedback in momentum-driven winds by supermassive black holes cannot explain $M_{\text{bh}} - \sigma$ correlation



Back to the drawing board...

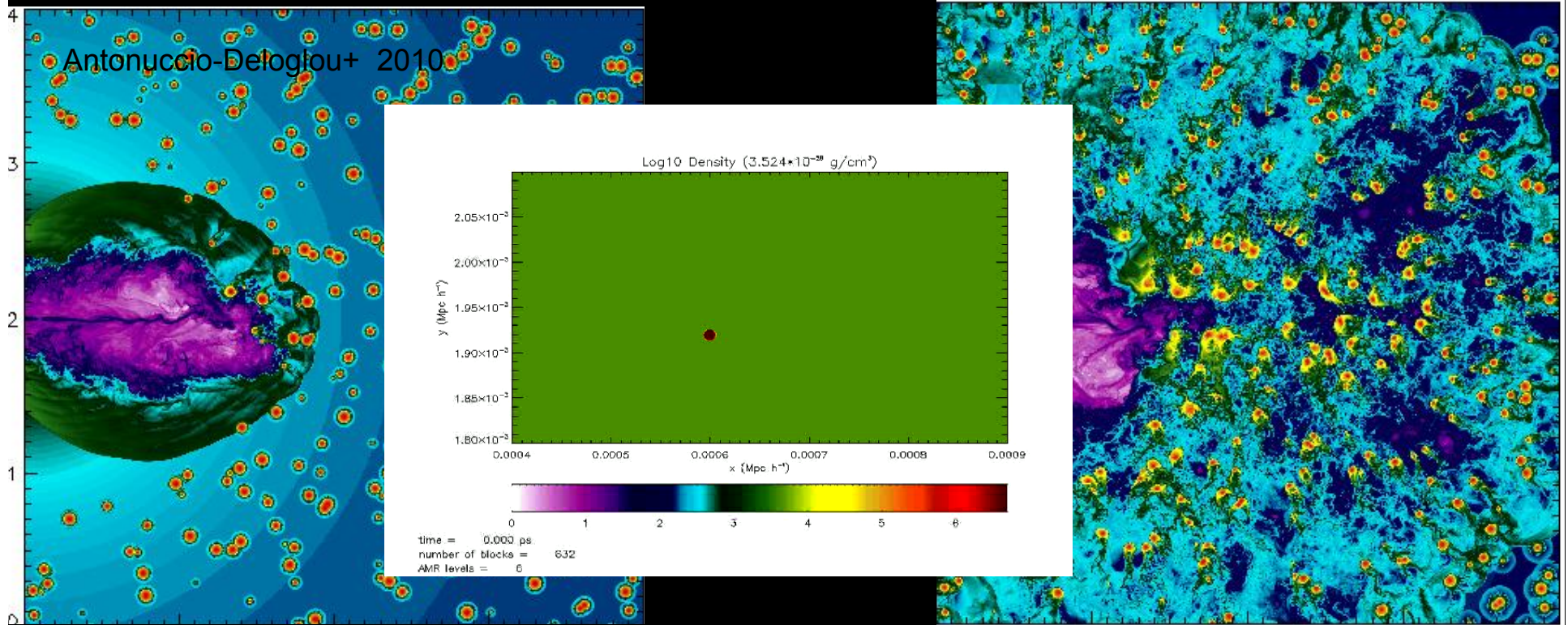
Its not supernovae,
its not active galactic
nuclei...
maybe its both!

start formation rate is boosted by AGN

If AGN-driven outflows trigger star formation,

JS + C. Norman 2008

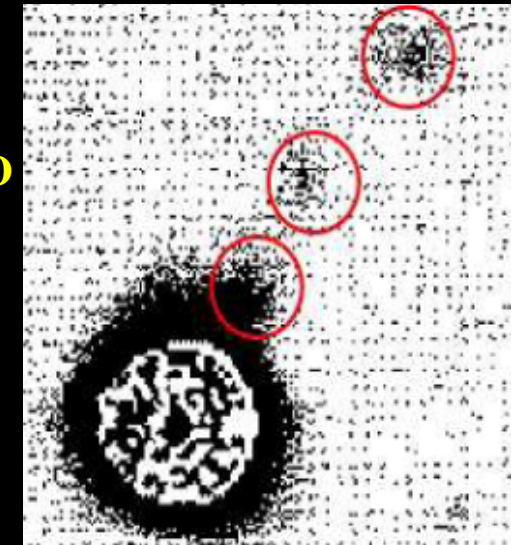
**the star formation rate is boosted by factor $t_{\text{dyn}}/t_{\text{jet}}$
and the outflow momentum is amplified by supernovae**



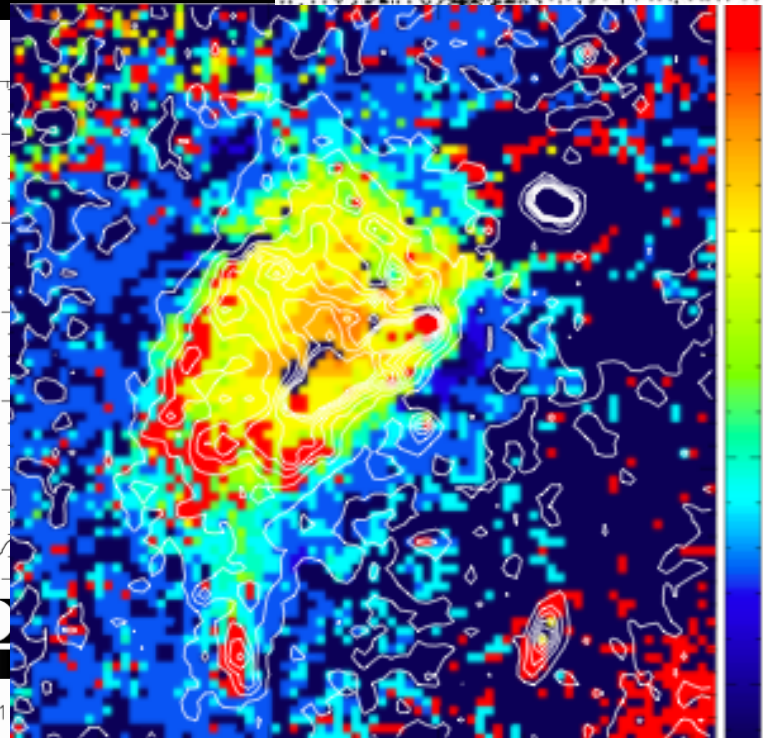
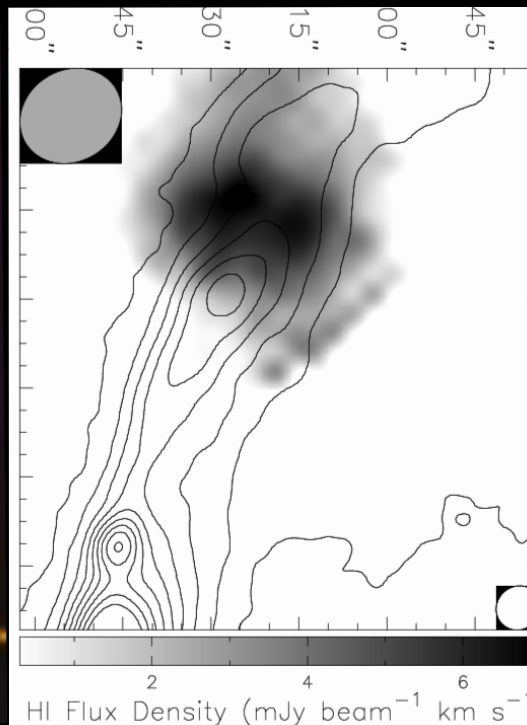
momentum boost by AGN + star formation

JS + A. Nusser 2010

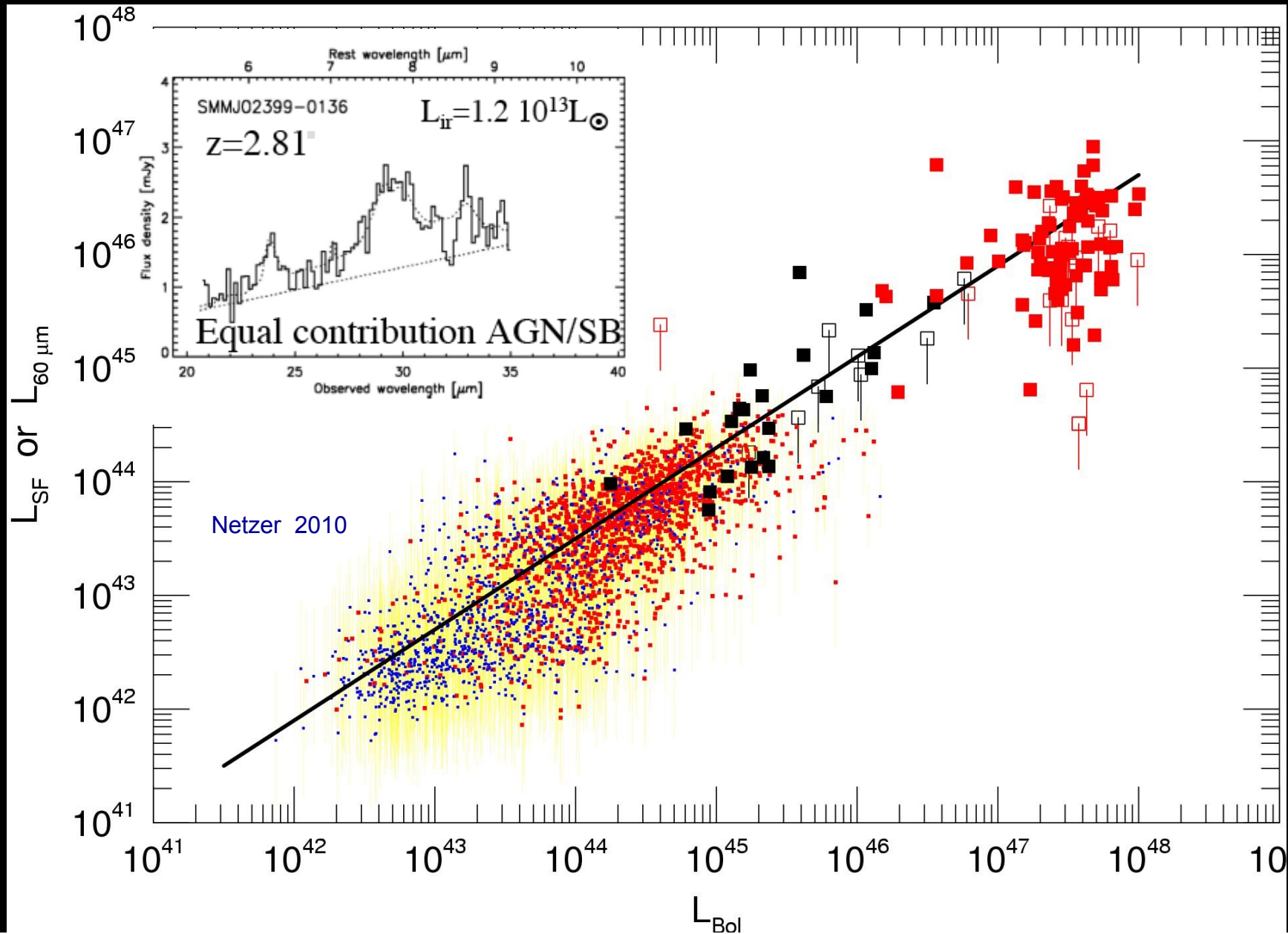
$z = 4.7$ quasar + CO



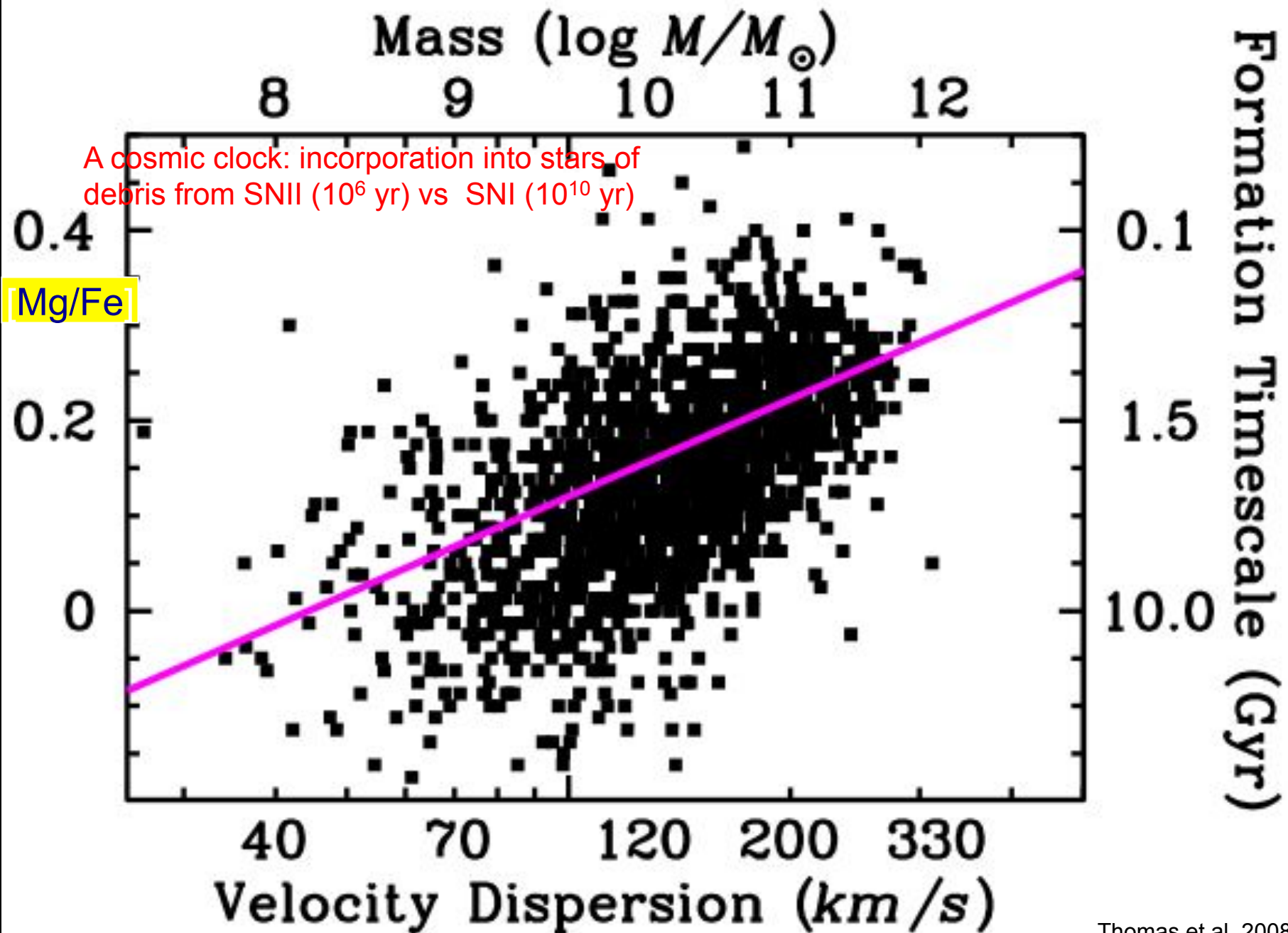
Momentum boost =
supernova momentum/Eddington momentum
 ~ 10



connection between AGN and starbursts

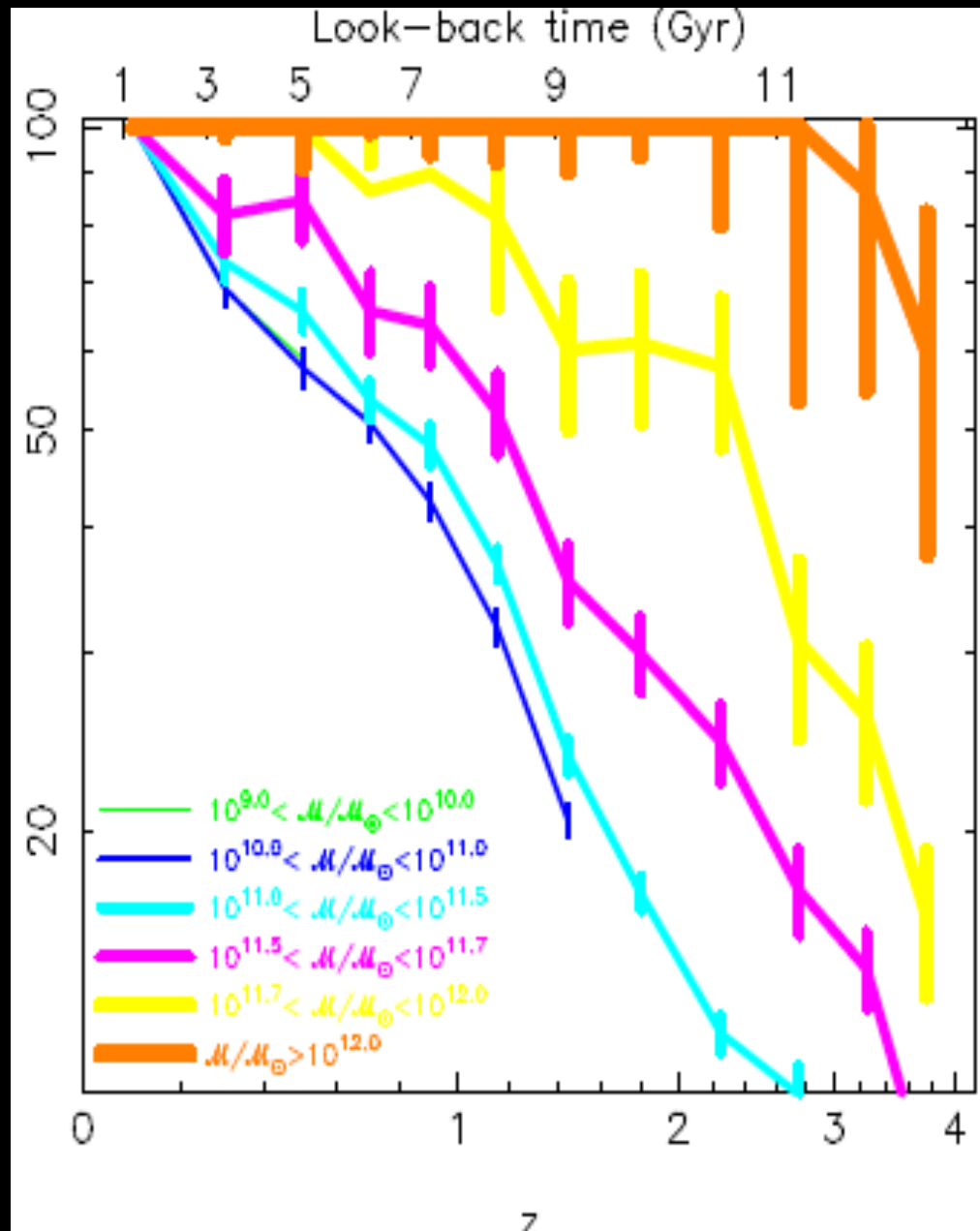


DOWNSIZING INFERRED: A SURPRISE

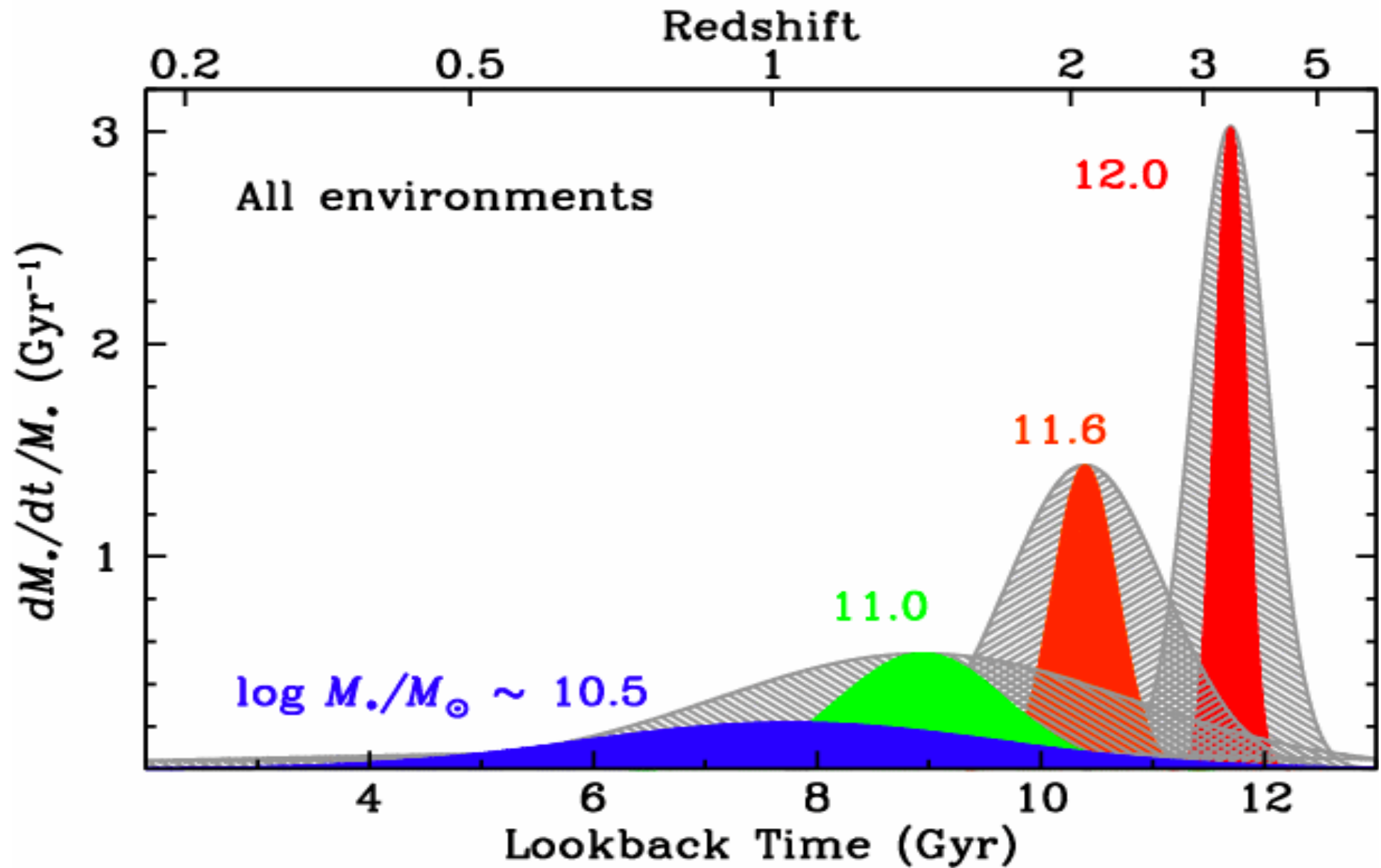


Downsizing in Stellar Mass

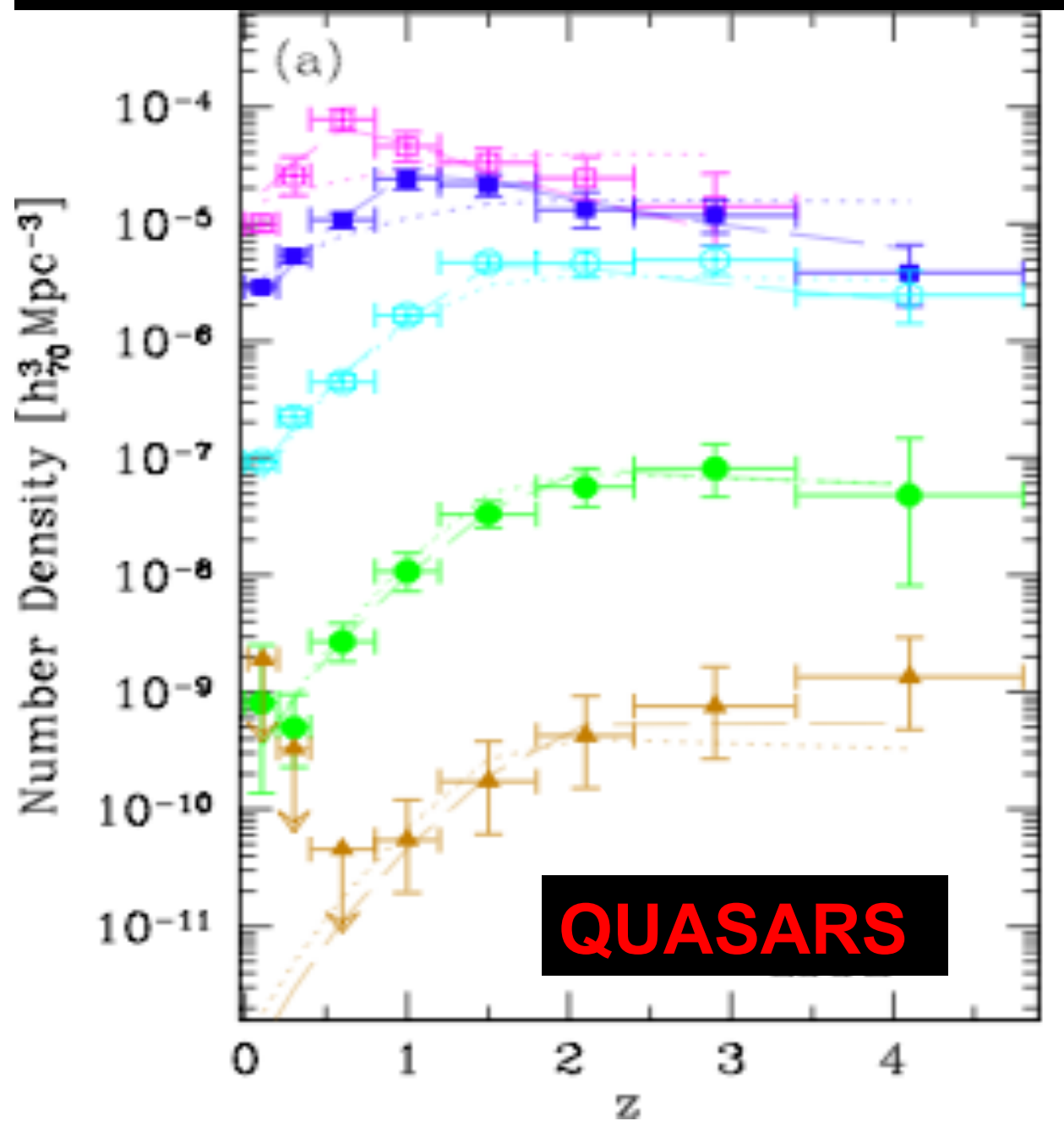
Perez-Gonzalez et al. 2007



Downsizing Summarized



DOWNSIZING IN BLACK HOLE MASS



Log L Range

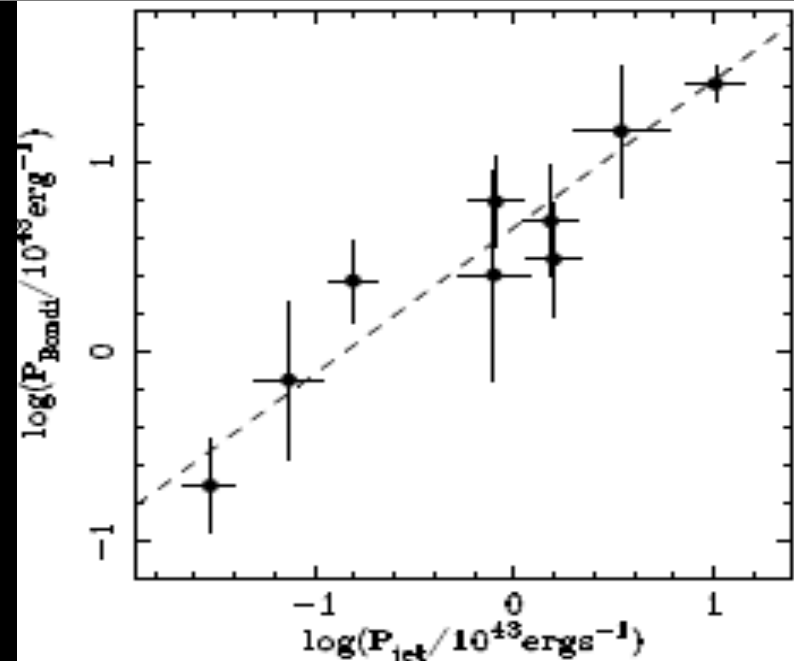
▲	>46.0
●	$45.0-46.0$
○	$44.0-45.0$
■	$43.0-44.0$
□	$42.0-43.0$

Why efficient formation of massive spheroids ?

- assume AGN-driven outflows trigger/regulate star formation

Why downsizing of massive spheroids and AGN?

- Need SMBH feeding and outflow to be nonlinear function of SMBH mass

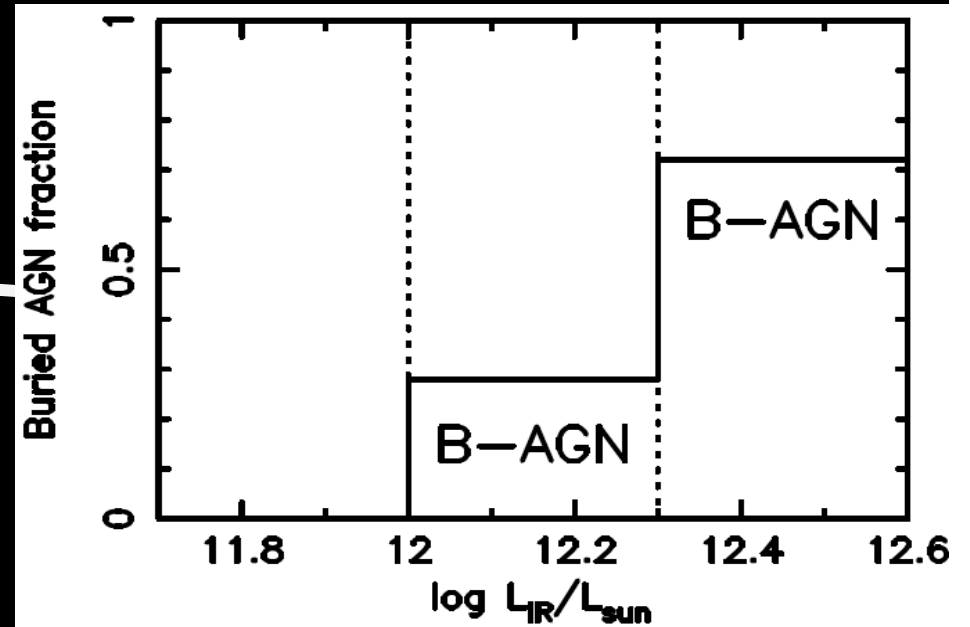
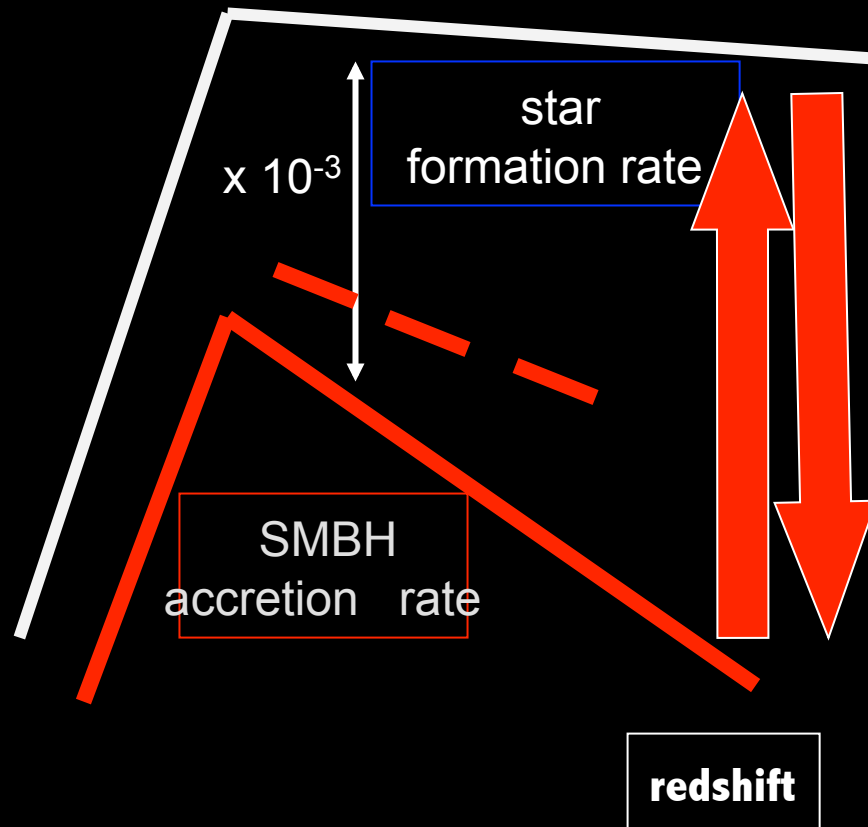


Active galactic nuclei-

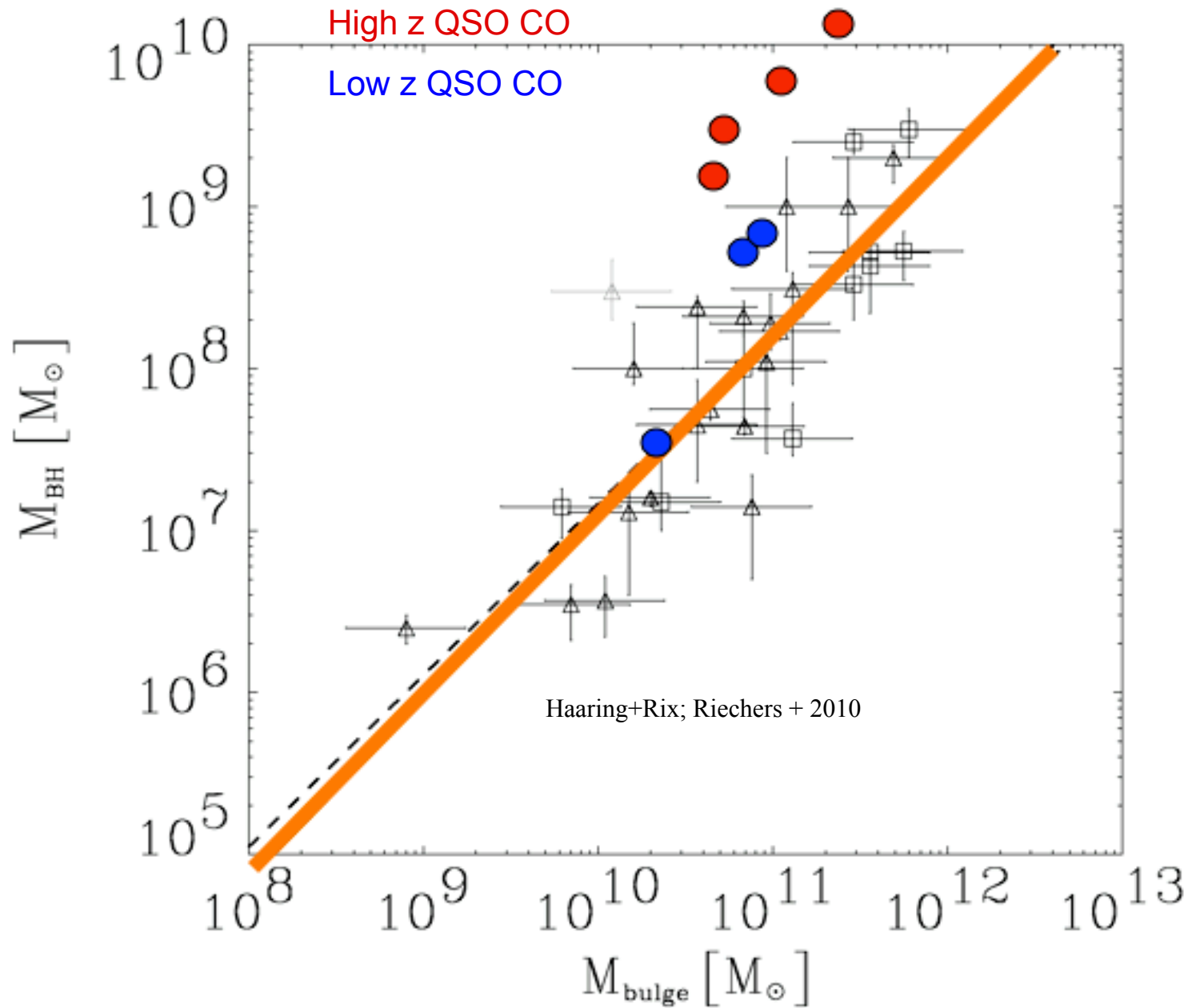
an aftermath or precursor to star formation?

gravity-induced
star formation

quenching and/or
triggering



SMBH in $z \sim 6$ quasars lie high....



4. Star formation

two modes of star formation:

A: without active galactic nuclei

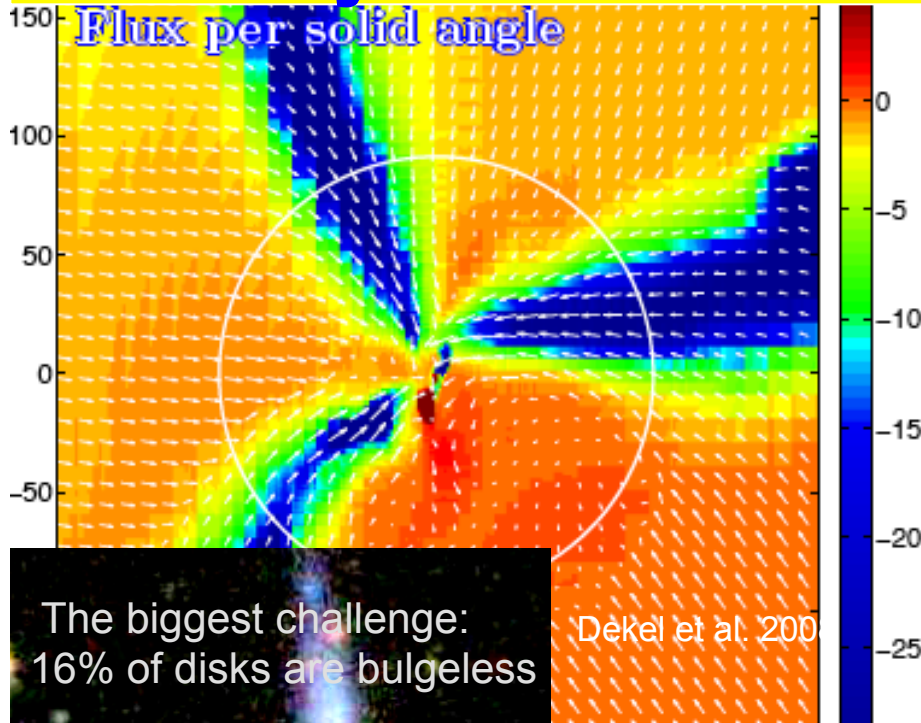
B: with active galactic nuclei

two modes of gas accretion:

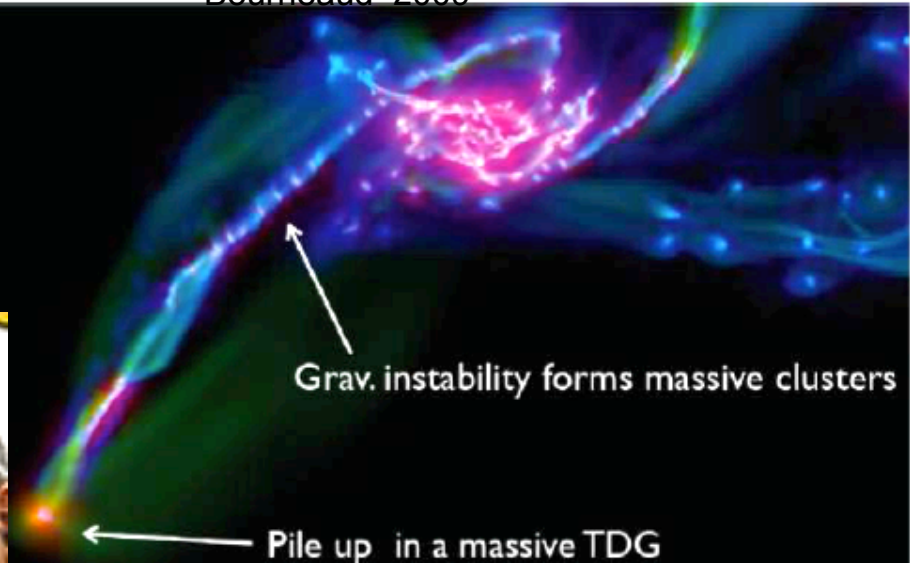
A: cold streams/minor mergers

B: major mergers/cooling flows

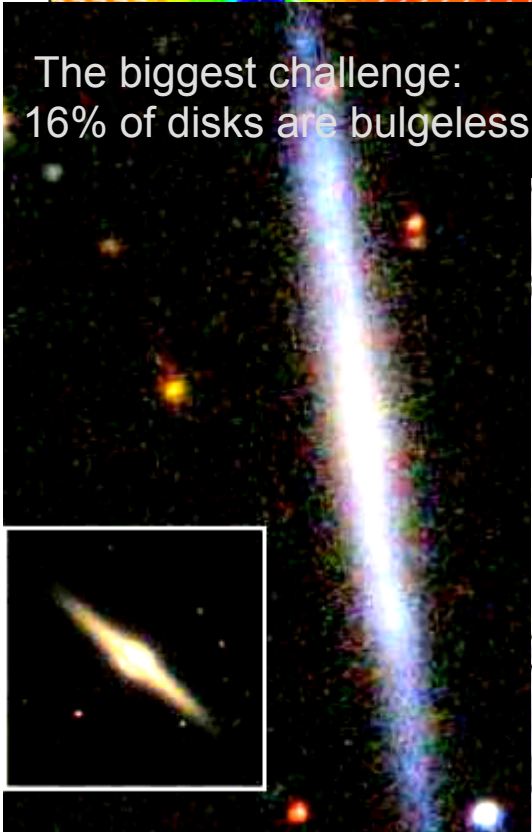
Infall by cold streams



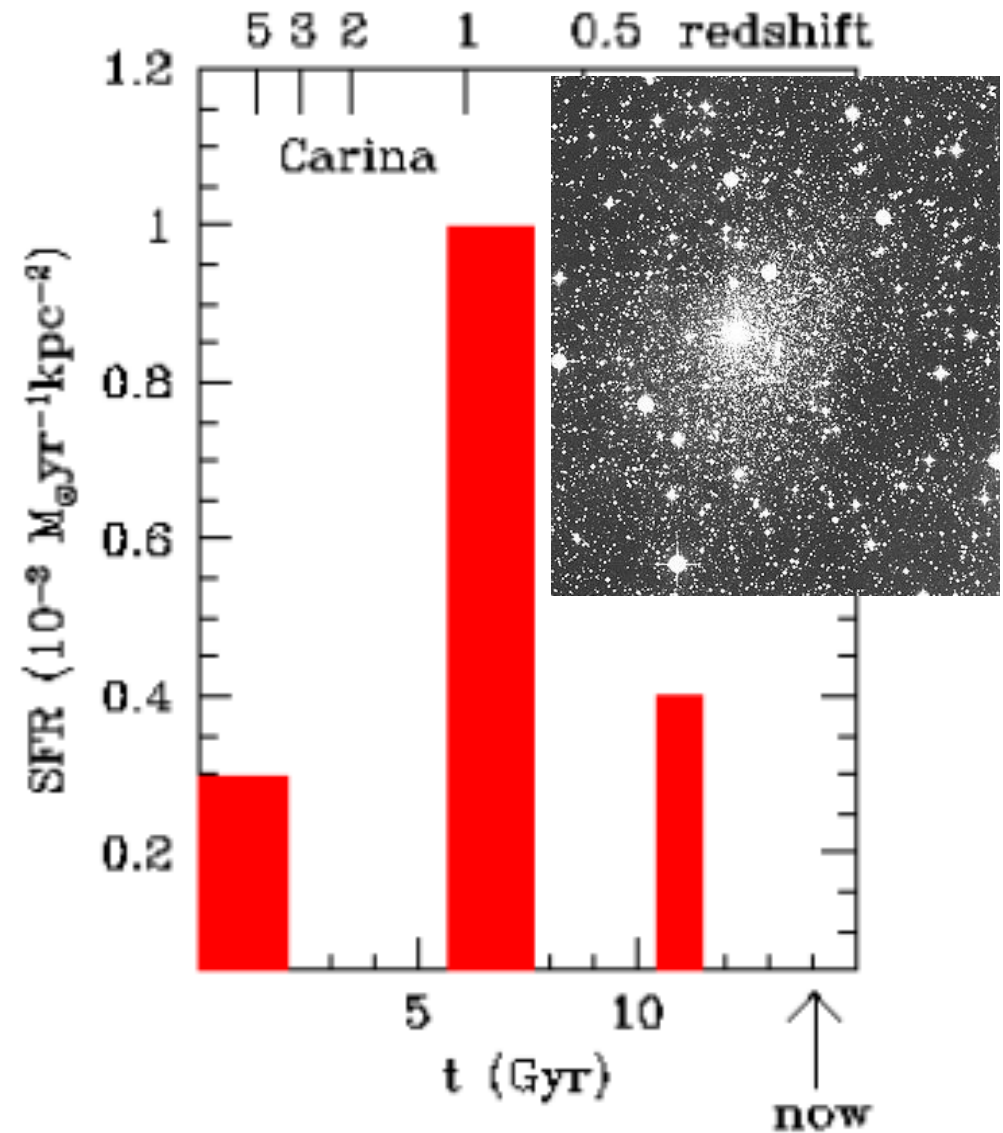
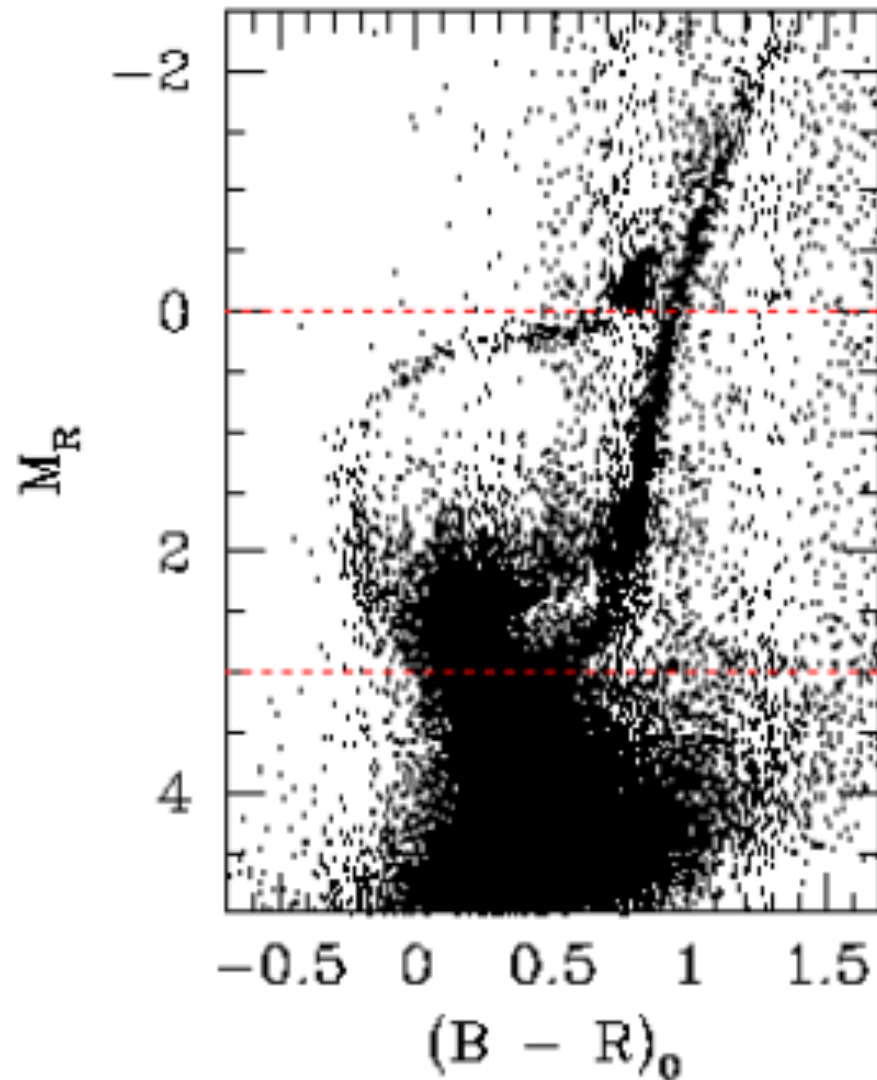
major mergers



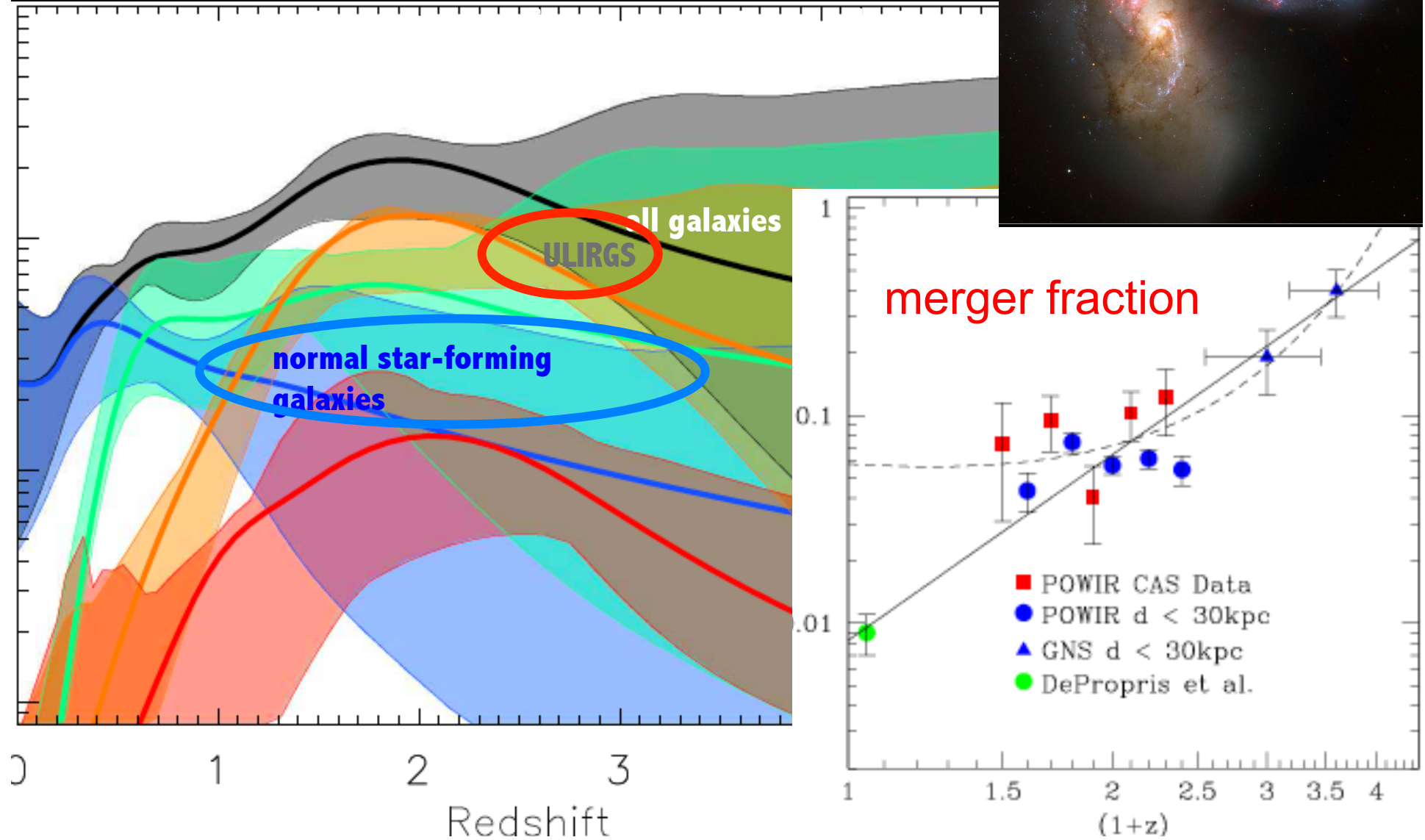
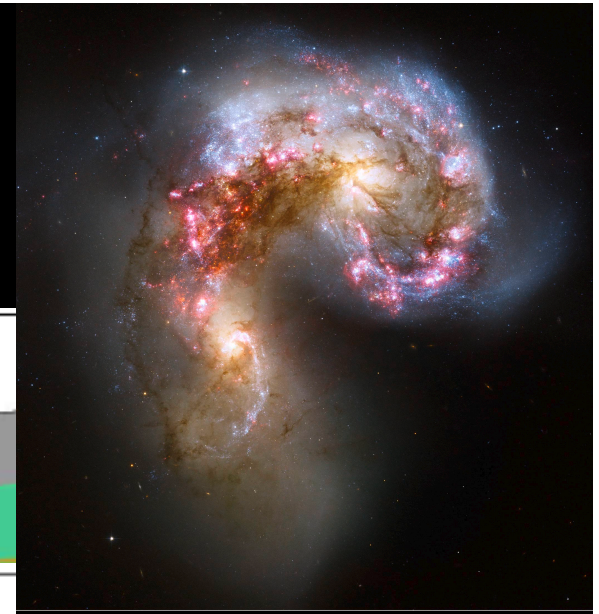
The biggest challenge:
16% of disks are bulgeless



cold accretion dominates for disks and dwarfs

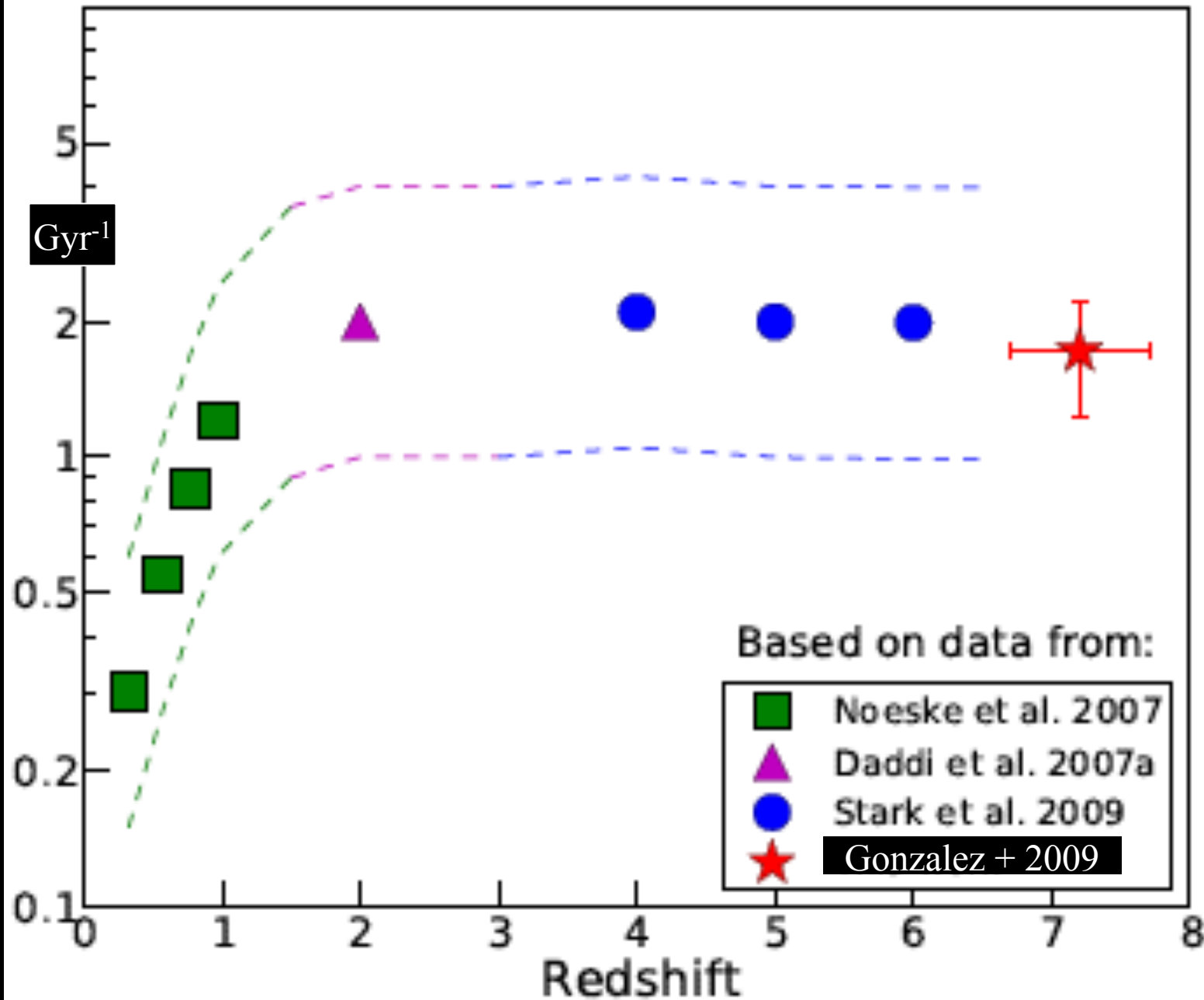


Cosmic star formation rate



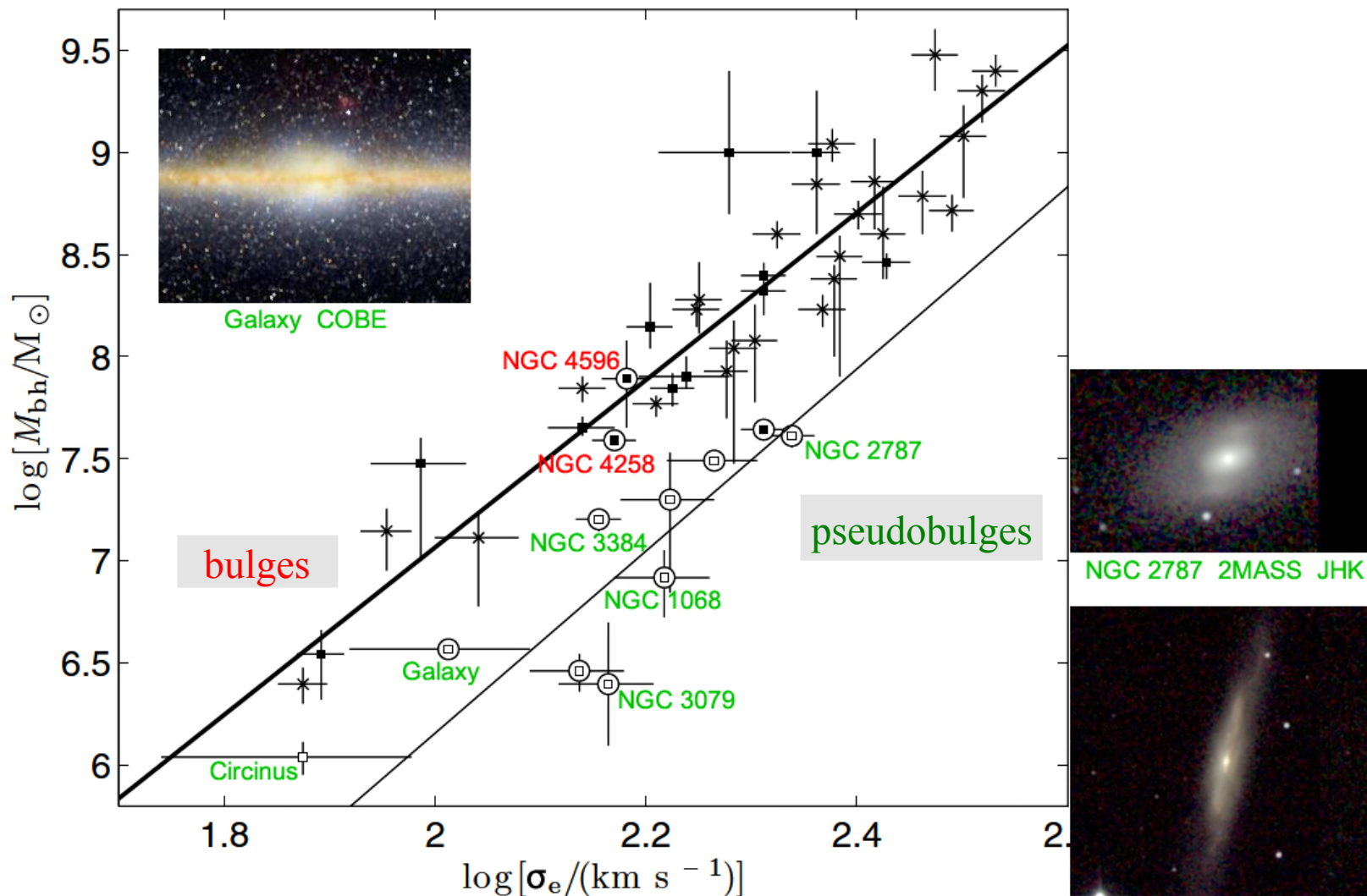
specific star formation rate

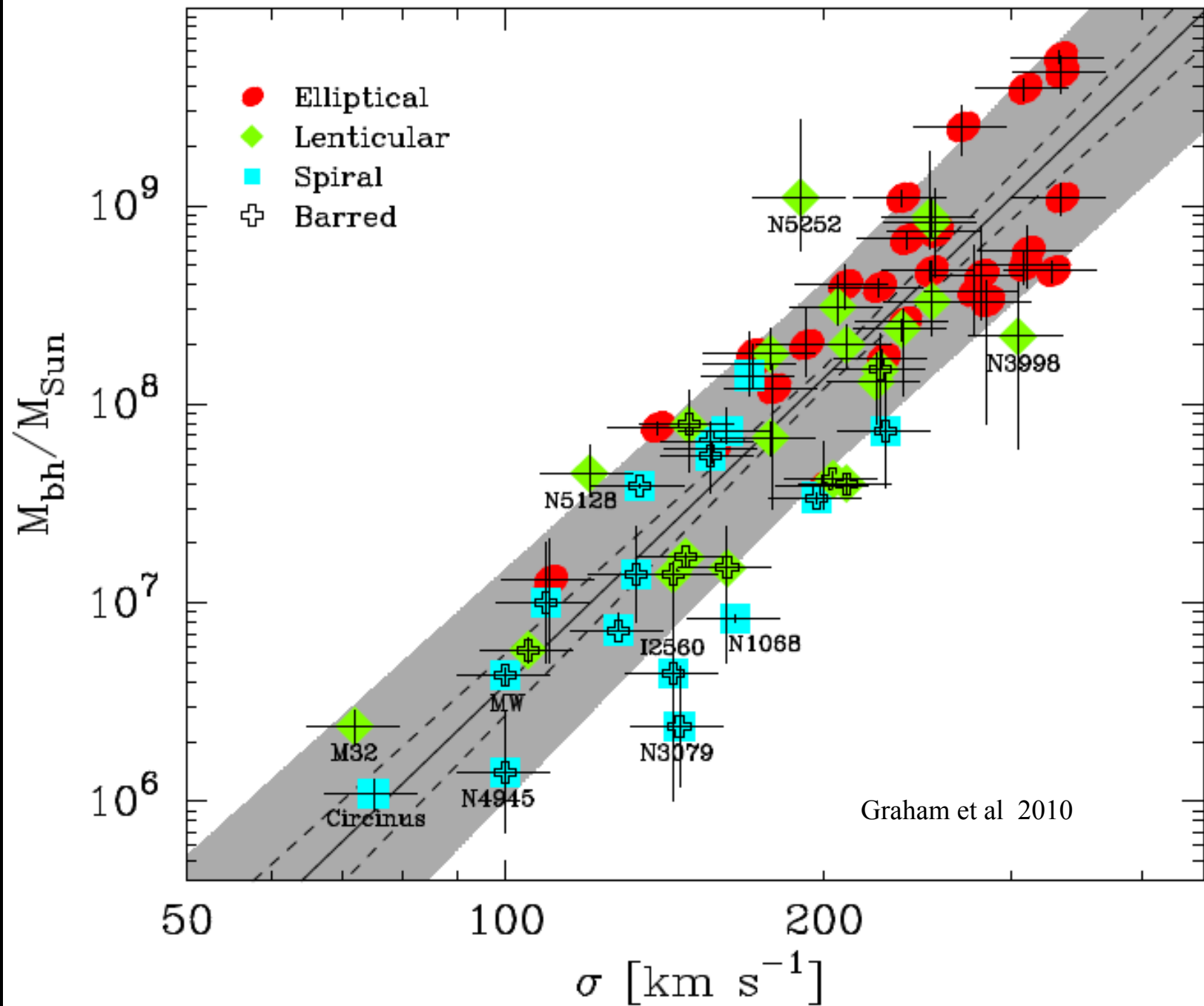
$$= \text{SFR} / \text{stellar mass}$$



SMBH in pseudobulges lie low....

Suggests disk mode formation of SMBH is distinct from spheroid mode



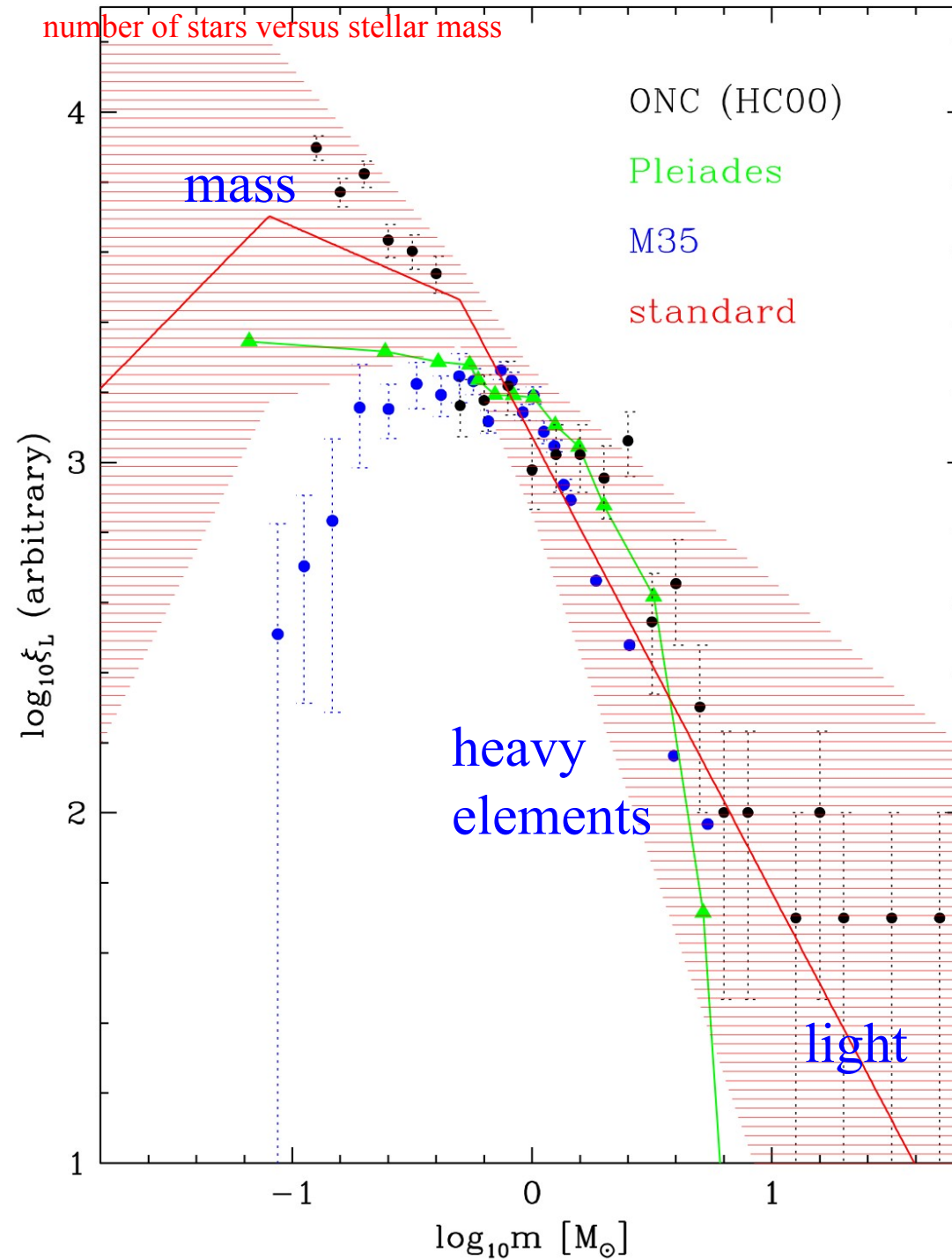


What's missing?

A robust theory of star
formation

Kroupa 2003

Initial mass
function
of stars:
a fundamental
theory is
lacking



Feedback: star formation

fragmentation theory applied to collapsing interstellar cloud implies minimum fragment mass.....a robust but wrong result

$$\sim \alpha_g^{-3/2} m_p \sim 0.01 M_\odot$$

resolution: **accretion** of cold gas

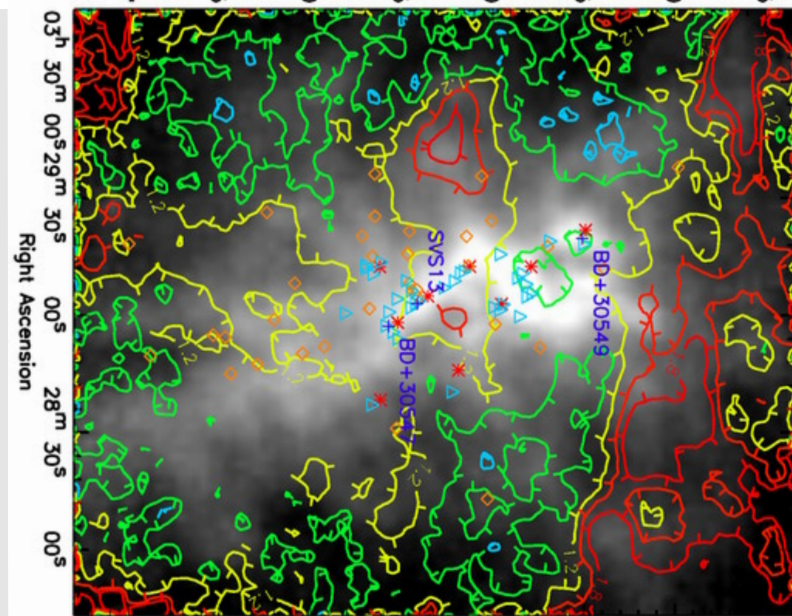
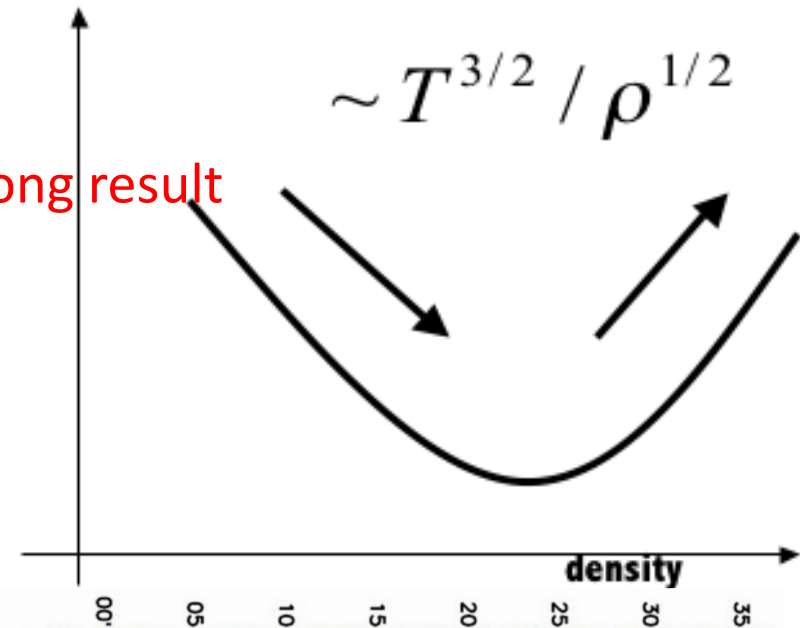
$$\dot{M}_{gas} \sim V_s^3 / G \Rightarrow$$

$\sim 10^{-6} M_{sun}/yr$ today at $T=10K$
but $\sim 10^{-3} M_{sun}/yr$ at $T=1000K$

first stars were massive

halted by **feedback**: taps stellar energy via MHD turbulence

Jeans mass

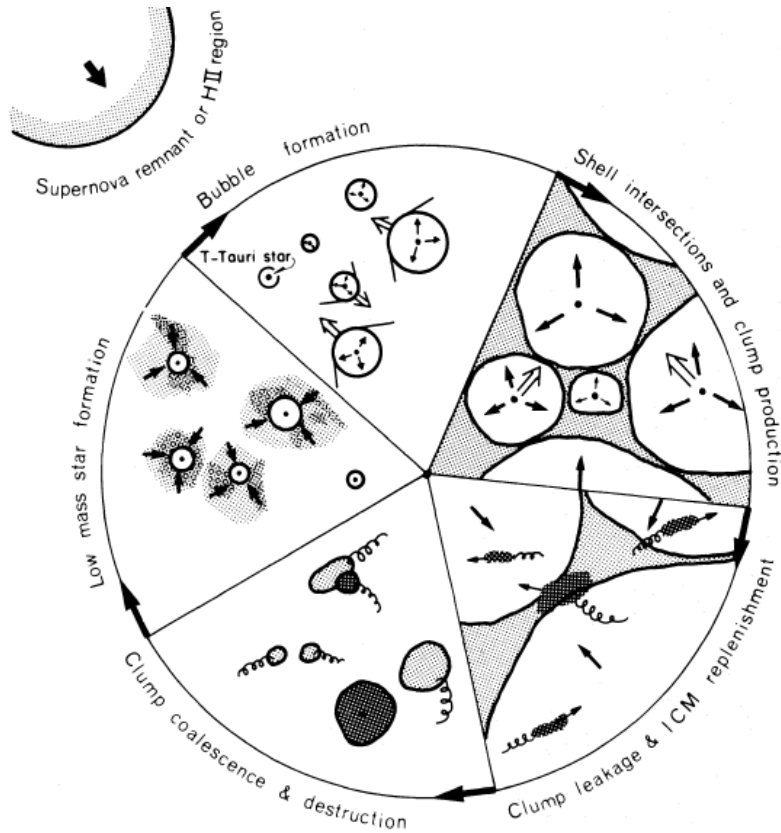


NGC1333: Quillen+2006

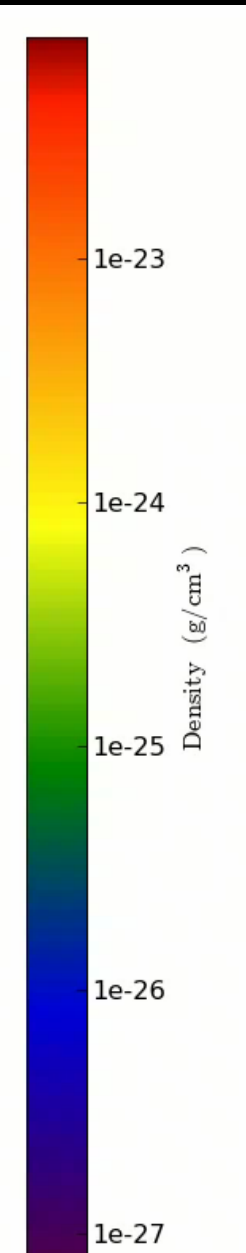
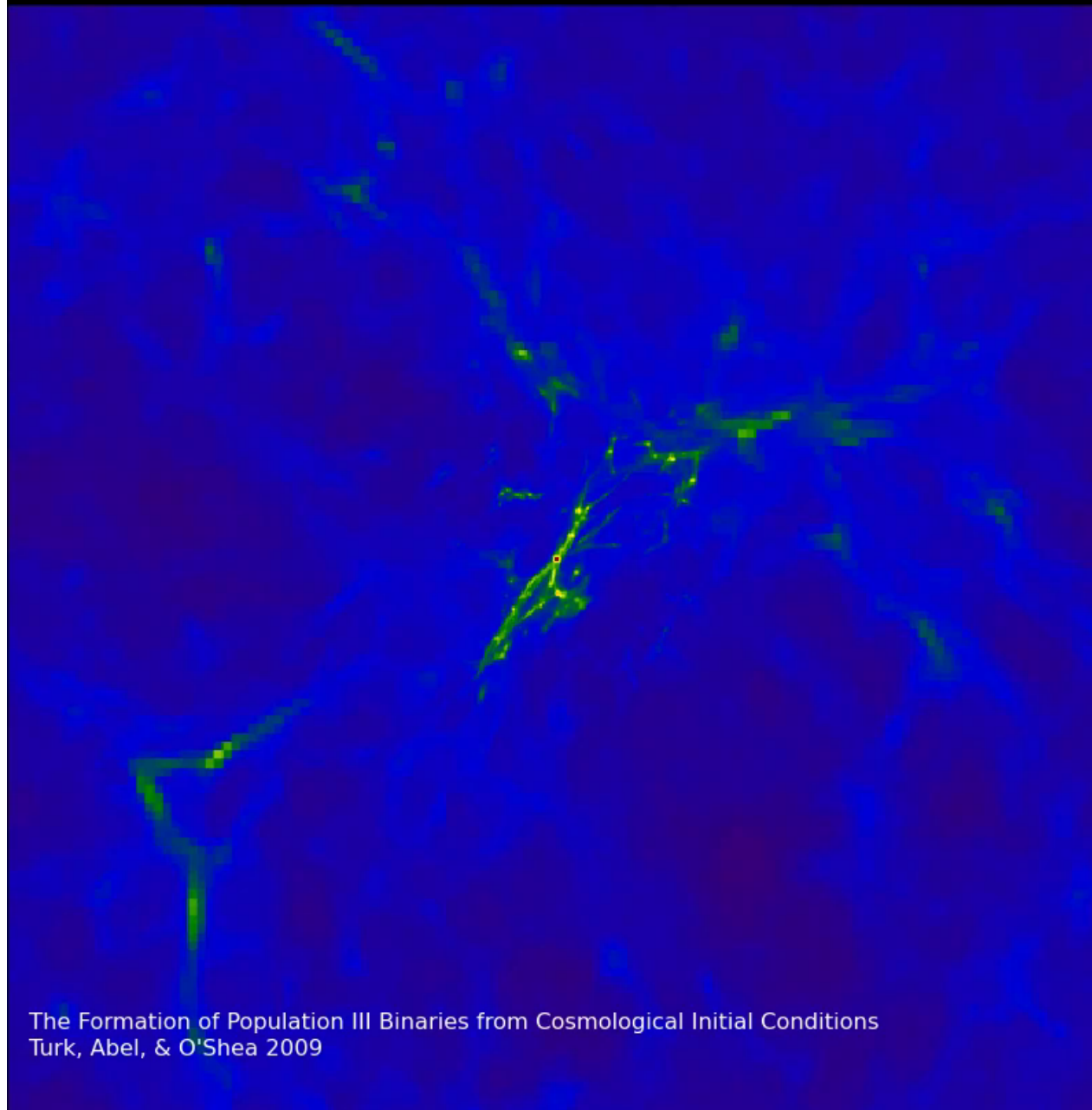
current epoch feedback: cloud cores

Silk & Norman 1980
Li & Nakamura 2008
Abel & Nakamura 2010

Protostellar outflows
(in magnetised clouds)



The first stars: fragmentation



a hybrid model

cold gas flows via filaments/minor mergers lead to disk formation/bulges

at low efficiency

major mergers + hot infall + cooling forms massive spheroids

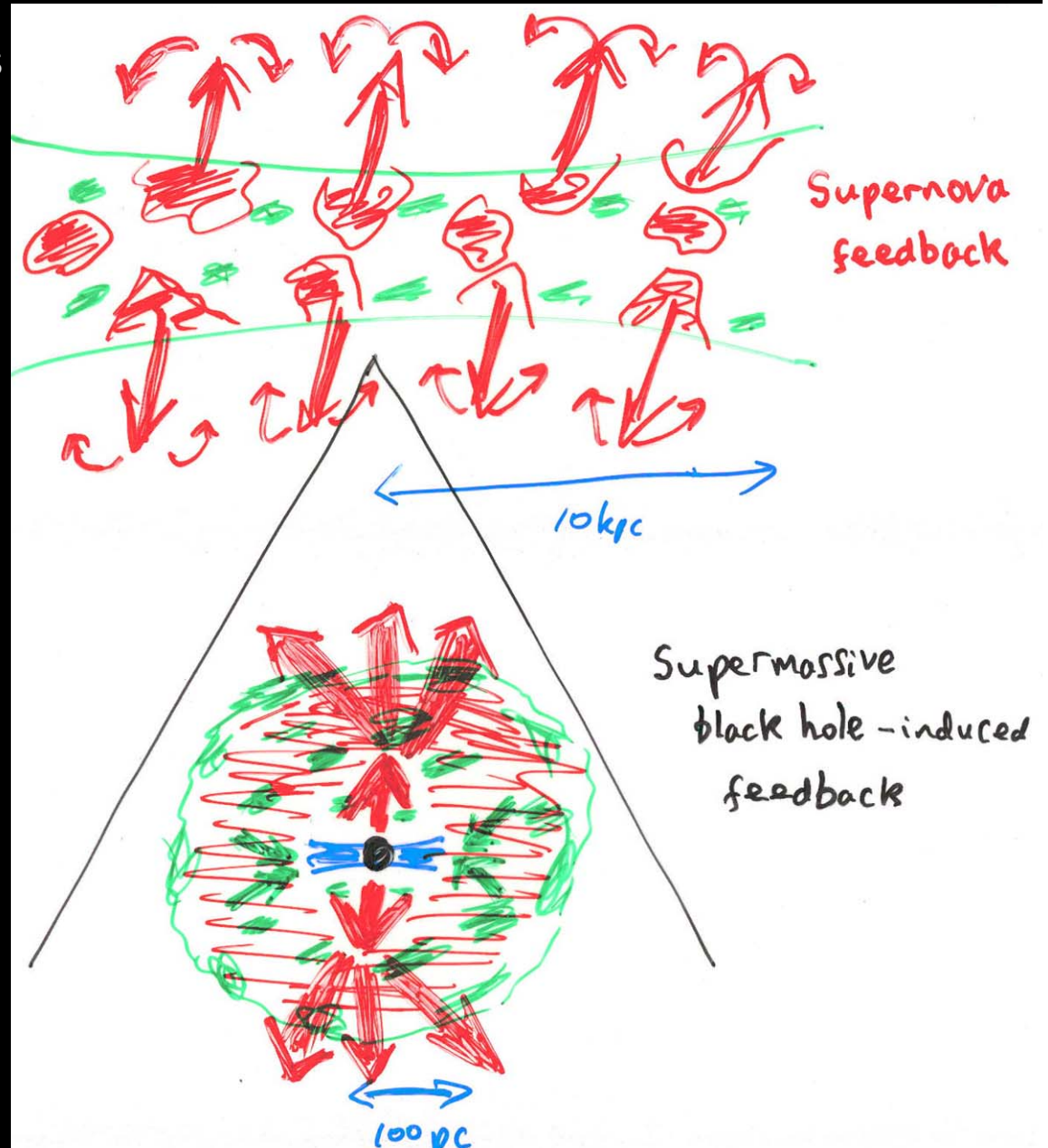
at high efficiency

role of supermassive black holes

quenching of star formation

Intracluster gas heating

triggering of star formation



GALAXY FORMATION CHALLENGES COSMOLOGY

RESURRECTION VIA FUNDAMENTAL PHYSICS

- MODIFYING THE NATURE OF DARK MATTER?
- MODIFYING GRAVITY?

RESURRECTION VIA ASTROPHYSICS

- FEEDBACK via SUPERNOVAE and AGN
- UNDERSTANDING STAR FORMATION

IMPROVED RESOLUTION IN THEORY AND IN OBSERVATION
IS ESSENTIAL FOR CREDIBILITY