



SUSY FCNC at the LHC

Siannah Peñaranda-Rivas

Departamento de Física Teórica
Universidad de Zaragoza

Flavour Physics at LHC run II

Outline

1 Introduction

- Flavour Changing Neutral Currents
- Flavour-changing interactions in SUSY
- R-parity conserving and violating SUSY models

2 Approaches for probing SUSY

3 New Physics Beyond the SM at the LHC

- $B-$ anomalies and LFV (SUSY)

4 Other relevant flavour changing interactions in SUSY

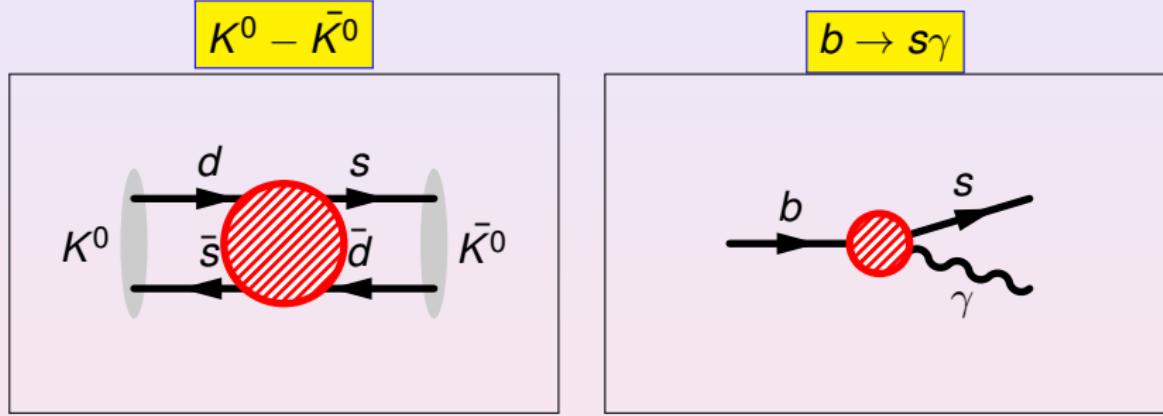
- Direct FCNC production @ LHC
- Higgs FCNC @ LHC

5 Conclusions

Introduction

Flavour Changing Neutral Currents (FCNC)

- FCNC are processes in which one up-type (or down-type) quark is converted into another one of the same type.



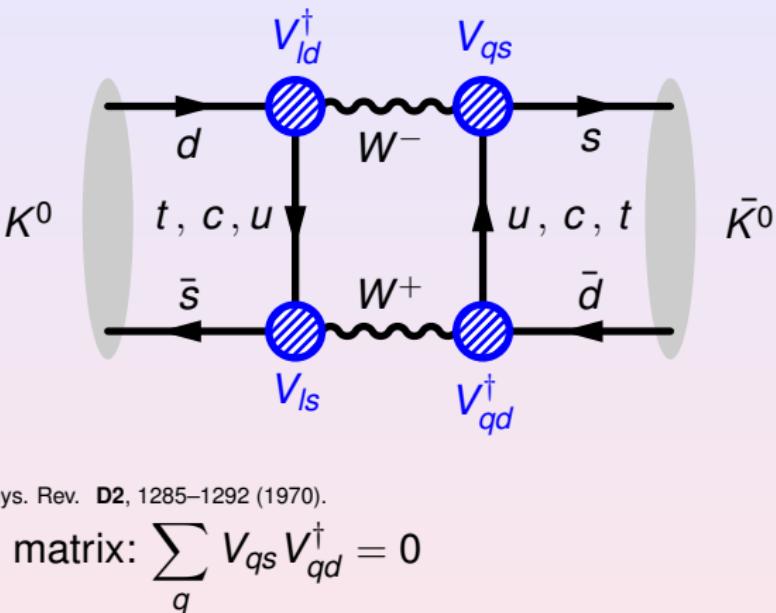
- Effect of mixing: Mass matrix

$$\begin{pmatrix} m_K^2 & A \\ A & m_K^2 \end{pmatrix} \Rightarrow \text{Diagonalize} \Rightarrow m_{K_{1,2}}^2 = m_K^2 \pm A$$

- Mass difference $\Delta m_K \equiv |m_{K_1} - m_{K_2}|$: signal of FCNC

Standard Model

- FCNC absent at the tree-level
- Produced at one-loop by
 - charged currents
 - Cabibbo-Kobayashi-Maskawa mixing matrix (V)
- GIM Mechanism



S. L. Glashow, J. Iliopoulos and L. Maiani, Phys. Rev. **D2**, 1285–1292 (1970).

$$\Rightarrow \text{Unitarity of the CKM matrix: } \sum_q V_{qs} V_{qd}^\dagger = 0$$

- Loop induced \oplus GIM \implies FCNC processes have very small rates

New Physics (NP)

New Physics \rightarrow New FCNC Sources



– Strong constraints from low energy data

- Ideal place to get indirect evidence of NP
- So far, most of experimental results on flavor observables are consistent with SM expectations and lead to strong indirect constraints on NP models
- Increase the sensitivity of flavour experiments & More precision on theoretical determination
- Here we will focus on some SUSY contributions to flavour observables

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New Physics: Supersymmetry

Flavour-changing interactions in SUSY

- Minimal Flavor Violation (MFV) in the MSSM
 - FC phenomena is analogue to the SM case
 - supersymmetrization of the one-loop SM contributions
 - squarks are assumed to be aligned with quarks
 - originates from CKM matrix as the only source and proceeds via loop-contributions
 - the size is expected to be small
- Non Minimal Flavor Violation (NMFV) in the MSSM
 - Additional FC phenomena is due to misalignment between the rotations that diagonalize quark/squark sectors (beyond CKM)
 - arise at tree level
 - sizeable contributions are expected to occur

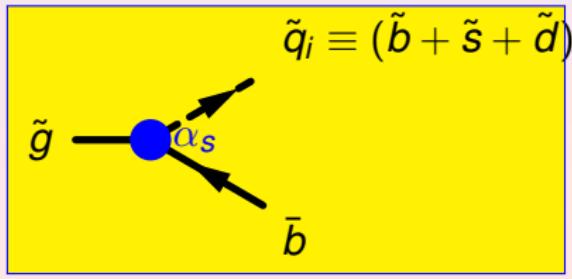
Flavour-changing interactions in SUSY

- Supersymmetry allows for flavour-mixing terms in the scalar-quark mass matrix
- Squark Mixing:
$$\mathcal{M}_{\tilde{q}}^2 = \begin{pmatrix} M_Q^2 + m_q^2 + \cos 2\beta (T_3 - Q_q \sin \theta_W^2) M_Z^2 & m_q(A - \mu \{\cot \beta, \tan \beta\}) \\ m_q(A - \mu \{\cot \beta, \tan \beta\}) & M_{U,D}^2 + m_q^2 + \cos 2\beta Q_q \sin \theta_W^2 M_Z^2 \end{pmatrix}$$
 $M_Q^2, A, M_{U,D}^2$ are 3×3 matrices in generation space
- Flavour Mixing parameters

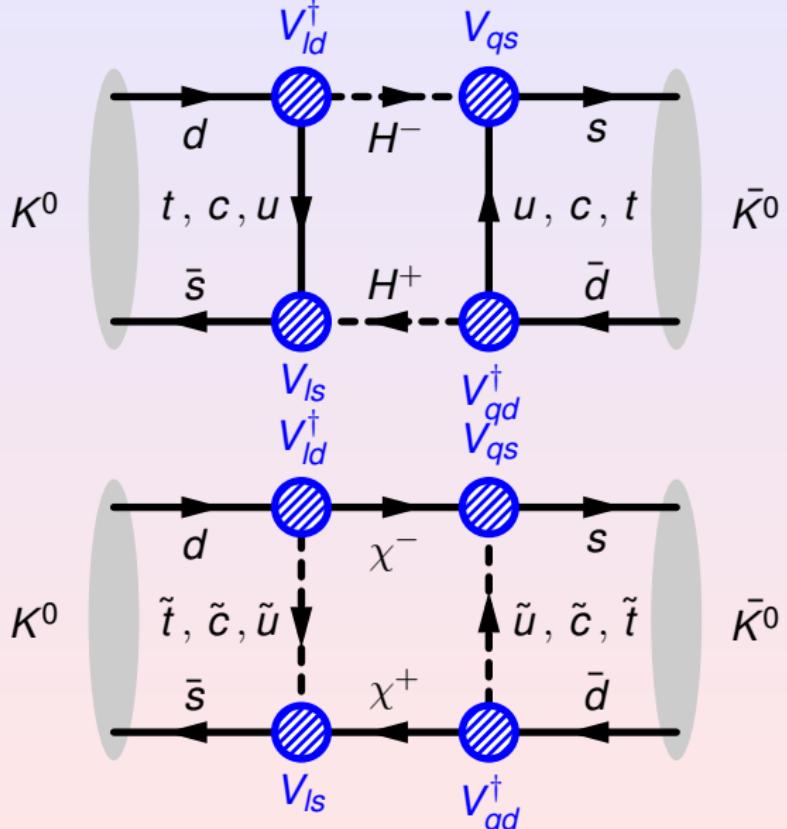
$$\delta_{ij} = \frac{\mathcal{M}_{ij}^2}{\mathcal{M}_{ii}\mathcal{M}_{jj}}$$

⇒ induces strong FCNC interactions!

SUSY-QCD tree-level gluino coupling



→ extra FCNC induced at one-loop and CKM by extra charged particles: H^\pm and χ^\pm



Constraints

- Limits from Low Energy Data

F. Gabbiani *et. al.* Nucl. Phys B 477, 321 (1996)

- $K^0 \bar{K}^0 (d\bar{s} \leftrightarrow \bar{d}s)$: $\Delta m_K \implies \delta_{12}^d$
- $D^0 \bar{D}^0 (c\bar{u} \leftrightarrow \bar{c}u)$: $\Delta m_D \implies \delta_{12}^u$
- $B^0 \bar{B}^0 (b\bar{d} \leftrightarrow \bar{b}d)$: $\Delta m_B \implies \delta_{13}^d$

δ_{12}	\lesssim	$.1 \sqrt{m_{\tilde{u}} m_{\tilde{c}}} / 500 \text{ GeV}$
δ_{13}	\lesssim	$.098 \sqrt{m_{\tilde{u}} m_{\tilde{t}}} / 500 \text{ GeV}$
δ_{23}	\lesssim	$8.2 m_{\tilde{c}} m_{\tilde{t}} / (500 \text{ GeV})^2$

- New B -data from Belle, Babar, LHC: additional constraints
- The Flavour-Changing terms are communicated from the up- to the down-sector by CKM e.g. M.Misiak *et. al.*, hep-ph/9703442

Due to $SU(2)_L$ gauge invariance

$$\begin{array}{ccc} (M_{LL}^d)^2 & = & CKM^\dagger \times (M_{LL}^u)^2 \times CKM \\ \cancel{} & & \cancel{} \\ (M_{LL}^d)_{\text{DIAG}}^2 & & \cancel{1} \tilde{M}^2 \end{array}$$

- ⇒ the bounds are transferred to the up-quark sector
- ⇒ top-charm FCNC are constrained by b -sector measurements.

Flavour-changing interactions in SUSY

- R-parity conserving SUSY models: **MFV** and **NMFV**
 - Provides elegant solutions to the dark matter and hierarchy problems.
 - Leads to natural GUT
- R-parity violating (RPV) SUSY:
 - RPV terms are allowed in the superpotential:

$$W = W_{MSSM} + \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + k_i L_i H_u + \lambda''_{ijk} \bar{U}_i \bar{D}_j \bar{D}_k$$

Lepton number violating

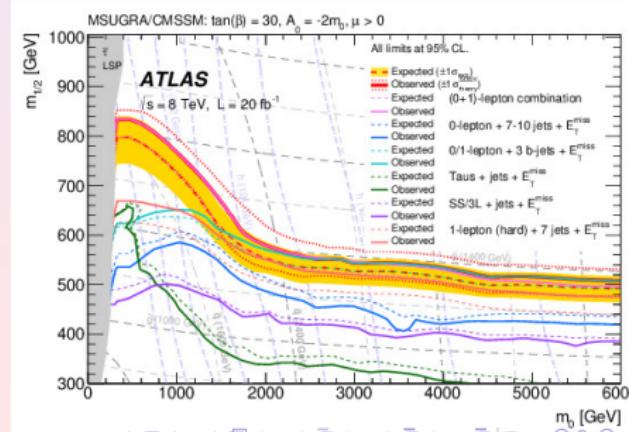
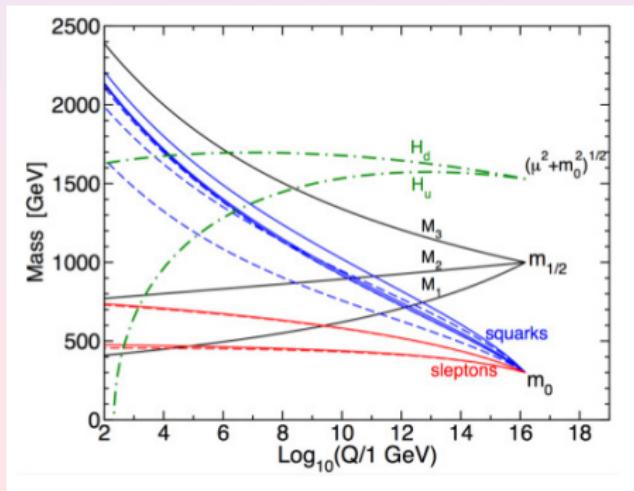
Baryon number viol.

- Resonant/associated single SUSY particle production is possible
- Could explain large mixing angles and hierarchical masses of neutrinos
- The lightest SUSY particle (LSP) is no longer stable
- No dark matter candidate :-(
- Other non-minimal extensions exist, e.g. one extra Higgs (NMSSM), extra $U(1)$ groups (Z'), extra neutrinos (see-saw models), etc.

Approaches for probing SUSY

Most common approaches for probing SUSY:

- **Concrete models:** e.g. mSUGRA/CMSSM, GMSB:
 - easy interpretability as a full theory
 - but, rigid relationships of parameters, not necessarily realistic:
all masses related to few parameters: m_0 , $m_{1/2}$, μ
- **Simplified models:** very reduced, accessible spectrum
 - focus on decay chains to which LHC is sensitive, easier to reinterpret
 - not a full SUSY model



SUSY searches in simplified models

ATLAS SUSY Searches* - 95% CL Lower Limits

May 2017

ATLAS Preliminary

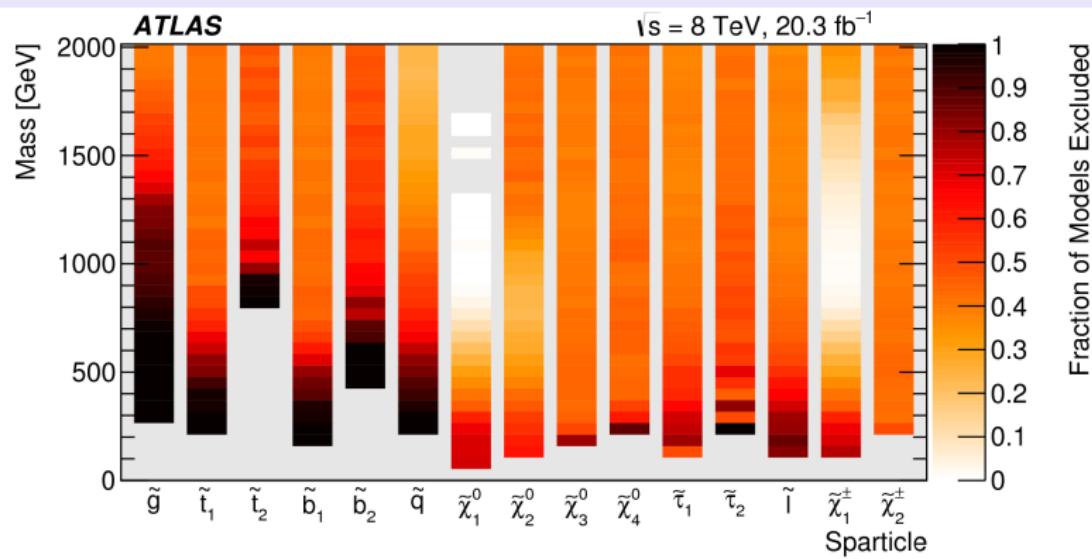
$\sqrt{s} = 7, 8, 13 \text{ TeV}$

Reference

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt (\text{fb}^{-1})$	Mass limit	$\sqrt{s} = 7, 8 \text{ TeV}$	$\sqrt{s} = 13 \text{ TeV}$	Reference
Inclusive Searches	MSUGRA/CMSSM $\tilde{q}\tilde{q}, \tilde{q}\tilde{q}^0$	0-3 e, μ, τ : 2-10 jets; 3 b	Yes	20.3	\tilde{q}, \tilde{q}	1.85 TeV	$m_{\tilde{q}}^2=m_{\tilde{q}^0}$	1507.05525
		0 2-6 jets	Yes	36.1	\emptyset	1.57 TeV	$m_{\tilde{q}}^2=m_{\tilde{q}^0}^2=5 \text{ GeV}$	ATLAS-CONF-2017-022
	$\tilde{q}\tilde{q}, \tilde{q}\tilde{q}^0$ (compressed)	mono-jet	1-3 jets	Yes	3.2	\emptyset	$m_{\tilde{q}}^2>200 \text{ GeV}, m_{\tilde{q}}(1^{\text{st}}, \text{min. d})=m_{\tilde{q}}(2^{\text{nd}}, \text{gen. 4})$	1604.07773
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g}^0$	0 2-6 jets	Yes	36.1	\emptyset	2.02 TeV	$m_{\tilde{g}}^2>200 \text{ GeV}$	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g}^0$ ($\tilde{q}\tilde{q}\tilde{q}\tilde{q}$)	0 2-6 jets	Yes	36.1	\emptyset	2.01 TeV	$m_{\tilde{g}}(1^{\text{st}})>200 \text{ GeV}, m_{\tilde{g}}(2^{\text{nd}})>0.5(m_{\tilde{q}}^2+m_{\tilde{q}^0}^2)$	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g}^0$ ($\ell\ell/\nu\tau\bar{\tau}$)	3 e, μ 4 jets	-	36.1	\emptyset	1.825 TeV	$m_{\tilde{g}}^2>400 \text{ GeV}$	ATLAS-CONF-2017-030
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g}^0$ ($WZ\tilde{Z}$)	0 7-11 jets	Yes	36.1	\emptyset	1.8 TeV	$m_{\tilde{g}}^2>400 \text{ GeV}$	ATLAS-CONF-2017-030
	GMSB (\tilde{f}, \tilde{L} NLSP)	1-2 $e + 0-1 \ell$ 0-2 jets	Yes	3.2	\emptyset	2.0 TeV	$m_{\tilde{f}}^2>100 \text{ GeV}$	1604.09100
	GGM (higgsino-bino NLSP)	2 γ	Yes	3.2	\emptyset	1.85 TeV	$m_{\tilde{N}}^2<1 \text{ mm}$	1507.05683
	GGM (higgsino-bino NLSP)	γ 2 jets	Yes	20.3	\emptyset	1.37 TeV	$m_{\tilde{N}}^2<500 \text{ GeV}, \text{cr}(N_\text{NLSP})<0.1 \text{ mm}, \mu<0$	ATLAS-CONF-2016-066
	GGM (higgsino-bino NLSP)	γ 2 jets	Yes	13.3	\emptyset	1.6 TeV	$m_{\tilde{N}}^2<500 \text{ GeV}, \text{cr}(N_\text{NLSP})<0.1 \text{ mm}, \mu>0$	1503.03290
	Gravitino LSP	2 e, μ (Z)	2 jets	Yes	20.3	\emptyset	$m_{\tilde{g}}^2>1\times 10^{-3} \text{ GeV}, m(g)=m(\tilde{g})=1.5 \text{ TeV}$	1502.01518
3 rd gen. \tilde{b} med.	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0 3 b	Yes	36.1	\emptyset	1.85 TeV	$m_{\tilde{b}}^2<500 \text{ GeV}$	ATLAS-CONF-2017-021
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0-1 e, μ 3 b	Yes	36.1	\emptyset	1.92 TeV	$m_{\tilde{b}}^2<200 \text{ GeV}$	ATLAS-CONF-2017-021
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0-1 e, μ 3 b	Yes	20.1	\emptyset	1.97 TeV	$m_{\tilde{b}}^2<300 \text{ GeV}$	1407.0605
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0	2 b	Yes	36.1	\emptyset	$m_{\tilde{b}}^2>600 \text{ GeV}$	ATLAS-CONF-2017-038
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0-1 e, μ	3 b	Yes	36.1	\emptyset	$m_{\tilde{b}}^2>200 \text{ GeV}$	ATLAS-CONF-2017-030
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{b}}^2>100 \text{ GeV}$	1209.2102, ATLAS-CONF-2016-077
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{b}}^2=2(m_{\tilde{g}}^2), m_{\tilde{b}}^2=55 \text{ GeV}$	1506.08616, ATLAS-CONF-2017-020
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$ (natural GMSB)	0	mono-jet	Yes	3.2	\emptyset	$m_{\tilde{b}}^2>5 \text{ GeV}$	1604.07773
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\emptyset	$m_{\tilde{b}}^2>150 \text{ GeV}$	1403.5222
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$ (natural GMSB)	3 e, μ (Z)	1 b	Yes	36.1	\emptyset	$m_{\tilde{b}}^2>0 \text{ GeV}$	ATLAS-CONF-2017-019
	$\tilde{g}, \tilde{g}, \tilde{b}\tilde{b}^0$ (natural GMSB)	1-2 e, μ	4 b	Yes	36.1	\emptyset	$m_{\tilde{b}}^2>0 \text{ GeV}$	ATLAS-CONF-2017-019
3 rd gen. squares direct production	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0 2 b	Yes	36.1	\emptyset	950 GeV	$m_{\tilde{t}}^2>420 \text{ GeV}$	ATLAS-CONF-2017-038
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	2 e, μ (SS)	1 b	Yes	36.1	\emptyset	$m_{\tilde{t}}^2>200 \text{ GeV}$	ATLAS-CONF-2017-030
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	1 b	Yes	4.7/13.3	\emptyset	$m_{\tilde{t}}^2=2(m_{\tilde{g}}^2), m_{\tilde{t}}^2=55 \text{ GeV}$	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>5 \text{ GeV}$	1506.08616, ATLAS-CONF-2017-020
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>100 \text{ GeV}$	1604.07773
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>200 \text{ GeV}$	1403.5222
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>500 \text{ GeV}$	1507.05493
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>800 \text{ GeV}$	1507.05493
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>1000 \text{ GeV}$	1604.07773
	$\tilde{t}_1, \tilde{t}_1, \tilde{b}_1, \tilde{b}_1$	0-2 e, μ 0-2 jets	2 b	Yes	20.3/36.1	\emptyset	$m_{\tilde{t}}^2>1500 \text{ GeV}$	1405.05620
EW direct	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	96.1	\emptyset	$m_{\tilde{e}}^2>90-440 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	96.1	\emptyset	$m_{\tilde{e}}^2>710 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	96.1	\emptyset	$m_{\tilde{e}}^2>780 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	36.1	\emptyset	$m_{\tilde{e}}^2>117-170 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	90-198	\emptyset	$m_{\tilde{e}}^2>170-275 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	90-198	\emptyset	$m_{\tilde{e}}^2>205-270 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	90-198	\emptyset	$m_{\tilde{e}}^2>275-700 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	90-198	\emptyset	$m_{\tilde{e}}^2>275-700 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	90-198	\emptyset	$m_{\tilde{e}}^2>275-700 \text{ GeV}$	ATLAS-CONF-2017-039
	$\tilde{e}_R, \tilde{e}_R, \tilde{L}_R, \tilde{L}_R$	2 e, μ	0	Yes	90-198	\emptyset	$m_{\tilde{e}}^2>275-700 \text{ GeV}$	ATLAS-CONF-2017-039
Long-lived particles	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	36.1	\emptyset	$m_{\tilde{g}}^2>150 \text{ MeV}, m_{\tilde{g}}^2>0.5(m_{\tilde{g}}^2+m_{\tilde{g}^0}^2)$	ATLAS-CONF-2017-017
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	18.4	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, m_{\tilde{g}}^2>5(m_{\tilde{g}}^2+m_{\tilde{g}^0}^2)$	1506.05332
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	27.9	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1394.075
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	33.2	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1606.05109
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	36.1	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1606.04530
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	36.1	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1411.6795
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	36.1	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1409.5542
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	36.1	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1504.05162
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	36.1	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1504.05162
	$\tilde{g}, \tilde{g}, \tilde{g}, \tilde{g}$	0	1 jet	Yes	20.3	\emptyset	$m_{\tilde{g}}^2>100 \text{ MeV}, \tau>10 \text{ ns}$	1607.08070
RPV	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	0	1 jet	Yes	3.2	\emptyset	$m_{\tilde{e}}^2>-1.1, m_{\tilde{\nu}}^2>0.02$	1404.2500
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	2 e, μ (SS)	0-3 b	Yes	20.3	\emptyset	$m_{\tilde{e}}^2>400 \text{ GeV}, \Delta_{\tilde{e}} k = 1, 2$	ATLAS-CONF-2016-075
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	4 e, μ	-	Yes	13.3	\emptyset	$m_{\tilde{e}}^2>200 \text{ GeV}, \Delta_{\tilde{e}} k = 1, 2$	1405.5084
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	3 e, μ + 1 τ	-	Yes	20.3	\emptyset	$m_{\tilde{e}}^2>200 \text{ GeV}, \Delta_{\tilde{e}} k = 1, 2$	ATLAS-CONF-2016-057
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	0-4 e, μ large- R jets	-	14.8	\emptyset	$m_{\tilde{e}}^2>800 \text{ GeV}$	ATLAS-CONF-2016-057	
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	0-4 e, μ large- R jets	-	14.8	\emptyset	$m_{\tilde{e}}^2>800 \text{ GeV}$	ATLAS-CONF-2016-057	
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	1 e, μ 8-10 jets+0-4 b	-	36.1	\emptyset	$m_{\tilde{e}}^2>800 \text{ GeV}$	ATLAS-CONF-2016-013	
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	1 e, μ 8-10 jets+0-4 b	-	36.1	\emptyset	$m_{\tilde{e}}^2>800 \text{ GeV}$	ATLAS-CONF-2017-013	
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	0	2 jets + 2 b	-	15.4	\emptyset	$m_{\tilde{e}}^2>400 \text{ GeV}$	ATLAS-CONF-2016-022, ATLAS-CONF-2016-084
	$\tilde{e}_R, \tilde{e}_R, \tilde{\nu}_R, \tilde{\nu}_R$	2 e, μ	2 b	-	36.1	\emptyset	$m_{\tilde{e}}^2>400 \text{ GeV}, m_{\tilde{e}}^2>1.7 \text{ TeV}$	ATLAS-CONF-2017-036
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325
	Scalar charm, $\tilde{c} \rightarrow c\tilde{d}^0$	0	2 c	Yes	20.3	\emptyset	$m_{\tilde{c}}^2>200 \text{ GeV}$	1501.01325

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

pMSSM: 19 parameters (R_p -conserving)



MSUGRA	pMSSM
$m_{\tilde{g}} > 1.85 \text{ TeV}$	50% of models with $m_{\tilde{g}} > 1.4 \text{ TeV}$ allowed
$m_{\tilde{q}} > 1.85 \text{ TeV}$	50% of models with $m_{\tilde{q}} > 0.6 \text{ TeV}$ allowed
$m_{\tilde{\chi}_1^+} > 0.71 \text{ TeV}$	50% of models with $m_{\tilde{\chi}_1^+} > 0.1 \text{ TeV}$ allowed

pMSSM scans provide powerful way to reinterpret existing searches in context of full SUSY models

New Physics Beyond the SM at the LHC

There are few anomalies show in data hinting the need of new physics

$$B_s \rightarrow \mu^+ \mu^-$$

$\sim 2\sigma$

$$B \rightarrow D^{(*)} \tau \nu$$

$\sim 4\sigma$

$$B_s \rightarrow K^* l^+ l^- \quad (R_{K^*})$$

$\sim 2.1 - 2.5\sigma$

$$h \rightarrow \mu \tau$$

$\sim 2\sigma$ (Run1) - No excess any more

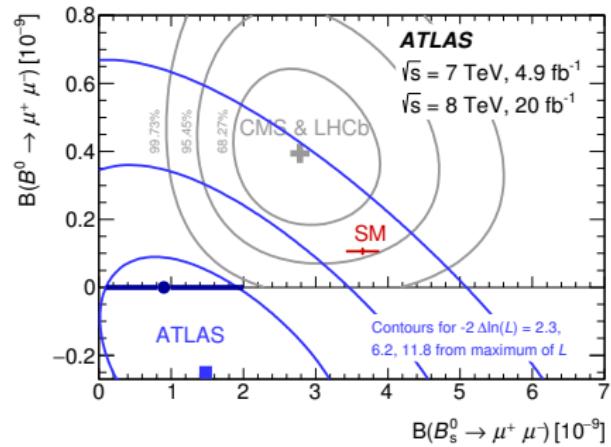
(B anomalies – See talks by Mescia, Martinez-Vidal, Kosnik and Capdevila on Wednesday, May 24)

(LFV – See talk by Fiorini on Monday, May 22)

How SUSY could accomodate these results?

$B_s \rightarrow \mu^+ \mu^-$

ATLAS, Eur. Phys. J. C 76 (2016) 513, arXiv:1604.04263



- ATLAS is consistent with the LHCb and CMS
- ATLAS consistency with SM is 2.0
- Room for NP destructively interfering with the SM

- CMS and LHCb observation:

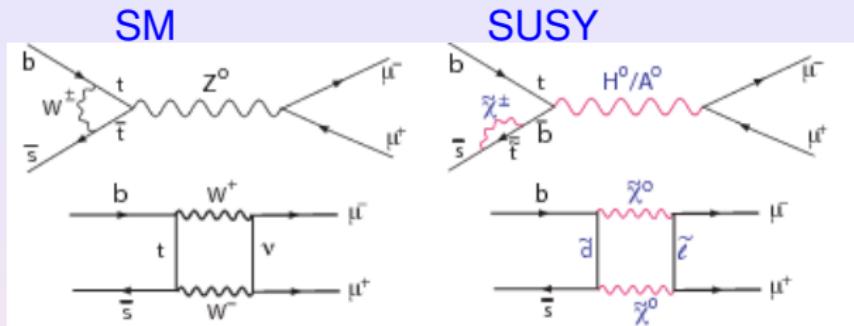
$$BR(B_d^0 \rightarrow \mu^+ \mu^-) = (3.9^{+1.6}_{-1.4}) \times 10^{-10}, BR(B_s^0 \rightarrow \mu^+ \mu^-) = (2.8^{+0.7}_{-0.6}) \times 10^{-9}$$

- ATLAS: $BR(B_d^0 \rightarrow \mu^+ \mu^-) < 4.2 \times 10^{-10}, 95\% C.L.,$

$$BR(B_s^0 \rightarrow \mu^+ \mu^-) < 3 \times 10^{-9}, 95\% C.L.,$$

$$BR(B_s^0 \rightarrow \mu^+ \mu^-) = (0.9^{+1.1}_{-0.8}) \times 10^{-9}$$

Rare decays: $B_s \rightarrow \mu^+ \mu^-$



- **SM prediction accurate:** Bobeth, arXiv:1405.4907

$$BR(B_s^0 \rightarrow \mu^+ \mu^-) = (3.65 \pm 0.23) \times 10^{-9}$$

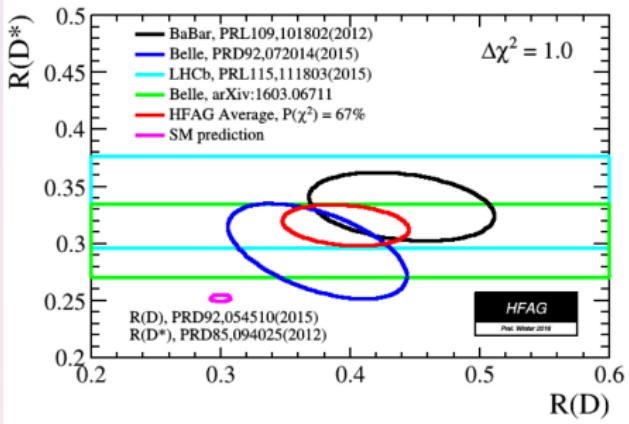
• SUSY

- Enhancement by a factor of order $\tan^6 \beta$
- Implications on the viability of SUSY (constrained and unconstrained models): Arbey, Battaglia, et al.
 - BR in the MSSM does not deviate from its SM prediction
 - LHC results remove 10% of the scan points in the CMSSM and a few % in the pMSSM
- Room for NP (SUSY) opened

$$B \rightarrow D^{(*)}\tau\nu$$

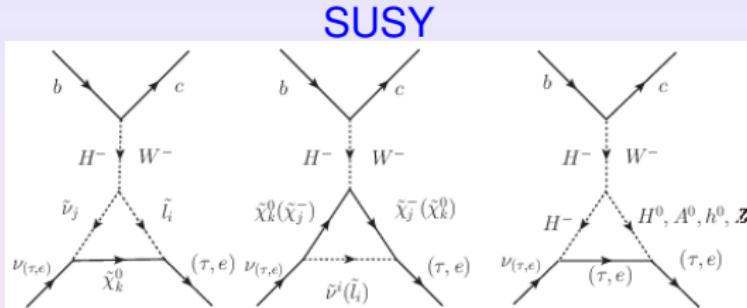
$$R(D) = \frac{BR(\bar{B}^0 \rightarrow D^+ \tau^- \bar{\nu}_\tau)}{BR(\bar{B}^0 \rightarrow D^+ l^- \bar{\nu}_l)}, \quad R(D^*) = \frac{BR(\bar{B}^0 \rightarrow D^{*+} \tau^- \bar{\nu}_\tau)}{BR(\bar{B}^0 \rightarrow D^{*+} l^- \bar{\nu}_l)}$$

PRL 115, 111803 (2015)



- The combined results disagree with the SM expectations at $\sim 4\sigma$
- In the SM, the only difference between the numerator and the denominator is the lepton mass
- Sensitive test to NP at tree level
- Inconsistent with Type II THDM ...
- SUSY has the potential to explain recent data

Boubaa et al, 1604.0341

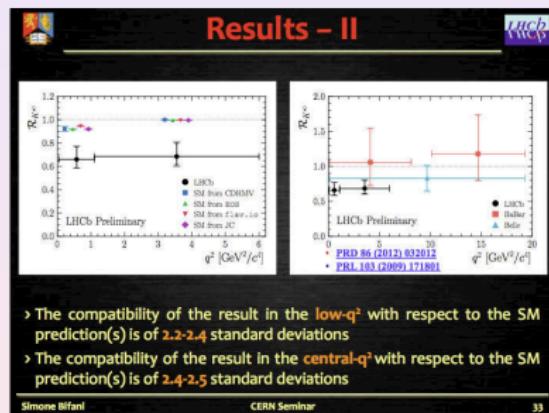


What needed to solve the anomaly? (Xiao-Gang He, SUSY 2016)

- Exp: More precise measurements!
- Theor: New Physics modify charged current interaction... in a way that
 - The first two and third generations interact differently;
 - Have P-parity conserving and violating ones differently!
THDM-II cannot explain both
SUSY OK (clear signal), Boubaa et al, 1604.0341.
- This is a striking hint of violation of the lepton flavour universality which clearly needs also to be checked in other modes.

$B \rightarrow K^* l^+ l^-$ (R_{K^*})

- Experimental: At present no serious problems at colliders
Deviations $2 - 3\sigma$ appear from time to time:
- LHCb: CERN seminar 18/04/2017:



$$\mathcal{R}_{K^{*0}} = \frac{\mathcal{B}(B^0 \rightarrow K^{*0} \mu^+ \mu^-)}{\mathcal{B}(B^0 \rightarrow K^{*0} e^+ e^-)} =$$

$$\begin{cases} 0.660^{+0.110}_{-0.070} \pm 0.024 & \text{low } q^2 \\ 0.685^{+0.113}_{-0.069} \pm 0.047 & \text{central } q^2 \end{cases}$$

- SM prediction: 1
(Lepton Flavour Universality)
- Papers explanation:
NP: Z' , leptoquarks

- ? SUSY with RPV:

R_p -violation superpotential: $W = \dots + \lambda'_{ijk} L_i Q_j \bar{D}_k + \dots$

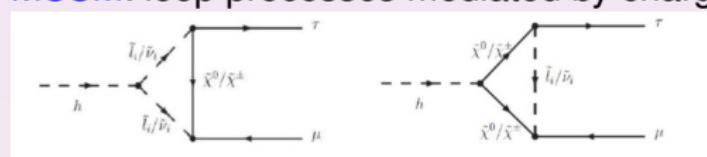
???... sfermions behave as leptoquarks... (Work in progress)

LFV Higgs decay $h \rightarrow \mu\tau$?

- The CMS and ATLAS collaborations reported the first signal of LFV Higgs $h \rightarrow \mu\tau$ (Run1)

$$BR(h \rightarrow \mu\tau) = 8.4_{-3.7}^{+3.9} \times 10^{-3} \text{ (CMS)}$$
$$BR(h \rightarrow \mu\tau) = (7.7 \pm 6.2) \times 10^{-3} \text{ (ATLAS)}$$

- The SM predicts no tree-level LFV Higgs coupling
- MSSM:** loop processes mediated by charginos or neutralinos.



- However, SUSY models with non-zero family mixing in the sleptons also result in enhancement in other LFV processes such as $\mu \rightarrow e\gamma$, $\tau \rightarrow e\gamma$, and $\tau \rightarrow \mu\gamma$ \Rightarrow Important correlation between observables
- Run2: The limits can be used to constrain the corresponding flavour violating Yukawa couplings, absent in the standard model.

Model Setup

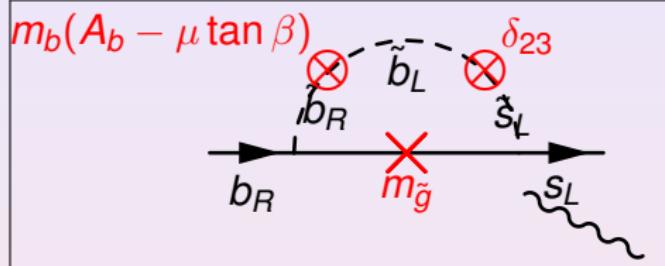
- Generic Minimal Supersymmetric Standard Model
- Scan over the parameters space:
 - ⇒ works done some years ago
 - ⇒ similar to present pMSSM scenarios
 - ⇒ updated
- Flavour violation terms only in the Left-Left sector
 - ⇒ Naturally generated by Renormalization Group Equations
 - ⇒ (results similar with Right-Right and Left-Right mixing)

$B(b \rightarrow s\gamma)$

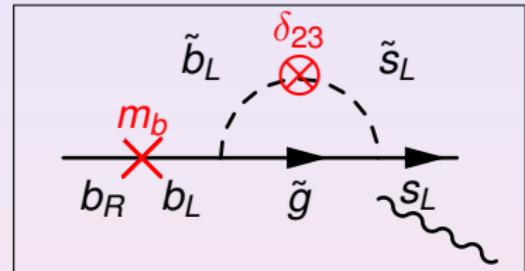
Borzumati **ZPC63** (1994) 291, hep-ph/9310212; Borzumati, Greub, Hurt, Wyler, **PRD62** (2000) 075005, hep-ph/9911245.

[True for Left-Left mixing only!]

- FCNC SUSY-QCD contribution:



Leading, Double insertion



Sub-Leading, Single insertion

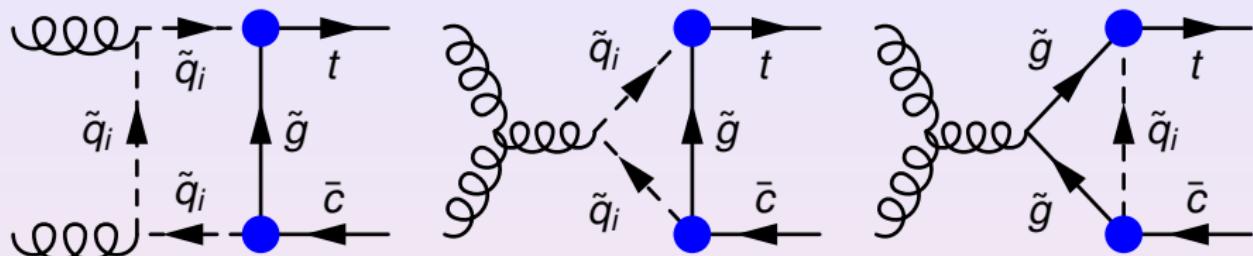
- The Feynman Amplitude:

$$A^{\text{SUSY-QCD}}(b \rightarrow s\gamma) \sim \delta_{23} \frac{m_b(A_b - \mu \tan \beta)}{M_{\text{SUSY}}^2} \times \frac{1}{m_{\tilde{g}}}$$

- Similar coupling structure in $pp \rightarrow t\bar{c}$ ($\sim m_t(A_t - \mu/\tan \beta)$) but different in Hqq' ($\sim \mu$)
- Relevant interplay between observables

Direct FCNC production @ LHC

$pp[gg] \rightarrow tc$



- Some previous works

J.J. Lui *et al.*, **Nucl. Phys. B** 705 (2005) 3, hep-ph/0404099

G. Eilam, M. Frank, I. Turan, **Phys. Rev. D74** (2006) 035012, hep-ph/0601253

J. Guasch, W. Hollik, S.P., J. Sola, **Nucl. Phys. Proc. Suppl.** 157 (2006) 152, hep-ph/0601218

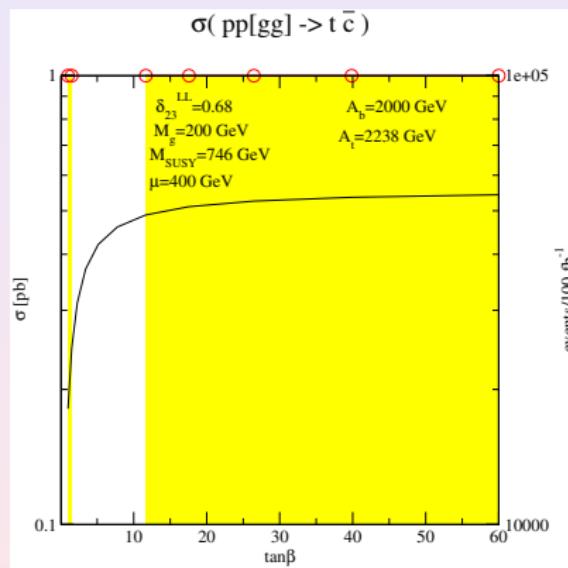
- Leading terms from Left-Left sector: similar structure to $b \rightarrow s\gamma$

$$A(gg \rightarrow t\bar{c}) \sim \delta_{23} \frac{m_t(A_t - \mu/\tan\beta)}{M_{SUSY}^2} \times \frac{1}{m_{\tilde{g}}}$$

- Large rates \Rightarrow Large δ_{23} and Large $(A_t - \mu/\tan\beta)$
 - high sensitivity to A_t

$$pp[gg] \rightarrow t\bar{c}$$

$$\sigma(pp[gg] \rightarrow t\bar{c}) \sim (\delta_{23})^2 \frac{m_t^2(A_t - \mu/\tan\beta)^2}{M_{SUSY}^4} \frac{1}{m_{\tilde{g}}^2}$$



- For small $\tan\beta$ there are no restrictions from $b \rightarrow s\gamma$ and σ increases as $\sim A_t^2$
 - $\Rightarrow \sigma^{max}(pp[gg] \rightarrow t\bar{c} + \bar{t}c) \simeq 1 \text{ pb}$ and $\sim 10^5$ events for 100 fb^{-1}
- Cross-section decays significantly with M_{SUSY} and very fast with $M_{\tilde{g}}$
 - For $M_{\tilde{g}} \sim 500 \text{ GeV}$:
 $\sigma^{max}(pp[gg] \rightarrow tc) \simeq 0.04 \text{ pb}$
 - Cross-sections $\sim 0.5 \text{ pb}$ possible
 - $\sim 100,000$ events/ 100 fb^{-1} for $t\bar{c} + \bar{t}c$ processes

$$BR^{exp}(b \rightarrow s\gamma) \sim (2.1 - 4.5) \times 10^{-4} \text{ (within } 3\sigma)$$

$$\sigma^{SM}(pp[gg] \rightarrow t\bar{c}) \sim 3.6 \times 10^{-7} \text{ pb} \rightarrow 6 \text{ orders of magnitude larger than SM!}$$

Comparison with Higgs FCNC

- Take parameters of maximum $\sigma(pp \rightarrow h \rightarrow t\bar{c})$: Large M_{SUSY} and $m_{\tilde{g}}$

$$M_{SUSY} \simeq m_{\tilde{g}} \simeq 880 \text{ GeV}, \mu \simeq -700 \text{ GeV}, \delta_{23} \simeq 10^{-0.1} \simeq 0.79$$

$$\sigma(pp \rightarrow H^0 \rightarrow t\bar{c} + \bar{t}c) \simeq 2.5 \times 10^{-3} \text{ pb } [\tan \beta = 5]$$

$$\sigma(pp[gg] \rightarrow t\bar{c}) \simeq 1.8 \times 10^{-3} \text{ pb}$$

⇒ Same order of magnitude as Higgs-mediated FCNC !?

Higgs FCNC @ LHC

$$\begin{aligned}\sigma(pp \rightarrow h \rightarrow q\bar{q}') &\equiv \sigma(pp \rightarrow hX)B(h \rightarrow q\bar{q}') \\ &\equiv \sigma(pp \rightarrow hX)\frac{\Gamma(h \rightarrow q\bar{q}' + \bar{q}'q')}{\sum_i \Gamma(h \rightarrow X_i)} \quad (q\bar{q}' \equiv bs \text{ or } tc).\end{aligned}$$

- Computation:

- $\sigma(pp \rightarrow hX)$: HIGLU and HQQ packages

M. Spira, hep-ph/9510347; <http://people.web.psi.ch/~spira/higlu/>, and ...~spira/hqq/.

- $\Gamma(h \rightarrow X)$: FCNC FCHDECAY

S. Béjar, J. Guasch; <http://fchdecay.googlepages.com>

- $\Gamma(h \rightarrow q\bar{q}')$: SUSY-QCD contributions

- Don't assume alignment
- Exact diagonalization of 6×6 squark mass matrix
- Assume mixing only in the LL sector

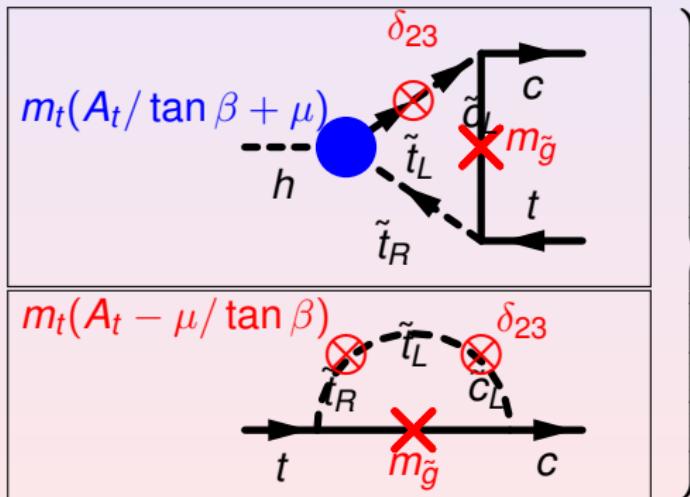
- SM values:

$$\begin{array}{lll} BR(H^{SM} \rightarrow b\bar{s}) & \lesssim & 10^{-7} \quad (m_H < 2M_W) \\ & \lesssim & 10^{-10} \quad (m_H > m_t) \end{array} \quad \Big| \quad \begin{array}{lll} BR(H^{SM} \rightarrow t\bar{c}) & \lesssim & 10^{-13} \end{array}$$

Leading contributions

- Diagrams with a chirality flip are enhanced by $m_{\tilde{g}}$: mass-insertion approximation

$H \rightarrow tc$



- The terms proportional to A_t cancel in the sum.
J.Guasch, P.Häfliger, M.Spira
PRD68 115001, 2003, hep-ph/0305101
- Equivalent structure for bs -channel

Leading contributions

- We can write an effective Lagrangian:

$$G_{Hqq'} \sim \delta_{23} \frac{m_{\tilde{g}} \mu}{M_{SUSY}^2} \left\{ \begin{array}{lll} \cos(\beta - \alpha_{\text{eff}}) & (h^0) & M_{A^0} \gg M_Z \\ \sin(\beta - \alpha_{\text{eff}}) & (H^0) & \xrightarrow{\quad} \xrightarrow{\quad} \xrightarrow{\quad} \xrightarrow{\quad} \\ 1 & (A^0) & \alpha_{\text{eff}} \rightarrow \beta - \pi/2 \end{array} \right. \begin{array}{lll} 0 & (h^0) \\ 1 & (H^0) \\ 1 & (A^0) \end{array} \right\}$$

D. A. Demir, Phys. Lett. **B571**, 193–208 (2003), hep-ph/0303249.

- Different coupling structure in Hqq' ($\sim \mu$) and $bs\gamma$ ($\sim A_b - \mu \tan \beta$)
⇒ Possibility of small contribution to $A(b \rightarrow s\gamma)$ and large contribution to $BR(H \rightarrow qq')$
- Numerical results $BR(h \rightarrow qq')$

Find the maximum BR : MSSM parameter space scan:

$$BR(h \rightarrow bs) \lesssim 10^{-3}$$

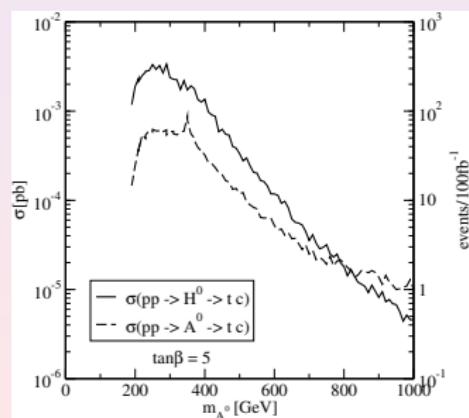
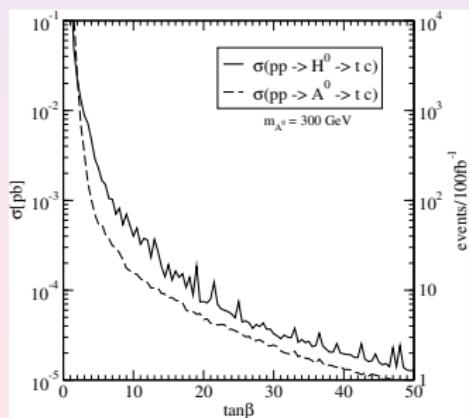
$$BR(h \rightarrow tc) \sim 10^{-3}$$

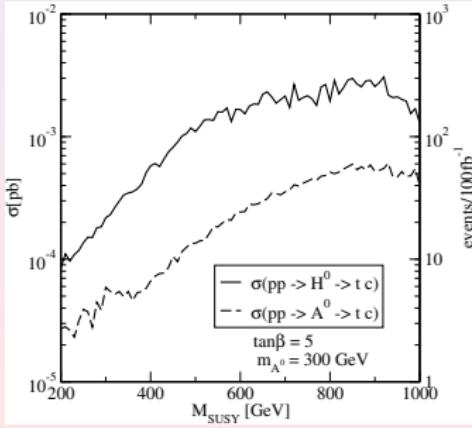
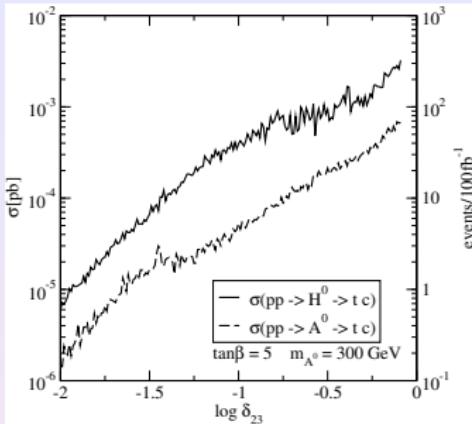
⇒ several orders of magnitude larger than in the SM!!

Combination with production

Combined analysis $\sigma(pp \rightarrow h \rightarrow tc)$

- Only H^0/A^0 possible
- Large at small $\tan\beta$
- differences at small M_{A^0} :
 - Near threshold for $H^0 \rightarrow \tilde{q}_1 \tilde{q}_1$
 - not possible for A^0





The best situation

- Maximum at maximal δ_{23}
- Maximum at maximal M_{SUSY}

$$M_{A^0} = 300 \text{ GeV}, \tan\beta = 5$$

h	H^0	A^0
$\sigma(pp \rightarrow h \rightarrow tc)$	2.4×10^{-3} pb	5.8×10^{-4} pb
events/100 fb $^{-1}$	240	58
$B(h \rightarrow tc)$	1.9×10^{-3}	5.7×10^{-4}
$\Gamma(h \rightarrow X)$	0.41 GeV	0.39 GeV
δ_{23}	0.79	0.83
$m_{\tilde{q}}$	880 GeV	850 GeV
A_t	-2590 GeV	2410 GeV
μ	-700 GeV	-930 GeV
$B(b \rightarrow s\gamma)$	4.13×10^{-4}	4.47×10^{-4}

$\tan\beta$	4	3	2
$\sigma(pp \rightarrow H^0 \rightarrow tc)$	5 fb	9 fb	20 fb
events/100 fb $^{-1}$	500	900	2000

- ⇒ increases fast at low $\tan\beta$
- ⇒ several thousand events could be produced

Comparison with direct FCNC production

- Not taking parameters of maximum $\sigma(pp \rightarrow h \rightarrow t\bar{c})$: small $m_{\tilde{g}}$

$$m_{\tilde{g}} \simeq 200 \text{ GeV}, M_{\text{SUSY}} \simeq 800 \text{ GeV}, |\mu| \simeq 700 \text{ GeV}, \delta_{23} \simeq 0.7$$

$$\sigma(pp \rightarrow H^0 \rightarrow t\bar{c} + \bar{t}c) \simeq 10^{-3} \text{ pb } [\tan \beta = 5]$$

$$\sigma(pp[gg] \rightarrow t\bar{c}) \simeq 0.5 \text{ pb}$$

⇒ 2-3 orders of magnitude larger than Higgs-mediated FCNC

SUSY FCNC at the LHC

- Effects at LHC: $\sigma(pp[gg] \rightarrow t\bar{c})$, $\sigma(pp \rightarrow h \rightarrow t\bar{c})$
- Direct production is competitive to Higgs-mediated processes
⇒ Direct process can give much larger rates

Parameter	Higgs-mediated	Direct production
$\tan \beta$	Decreases fast	insensitive
M_{A^0}	Decreases fast	insensitive
M_{SUSY}	Prefers large	Decreases fast
A_t	insensitive	very sensitive
δ_{23}	Moderate	Moderate

- Left-Left flavour mixing gives large rates
- Experimental issues:
 - Signal: single top-quark + light c -jet
⇒ Evidence of new physics

CONCLUSIONS

FCNC processes can be a helpful signature of SUSY physics at the LHC

- FCNCs are part of SUSY
- Constrained by low energy data
- Constrained by B-anomalies and LFV at the LHC
 - The results remove ONLY between 10% – 2% of the scan points in general SUSY models
- Room for NP (SUSY) opened
- Run II data expected to increase precisions ...

Some works

- $H \rightarrow bs, H \rightarrow tc$ MSSM

A. M. Curiel, M. J. Herrero, W. Hollik, F. Merz, S. P., Phys. Rev. D **69** (2004) 075009 [hep-ph/0312135].

A. M. Curiel, M. J. Herrero, D. Temes, Phys. Rev. D **67** (2003) 075008 [hep-ph/0210335].

- $H \rightarrow bs + b \rightarrow s\gamma$

S. Béjar, F. Dilmé, J. Guasch and J. Solà, JHEP **0408** (2004) 018 [hep-ph/0402188].

T. Hahn, W. Hollik, J.I. Illana, S. P., hep-ph/0512315.

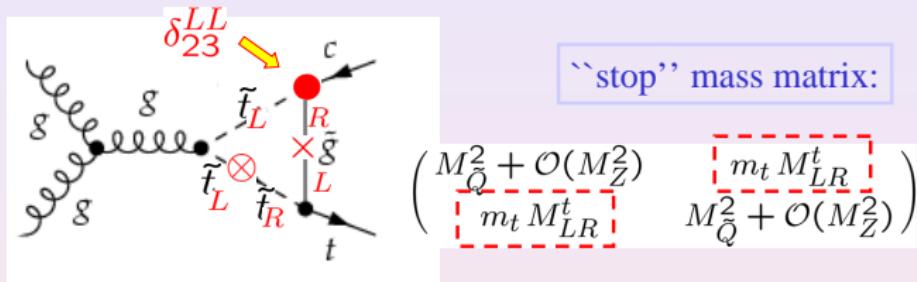
- $pp \rightarrow H + H \rightarrow bs + H \rightarrow tc + b \rightarrow s\gamma$

S. Béjar, J. Guasch, J. Solà, JHEP **0510** (2005) 113, hep-ph/0508043

- ...

Leading terms

Typical behavior of the cross-section



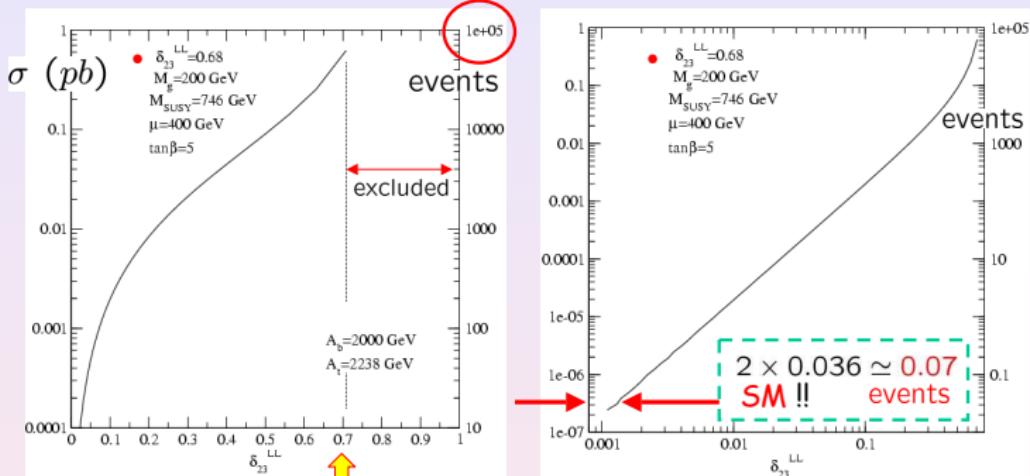
Amplitude:

$$(M_{LR}^t = A_t - \mu/\tan\beta)$$

$$A(pp[gg] \rightarrow t\bar{c}) \sim \delta_{23}^{LL} \times \frac{m_t(A_t - \mu/\tan\beta)}{M_{SUSY}^2} \times \frac{1}{m_{\tilde{g}}}$$

$$(\sigma \sim |A(pp[gg] \rightarrow t\bar{c})|^2)$$

$$\sigma(pp[gg] \rightarrow t\bar{c}) \sim \left(\delta_{23}^{LL}\right)^2 \frac{m_t^2(A_t - \mu/\tan\beta)^2}{M_{SUSY}^4} \frac{1}{M_g^2}$$



$$\delta_{23}^{LL} \simeq 0.7$$

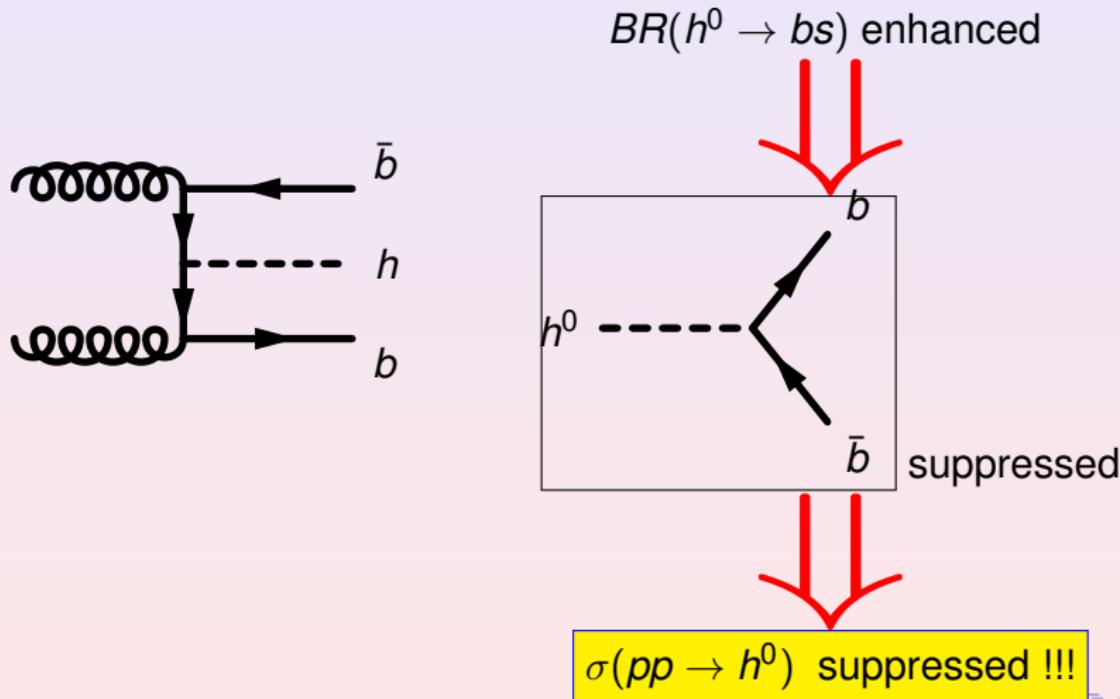
(Not even a single event in the SM
for the entire lifetime of the LHC !!)

$$(\sigma_{SM} = 3.6 \times 10^{-7} \text{ pb})$$

⇒ only the presence of new physics could be an explanation for these events, if they are ever detected

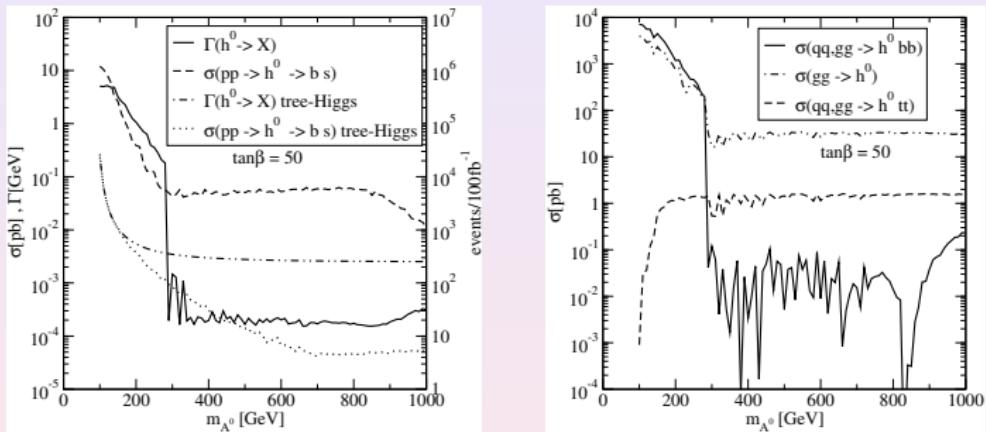
Combination with production

- At large $\tan \beta$ the main production channel for h^0 is associated production: $\sigma(pp \rightarrow h^0 b\bar{b})$



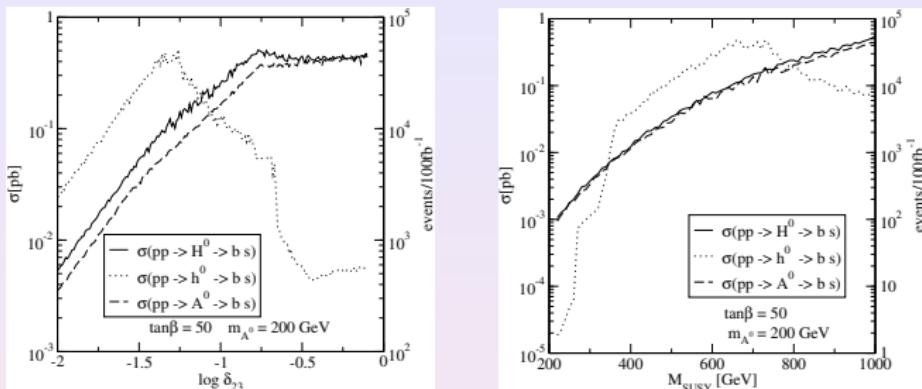
Combined analysis $\sigma(pp \rightarrow h \rightarrow qq')$

$$\sigma(pp \rightarrow h \rightarrow bs)$$



- Maximized production rates for h^0
- $M_{A^0} < 300$ GeV: enhancement of $\sigma(pp \rightarrow h^0)$ dominates
- $M_{A^0} > 300$ GeV: suppression of $\Gamma(h^0 \rightarrow X)$ dominates

Combined analysis $\sigma(pp \rightarrow h \rightarrow bs)$



- h^0 : Maximum attained for small δ_{23} , $M_{\text{SUSY}} \sim 700 \text{ GeV}$
 - ⇒ Larger $\delta_{23} \Rightarrow$ smaller μ ($b \rightarrow s\gamma$)
 - ⇒ Small $M_{\text{SUSY}} \Rightarrow$ small δ_{23} ($b \rightarrow s\gamma$)
- H^0/A^0 : Maximum at large M_{SUSY}
 - ⇒ Large $M_{\text{SUSY}} \implies$ small $B(b \rightarrow s\gamma) \implies$ larger δ_{23} allowed
 - Large $\delta_{23} \Rightarrow \mu$ has to decrease to obtain acceptable $B(b \rightarrow s\gamma) \Rightarrow BR(H^0/A^0 \rightarrow bs)$ can not grow.
- Maximum values: $\sigma(pp \rightarrow h \rightarrow bs) \sim 0.4 \text{ pb}$ and 10^4 events/100fb $^{-1}$

Finding the maximum

- Define: $\delta_{33}^{LR} = \frac{m_t(A_t - \mu/\tan\beta)}{M_{SUSY}^2}$
- $\sigma = (\delta_{23})^2(\delta_{33}^{LR})^2 = \text{constant}$ defines an hyperbola in the $\delta_{23} - \delta_{33}^{LR}$ plane
- Mass (approximation):

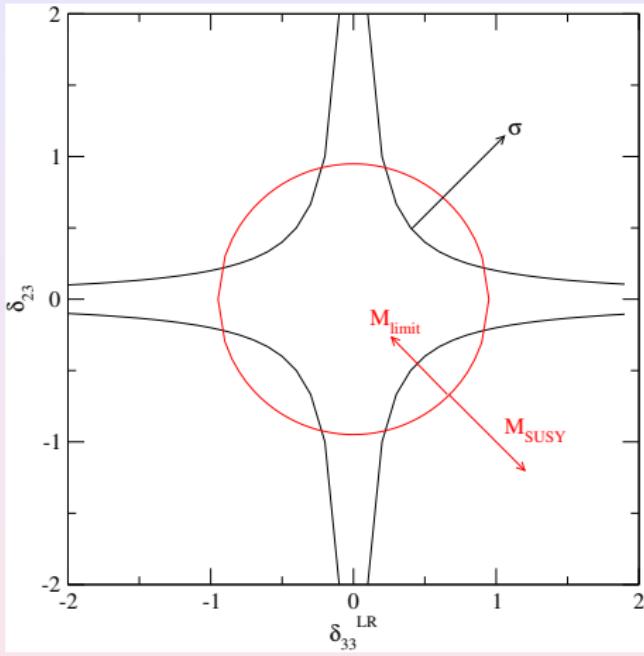
$$m_{\tilde{q}}^2 = M_{SUSY}^2 \begin{pmatrix} & c_L & t_L & t_R \\ \hline c_L & 1 & \delta_{23} & 0 \\ t_L & \delta_{23} & 1 & \delta_{33}^{LR} \\ t_R & 0 & \delta_{33}^{LR} & 1 \end{pmatrix}$$

- lightest mass:

$$m_q^2 = M_{SUSY}^2 \left(1 - \sqrt{(\delta_{23})^2 + (\delta_{33}^{LR})^2} \right) > M_{\text{limit}}^2 .$$

Experimental limit defines a circle in $\delta_{23} - \delta_{33}^{LR}$ plane:

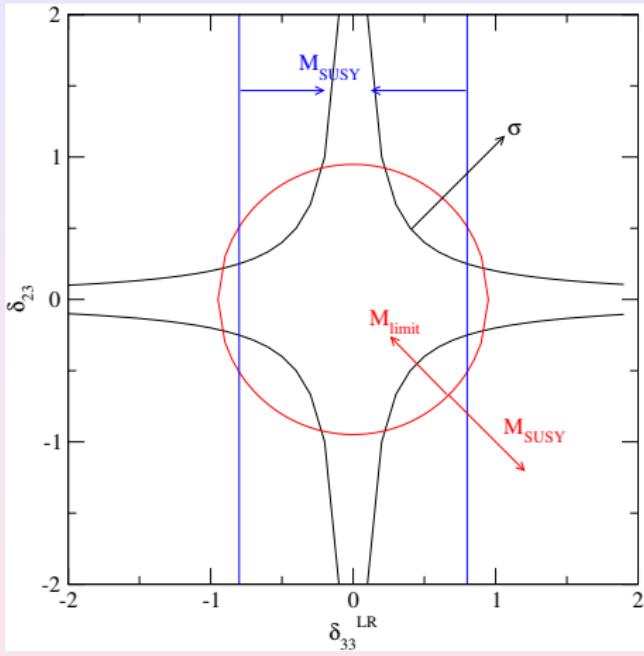
$$(\delta_{23})^2 + (\delta_{33}^{LR})^2 < \left(1 - \frac{M_I^2}{M_{SUSY}^2} \right)^2 \equiv R^2$$



Maximum at:

$$\delta_{23} = \delta_{33}^{LR} = \frac{R}{\sqrt{2}} = \frac{1}{\sqrt{2}} \left(1 - \frac{M_I^2}{M_{SUSY}^2} \right) \rightarrow \frac{1}{\sqrt{2}} .$$

$$M_{SUSY} \rightarrow \infty$$



non-colour breaking vacua:
 $|A_t| \sim < 3M_{SUSY}$

$$|\delta_{33}^{LR}| < \sim \frac{3m_t}{M_{SUSY}}$$

The maximum is obtained when:

- the diagonal: $\delta_{23} = \delta_{33}^{LR}$
- the limit mass circle
- the limit from A_t

cross in a single point

Exact equations

$$\delta_{23} = \delta_{33}^{LR}$$

$$\delta_{33}^{LR} = \frac{m_t (3M_{SUSY} - \mu / \tan \beta)}{M_{SUSY}^2}$$

$$(\delta_{23})^2 + (\delta_{33}^{LR})^2 = \left(1 - \frac{M_I^2}{M_{SUSY}^2}\right)^2$$

setting:

$m_t = 175 \text{ GeV}$, $M_I = 150 \text{ GeV}$, $\mu = 400 \text{ GeV}$, $\tan \beta = 5$ (by $b \rightarrow s\gamma$)

$$\delta_{33}^{LR} = \delta_{23} = 0.678525$$

$$M_{SUSY} = 746.082 \text{ GeV}$$

$$A_t = 2238.25 \text{ GeV}$$