

## THE NEUTRINO MASS PROBLEM

- Neutrinos  $\rightarrow$  Unique Probe  
 $\rightarrow$  Intrinsic Properties
- SM choice of Chiral Fields
- Without  $\nu_R \dots$ 
  - SM  $\Rightarrow m_\nu = 0$
  - dim. 5 WINDOW TO BSM
- With  $\nu_R \dots$ 
  - Both  $m_D$  and  $m_R$  Masses
  - Global  $U(1) \leftrightarrow m_R = 0$  ?
- Facts ... and Opinions

## Neutrinos as a Unique Probe: $10^{-33} - 10^{+28}$ cm

### • Particle Physics

- $\nu N, \mu N, eN$  scattering: existence/ properties of quarks, QCD
- Weak decays ( $n \rightarrow p e^- \bar{\nu}_e, \mu^- \rightarrow e^- \nu_\mu \bar{\nu}_e$ ): Fermi theory, parity violation, mixing
- Neutral current, Z-pole, atomic parity: electroweak unification, field theory,  $m_t$ ; severe constraint on physics to TeV scale
- Neutrino mass: constraint on TeV physics, grand unification, superstrings

### • Astrophysics/Cosmology

- Core of Sun
- Supernova dynamics
- Atmospheric neutrinos (cosmic rays)
- AGNs, cosmic rays, violent events, GRBs
- Large scale structure (dark matter)
- Nucleosynthesis (big bang - small  $A$ ; stellar - to iron; supernova - large  $N$ )
- Baryogenesis
- Simultaneous probes of  $\nu$  and astrophysics

② There was a time ...

1957 → Two-component  $\nu$  theory

↓      ↓      ↓

Only  $\nu_L$  in weak interactions



Argument in favour for  $m_\nu = 0$

1958 → (V-A) theory  $\Rightarrow$  CHIRAL fields enter  
for ALL fermions

Nothing special for  $\nu$ 's  $\leftrightarrow \nu_L, e_L, N_L$

No reason why  $m_\nu = 0$

BUT : -  $e_R, q_R$  needed }  $\Rightarrow$  Dirac Mass  $\neq 0$   
QED, QCD }

-  $\nu_R$ ? at will  $\leftrightarrow$  Sterile

1967 → Standard Model "choice":  $\nu_R'$   
Dirac Mass = 0

{ No  $SU(2) \times U(1)$  gauge invariant term  
for Majorana Mass

↓      ↓

$m_\nu = 0$

⊕ Lepton-Flavour Conservation }

BUT : - SM with total symmetry (?)

QUARKS  $\longleftrightarrow$  LEPTONS

$$SU(2) \times U(1)$$

## Chiral Fields

### Quarks

	$I_3$	$Y$	$Q$
$u_L$	$+1/2$	$1/6$	$2/3$
$d_L$	$-1/2$	$1/6$	$-1/3$
-----			
$u_R$	0	$2/3$	$2/3$
$d_R$	0	$-1/3$	$-1/3$

Have to...

### Leptons

	$I_3$	$Y$	$Q$
$\nu_L$	$+1/2$	$-1/2$	0
$e_L$	$-1/2$	$-1/2$	-1
-----			
$e_R$	0	-1	-1

$$??? \quad \nu_R \quad 0 \quad 0 \quad 0$$

- Follow the fathers of SM :

(AT PRESENT ENERGIES)  $SU(2) \times U(1)$

with a particle content WITHOUT  $\nu_R$

- Is it possible (with only  $\nu_L$ ) to have  $m_\nu \neq 0$  ?

A priori, YES  $\longleftrightarrow$  Majorana

$$\mathcal{L}_{\text{mass}}^{\text{Maj}} = -\frac{1}{2} \bar{\nu}_L M \nu_L^c + \text{h.c.}$$

$SU(3)_\text{colour} \otimes U(1)_\text{e.m.}$  invariant  $\Rightarrow$  legal

$$\mathcal{L}^{\text{Maj}} + \text{Pauli} \Rightarrow M^T = M$$

$$\text{Diagonalization} \rightarrow M = U m U^T$$

$$\mathcal{L}_{\text{mass}}^{\text{Maj}} = -\frac{1}{2} \bar{X} m X$$

$$X \equiv U^+ \nu_L + (\bar{U}^+ \nu_L)^c$$

Majorana Field

$$X = X^c = C \bar{X}^T$$

No phase-freedom to rotate away

$\Rightarrow$  Neutrinos would be "true" neutrals

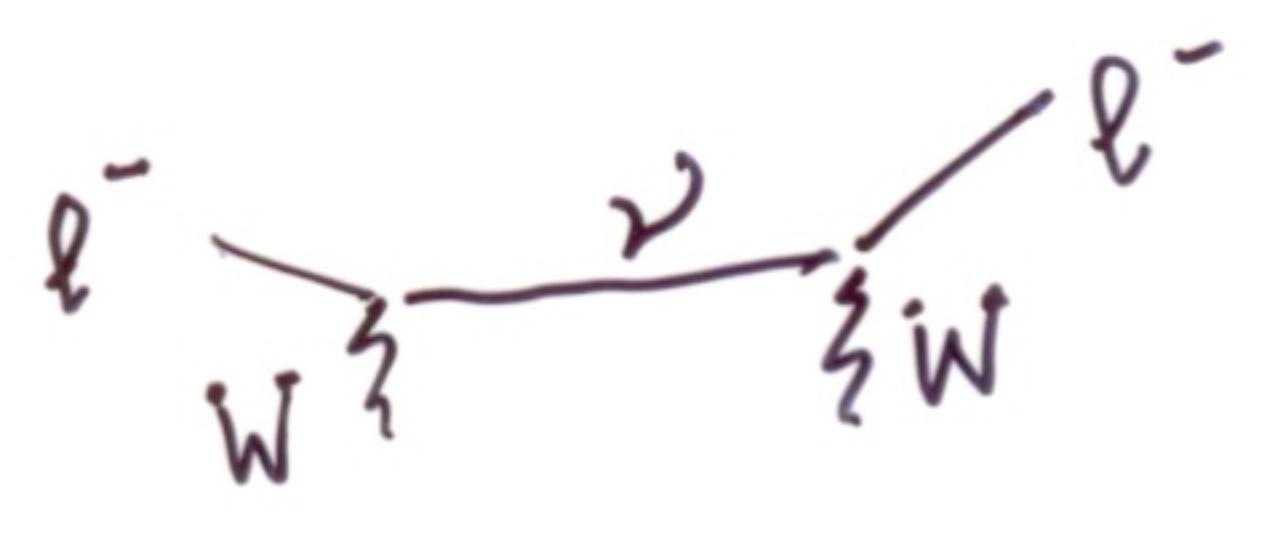
- NO conserved GLOBAL LEPTON NUMBER

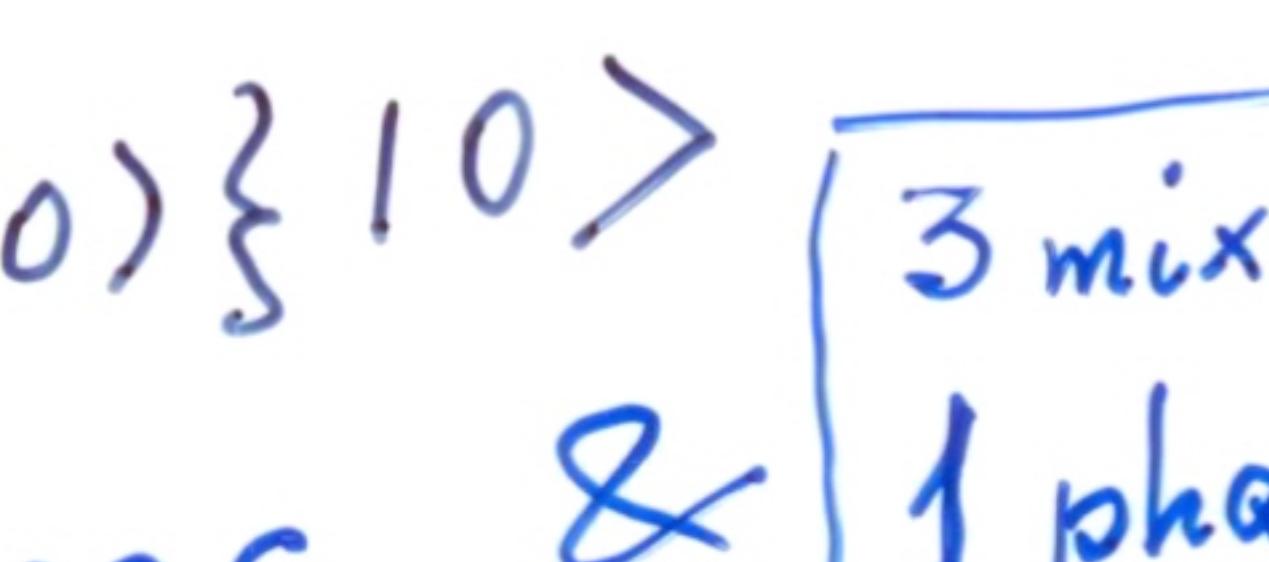
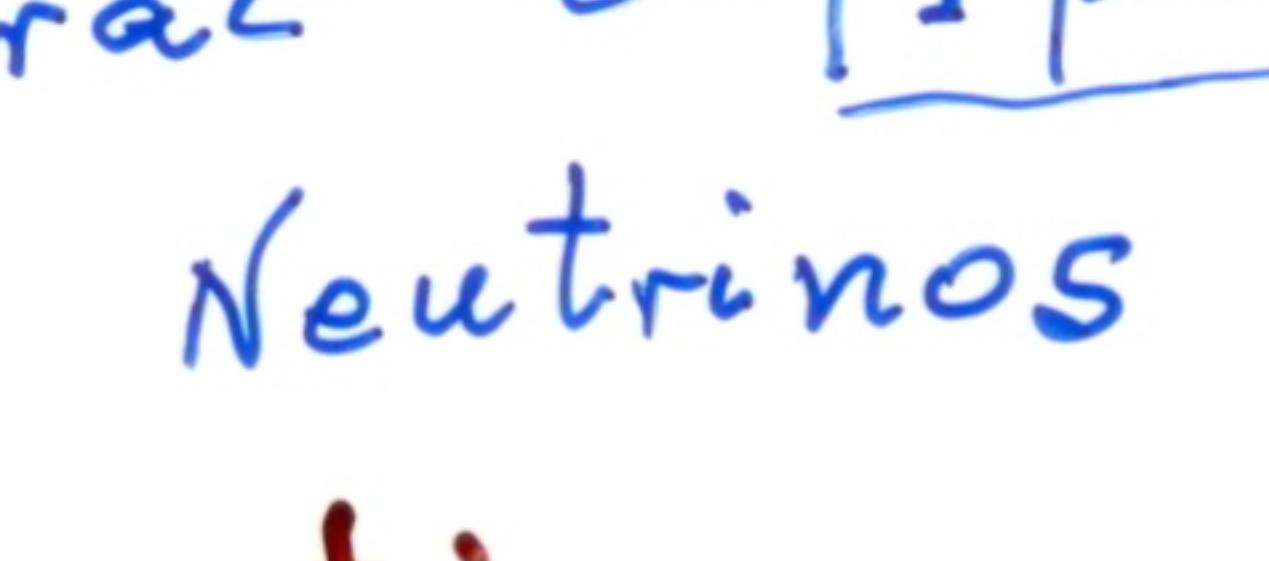
$\Rightarrow \beta\beta_{-\alpha\nu}$  allowed

BUT ...

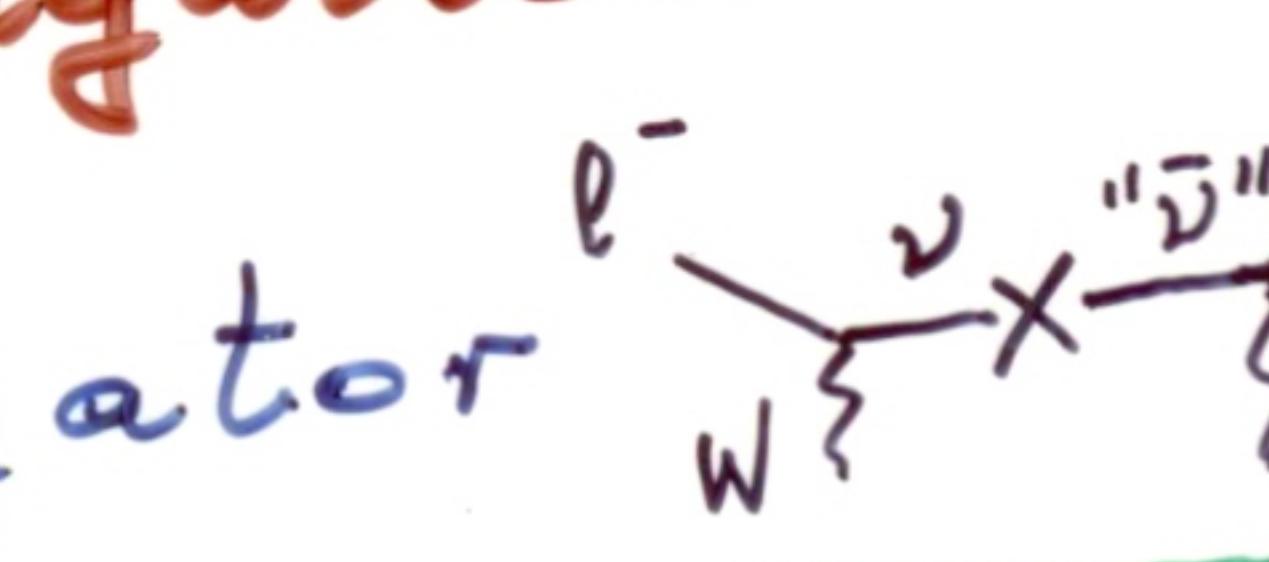
- Fermion Field

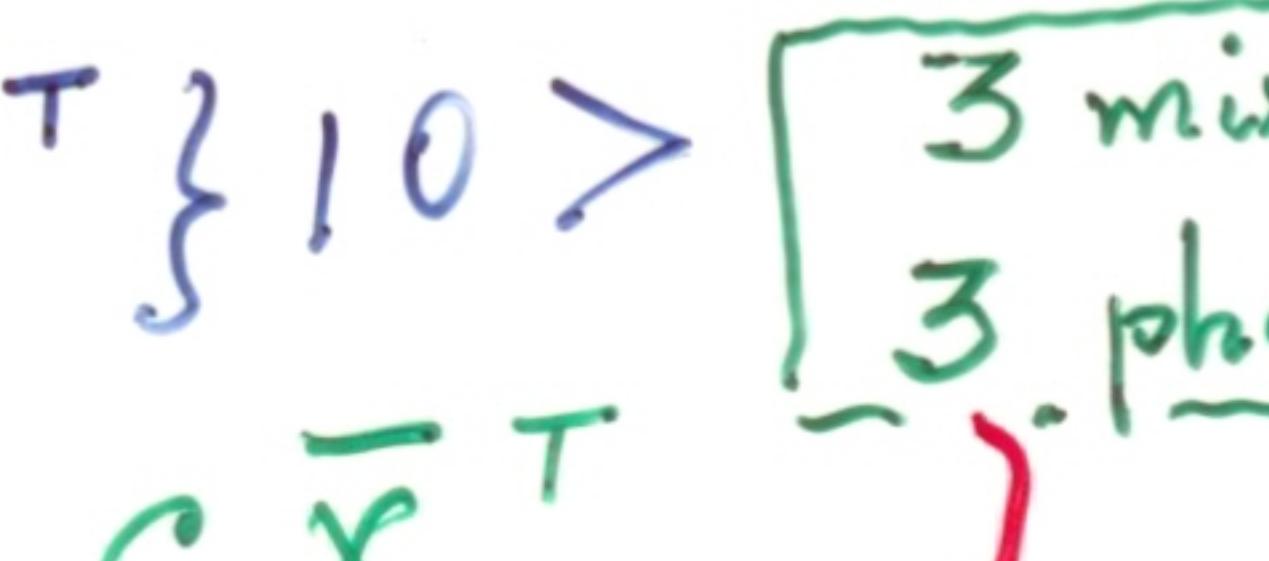
$$\psi(x) = \int \frac{d^3 p}{(2\pi)^{3/2}} \frac{1}{\sqrt{2p^0}} \sum_{\lambda} \left\{ c_{\lambda}(p) u_{\lambda}(p) e^{-ipx} + d_{\lambda}^{+}(p) G [\bar{u}_{\lambda}(p)]^T e^{ipx} \right\}$$

- Dirac Propagator 

Flavour Oscillations  $\nLeftarrow \langle 0 | T \{ \psi(x) \bar{\psi}(0) \} | 0 \rangle$     
 ∃ for both Dirac & Majorana Neutrinos 

### ≡ Neutrino propagation

- Majorana Propagator 

$$\langle 0 | T \{ \psi(x) \psi(0)^T \} | 0 \rangle$$
 

2 additional phases  $\Leftrightarrow \psi \equiv X = C \bar{X}^T$

$$c_{\lambda}(p) = d_{\lambda}(p)$$

∃ only for Majoranas

### ≡ "Neutrino-antineutrino" propagation

#

- $\mathcal{L}^{\text{Maj}}$  cannot be GENERATED by mass SSB of a RENORMALIZABLE (dim. 4) Yukawa interaction! Higgs triplet?

1970's → 2000's: Requirement of RENORMALIZABILITY has a different PHILOSOPHY:

- Take particle content of SM and ask WHAT IS THE LOWEST DIMENSION (NON-RENORMALIZABLE) OPERATOR WITH  $SU(2) \times U(1)$  GAUGE INVARIANCE?

UNIQUE  
dim. 5

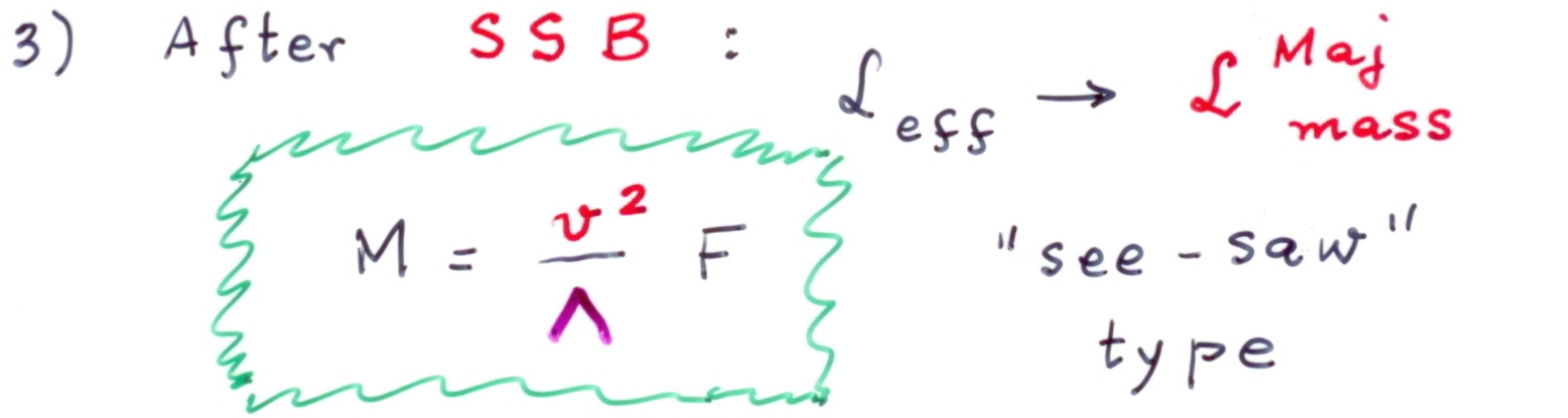
$$\mathcal{L}_{\text{eff}} = -\frac{1}{2\Lambda} (\tilde{\bar{l}}_L \varphi) F (\tilde{\varphi}^+ \bar{l}_L)$$

$$\tilde{\bar{l}} = i \gamma_2 l^c = i \gamma_2 c \bar{l}^T$$

- The FIRST window to BEYOND SM
- $F = F^T$  matrix in flavour-space
- Coupling  $\frac{1}{\Lambda} \Leftrightarrow \Lambda \equiv \text{New Physics Scale}$

$$\Rightarrow 1) \Delta L = 2$$

$$2) F \Rightarrow \text{Mixing, LFV}$$



### Conclusion

$L_{\text{mass}}^{\text{Maj}}$  for "light" neutrinos GENERATED from  $L_{\text{eff}}$  (at present energies).

- Origin of  $\Lambda$  ?

1) Heavy mass  $\nu_R$  ?

2) Fierz - reordering

$$L_{\text{eff}} = -\frac{1}{4\Lambda} (\bar{l}_L F \vec{\epsilon} l_L) (\tilde{\phi}^+ \vec{\epsilon} \phi)$$

Higgs triplet ?

- A particular example : SO(10) GUT

$\Rightarrow$  As part of SB,  $\nu_R$  obtains a large Majorana mass  $\Lambda$ .

- The Dirac mass terms  $\nu_R \leftrightarrow \nu_L m_D$  produce mixing of heavy  $\nu_R$  into  $\nu_L \Rightarrow m_\nu = \frac{m_D^2}{\Lambda}$

-  $\nu_R$  and Yukawa interaction



$$\mathcal{L}_Y = -\frac{\sqrt{2}}{v} \bar{\ell}_L M \nu_R \tilde{\phi} + h.c.$$

$$\tilde{\phi} = i \gamma_2 \phi^* \xrightarrow{\text{SSB}} \begin{cases} (v + H)/\sqrt{2} \\ 0 \end{cases}$$

⇒ Total analogy with quarks,

CKM-matrix, ---

**U**

↓ Pontecorvo-MNS

-  $\nu_L$ 's are Dirac particles

-  $\nu_R$  only in Mass-terms ⇒

⇒ Masses and Mixings

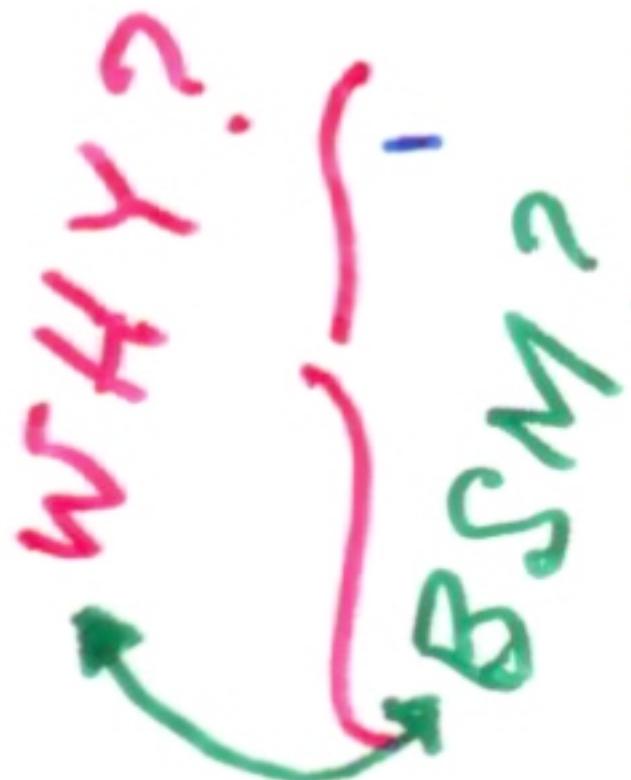
- Mixing relevant for CC

- Mixing irrelevant for NC-Lagrangian

⇒ GIM-suppressed FCNC

-  $\nu_R$ 's do not appear elsewhere than H

L.F.V., but  $\mathcal{L}$  still invariant under  
GLOBAL U(1)-GAUGE TRANSFORMATION  
TOTAL LEPTON NUMBER



③ Is it possible (with only  $\nu_L$ ) to generate  $m_\nu \neq 0$ ?

YES (a priori) ↔ Majorana

## The Pontecorvo MNS Matrix

$$\begin{Bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{Bmatrix} = U \begin{Bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{Bmatrix}$$

For Flavour oscillations

$U$ : 3 mixings, 1 phase

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$\approx$  Atmospheric  
LBL-beams

? ?

Atmospheric  
(through Matter)?  
Reactor?

$\approx$  Solar  
Reactor (if LMA-MSW)

Appearance  $\nu_\mu \rightarrow \nu_e$ !

"Extended" SM

- $SU(2) \times U(1)$  gauge invariance }
- SM +  $\nu_R$  in particle content }

$$\mathcal{L}(x) = -\frac{\sqrt{2}}{v} \bar{\ell}_L m_D \nu_R \tilde{\phi} + h.c.$$

\$SU(2) \times U(1)\$ invariant

$$-\frac{1}{2} \bar{\nu}_R m_R (\nu_R)^c + h.c.$$

$\Rightarrow$  Not only L-R Dirac Mass,  
R-Majorana too

$$\mathcal{L}^{D-M} = -\bar{\nu}_R m_D \nu_L - \frac{1}{2} \bar{\nu}_R m_R (\nu_R)^c$$

$$+ h.c. \equiv -\frac{1}{2} (\bar{n}_L)^c M n_L + h.c.$$

$$n_L \equiv \begin{Bmatrix} \nu_L \\ (\nu_R)^c \end{Bmatrix}, \quad M = \begin{pmatrix} 0 & m_D \\ m_D & m_R \end{pmatrix}$$

$\Rightarrow$  Physical neutrinos  $\equiv$  MAJORANA  
Unless  $m_R = 0$ ? Why? Global  $U(1)$ ?  
 $\Leftrightarrow$  BSM

FACTS

- SM  $\leftrightarrow \nu_L$  only

$$\Rightarrow m_\nu = 0$$

- BSM +  $\nu_L$  only at present

$\Rightarrow$  Unique window  $\equiv$  dim. 5

$\Rightarrow m_\nu \neq 0$  MAJORANA

$\Delta L = 2$  interactions at  
higher energies

- SM +  $\nu_R$  (light)

$\Rightarrow m_\nu \neq 0$  MAJORANA

$\Delta L = 2$  at present energies

-  $m_R = 0$  ?  $\Leftrightarrow$  U(1) - global symmetry  
 $\Delta L = 0$  ?

$m_\nu \neq 0$  DIRAC