Numerical Schemes & Simulation

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Fluid-Particles Interaction Models Asymptotic Models and Simulation

J. A. Carrillo in collaboration with T. Goudon and P. Lafitte

(CPDE 2005) & Work in Progress

ICREA - Universitat Autònoma de Barcelona

Benasque, 31/08/07

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- Kinetic Modelling
- Fluid-Particles Interaction

2 Asymptotics

- Dimensionless Formulation
- Asymptotic Limits
- Asymptotic Systems

3 Numerical Schemes & Simulation

- Asymptotic Preserving Kinetic Schemes: Bubbling
- Numerical Simulation
- Conclusions & Perspectives

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Kinetic Modelling

Atmospheric Pollution Modelling





Pollution in Los Ángeles, Madrid and Beijing.

Asymptotics

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Statistical description



- Particles Description: impossible due to their huge number.
- Kinetic Description: $f(t, x, \xi)$ represents the number density of particles at time *t* in position *x* with velocity ξ .
- Hydrodynamic Description: Continuum mechanics approach based on balance equations for density, momentum and temperature.

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Macroscopic Quantities: Moments

• Particle density:

$$\rho(t,x) = \int_{\mathbb{R}^3} f(t,x,\xi) \, d\xi$$

• Momentum:

$$J(t,x) = \rho U(t,x) = \int_{\mathbb{R}^3} \xi f(t,x,\xi) \, d\xi$$

• Temperature:

$$3\rho\theta(t,x) = \int_{\mathbb{R}^3} |\xi - U(t,x)|^2 f(t,x,\xi) \, d\xi$$

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Fluid-Particles Interaction

Formation of Aerosols in the atmosphere

Sources	Estimation	Catégorie de taille des parti-
	des émissions	cules
	$(10^{12}g/an)$	
Sources naturelles		
Croûte terrestre et érosion	1500	grosses particules
éolienne		
Océans (sel)	1300	accumulation et grosses parti-
		cules
Volcans	30	grosses particules
Débris biologiques	50	grosses particules
Sulphates dérivés des gaz	130	particules très fines
biogéniques		
Sulphates dérivés des gaz vol-	20	particules très fines
caniques		
Matières organiques	60	particules très fines
Nitrates dérivés des NO_x	30	particules très fines
Total des sources naturelles	3100	
Sources anthropiques		
Poussières industrielles	100	particules très fines et grosses
Suie	10	particules très fines
Sulphates dérivés de SO_2	190	particules très fines
Feux	90	particules très fines
Nitrates dérivés des NO_x	50	grosses particules
Matières organiques	10	particules très fines
Total des sources anthro-	450	
piques		
Total des deux sources	3600	



Pandis & Seinfeld (1998), Madelaine (1982)

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Two phase flow:

• [Dense Phase] Fluid: continuum mechanics description in terms of density of the fluid n(t, x) and velocity field u(t, x).

Let $\rho_{\rm F}$ a typical value of the fluid mass per unit volume. Fluid Equations: Compressible Euler.

• [Dispersed Phase] Particles: kinetic description in terms of the number density of particles $f(t, x, \xi)$ in phase space (x, ξ) to compute velocity fluctuations around the fluid velocity u(t, x).

Particles are spheres of radius a > 0 with mass given by $m_{\rm p} = \frac{4}{3} \rho_{\rm p} \pi a^3$, $\rho_{\rm p}$ being the particle mass per unit volume.

Particles are assumed to follow a kind of Brownian motion:

$$m_{\rm P} x'' + F(t, x, x') = \Gamma(t).$$

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Forces to be considered

Forces F(t, x, x'):

• Friction: The fluid produces a friction force on the particles

 $6\pi\mu a\big(u(t,x)-\xi\big),$

with $\mu>0$ being the dynamic viscosity of the fluid. Accordingly, the force exerted by the particles on the fluid is given by the sum

$$6\pi\mu a\int_{\mathbb{R}^3} (\xi - u(t,x))f\,\mathrm{d}\xi.$$

• Gravity+Buoyancy: External forces per unit volume acting on the particles $-m_{\rm P}\nabla_x \Phi$ and on the fluid $\alpha \rho_{\rm F} \nabla_x \Phi$.

 $\alpha \in \mathbb{R}$ is a dimensionless parameter which measures the ratio of the strength of the external force on each phase:

$$\Phi(x) = (1 - \rho_{\rm F}/\rho_{\rm P})gx_3 \qquad \alpha = \frac{1}{1 - \rho_{\rm F}/\rho_{\rm P}},$$

by Archimedes rule.

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Disperse Phase

Vlasov-Fokker-Planck equation:

$$\partial_t f + \xi \cdot \nabla_x f - \nabla_x \Phi \cdot \nabla_\xi f = \frac{9\mu}{2a^2\rho_{\rm P}} \operatorname{div}_\xi \Big((\xi - u)f + \frac{k\theta_0}{m_{\rm P}} \nabla_\xi f \Big).$$

where k stands for the Boltzmann constant, and $\theta_0 > 0$ controls the noise strength.

Fokker-Planck term:

The Fokker-Planck term implies a relaxation in velocity towards equilibrium densities of the form

$$\rho(t,x)M(\xi) = \rho(t,x) \left(2\pi \frac{k\theta_0}{m_{\rm p}}\right)^{-3/2} \exp\left\{-m_{\rm p}|\xi - u(t,x)|^2/2k\theta_0\right\},$$

with typical Stokes relaxation time given by

$$T_S = \frac{m_{\rm P}}{6\pi\mu a} = \frac{2\rho_{\rm P}a^2}{9\mu}$$

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Final PDE Model

Vlasov-Euler-Fokker-Planck system:

We arrive at the system:

$$\partial f + \xi \cdot \nabla_x f - \nabla_x \Phi \cdot \nabla_{\xi} f = \frac{9\mu}{2a^2\rho_{\rm P}} \operatorname{div}_{\xi} \left((\xi - u)f + \frac{k\theta_0}{m_{\rm P}} \nabla_{\xi} f \right), \qquad (1)$$

$$\partial_t n + \operatorname{div}_x(nu) = 0,$$
 (2)

$$\rho_{\rm F} \Big(\partial_t (nu) + {\rm Div}_x (nu \otimes u) + \alpha n \nabla_x \Phi \Big) + \nabla_x p(n) = 6\pi \mu a \int_{\mathbb{R}^3} (\xi - u) f \, \mathrm{d}\xi. \tag{3}$$

where *k* stands for the Boltzmann constant, and $\theta_0 > 0$ controls the noise strength and p(n) is a general pressure law, for instance $p(n) = C_{\gamma} n^{\gamma}, \gamma \ge 1, C_{\gamma} > 0$.

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Numerical Schemes & Simulation

DimensionLess PDE Model

DimensionLess Vlasov-Euler-Fokker-Planck system:

$$\partial_t f + \beta \xi \cdot \nabla_x f - \eta' \nabla_x \Phi \cdot \nabla_{\xi} f = \frac{1}{\epsilon} \nabla_{\xi} \cdot \left(\left(\xi - \frac{1}{\beta} u \right) f + \nabla_{\xi} f \right), \tag{4}$$

$$\partial_t n + \operatorname{div}_x(nu) = 0,$$
 (5)

$$\partial_t(nu) + \operatorname{Div}_x(nu \otimes u) + \nabla_x p(n) + \eta \, n \nabla_x \Phi = \frac{1}{\epsilon} \frac{\rho_{\rm P}}{\rho_{\rm F}} (J - \rho u). \tag{6}$$

where

$$\rho(t,x) = \int_{\mathbb{R}^3} f(t,x,\xi) \,\mathrm{d}\xi, \qquad J(t,x) = \beta \,\int_{\mathbb{R}^3} \xi f(t,x,\xi) \,\mathrm{d}\xi.$$

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Entropy Decay:

Assumme the scaling:

$$\frac{\rho_{\rm P}}{\rho_{\rm F}} \beta^2 = 1, \qquad \eta' = \varsigma \beta, \qquad \text{with } \varsigma = \pm 1.$$

Defining the free energies associated respectively to the particles and the fluid as:

$$\begin{split} \mathcal{F}_{\mathrm{P}}(t) &= \int_{\mathbb{R}^3}\!\!\int_{\mathbb{R}^3} \left(f \ln(f) + \frac{\xi^2}{2} f + \varsigma \Phi f \right) \mathrm{d}\xi \,\mathrm{d}x, \\ \mathcal{F}_{\mathrm{F}}(t) &= \int_{\mathbb{R}^3} \left(n \frac{|u|^2}{2} + \Pi(n) + \eta \Phi n \right) \mathrm{d}x, \end{split}$$

where $\Pi : \mathbb{R}^+ \longrightarrow \mathbb{R}^+$ is defined by $s\Pi''(s) = p'(s)$. Then, we have the crucial dissipation:

$$\frac{\mathrm{d}}{\mathrm{d}t}\Big(\mathcal{F}_{\mathrm{P}}+\mathcal{F}_{\mathrm{F}}\Big)+\frac{1}{\varepsilon}\int_{\mathbb{R}^{3}}\int_{\mathbb{R}^{3}}\left|(\xi-\beta^{-1}u)\sqrt{f}+2\nabla_{\xi}\sqrt{f}\right|^{2}\mathrm{d}\xi\,\mathrm{d}x\leq0.$$

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$$\frac{\rho_{\rm P}}{\rho_{\rm F}} \beta^2 = 1, \qquad \eta' = \varsigma \beta, \qquad \text{with } \varsigma = \pm 1.$$

Defining the free energies associated respectively to the particles and the fluid as:

$$\begin{split} \mathcal{F}_{\mathrm{P}}(t) &= \int_{\mathbb{R}^3}\!\int_{\mathbb{R}^3} \left(f \ln(f) + \frac{\xi^2}{2} f + \varsigma \Phi f \right) \mathrm{d}\xi \,\mathrm{d}x, \\ \mathcal{F}_{\mathrm{F}}(t) &= \int_{\mathbb{R}^3} \left(n \frac{|u|^2}{2} + \Pi(n) + \eta \Phi n \right) \mathrm{d}x, \end{split}$$

where $\Pi : \mathbb{R}^+ \longrightarrow \mathbb{R}^+$ is defined by $s\Pi''(s) = p'(s)$. Then, we have the crucial dissipation:

$$\frac{\mathrm{d}}{\mathrm{d}t}\Big(\mathcal{F}_{\mathrm{P}}+\mathcal{F}_{\mathrm{F}}\Big)+\frac{1}{\varepsilon}\int_{\mathbb{R}^{3}}\int_{\mathbb{R}^{3}}\left|(\xi-\beta^{-1}u)\sqrt{f}+2\nabla_{\xi}\sqrt{f}\right|^{2}\mathrm{d}\xi\,\mathrm{d}x\leq0.$$

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Dissipation Properties 2

Comments:

• Entropy Dissipation: This claim helps in understanding the asymptotic regime $\varepsilon \ll 1$: we infer that *f* has essentially a hydrodynamic behavior

 $f(t,x,\xi) \simeq \rho(t,x) (2\pi)^{-3/2} \exp\left(-|\xi - \beta^{-1}u(t,x)|^2/2\right) = \rho(t,x) M_{u(t,x)/\beta}(\xi).$
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Asymptotic Linnis

Bubbling Regime

We set

$$eta = rac{1}{\sqrt{arepsilon}}, \qquad |\eta'| = rac{1}{\sqrt{arepsilon}}, \qquad rac{
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ho_{ extsf{F}}} = arepsilon.$$

meaning that:

Stokes velocity \simeq Typical velocity of the fluid \ll Thermal velocity.

The dispersed phase is buoyancy driven while the flow is gravity driven. Here, $\eta' < 0$ and the external forces act in opposite directions on the particles and on the fluid.

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Asymptotic Limits

Flowing Regime

We set

$$\beta^2 \frac{\rho_{\rm P}}{\rho_{\rm F}} = 1, \qquad \beta = |\eta'|$$
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- Kinetic Modelling
- Fluid-Particles Interaction

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- Asymptotic Limits
- Asymptotic Systems

Numerical Schemes & Simulation

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- Numerical Simulation
- Conclusions & Perspectives

Asymptotics

Numerical Schemes & Simulation

Asymptotic Systems

Diffusion Asymptotics: Bubbling Regime

DimensionLess Bubbling Regime:

$$\partial_t f + \frac{1}{\sqrt{\varepsilon}} \left(\xi \cdot \nabla_x f + \nabla_x \Phi \cdot \nabla_\xi f \right) = \frac{1}{\epsilon} \nabla_\xi \cdot \left((\xi - \sqrt{\varepsilon}u) f + \nabla_\xi f \right),$$
$$\partial_t n + \operatorname{div}_x(nu) = 0,$$
$$\partial_t (nu) + \operatorname{Div}_x(nu \otimes u) + \nabla_x p(n) + \eta n \nabla_x \Phi = J - \rho u.$$

where

$$\rho(t,x) = \int_{\mathbb{R}^3} f(t,x,\xi) \,\mathrm{d}\xi, \qquad J(t,x) = \frac{1}{\sqrt{\varepsilon}} \int_{\mathbb{R}^3} \xi f(t,x,\xi) \,\mathrm{d}\xi.$$

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Asymptotics

Numerical Schemes & Simulation

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Asymptotic Systems

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Numerical Schemes & Simulation

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Diffusion Asymptotics: Bubbling Regime

- As ε tends to 0, we should have that $f(t, x, \xi) \simeq \rho(t, x) M(\xi)$.
- Hilbert expansion: we plug the ansatz

$$f_{\varepsilon} = f^{(0)} + \sqrt{\varepsilon} f^{(1)} + \varepsilon f^{(2)} + \dots$$

into the kinetic equation.

DimensionLess Bubbling Regime:

We end up with the limiting system

 $\begin{aligned} \partial_t \rho + \operatorname{div}_x \left(\rho(u + \nabla_x \Phi) - \nabla_x \rho \right) &= 0, \\ \partial_t n + \operatorname{div}_x(nu) &= 0, \\ \partial_t(nu) + \operatorname{Div}_x(nu \otimes u) + \nabla_x(p(n) + \rho) + (\eta n - \rho) \nabla_x \Phi &= 0 \end{aligned}$

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Numerical Schemes & Simulation

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Numerical Schemes & Simulation

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Motivation & Modelling 0000000000 Asymptotic Systems

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Asymptotic Systems

Hydrodynamic Asymptotics: Flowing Regime

DimensionLess Flowing Regime:

We end up with the limiting system

 $\begin{cases} \partial_t \rho + \operatorname{div}_x(\rho u) = 0, \\\\ \partial_t n + \operatorname{div}_x(nu) = 0 \\\\ \partial_t \left((n + \beta^{-2} \rho) u \right) + \operatorname{Div}_x \left((n + \beta^{-2} \rho) u \otimes u \right) \\\\ + \nabla_x \left(\rho + p(n) \right) + (\eta n + \varsigma \rho) \nabla_x \Phi = 0. \end{cases}$

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Asymptotic Preserving Kinetic Schemes: Bubbling

Bubbling Regime: Coupling of density and fluctuations

The scheme is based on the expansion

$f_{\varepsilon}(t, x, \xi) = \rho_{\varepsilon}(t, x) M(\xi) + \sqrt{\varepsilon} r_{\varepsilon}(t, x, \xi)$

with the "fluctuations" r_{ε} bounded in L^2 by entropy dissipation.

We rewrite the scaled kinetic equation as

$$\partial f_{\varepsilon} + \xi \cdot \nabla_{x} r_{\varepsilon} + (u_{\varepsilon} + \nabla_{x} \Phi) \nabla_{\xi} r_{\varepsilon} = \frac{1}{\varepsilon} L f_{\varepsilon} + \frac{1}{\sqrt{\varepsilon}} M(\xi) S_{\varepsilon}(t, x, \xi),$$

where

$$S_{\varepsilon}(t,x,\xi) = -\xi \cdot \nabla_x \rho_{\varepsilon} - \xi \cdot (u_{\varepsilon}(t,x) + \nabla_x \Phi) \rho_{\varepsilon},$$

$$\partial_t r_{\varepsilon} = \frac{1}{\varepsilon} L r_{\varepsilon} - \frac{1}{\varepsilon} M S_{\varepsilon} - \frac{1}{\sqrt{\varepsilon}} \left[\xi \cdot \nabla_x r_{\varepsilon} + (u_{\varepsilon} + \nabla_x \Phi) \nabla_{\xi} r_{\varepsilon} - M \nabla_x \cdot \left(\int_{\mathbb{R}^3} \xi_{\star} r_{\varepsilon} \, \mathrm{d}\xi_{\star} \right) \right]$$

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Given n^k, u^k, f^k, r^k , evaluation of n, u, f, r at time $k\Delta t$:

Step 0.- Solve the Euler equations for the fluid density *n* and velocity *u*.

O Particles density is constant for this step:

$$\int_{\mathbb{R}^3} \xi r^k \,\mathrm{d}\xi - u \int_{\mathbb{R}^3} f^k \,\mathrm{d}\xi.$$

- Numerical method: Després & Lagoutière 99'-04' which preserves with accuracy the shock structure of the hyperbolic system. This defines n^{k+1} and u^{k+1} .
- So Different stability conditions: We perform Step 0 on a time interval $(k\Delta t_h, (k+1)\Delta t_h)$, and then we make several sub-cycles (Step 1-Step 2) below on time intervals $(k'\Delta t_p, (k'+1)\Delta t_p)$, for some $\Delta t_p < \Delta t_h$. Typically, the space mesh size Δx being given, we have $\Delta t_p = \mathcal{O}(\Delta x^2)$ but $\Delta t_h = \mathcal{O}(\Delta x)$.

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Splitting Method 1

Given n^k , u^k , f^k , r^k , evaluation of n, u, f, r at time $k\Delta t$:

Step 0.- Solve the Euler equations for the fluid density *n* and velocity *u*.

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$$\int_{\mathbb{R}^3} \xi r^k \,\mathrm{d}\xi - u \int_{\mathbb{R}^3} f^k \,\mathrm{d}\xi.$$

- Numerical method: Després & Lagoutière 99'-04' which preserves with accuracy the shock structure of the hyperbolic system. This defines n^{k+1} and u^{k+1} .
- Different stability conditions: We perform Step 0 on a time interval $(k\Delta t_h, (k+1)\Delta t_h)$, and then we make several sub-cycles (Step 1-Step 2) below on time intervals $(k'\Delta t_p, (k'+1)\Delta t_p)$, for some $\Delta t_p < \Delta t_h$. Typically, the space mesh size Δx being given, we have $\Delta t_p = \mathcal{O}(\Delta x^2)$ but $\Delta t_h = \mathcal{O}(\Delta x)$.

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Motivation & Modelling 000000000 Asymptotic Preserving Kinetic Schemes: <u>Bubbling</u>

Asymptotics

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Splitting Method 2

Step 1.- Solve the stiff equations

$$\partial_t f = \frac{1}{\varepsilon} L f, \qquad \partial_t r = \frac{1}{\varepsilon} L r + \frac{1}{\varepsilon} M S,$$

where

$$S = -\xi \cdot \nabla_x \rho + \xi \cdot (u^{k+1} + \nabla_x \Phi) \rho.$$

Note that $\rho = \int f \,d\xi$ is not modified by the first equation so that the source term in the second equation can be treated as constant in time.

Step 2.- Solve the transport part

$$\partial_t f + \xi \cdot \nabla_x r + (u^{k+1} + \nabla_x \Phi) \cdot \nabla_\xi r = 0, \qquad \partial_t r = 0$$

Note that the convection term has characteristic speed ξ and not $\xi/\sqrt{\varepsilon}$) which defines $f^{k'+1}$ and $\rho^{k'+1} = \int f^{k'+1} d\xi$.

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Numerical Schemes & Simulation

Splitting Method 2

Step 1.- Solve the stiff equations

$$\partial_t f = \frac{1}{\varepsilon} L f, \qquad \partial_t r = \frac{1}{\varepsilon} L r + \frac{1}{\varepsilon} M S,$$

where

$$S = -\xi \cdot
abla_x
ho + \xi \cdot (u^{k+1} +
abla_x \Phi)
ho.$$

Note that $\rho = \int f d\xi$ is not modified by the first equation so that the source term in the second equation can be treated as constant in time.

Step 2.- Solve the transport part

$$\partial_t f + \xi \cdot \nabla_x r + (u^{k+1} + \nabla_x \Phi) \cdot \nabla_\xi r = 0, \qquad \partial_t r = 0$$

Note that the convection term has characteristic speed ξ and not $\xi/\sqrt{\varepsilon}$) which defines $f^{k'+1}$ and $\rho^{k'+1} = \int f^{k'+1} d\xi$.

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Asymptotic Preserving Kinetic Schemes: Bubbling

Approximation of $e^{\Delta t L/\varepsilon}$

Solutions of

$$\partial_t F = \frac{1}{\varepsilon} LF + H$$

can be explicitly computed since the fundamental solution of L is given by

$$\mathcal{G}(t,\xi,\xi_{\star}) = D_{\gamma(t)} \exp\left(-\frac{|\xi-\gamma(t)\xi_{\star}|^2}{2(1-\gamma(t)^2)}\right), \qquad \gamma(t) = e^{-t} \ , \ D_{\gamma(t)} = \frac{1}{(2\pi(1-\gamma(t)^2))^{N/2}}.$$

$$F(t,\xi) = \int_{\mathbb{R}^3} \mathcal{G}\Big(\frac{t-s}{\varepsilon},\xi,\xi_\star\Big) F(s,\xi_\star) \,\mathrm{d}\xi_\star + \int_s^t \mathcal{G}\Big(\frac{t-\sigma}{\varepsilon},\xi,\xi_\star\Big) H(\sigma,\xi_\star) \,\mathrm{d}\xi_\star \,\mathrm{d}\sigma.$$

$$D_{\gamma} \int_{\mathbb{R}^3} \exp\left(-\frac{|\xi - \gamma \xi_{\star}|^2}{2(1 - \gamma^2)}\right) F(\xi_{\star}) \,\mathrm{d}\xi_{\star} = M(\xi) \left(\int_{\mathbb{R}^3} F(\xi_{\star}) \,\mathrm{d}\xi_{\star} + \gamma \xi \int_{\mathbb{R}^3} \xi_{\star} F(\xi_{\star}) \,\mathrm{d}\xi_{\star}\right).$$

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Duhamel formula:

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Since it involves the quantity $e^{-t/\varepsilon}$ with $0 < \varepsilon \ll 1$, then

$$D_{\gamma} \int_{\mathbb{R}^3} \exp\left(-\frac{|\xi - \gamma\xi_\star|^2}{2(1-\gamma^2)}\right) F(\xi_\star) \,\mathrm{d}\xi_\star = M(\xi) \left(\int_{\mathbb{R}^3} F(\xi_\star) \,\mathrm{d}\xi_\star + \gamma\xi \int_{\mathbb{R}^3} \xi_\star F(\xi_\star) \,\mathrm{d}\xi_\star\right).$$

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Approximation of $e^{\Delta t L/\varepsilon}$

We use this expansion to approximate the Duhamel formula with $H(t, x, \xi) = -\frac{1}{\varepsilon}M(\xi)S(k'\Delta t, x, \xi)$ which is not modified during the time step. Accordingly, we make appear

$$\frac{1}{\varepsilon} \int_{k'\Delta t}^{(k'+1)\Delta t} e^{(\sigma-k'\Delta t)/\varepsilon} \,\mathrm{d}\sigma = 1 - e^{-\Delta t/\varepsilon}.$$

$$\begin{cases} f^{k'+1/2}(\xi) &= M(\xi) \left(\rho^{k'} + e^{-\Delta t/\varepsilon} \xi \int_{\mathbb{R}^N} \xi_* f^{k'} \, \mathrm{d}\xi_* \right), \\ r^{k'+1/2}(\xi) &= e^{-\Delta t/\varepsilon} M(\xi) \left(\xi \int_{\mathbb{R}^N} \xi_* r^{k'} \, \mathrm{d}\xi_* \right) + (1 - e^{-\Delta t/\varepsilon}) \, M(\xi) S^{k'}. \end{cases}$$

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Therefore, Step 1 of the method reduces to:

$$\begin{cases} f^{k'+1/2}(\xi) &= M(\xi) \left(\rho^{k'} + e^{-\Delta t/\varepsilon} \xi \int_{\mathbb{R}^N} \xi_\star f^{k'} \, \mathrm{d}\xi_\star \right), \\ r^{k'+1/2}(\xi) &= e^{-\Delta t/\varepsilon} M(\xi) \left(\xi \int_{\mathbb{R}^N} \xi_\star r^{k'} \, \mathrm{d}\xi_\star \right) + (1 - e^{-\Delta t/\varepsilon}) \, M(\xi) S^{k'}. \end{cases}$$

Properties Numerical Method

- Asymptotic Preserving: taking the completely relaxed model, i.e., $\varepsilon = 0$, yields an approximation scheme of the Smoluchowski-Euler model.
- Full discretization: A downwind discretization for $-\xi \partial_x \rho$ in Step 1 and an upwind discretization for $-\xi \partial_x r$ in Step 2. For the discretization of $(u^{k+1} \partial_x \Phi) \partial_{\xi} r$ in Step 2, we choose a centered discretization in velocity.
- Boundary conditions: At the convection Step 2 specular reflections for the fluxes associated to the convection in space. At the end of Step 1, we impose a boundary condition on the fluctuations coherent with specular reflection for *f*.
- Well-Balanced: Equilibrium states for the kinetic equation are kept up to consistency error.

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- Kinetic Modelling
- Fluid-Particles Interaction

2 Asymptotics

- Dimensionless Formulation
- Asymptotic Limits
- Asymptotic Systems

3 Numerical Schemes & Simulation

- Asymptotic Preserving Kinetic Schemes: Bubbling
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Conclusions & Perspectives

- Excellent Asymptotic Preserving Schemes for the simplified coupled kinetic-fluid model for atmospheric pollutants.
- Improvements of the model: coagulation in size of pollutants, different species, chemical reactions,...
- Use of these models in realistic situations: contacts underway with mechanicists in Lyon and Lille.

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