

## Observational cosmology – Exercise

### Taller de Altas Energías 2010

You are designing a space mission to map the cosmic microwave background radiation over the full sky with a sensitivity of 5 microK per  $10'$  resolution element. The instrument uses an array of focal plane detectors with equivalent system temperature of 12K and bandwidth 20 GHz. The cryogenic system allows a detectors lifetime of 2 years.

How many detectors do you need to implement in the focal plane? How many would you need if the sensitivity requirement was 5 times more stringent? (Neglect foreground emissions and systematic effects).

SOLUTION:

The white noise of the array is given by (assuming radiometer constant equal to unity):

$$\delta T_{rms} \approx \frac{T_{sys} + T_{sky}}{\sqrt{\Delta\nu \cdot (N\tau)}}$$

Where  $N$  is the number of detectors to be found. The time necessary to reach the desired sensitivity in each 10' pixel is:

$$\begin{aligned} \tau_{pix} &\approx \frac{1}{N} \left( \frac{T_{sys} + T_{sky}}{\sqrt{\Delta\nu} \cdot \delta T_{rms}} \right)^2 = \frac{1}{N\Delta\nu} \left( \frac{T_{sys} + T_{sky}}{\delta T_{rms}} \right)^2 = \\ &= \frac{1}{N} \frac{1}{20 \cdot 10^9 \text{ Hz}} \left( \frac{(12 + 2.7)\text{K}}{5 \cdot 10^{-6} \text{ K}} \right)^2 = \frac{1}{N} 432.2 \text{ sec} \end{aligned}$$

The number 10' square pixels in the full sky is:

$$n_{pix} = \frac{\Omega_{sky}}{\Omega_{beam}} = \frac{4\pi}{(10/60)^2 \pi^2} \frac{180^2}{\pi^2} \approx 1.5 \times 10^6,$$

and the total observing time must be:

$$\tau_{tot} \approx \tau_{pix} n_{pix} = \frac{1}{N} 432.2 \cdot 1.5 \times 10^6 \text{ sec}$$

The lifetime of the mission is  $\tau_{tot} = 2 \text{ yrs} = 6.3 \times 10^7 \text{ sec}$ , so we have

$$N = \frac{432.2 \cdot 1.5 \times 10^6 \text{ sec}}{6.3 \times 10^7 \text{ sec}} \approx 10$$

A requirement  $\delta T_{rms} = 1 \text{ microK}$  (instead of 5microK) per 10' pixel, implies that  $\tau_{pix}$  is 25 times larger, and therefore  $N = 10 \times 25 = 250$ .