

Rare B decays: The Terminator for New Physics?

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PLAN of the TALK

- I. Methodology to obtain Rare B decay constraints in the space of correlations between Wilson Coefficients.
- II. Angular distribution of $B \rightarrow K^*(\rightarrow K\pi)l^+l^-$.
- III. Isospin asymmetry of $B \rightarrow K^*\mu^+\mu^-$

- $B \rightarrow \tau \nu$ & CKM fit (BaBar & Belle)
- $B_s \rightarrow \mu^+ \mu^-$ (CDF excess)
- ϕ_s (CDF and D0 hints of large value)
- A_{f_s} (D0 evidence)
- $A_{CP}(B \rightarrow K\pi)$ puzzle (BaBar and Belle)
- $A_{FB}(B \rightarrow K^* \mu^+ \mu^-)$ (BaBar, Belle & CDF hints)
- $A_I(B \rightarrow K^{(*)} \mu^+ \mu^-)$ (BaBar, Belle & CDF hints)

- $B \rightarrow \tau \nu$ & CKM fit (BaBar & Belle)
- $B_s \rightarrow \mu^+ \mu^-$ (CDF excess) consistent with SM
- ϕ_s (CDF and D0 hints of large value) consistent with SM
- A_{f_s} (D0 evidence)
- $A_{CP}(B \rightarrow K\pi)$ puzzle (BaBar and Belle)
- $A_{FB}(B \rightarrow K^* \mu^+ \mu^-)$ (BaBar, Belle & CDF hints) consistent SM
- $A_I(B \rightarrow K^{(*)} \mu^+ \mu^-)$ (BaBar, Belle & CDF hints) end of seminar

Conclusion: "New Physics will be extremely subtle"

- Model Building: **The Era of "Order of Magnitude" NP in Flavour (and check only $B \rightarrow X_s \gamma$) is gone**
- Flavour Physicists: **Redefine strategies to Focus when possible on Precise and Clean Observables.**

What type of New Physics (by category not by model) to look for?

- isospin violating
- right handed currents
- new scalars/tensors
- CP violating NP (Wilson coefficients) in decay.

Experimentalists: Where (which processes) to look for?

- penguin dominated (controlled IR div) and specific WC info.
- d/u spectator different processes (isospin) and d/s (U-spin)
- q^2 -observables (enhanced kin.): angular distribution
- F_L (less clean, many low values unexplained, interesting)

For Rare B decays: UT \rightarrow Wilson Coefficient planes.

Rare decays constraints: from UT to WC correlations

Discussion on constraints on WC from radiative and leptonic B decays should be addressed in a given framework, specific scenarios & observables

S. Descotes, D. Ghosh, JM., M. Ramon, [hep-ph/1104.3342](https://arxiv.org/abs/hep-ph/1104.3342)

- **Framework:** NP in C_7, C_9, C_{10} and $C_{7'}, C_{9'}, C_{10'}$
[chirally-flipped operators $\gamma_5 \rightarrow -\gamma_5$] as a real shift in the Wilson coefficients
- **Scenarios** (from the more specific to the more general)
 - A : NP in 7,7' only
 - B : NP in 7,7', 9,10 only
 - C : NP in 7,7',9,10,9',10' only
- **Classes within a Framework**
 - I: observables sensitive only to 7,7'
 - II: observables sensitive only to 7,7',9,9',10,10'
 - III: observables sensitive to 7,7',9,9',10,10' and more

The effective Hamiltonian describing the $b \rightarrow s l^+ l^-$ transition

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \sum_{i=1}^{10} [\mathbf{C}_i(\mu) \mathcal{O}_i(\mu) + \mathbf{C}'_i(\mu) \mathcal{O}'_i(\mu)],$$

$\mathbf{C}_i^{(\prime)}(\mu)$ are Wilson coefficients and $\mathcal{O}_i^{(\prime)}(\mu)$ are local operators.

We concentrate on *Electromagnetic dipole+ semileptonic operators*:

$$\begin{aligned} \mathcal{O}_7 &= \frac{e}{16\pi^2} m_b (\bar{s} \sigma_{\mu\nu} P_R b) F^{\mu\nu}, & \mathcal{O}_9 &= \frac{e^2}{16\pi^2} (\bar{s} \gamma_\mu P_L b) (\bar{l} \gamma^\mu l), \\ \mathcal{O}_{10} &= \frac{e^2}{16\pi^2} (\bar{s} \gamma_\mu P_L b) (\bar{l} \gamma^\mu \gamma_5 l), \end{aligned}$$

where $P_{L,R} = (1 \mp \gamma_5)/2$ and **primed counterpart operators**

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Limited sensitivity to hadronic inputs, or strong impact on analysis

- Class-I

- $\mathcal{B}(B \rightarrow X_s \gamma)$ with $E_\gamma > 1.6 \text{ GeV}$ [Misiak, Steinhauser, Haisch]
- exclusive time-dependent CP asymmetry $S_{K^* \gamma}$
- isospin asymmetry $A_I(B \rightarrow K^* \gamma)$ [Beneke, Feldman, Seidel]
[Kagan, Neubert, Feldman, J.M.]

- Class-II

- Integrated transverse asym. \tilde{A}_T^2 in $B \rightarrow K^* l^+ l^-$ over low- q^2 region [Kruger and J.M.]

- Class-III

- $\mathcal{B}(B \rightarrow X_s l^+ l^-)$ [Bobeth et al., Huber, Lunghi et al.]
- Integrated \tilde{F}_L and \tilde{A}_{FB} in $B \rightarrow K^* l^+ l^-$ [1-6 GeV²] [Beneke, Feldman]

For each observable

- Simple numerical parametrisation as $\delta C_i = C_i(\mu_b) - C_i^{SM}(\mu_b)$
- More statistically significant treatment of constraints.
- Uncertainties ΔX_{th} from SM analysis.

Class-I observables: inclusive $\mathcal{B}(\bar{B} \rightarrow X_s \gamma)$

Class-I : only depending on $C_7, C_{7'}$, related to radiative decays

[Misiak, Gambino, Steinhauser...]

$$\mathcal{B}(\bar{B} \rightarrow X_s \gamma)_{E_\gamma > 1.6 \text{ GeV}}^{\text{exp}} = (3.55 \pm 0.24 \pm 0.09) \times 10^{-4}$$

$$\mathcal{B}(\bar{B} \rightarrow X_s \gamma)_{E_\gamma > 1.6 \text{ GeV}}^{\text{th}} = [a_{(0,0)} + a_{(7,7)} [(\delta C_7)^2 + (\delta C_{7'})^2] + a_{(0,7)} \delta C_7 + a_{(0,7')} \delta C_{7'}] \times 10^{-4}$$

$$\mathcal{B}(\bar{B} \rightarrow X_s \gamma)_{E_\gamma > 1.6 \text{ GeV}}^{\text{SM}} = (3.15 \pm 0.23) \times 10^{-4}$$

- SM value $[a_{(0,0)}]$ expressed as

$$\mathcal{B}(B \rightarrow X_s \gamma)_{E_\gamma > E_0}^{\text{SM}} = \mathcal{B}(B \rightarrow X_c e \bar{\nu}) \left| \frac{V_{ts}^* V_{tb}}{V_{cb}} \right|^2 \frac{6\alpha_{\text{em}}}{\pi} [P(E_0) + N(E_0)]$$
$$P(E_0) = \sum_{i,j=1\dots 8} C_i^{\text{eff}}(\mu) C_j^{\text{eff}*}(\mu) K_{ij}(E_0, \mu)$$

- numerical a 's reproducing [Misiak, Steinhauser, Haisch]

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Class-I observables: isospin asymmetry in $B \rightarrow K^* \gamma$

$$A_I(B \rightarrow K^* \gamma) = \frac{\Gamma(\bar{B}^0 \rightarrow \bar{K}^{*0} \gamma) - \Gamma(B^- \rightarrow K^{*-} \gamma)}{\Gamma(\bar{B}^0 \rightarrow \bar{K}^{*0} \gamma) + \Gamma(B^- \rightarrow K^{*-} \gamma)}$$

- NLO QCDF: isospin asymmetry from nonfactorisable contributions: spectator quark emits the photon
- thus no change once chirally-flipped operators included, apart from normalisation to isospin-averaged branching ratio
- Strong discriminator of the sign of C_7 [Descotes, D. Ghosh, JM, M. Ramon 2011]. Excellent agreement SM-experiment

$$A_I(B \rightarrow K^* \gamma)^{exp} = 0.052 \pm 0.026$$

$$A_I(B \rightarrow K^* \gamma)^{th} = c \times \frac{\sum_k d_k (\delta C_7)^k}{\sum_{k,l} e_{k,l} (\delta C_7)^k (\delta C_{7'})^l} \pm \delta c.$$

$$A_I(B \rightarrow K^* \gamma)^{SM} = 0.041 \pm 0.025 \text{ (updates Feldman\&JM'03)}$$

- c, d, e determined numerically.

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Class-I observables: $B \rightarrow K^* \gamma$ CP-asymmetry

[Beneke Feldmann Seidel, Ball and Zwicky]

$$\frac{\Gamma(\bar{B}^0(t) \rightarrow \bar{K}^{*0} \gamma) - \Gamma(B^0(t) \rightarrow K^{*0} \gamma)}{\Gamma(\bar{B}^0(t) \rightarrow \bar{K}^{*0} \gamma) + \Gamma(B^0(t) \rightarrow K^{*0} \gamma)} = S_{K^* \gamma} \sin(\Delta m_B t) - C_{K^* \gamma} \cos(\Delta m_B t)$$

- Probe of photon helicity

$$S_{K^* \gamma} = \frac{2 \operatorname{Im} [e^{-2i\beta} (\mathcal{A}_L^* \bar{\mathcal{A}}_L + \mathcal{A}_R^* \bar{\mathcal{A}}_R)]}{|\mathcal{A}_L|^2 + |\mathcal{A}_R|^2 + |\bar{\mathcal{A}}_L|^2 + |\bar{\mathcal{A}}_R|^2}$$

- Computed at NLO in QCD factorisation. At LO,

$$S_{K^* \gamma}^{(\text{LO})} = \frac{-2 |C_{7'}/C_7|}{1 + |C_{7'}/C_7|^2} \sin(2\beta - \arg(C_7 C_{7'}))$$

[Grinstein et al, Bobeth et al]

$$S_{K^* \gamma}^{\text{exp}} = -0.16 \pm 0.22 (\text{HFAG})$$

$$S_{K^* \gamma} = f \begin{matrix} +\delta_f^u \\ -\delta_f^d \end{matrix} + \frac{\sum_{k,l} g_{k,l} (\delta C_7)^k (\delta C_{7'})^l}{\sum_{k,l} h_{k,l} (\delta C_7)^k (\delta C_{7'})^l}$$

$$S_{K^* \gamma}^{\text{SM}} = -0.03 \pm 0.01$$

- f, g, h fitting coefficients

Class-I observables: $B \rightarrow K^* \gamma$ CP-asymmetry

[Beneke Feldmann Seidel, Ball and Zwicky]

$$\frac{\Gamma(\bar{B}^0(t) \rightarrow \bar{K}^{*0} \gamma) - \Gamma(B^0(t) \rightarrow K^{*0} \gamma)}{\Gamma(\bar{B}^0(t) \rightarrow \bar{K}^{*0} \gamma) + \Gamma(B^0(t) \rightarrow K^{*0} \gamma)} = S_{K^* \gamma} \sin(\Delta m_B t) - C_{K^* \gamma} \cos(\Delta m_B t)$$

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Class-II observables: A_T^2 asymmetry

Class-II : depending only on dipole and semileptonic operators

$B \rightarrow K^* \ell^+ \ell^-$ asymmetry $A_T^2(q^2) = \frac{|A_\perp|^2 - |A_\parallel|^2}{|A_\perp|^2 + |A_\parallel|^2}$, [Kruger and J.M.]

- A_\perp and A_\parallel depend only on $C_{7,7',9,9',10,10'}$ (no tensors/scalars)
- strong potential to discriminate inside C_7^l allowed space.

At low q^2 , at NLO QCD factorisation $A_T^2(q^2) = A_T^{(2), CV}(q^2) \begin{matrix} +\delta_u(q^2) \\ -\delta_d(q^2) \end{matrix}$
with fitting q^2 -polynomials for errors δ_u, δ_d and central value

$$A_T^{(2), CV}(q^2) = \frac{\sum_{i=0,7,7',9,9',10,10'} \sum_{j=i,\dots,10'} k(q^2) F_{(i,j)}(q^2) \delta C_i \delta C_j}{\sum_{i=0,7,7',9,9',10,10'} \sum_{j=i,\dots,10'} k(q^2) G_{(i,j)}(q^2) \delta C_i \delta C_j}$$

[$\delta C_0 = 1$ to deal with constant, linear and quadratic terms]

Longer list of new clean class-II observables P_i in part II...

Class-II observables: A_T^2 asymmetry

Class-II : depending only on dipole and semileptonic operators

$B \rightarrow K^* \ell^+ \ell^-$ asymmetry $A_T^2(q^2) = \frac{|A_\perp|^2 - |A_\parallel|^2}{|A_\perp|^2 + |A_\parallel|^2}$, [Kruger and J.M.]

- A_\perp and A_\parallel depend only on $C_{7,7',9,9',10,10'}$ (no tensors/scalars)
- strong potential to discriminate inside C_7^I allowed space.

At low q^2 , at NLO QCD factorisation $A_T^2(q^2) = A_T^{(2), CV}(q^2)_{-\delta_d(q^2)}^{+\delta_u(q^2)}$
with fitting q^2 -polynomials for errors δ_u, δ_d and central value

$$A_T^{(2), CV}(q^2) = \frac{\sum_{i=0,7,7',9,9',10,10'} \sum_{j=i,..10'} k(q^2) F_{(i,j)}(q^2) \delta C_i \delta C_j}{\sum_{i=0,7,7',9,9',10,10'} \sum_{j=i,..10'} k(q^2) G_{(i,j)}(q^2) \delta C_i \delta C_j}$$

[$\delta C_0 = 1$ to deal with constant, linear and quadratic terms]

Longer list of new clean class-II observables P_i in part II...

Class-III observables: $\bar{B} \rightarrow X_s \mu^+ \mu^-$

Class-III: depending on dipole and semileptonic operators, but also others (scalar, tensors) \implies most of semileptonic observables

- $\bar{B} \rightarrow X_s \mu^+ \mu^-$ at low q^2 [1-6 GeV²]

$$\mathcal{B}(\bar{B} \rightarrow X_s \mu^+ \mu^-)^{\text{exp}} = (1.60 \pm 0.50) \times 10^{-6}$$

$$\mathcal{B}(B \rightarrow X_s \mu^+ \mu^-) = 10^{-7} \times \sum_{i,j=0,7,7',9,9',10,10'} b_{(i,j)} \delta C_i \delta C_j$$

$$\mathcal{B}(\bar{B} \rightarrow X_s \mu^+ \mu^-)^{\text{SM}} = (1.59 \pm 0.15) \times 10^{-6}$$

- $\delta C_7, \delta C_9, \delta C_{10}$ -only contributions known up to NNLO including e.m. corrections [Bobeth et al, Huber et al]
- $\delta C_{7'}, \delta C_{9'}, \delta C_{10}'$ -only contributions with similar structure ($\gamma_5 \rightarrow -\gamma_5$)
- crossed terms (primed-unprimed) only at LO in α_s , and are suppressed by m_s/m_b [Guetta Nardi]
- b coefficients determined numerically agreeing with [Huber et al]

Class-III observables: \tilde{A}_{FB} and \tilde{F}_L

$$A_{FB} = \left(\int_0^1 d(\cos\theta_l) \frac{d^2\Gamma}{dq^2 d\cos\theta_l} - \int_{-1}^0 \dots \right) / \frac{d\Gamma}{dq^2} \quad F_L = |A_0|^2 / \frac{d\Gamma}{dq^2}$$

The average forward-backward asymmetry \tilde{A}_{FB} (same for \tilde{F}_L) is

$$\tilde{A}_{FB} = \int_{1\text{GeV}^2}^{6\text{GeV}^2} \frac{d\Gamma}{dq^2} A_{FB}(q^2) dq^2 / \int_{1\text{GeV}^2}^{6\text{GeV}^2} \frac{d\Gamma}{dq^2}$$

$$\tilde{A}_{FB} = \frac{\int_{1\text{GeV}^2}^{6\text{GeV}^2} \sum_{i=0,7,7',9,9',10,10'} \sum_{j=i,\dots,10'} k(q^2) H_{(i,j)}(q^2) \delta C_i \delta C_j dq^2}{\int_{1\text{GeV}^2}^{6\text{GeV}^2} \sum_{i=0,7,7',9,9',10,10'} \sum_{j=i,\dots,10'} k(q^2) l_{(i,j)}(q^2) \delta C_i \delta C_j dq^2} \begin{matrix} +\tilde{\delta}_u \\ -\tilde{\delta}_d \end{matrix}$$

computed at NLO in QCD factorisation with fitting q^2 -polynomials for central value and errors (same for \tilde{F}_L)

$$\tilde{A}_{FB}^{SM} = -0.049 \pm 0.046 \quad \tilde{F}_L^{SM} = 0.721 \pm 0.043$$

Class-III observables: \tilde{A}_{FB} and \tilde{F}_L

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$$\tilde{A}_{FB} = \int_{1\text{GeV}^2}^{6\text{GeV}^2} \frac{d\Gamma}{dq^2} A_{FB}(q^2) dq^2 / \int_{1\text{GeV}^2}^{6\text{GeV}^2} \frac{d\Gamma}{dq^2}$$

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Standard Model values

In the SM, NNLO in $\overline{\text{MS}}$ including electromagnetic corrections
[Chetyrkin, Misiak and Münz, Bobeth et al., Huber et al.]

$C_1(\mu_b)$	$C_2(\mu_b)$	$C_3(\mu_b)$	$C_4(\mu_b)$	$C_5(\mu_b)$
-0.263	1.011	-0.006	-0.081	0.000
$C_6(\mu_b)$	$C_7^{\text{eff}}(\mu_b)$	$C_8^{\text{eff}}(\mu_b)$	$C_9(\mu_b)$	$C_{10}(\mu_b)$
0.001	-0.292	-0.166	4.075	-4.308

- High-scale $\mu_0 = 2M_W$ [uncertainty: varied from M_W to $4M_W$]
- Low-scale $\mu_b = 4.8$ GeV [uncertainty: varied from 2.4 to 9.6 GeV]

For the chirally-flipped operators, we have the SM values

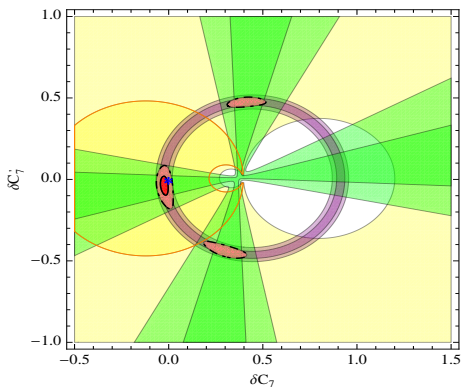
$$C_{7'}^{SM} = \frac{m_s}{m_b} C_7^{SM}, \quad C_{9',10'}^{SM} = 0$$

Exploring New Physics Constraints

on $C_i^{(\prime)}$:

Scenario A, B and C

$\delta C_7 - \delta C_{7'}$ plane : constraints at 1 and 2 σ



Class I observables (only $O_{7,7'}$)
dark 1 σ , light 2 σ

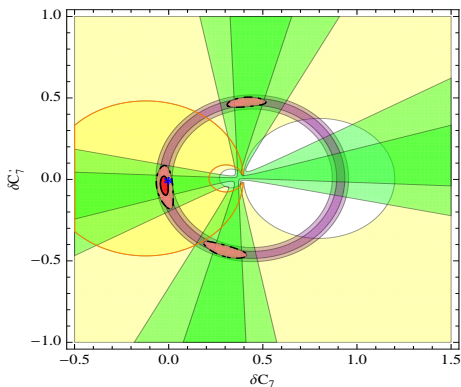
- A_I (yellow)
- $B(B \rightarrow X_s \gamma)$ (purple)
- $S_{K^* \gamma}$ (green)

Overlap regions (red dark and light)

- SM region at 1 σ solid black contour red dark $(C_7, C_{7'}) \sim (C_7^{SM}, 0)$.
- two non-SM solutions also allowed at 2 σ $(C_7, C_{7'}) \simeq (0, \pm 0.4)$ (dashed)

- A_I disfavours flipped-sign solution $(C_7, C_{7'}) = (-C_7^{SM}, 0)$
 \implies Same conclusion as [Gambino, Haisch, Misiak],
 without using Class-III $B \rightarrow X_s \ell^+ \ell^-$
- Flipped sign solution disfavored in Scenario A by $> 2\sigma$.

$\delta C_7 - \delta C_{7'}$ plane : constraints at 1 and 2 σ



Class I observables (only $O_{7,7'}$)
dark 1 σ , light 2 σ

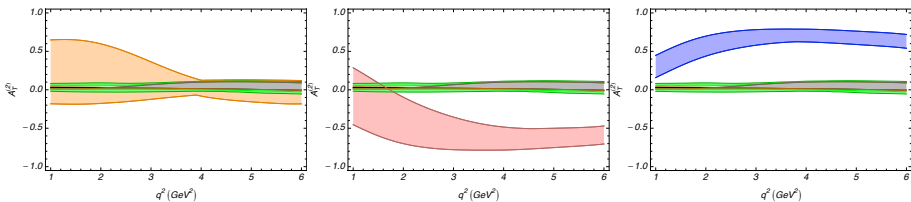
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- Flipped sign solution disfavored in Scenario A by $> 2\sigma$.

Scenario A ($C_{7,7'}$): prediction for class-II observable A_T^2

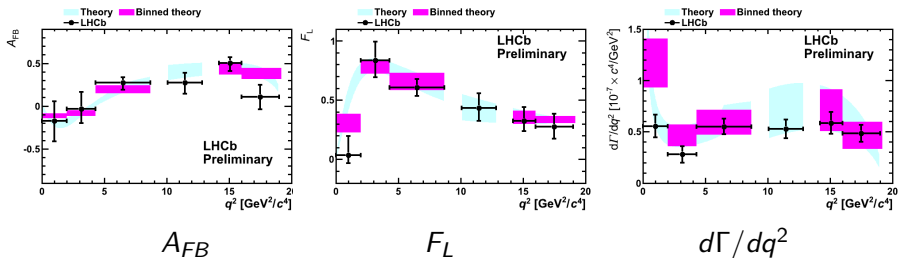


- Mapping of the three black allowed regions under Scenario A ($C_7, C_{7'}$) into A_T^2 only restricted by class-I observables: $B \rightarrow X_s \gamma$, A_I , $S_{K^* \gamma}$.
- $A_T^2(q^2)$ for $q^2 = 1 \dots 6$ GeV² has different shapes for the three regions in ($C_7, C_{7'}$)
 - $(\delta C_7, \delta C_{7'}) \simeq (0, 0)$ (left)
 - $(\delta C_7, \delta C_{7'}) \simeq (0.3, -0.4)$ (center)
 - $(\delta C_7, \delta C_{7'}) \simeq (0.3, 0.4)$ (right)

Notice that for the two non-SM regions $C_7 \sim 0$. In this case (opposite to $C_7 \sim C_7^{SM}$ case) positive A_T^2 is for $C_7' > 0$ and negative for $C_7' < 0$.

LHCb results EPS11 \rightarrow LHCb Moriond

At EPS11, new results from LHCb on $B_s \rightarrow \mu\mu$ and $B \rightarrow K^* \ell\ell$



Including LHCb in the world average

$$\tilde{A}_{FB} = 0.33_{-0.24}^{+0.22} \rightarrow 0.04 \pm 0.12 \rightarrow -0.130_{-0.078}^{+0.068}$$

$$\tilde{F}_L = 0.60_{-0.19}^{+0.18} \rightarrow 0.60 \pm 0.09 \rightarrow 0.622_{-0.057}^{+0.059}$$

What is the impact of LHCb results alone from F_L and A_{FB} ?

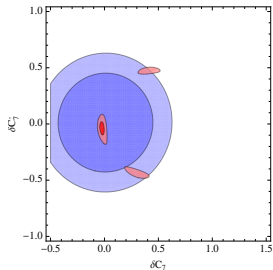
- Still same LHC constraints on $C_7, C_{7'}$ from $b \rightarrow s\gamma$
- Different impact for Scenarios A and B

Scenario A ($C_{7,7'}$): class-III observables (Moriond12) 1-2 σ

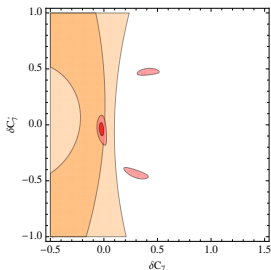
Scenario A:

\implies class-III observables constrain further the shifts $\delta C_7, \delta C_{7'}$

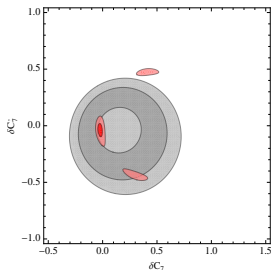
Correlations in $(\delta C_7, \delta C_{7'})$ plane:



$B(B \rightarrow X_s \mu^+ \mu^-)$



\tilde{A}_{FB}

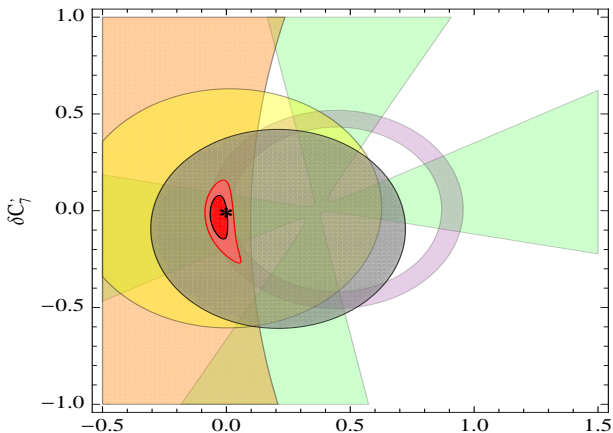


\tilde{F}_L

- $B(B \rightarrow X_s \mu^+ \mu^-)$ favors SM-like region (at 1σ) and non-SM regions (at 2σ). All three disfavors flipped sign solution at more than 2σ .
- \tilde{A}_{FB} selects SM region and \tilde{F}_L favours SM region and the lower non-SM region.

Scenario A ($C_{7,\gamma}$): class-III observables (Moriond12) 1-2 σ

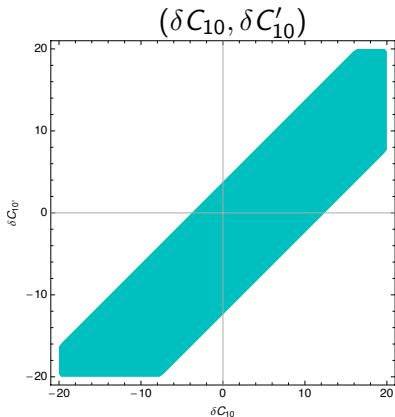
Scenario A: Combined constrain from class I & III observables



Purple ($B \rightarrow X_s \gamma$), Grey (F_L), Orange (A_{FB}), Yellow ($B \rightarrow X_s II$), Green ($S_{K^* \gamma}$). All constraints shown here at 2 sigma. Good agreement of all observables with SM region at 1 σ . No non-SM region allowed at < 2 σ .

Constraint on $C_{10}, C_{10'}$ from $B_s \rightarrow \mu^+ \mu^-$

$$\mathcal{B}(\bar{B}_s \rightarrow \mu^+ \mu^-) = \frac{G_F^2 \alpha^2}{16\pi^3} f_{B_s}^2 m_{B_s} \tau_{B_s} |V_{tb} V_{ts}^*|^2 m_\mu^2 \sqrt{1 - \frac{4m_\mu^2}{m_{B_s}^2}} |C_{10} - C_{10'}|^2$$



Using our inputs, we get

$$\mathcal{B}(\bar{B}_s \rightarrow \mu^+ \mu^-)^{\text{SM}} = (3.44 \pm 0.32) \cdot 10^{-9}$$

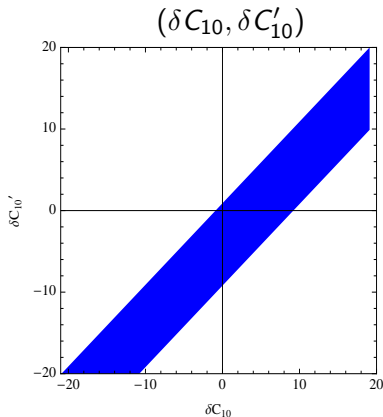
one order of magnitude smaller than
90% CL LHCb exp bound

$$\mathcal{B}(\bar{B}_s \rightarrow \mu^+ \mu^-)^{\text{exp}} < 1.2 \cdot 10^{-8}.$$

leading to weak constraints on C_{10}
(Scenario B) and $C_{10'}$ (Scenario C)

Constraint on $C_{10}, C_{10'}$ from $B_s \rightarrow \mu^+ \mu^-$

$$\mathcal{B}(\bar{B}_s \rightarrow \mu^+ \mu^-) = \frac{G_F^2 \alpha^2}{16\pi^3} f_{B_s}^2 m_{B_s} \tau_{B_s} |V_{tb} V_{ts}^*|^2 m_\mu^2 \sqrt{1 - \frac{4m_\mu^2}{m_{B_s}^2}} |C_{10} - C_{10'}|^2$$



Using our inputs ($f_{B_s} = 227.7 \text{ MeV}$)

$$\mathcal{B}(\bar{B}_s \rightarrow \mu^+ \mu^-)^{\text{SM}} = (3.2 \pm 0.3) \cdot 10^{-9}$$

near one order of magnitude smaller than previous exp bound

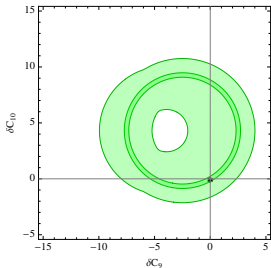
$$\mathcal{B}(\bar{B}_s \rightarrow \mu^+ \mu^-)^{\text{exp}} < 4.5 \cdot 10^{-9}.$$

leading to strong constraints on C_{10} (Scenario B) and $C_{10'}$ (Scenario C)

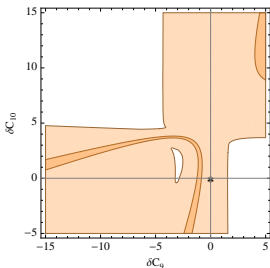
Scenario B ($C_{7,7',9,10}$): class-III constraints in $(\delta C_9, \delta C_{10})$

In Scenario B, NP in

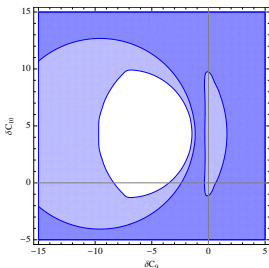
- $C_7, C_{7'}$: same constraints/plot as before from class-I obs., but NOW three red regions allowed at 2σ (class-III does no cut)
- Correlation $\delta C_9, \delta C_{10}$: to be fixed from class-III observables:



$B(B \rightarrow X_s \mu^+ \mu^-)$



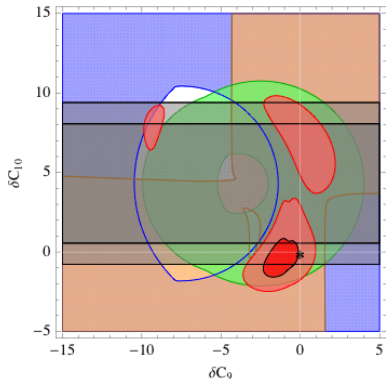
\tilde{A}_{FB}



\tilde{F}_L

- Small absolute values of (C_9, C_{10}) disfavoured by \tilde{F}_L .

Scenario B: overlap in $\delta C_9 - \delta C_{10}$ plane



- $\mathcal{B}(B \rightarrow X_s \mu^+ \mu^-)$ (green)
- A_{FB} (orange)
- F_L (blue)

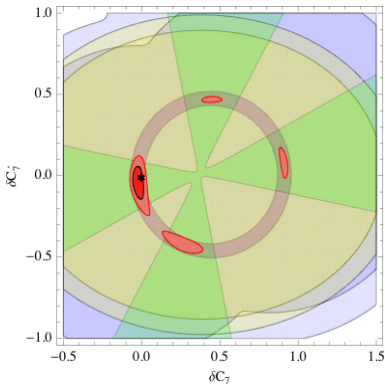
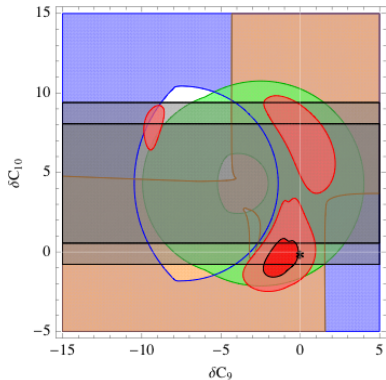
Two overlapped region (red)

- SM region around $(\delta C_9, \delta C_{10}) = (0, 0)$
- non-SM region $(C_9, C_{10}) \simeq (C_9^{\text{SM}}, -C_{10}^{\text{SM}})$

\implies Scenario B NP may alter (C_7, C_7') but also (C_9, C_{10}) from their SM values to get a tiny improvement with data and reproduce the experimental values. Non-SM regions allowed at 2 sigma.

$\implies \text{BR}(B_s \rightarrow \mu^+ \mu^-)$ cuts further in the allowed region.

Scenario B: overlap in $\delta C_9 - \delta C_{10}$ plane



\implies Scenario B NP may alter (C_7, C_7') but also (C_9, C_{10}) from their SM values to get a tiny improvement with data and reproduce the experimental values. Non-SM regions allowed at 2 sigma.

$\implies \text{BR}(B_s \rightarrow \mu^+ \mu^-)$ cuts further in the allowed region.

Angular distribution of

$$B \rightarrow K^* (\rightarrow K \pi) l^+ l^-:$$

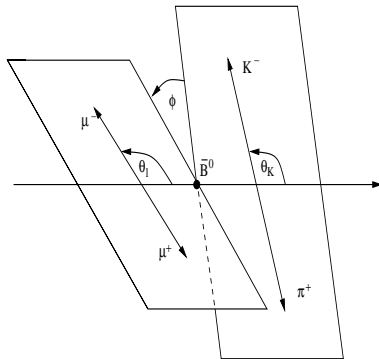
RH currents/new scalars/CPV in WC

Differential decay distributions

The decay $\bar{B}_d \rightarrow \bar{K}^{*0}(\rightarrow K^- \pi^+) l^+ l^-$ with the K^{*0} on the mass shell is described by s and three angles θ_l , θ_K and ϕ

$$\frac{d^4\Gamma}{dq^2 d\cos\theta_l d\cos\theta_K d\phi} = \frac{9}{32\pi} J(q^2, \theta_l, \theta_K, \phi)$$

- $q^2 = s$ square of the lepton-pair invariant mass.
- θ_l angle between $p_{l^-}^{\vec{}}$ in $l^+ l^-$ rest frame and dilepton's direction in rest frame of \bar{B}_d
- θ_K angle between $p_{K^-}^{\vec{}}$ in the \bar{K}^{*0} rest frame and direction of the \bar{K}^{*0} in rest frame of \bar{B}_d
- ϕ angle between the planes defined by the two leptons and the $K - \pi$ planes.



Differential decay distributions

The decay $\bar{B}_d \rightarrow \bar{K}^{*0}(\rightarrow K^- \pi^+) I^+ I^-$ with the K^{*0} on the mass shell is described by s and three angles θ_I , θ_K and ϕ

$$\frac{d^4\Gamma}{dq^2 d\cos\theta_I d\cos\theta_K d\phi} = \frac{9}{32\pi} J(q^2, \theta_I, \theta_K, \phi)$$

The differential distribution splits in J_i coefficients:

$$\begin{aligned} J(q^2, \theta_I, \theta_K, \phi) = & \\ & J_{1s} \sin^2 \theta_K + J_{1c} \cos^2 \theta_K + (J_{2s} \sin^2 \theta_K + J_{2c} \cos^2 \theta_K) \cos 2\theta_I + J_3 \sin^2 \theta_K \sin^2 \theta_I \cos 2\phi \\ & + J_4 \sin 2\theta_K \sin 2\theta_I \cos \phi + J_5 \sin 2\theta_K \sin \theta_I \cos \phi + (J_{6s} \sin^2 \theta_K + J_{6c} \cos^2 \theta_K) \cos \theta_I \\ & + J_7 \sin 2\theta_K \sin \theta_I \sin \phi + J_8 \sin 2\theta_K \sin 2\theta_I \sin \phi + J_9 \sin^2 \theta_K \sin^2 \theta_I \sin 2\phi. \end{aligned}$$

The information on

- the helicity/transversity amplitudes of the K^* ($H_{\pm 1,0}$ or $A_{\perp,\parallel,0}$) is inside the coefficients J_i .
- short distance physics C_i is encoded in ($H_{\pm 1,0}$ or $A_{\perp,\parallel,0}$)

$$J_{1s} = \frac{(2 + \beta_\ell^2)}{4} \left[|A_{\perp}^L|^2 + |A_{\parallel}^L|^2 + (L \rightarrow R) \right] + \frac{4m_\ell^2}{q^2} \text{Re} \left(A_{\perp}^L A_{\perp}^{R*} + A_{\parallel}^L A_{\parallel}^{R*} \right),$$

$$J_{1c} = |A_0^L|^2 + |A_0^R|^2 + \frac{4m_\ell^2}{q^2} \left[|A_t|^2 + 2\text{Re}(A_0^L A_0^{R*}) \right] + \beta_\ell^2 |A_S|^2,$$

$$J_{2s} = \frac{\beta_\ell^2}{4} \left[|A_{\perp}^L|^2 + |A_{\parallel}^L|^2 + (L \rightarrow R) \right], \quad J_{2c} = -\beta_\ell^2 \left[|A_0^L|^2 + (L \rightarrow R) \right],$$

$$J_3 = \frac{1}{2} \beta_\ell^2 \left[|A_{\perp}^L|^2 - |A_{\parallel}^L|^2 + (L \rightarrow R) \right], \quad J_4 = \frac{1}{\sqrt{2}} \beta_\ell^2 \left[\text{Re}(A_0^L A_{\parallel}^{L*}) + (L \rightarrow R) \right],$$

$$J_5 = \sqrt{2} \beta_\ell \left[\text{Re}(A_0^L A_{\perp}^{L*}) - (L \rightarrow R) - \frac{m_\ell}{\sqrt{q^2}} \text{Re}(A_{\parallel}^L A_S^* + A_{\parallel}^R A_S^*) \right],$$

$$J_{6s} = 2\beta_\ell \left[\text{Re}(A_{\parallel}^L A_{\perp}^{L*}) - (L \rightarrow R) \right], \quad J_{6c} = 4\beta_\ell \frac{m_\ell}{\sqrt{q^2}} \text{Re} \left[A_0^L A_S^* + (L \rightarrow R) \right],$$

$$J_7 = \sqrt{2} \beta_\ell \left[\text{Im}(A_0^L A_{\parallel}^{L*}) - (L \rightarrow R) + \frac{m_\ell}{\sqrt{q^2}} \text{Im}(A_{\perp}^L A_S^* + A_{\perp}^R A_S^*) \right],$$

$$J_8 = \frac{1}{\sqrt{2}} \beta_\ell^2 \left[\text{Im}(A_0^L A_{\perp}^{L*}) + (L \rightarrow R) \right], \quad J_9 = \beta_\ell^2 \left[\text{Im}(A_{\parallel}^{L*} A_{\perp}^L) + (L \rightarrow R) \right]$$

SCALARS: We have 8 complex amplitudes ($A_{\perp, \parallel, 0, (L, R) S, t}$) and 12 experimental inputs

NO SCALARS: We have 7 complex amplitudes ($A_{\perp, \parallel, 0, (L, R), t}$) and 11 experimental inputs

ZOO of observables in the market:

Observables strongly sensitive to hadronic soft form factors (SFFD):

A_{FB} (forward-backward asymmetry), F_L (longitudinal polarization fraction), A_{im} (Egede et al. 2008), $\frac{d\Gamma}{dq^2}$, J_i , S_i and A_i (Altmannshofer et al. 2009).

SFFI observables (at LO soft form factors cancels exactly):

A_T^2 Transverse asymmetry: Krueger-J.M (2005) A_T^3 and A_T^4 Asymmetries with Longitudinal sensitivity: Egede, Hurth, JM, Ramon, Reece (2008) A_T^5 Second Transverse asymmetry: Egede, Hurth, JM, Ramon, Reece (2010) A_T^{re} related to A_T^5 : Becirevic, Schneider (2011) A_T^{im} related to A^{im} : Becirevic, Schneider (2011)

Can one extract **all** the information from the angular distribution in an **efficient**, systematic and **clean** way?

Geometrical interpretation of angular distribution ($m_l = 0$)

Define

$$n_{\parallel} = (A_{\parallel}^L, A_{\parallel}^{R*}) \qquad m_1 = (H_{+1}^L, H_{-1}^{R*})$$

$$n_{\perp} = (A_{\perp}^L, -A_{\perp}^{R*}) \qquad \text{or} \qquad m_2 = (H_{-1}^L, H_{+1}^{R*})$$

$$n_0 = (A_0^L, A_0^{R*}) \qquad m_3 = (H_0^L, H_0^{R*})$$

Spin amplitudes

Helicity amplitudes

All physical information of the distribution encoded in 3 moduli + 3 relative angles (complex) - 1 constraint (**third relation**).

$$|n_{\parallel}|^2 = \frac{2}{3}J_{1s} - J_3, \quad |n_{\perp}|^2 = \frac{2}{3}J_{1s} + J_3, \quad |n_0|^2 = J_{1c}$$
$$n_{\perp}^{\dagger} n_{\parallel} = \frac{J_{6s}}{2} - iJ_9, \quad n_0^{\dagger} n_{\parallel} = \sqrt{2}J_4 - i\frac{J_7}{\sqrt{2}}, \quad n_0^{\dagger} n_{\perp} = \frac{J_5}{\sqrt{2}} - i\sqrt{2}J_8$$

Those are the building blocks of any observable (only 8 independent)

Constraint:

$$|(n_{\parallel}^{\dagger} n_{\perp})|n_0|^2 - (n_{\parallel}^{\dagger} n_0)(n_0^{\dagger} n_{\perp})|^2 = (|n_0|^2 |n_{\parallel}|^2 - |n_0^{\dagger} n_{\parallel}|^2)(|n_0|^2 |n_{\perp}|^2 - |n_0^{\dagger} n_{\perp}|^2)$$

This translates into an experimental test between J_i :

$$-J_{2c} = 6 \frac{(2J_{1s} + 3J_3)(4J_4^2 + J_7^2) + (2J_{1s} - 3J_3)(J_5^2 + 4J_8^2)}{16J_{1s}^2 - 9(4J_3^2 + J_{6s}^2 + 4J_9^2)} - 36 \frac{J_{6s}(J_4J_5 + J_7J_8) + J_9(J_5J_7 - 4J_4J_8)}{16J_{1s}^2 - 9(4J_3^2 + J_{6s}^2 + 4J_9^2)} \equiv \mathbf{f}$$

A second way of showing this is to count degrees of freedom and symmetries of the distribution.

Counting d.o.f. : Primary Observables

Experimental (J_i) \leftrightarrow theoretical (A_i) degrees of freedom

$$n_J - n_d = 2n_A - n_s$$

- n_J : # coefficients of differential distribution: J_i
- n_d : # relations between J_i
- n_A : # spin amplitudes
- n_s : # symmetries of the distribution

Case: Massless leptons with no scalars:

$n_J = 11$, $n_d = 3$ ($J_{1s} = 3J_{2s}$, $J_{1c} = -J_{2c}$ and the **third relation**),
 $n_A = 6$ (spin amplitudes), $n_s = 4$ **symmetries**.

Independent experimental inputs: $11-3=8=9$ b.b.-1 constraint

The most general case (massive leptons + scalars) is presented in
[J.M, F. Mescia, M.Ramon, J.Virto'12.](#)

Table: The dependencies between the coefficients in the differential distribution and the symmetries between the amplitudes in several cases.

Case	Coefficients	Dependencies	Amplitudes	Symmetries
$m_\ell = 0, A_S = 0$	11	3	6	4
$m_\ell = 0$	11	2	7	5
$m_\ell > 0, A_S = 0$	11	1	7	4
$m_\ell > 0$	12	0	8	4

All symmetries (massive with scalars) are known and described in J.M, F. Mescia, M.Ramon, J.Virto'12.

Is there a systematic way of extracting the maximally clean information from Angular Distributions?

or

Is there a basis of OBSERVABLES that covers all the information?

$$O = \left\{ \frac{d\Gamma}{dq^2}, A_{FB}, P_1, P_2, P_3, P_4, P_5, P_6 \right\}$$

- **SFFD Observables:** A_{FB} (or F_L) and $\frac{d\Gamma}{dq^2}$
- **Basis of Clean (SFFI) Observables:**

$$P_1 = \frac{|n_{\perp}|^2 - |n_{\parallel}|^2}{|n_{\perp}|^2 + |n_{\parallel}|^2} = \frac{J_3}{2J_{2s}} \quad P_4 = \frac{\text{Re}(n_0^{\dagger} n_{\parallel})}{\sqrt{|n_{\parallel}|^2 |n_0|^2}} = \frac{\sqrt{2}J_4}{\sqrt{-J_{2c}(2J_{2s} - J_3)}}$$
$$P_2 = \frac{\text{Re}(n_{\perp}^{\dagger} n_{\parallel})}{|n_{\perp}|^2 + |n_{\parallel}|^2} = \frac{J_{6s}}{8J_{2s}} \quad P_5 = \frac{\text{Re}(n_0^{\dagger} n_{\perp})}{\sqrt{|n_{\perp}|^2 |n_0|^2}} = \frac{J_5}{\sqrt{-2J_{2c}(2J_{2s} + J_3)}}$$
$$P_3 = \frac{\text{Im}(n_{\perp}^{\dagger} n_{\parallel})}{|n_{\perp}|^2 + |n_{\parallel}|^2} = \frac{-J_9}{4J_{2s}} \quad P_6 = \frac{\text{Im}(n_0^{\dagger} n_{\parallel})}{\sqrt{|n_{\parallel}|^2 |n_0|^2}} = \frac{-J_7}{\sqrt{-2J_{2c}(2J_{2s} - J_3)}}$$

- $P_i = 1..6$ form a basis for all clean observables.
- Two dirty + $P_{i=1..6}$ can generate any observable.
- If $J_{6c} \sim 0$ (no scalars) $P_{1,2,3,4,6}$ and P_5 fit $C_{7,7',9,10,9',10'}$.

Examples (clean ones in the clean basis):

$$\begin{aligned}
 A_T^{(2)} &= P_1 & A_T^{(re)} &= 2P_2 \\
 A_T^{(im)} &= -2P_3 & A_T^5 &= \sqrt{\frac{1 - P_1^2 - 4P_2^2 - 4P_3^2}{4}} \\
 A_T^3 &= f_1(P_i) & A_T^4 &= f_2(P_i)
 \end{aligned}$$

Examples of SFFD ones ($m_\ell = 0$):

$$A_{im} = -F_T P_3 \quad A_{FB} = -\frac{3}{2} F_T P_2 \quad F_L = 1 + \frac{2A_{FB}}{3P_2}$$

For the same reason better use $P_1 = A_T^2$ (FFI) than S_3 (FFD).

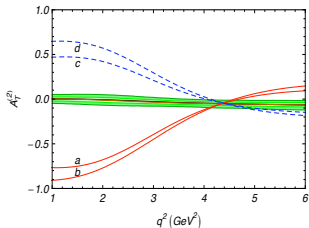
Why $A_T^2 = P_1$ is better than A_{FB} ? Why $A_{FB}|_{SM} \not\rightarrow A_T^2|_{SM}$?

Definition

Kruger, J.M. '05

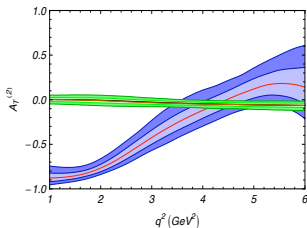
$$A_T^2 = \frac{|A_{\perp}|^2 - |A_{\parallel}|^2}{|A_{\perp}|^2 + |A_{\parallel}|^2} = -2 \frac{\text{Re}H_+^* H_-}{|H_+|^2 + |H_-|^2}$$

- Physics: Deviation from SM LH structure: $A_T^2|_{SM} \sim 0$ (from $A_{\perp} = -A_{\parallel}$).
- Absence of impact of RH currents in A_{FB} does not prevent a large A_T^2 .
- Domain: Low-Region $1 \leq q^2 \leq 6 \text{ GeV}^2$ (High region, see G. Hiller et al.)



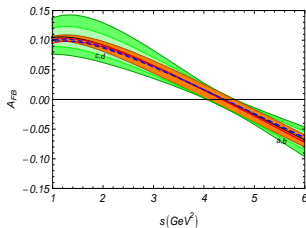
Susy [Lunghi, J.M. '07]

Λ/m_b : light(dark) green $\pm 5\%$ ($\pm 10\%$)



Exp. sens. susy (10fb^{-1})

light(dark) blue 1σ (2σ)



A_{FB} + RH currents

(Egede et al. '08)

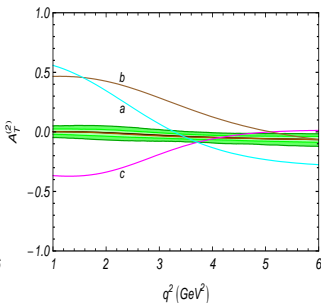
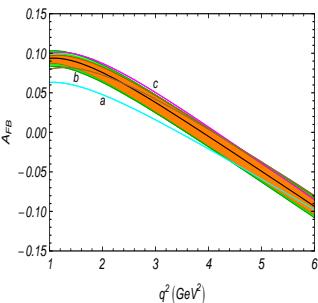
Other sensitivities of A_T^2 : CPV in $O_7 - O_7'$ and O_{10}'

- A_T^2 : CP violating phase (O_7') sensitivity BETTER than CP violating observables
- A_{FB} : **Mild** sensitivity to C_7' mod+phase A_T^2 : **Strong** sensitivity to C_7' mod+phase

$$\text{Num}(A_{FB}) \sim \frac{2m_b M_B}{q^2} C_7^{\text{eff}} + C_9 + \frac{2m_b M_B}{q^2} |C_7^{\text{NP}}| \cos \phi_7^{\text{NP}}$$

$$\text{Num}(A_T^2) = \frac{4m_b M_B}{q^2} \left[\left(\frac{2m_b M_B}{q^2} C_7^{\text{eff}} + C_9 \right) |C_7'| \cos \phi_7' + \frac{2m_b M_B}{q^2} |C_7'| |C_7^{\text{NP}}| \cos(\phi_7' - \phi_7^{\text{NP}}) \right]$$

- If only O_{10}' turned on A_T^2 has a different and characteristic q^2 -dependence for O_{10}' than for O_7 : no zero and maximal deviation around the AFB zero.



$C_7^{\text{NP}} e^{i\phi_7^{\text{NP}}}$	$C_7' e^{i\phi_7'}$
$0.26 e^{-i\frac{7\pi}{16}}$	$0.2 e^{i\pi}$ (a)
$0.07 e^{i\frac{3\pi}{5}}$	$0.3 e^{i\frac{3\pi}{5}}$ (b)
$0.03 e^{i\pi}$	0.07 (c)

Other sensitivities of $A_T^{(2)}$: CPV in $O_7 - O_7'$ and O_{10}'

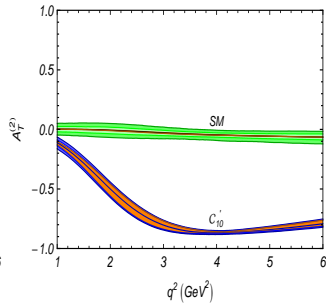
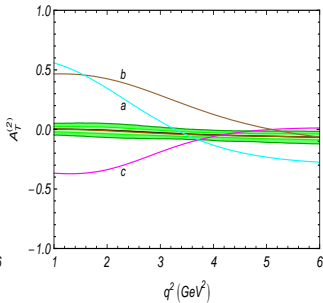
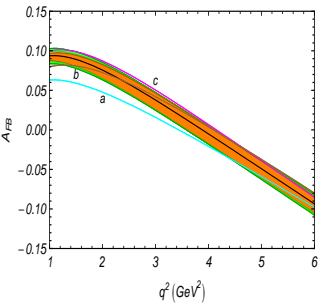
- $A_T^{(2)}$: CP violating phase (O_7') sensitivity BETTER than CP violating observables

A_{FB} : Mild sensitivity to C_7' mod+phase $A_T^{(2)}$: Strong sensitivity to C_7' mod+phase

$$\text{Num}(A_{FB}) \sim \frac{2m_b M_B}{q^2} C_7^{\text{eff}} + C_9 + \frac{2m_b M_B}{q^2} |C_7^{\text{NP}}| \cos \phi_7^{\text{NP}}$$

$$\text{Num}(A_T^{(2)}) = \frac{4m_b M_B}{q^2} \left[\left(\frac{2m_b M_B}{q^2} C_7^{\text{eff}} + C_9 \right) |C_7'| \cos \phi_7' + \frac{2m_b M_B}{q^2} |C_7'| |C_7^{\text{NP}}| \cos(\phi_7' - \phi_7^{\text{NP}}) \right]$$

- If only O_{10}' turned on $A_T^{(2)}$ has a different and characteristic q^2 -dependence for O_{10}' than for O_7 : no zero and maximal deviation around the AFB zero.



Measuring A_T^2

- Projection fits on each angle:

$$\frac{d\Gamma'}{d\phi} = \frac{\Gamma'}{2\pi} \left(1 + \frac{1}{2}(1 - F_L) \mathbf{A}_T^{(2)} \cos 2\phi + A_{\text{im}} \sin 2\phi \right),$$

$$\frac{d\Gamma'}{d\theta_l} = \Gamma' \left(\frac{3}{4} F_L \sin^2 \theta_l + \frac{3}{8} (1 - F_L) (1 + \cos^2 \theta_l) + A_{\text{FB}} \cos \theta_l \right) \sin \theta_l,$$

$$\frac{d\Gamma'}{d\theta_K} = \frac{3\Gamma'}{4} \sin \theta_K (2F_L \cos^2 \theta_K + (1 - F_L) \sin^2 \theta_K), \quad \Gamma' = d\Gamma/dq^2$$

Time schedule: during the first run ($1 - 2fb^{-1}$ enough).

- From full angular analysis with small bins, only two coefficients are necessary:

$$A_T^2 = \frac{|A_{\perp}|^2 - |A_{\parallel}|^2}{|A_{\perp}|^2 + |A_{\parallel}|^2} = \frac{\mathbf{J}_3}{2\mathbf{J}_{2s}} \text{ but indeed } \langle A_T^2 \rangle_{\text{bin}_i} = \frac{\int_{\text{bin}_i} F_T A_T^2 \Gamma'}{\int_{\text{bin}_i} F_T \Gamma'}$$

- Projection fits on each angle in the **new variables**:

$$\frac{d\Gamma'}{d\phi} = \frac{\Gamma'}{2\pi} \left(1 + \frac{1}{2} F_T \mathbf{P}_1 \cos 2\phi - F_T \mathbf{P}_3 \sin 2\phi \right),$$

$$\frac{d\Gamma'}{d\theta_l} = \Gamma' \sin \theta_l \left(\frac{3}{16} (3 - F_L) + \frac{3}{2} F_T \mathbf{P}_2 \cos \theta_l - \frac{3}{16} (2 - 3F_T) \cos 2\theta_l \right)$$

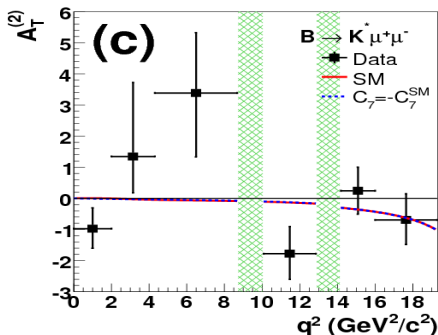
$$\frac{d\Gamma'}{d\theta_K} = \frac{3\Gamma'}{2} \sin \theta_K \left(F_L \cos^2 \theta_K + \frac{1}{2} F_T \sin^2 \theta_K \right), \quad \Gamma' = d\Gamma/dq^2$$

or if preferred $A_{FB} = -3F_T \mathbf{P}_2/2$.

- The full generalization of uniaxial distributions for massive leptons and scalars in JM, F. Mescia, M. Ramon, J. Virto'12.
- Also complete expressions of all J_i 's in terms of P_i .

CDF measurement of $A_T^{(2)}$

Finally, a first measurement from CDF has come out...



arXiv:1108.0695
[hep-ex] (more
precision soon)
and LHCb...

HOWEVER, it was obtained assuming zero isospin breaking between $B^0 \rightarrow K^{*0} \mu \mu$ and $B^+ \rightarrow K^{*+} \mu \mu$.

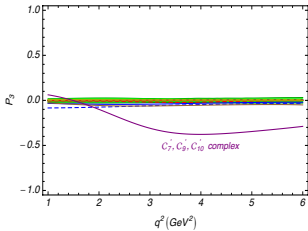
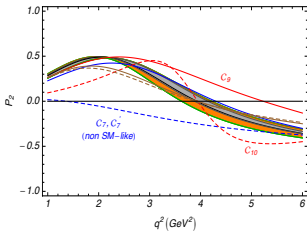
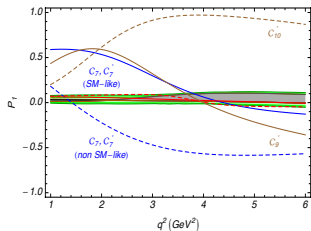
Sensitivities of $P_{1,2,3,4,5,6}$
and
massive $M_{1,2}$

$$P_i: \delta C_7, \delta C_7', \delta C_9, \delta C_{10}, \delta C_9', \delta C_{10}'$$

$$P_1 = A_T^2, C_{7,7',9',10'}$$

$$P_2 \text{ (Re)} C_{7,7',9,10}$$

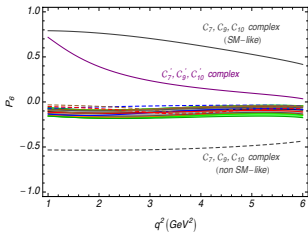
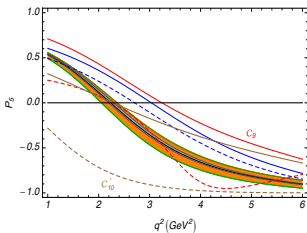
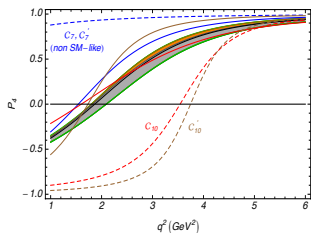
$$P_3 \text{ (Im)} \text{ Complex } C_{i'}$$



$$P_4 \text{ (Re)} C_{7,7',10,10'}$$

$$P_5 \text{ (Re), scalar } C_{9,10'}$$

$$P_6 \text{ (Im)} \text{ Complex } C_{i,i'}$$



Massive observables: $B \rightarrow K^*(\rightarrow K\pi)\tau^+\tau^-$

Massive case ($B \rightarrow K^*(\rightarrow K\pi)\tau^+\tau^-$) previous observables are trivially generalized with some prefactors (β_ℓ) in $P_{i=1\dots 6}$.

Two new specific observables comes out:

$$M_1 = \frac{4m_\ell^2}{q^2\beta_\ell^2} \frac{\text{Re}\left(A_\perp^L A_\perp^{R*} + A_\parallel^L A_\parallel^{R*}\right)}{[|A_\perp^L|^2 + |A_\parallel^L|^2 + (L \rightarrow R)]} = \frac{J_{1s} - J_{2s} \frac{(2+\beta_\ell^2)}{\beta_\ell^2}}{4J_{2s}}$$

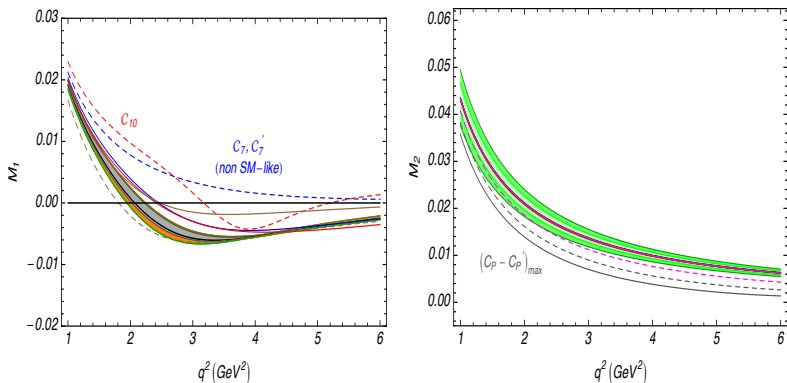
$$M_2 = \frac{4m_\ell^2}{q^2} \frac{[|A_t|^2 + 2\text{Re}(A_0^L A_0^{R*})]}{|A_0^L|^2 + |A_0^R|^2} = -\beta_\ell^2 \frac{J_{1c} + J_{2c} \frac{1}{\beta_\ell^2}}{J_{2c}}$$

Independent experimental inputs = 11-1 = **10**. Basis:

$$O = \left\{ \frac{d\Gamma}{dq^2}, A_{FB}, P_1, P_2, P_3, P_4, P_5, P_6, M_1, M_2 \right\}$$

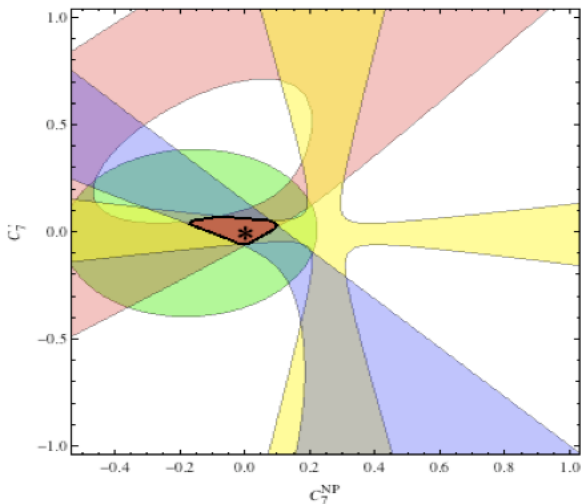
Sensitivities of massive observables $M_{1,2}$

In the case of muons we find for M_1^μ (left) and M_2^μ (right) [scalar]:



- M_1 exhibits some sensitivity to $C_{7,7'}$ and C_{10} even if very mild.
- M_2 is the only observable sensitive to Pseudoscalars.

In the not too far future we can expect to see this type of constraints from P_1 (yellow), P_2 (green), P_4 (red), P_5 (blue).



* SM solution.

Isospin violation

$$A_I(B \rightarrow K^* \Pi):$$

Definition:

[Feldmann, J.M '03]

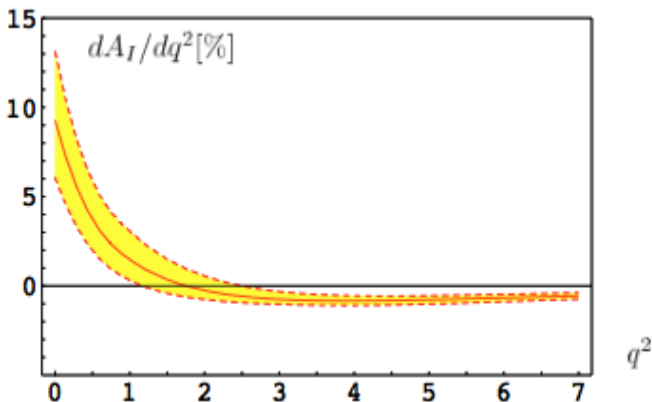
$$\frac{dA_I}{dq^2} \equiv \frac{d\Gamma[B^0 \rightarrow K^{*0} \ell^+ \ell^-]/dq^2 - d\Gamma[B^\pm \rightarrow K^{*\pm} \ell^+ \ell^-]/dq^2}{d\Gamma[B^0 \rightarrow K^{*0} \ell^+ \ell^-]/dq^2 + d\Gamma[B^\pm \rightarrow K^{*\pm} \ell^+ \ell^-]/dq^2} .$$

- Description for invariant mass of the lepton pair small:

$$1 \leq q^2 \leq 6 \text{ GeV}^2$$

- Systematic theoretical description using QCD factorization in the heavy quark limit.
- Sensitivity to NP via **spectator quark** in exclusive modes \neq inclusive counterparts in the short-distance dynamics.

For increasing values of q^2 the isospin-asymmetry decreases, and its central value becomes slightly negative above $q^2 = 2 \text{ GeV}^2$ and stays basically at a constant value of about -1%.



Non-factorizable (NF) graphs where a γ is radiated from spectator quark in **annihilation** or **spectator-scattering** diagrams.

- Contributions sensitive to the charge of the spectator quark.

$$C_9^\perp(q^2) = [1 + b_q^\perp(q^2)] C_9^{(0)\perp}(q^2), \quad C_9^\parallel(q^2) = [1 + b_q^\parallel(q^2)] C_9^{(0)\parallel}(q^2)$$

The functions $b_q^{\perp,\parallel}(q^2)$ parametrize NF effects from photon radiation from spectator quark.

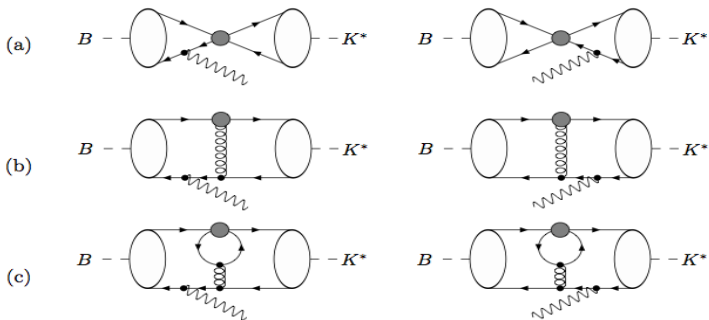
$$\frac{dA_I}{dq^2} [B \rightarrow K^* l^+ l^-] = \text{Re}(b_d^\perp - b_u^\perp) \times f(C_9^\perp, C_9^\parallel, C_{10}, b_{d,u}^{\perp,\parallel}, \xi_{\perp,\parallel})$$

- NF independent of the spectator quark drops out. It vanishes in naive factorization. NF effects tiny.
- In the limit $q^2 \rightarrow 0$ (photon pole in C_9^\perp dominates) and we recover KN:

$$A_I[B \rightarrow K^* \gamma] = \text{Re}[b_d^\perp(0) - b_u^\perp(0)]$$

For q^2 large longitudinal polarization dominates.

- Calculation requires to model IR divergences that are the main source of uncertainty.



(a) Annihilation topologies with operators \mathcal{O}_{1-6} , (b) Hard spectator interaction involving the gluonic penguin operator \mathcal{O}_8 , (c) Hard spectator interaction involving the operators \mathcal{O}_{1-6} . [T. Feldmann, J.M '02]

Main New Physics sensitivities of isospin asymmetry:

- The penguin operators O_{3-6} (dominant annihilation contribution) give the main effect to the isospin asymmetry in $B \rightarrow K^* \gamma$ and $B \rightarrow K^* I^+ I^-$ together with C_7 (its sign) and partially $C_{9,10}$.
 - $a_6^{(0)} = (\bar{C}_6 + \bar{C}_5/3)$ main impact at small values of q^2
 - $a_4^{(0)} = (\bar{C}_4 + \bar{C}_3/3)$ main impact at larger values of q^2
 - Sign of C_7^{eff} determines the half plain and has a large impact at $q^2 = 0$ on $A_I(B \rightarrow K^* \gamma)$.
 - The semi-leptonic operators $O_{9,10}$ are relevant for $B \rightarrow K^* I^+ I^-$ at not too small values of q^2 .

Main constraints:

- on QCD-penguin operators: non-leptonic B decays, life-time differences of B mesons, kaon decays.
- main constraint on semileptonic $C_{9,10}$ from A_{FB} . A change of sign in C_{10} or C_7 flips A_{FB} .

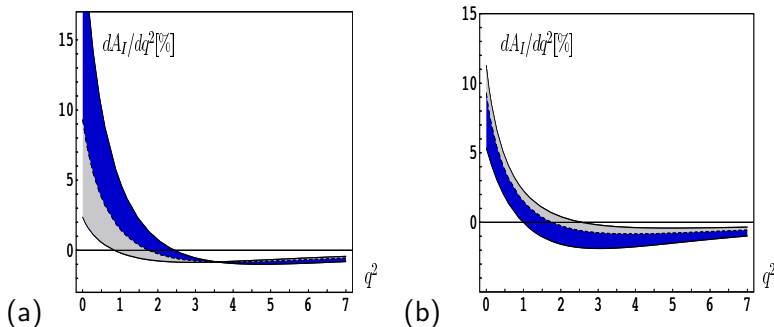
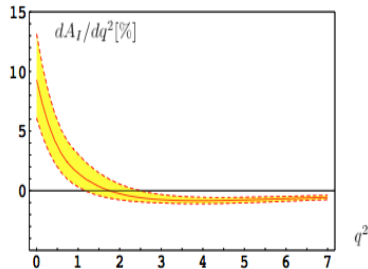


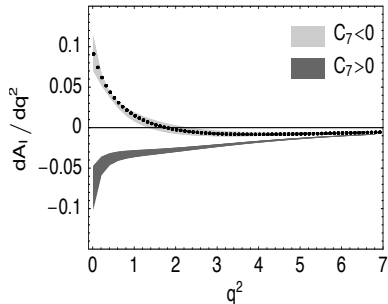
Figure: The isospin asymmetry dA_I/dq^2 for the decay $B \rightarrow K^* \ell^+ \ell^-$ as a function of q^2 . (a) The combination $a_6^{(0)} = (\bar{C}_6 + \bar{C}_5/3)$ in the function $K_1^{\perp(a)}$ is varied within a factor of two around its SM value (dark band = larger values, light band = smaller values). (b) The same for the combination $a_4^{(0)} = (\bar{C}_4 + \bar{C}_3/3)$ in the functions $K_1^{\parallel(a)}$ and $K_2^{\perp(a)}$.

MSSM (All susy particles are taken as heavy (about 1 TeV), except for charginos, sneutrinos, the light (mostly right-handed) stop, and charged Higgs fields. Flavor diagonal mass matrix in down sector. Regime for $\tan \beta = 2 - 40$). Impact on δC_7 and δC_8 .



SM

[Feldmann, J.M '03]



MSSM

Even if the flipped sign solution for C_7 is disfavored in some scenarios (see [S. Descotes et al. '11]) shows the difficulty to get -0.5! and also $A_I(B \rightarrow K^* \gamma)$

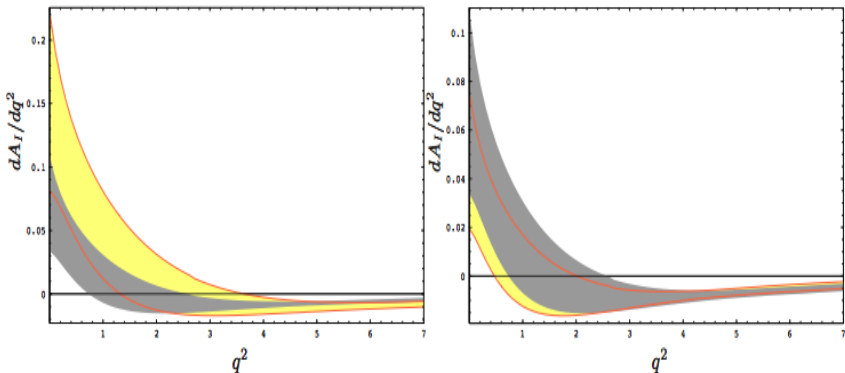
Other possibilities: NP contributions to the QCD-penguin operators C_3 , C_4 , C_5 and C_6 . In particular, we focus on the generic NP case in which

$$C_3^{NP} = C_5^{NP} = -\frac{1}{N_c} C_4^{NP} = -\frac{1}{N_c} C_6^{NP} \equiv \delta C$$

This arises in models of susy when gluino-squark dominates and in some models of extra dimensions.

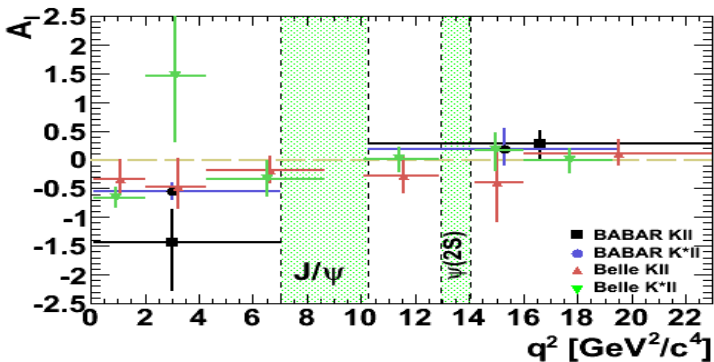
- Most important experimental bounds on the NP contributions to QCD-penguin operators from branching ratios like $B \rightarrow \phi K^0$, $B^+ \rightarrow \phi K^+$, $B^+ \rightarrow \pi^+ K^0$, etc. We compute these branching ratios in NLO QCD-factorization, in order to put bounds on δC .
- It turns out that by far the most important constraint comes from $B^+ \rightarrow \pi^+ K^0$.

The bounds considering the large hadronic uncertainties are in a range $-0.003 \leq \delta C \leq 0.012$:



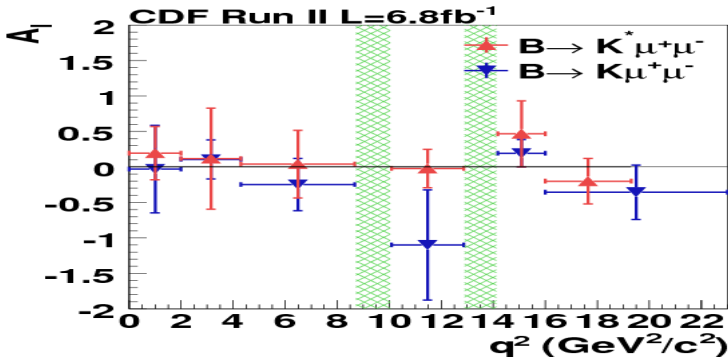
Left plot corresponds to $\delta C = 0.012$ and right plot $\delta C = -0.003$.
Negative δC preferred but still very far away from Babar data.

Babar/Belle measurement:



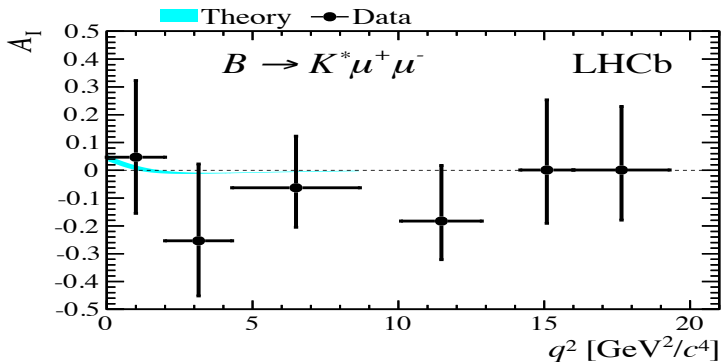
Babar finds 3.9σ , Belle is consistent with Babar for two of the bins. Above J/ψ Babar and Belle are consistent with SM.

CDF measurement PRL107, 201802 (2011) :



LHCb may give the last word on this. IS there a large isospin breaking or not?

LHCb measurement



Life is hard... but not yet the last word.

- Rare B decays phenomenology will play (is playing) a central role in slicing parameter space of models: Better to start taking into account **ALL constraints** coming from rare B decays (message for model building)
- Unitarity Triangle plane will be **complemented** by Wilson Coefficient correlations planes
- Concerning $\mathbf{B} \rightarrow \mathbf{K}^*(\rightarrow \mathbf{K}\pi)\mu^+\mu^-$ observables P_i **stay tuned** in the following years, since they will play a central role in signaling/constraining new New Physics regions.