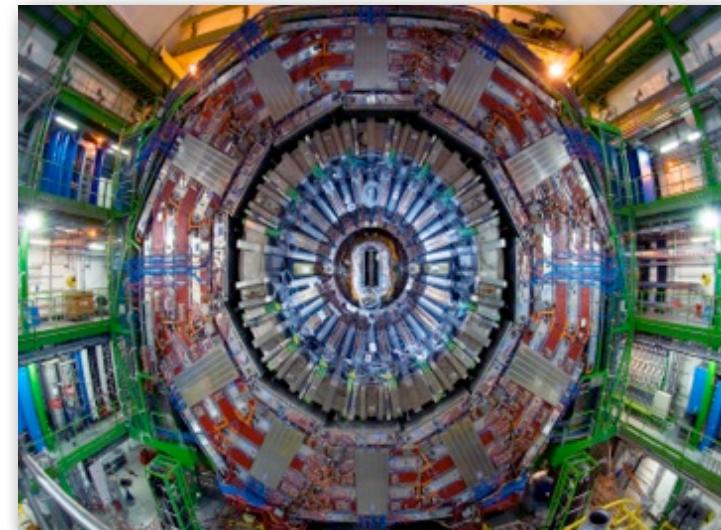
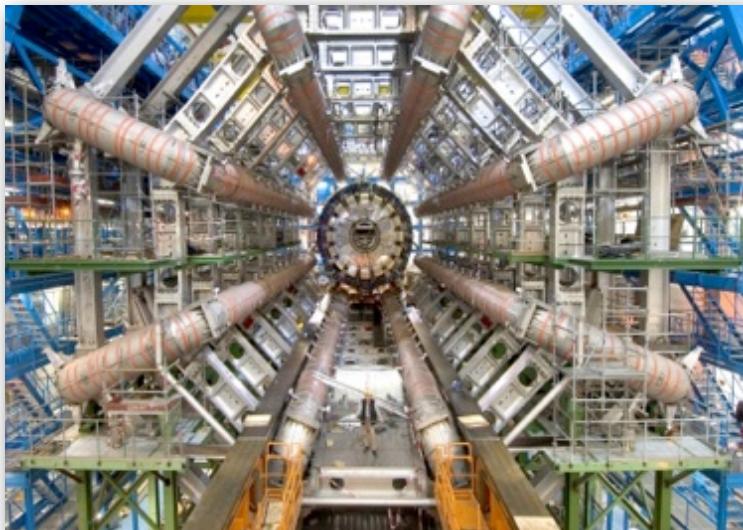


SEARCHES FOR EXOTIC PHYSICS



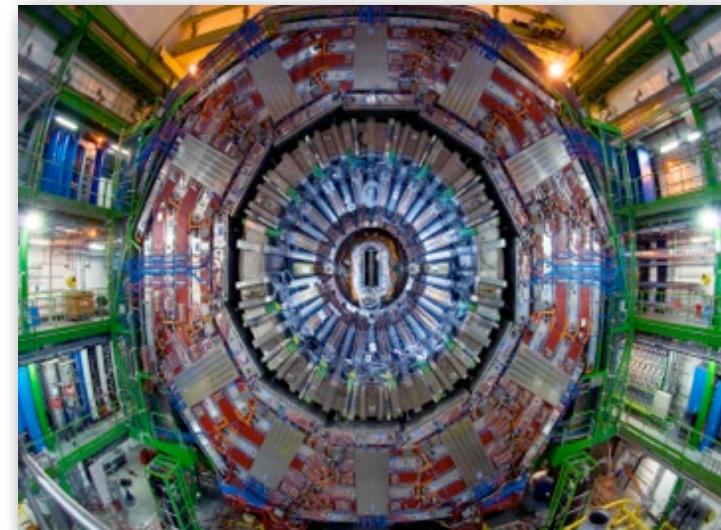
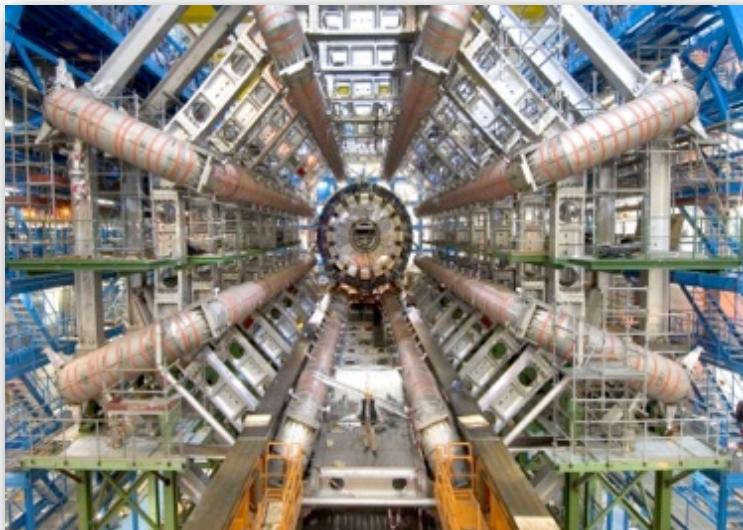
Steve Worm



Science & Technology Facilities Council
Rutherford Appleton Laboratory

Taller de Altas Energías
Banasque, 26 September 2013

SEARCHES FOR EXOTIC* PHYSICS



Steve Worm



Science & Technology Facilities Council
Rutherford Appleton Laboratory

Taller de Altas Energías
Banasque, 26 September 2013

*also known as “BSM (non - SUSY) Results I”

OUTLINE

- *What is “Exotic”?*
 - Standard Model, standard problems
 - New energy regime
 - Top example and detectors
- *Resonances and “standard stuff”*
 - Z' , ADD
 - W'
 - Dijet, dijet w/ b-tag
 - 3-jet, 3-jet w/ b-tag
 - black holes
- *Top-like BSM*
 - $W' \rightarrow tb$
 - vector-like T'
 - Q=5/3 top partners
 - $X \rightarrow tt$ semi-leptonic, hadronic
- *Di-boson resonance*
 - Boosted techniques
 - $W'/\rho_{TC} \rightarrow WZ \rightarrow 3l + \text{MET}$
 - $G_{\text{bulk}} \rightarrow WW \rightarrow l + \text{jet} + \text{MET}$
 - $G_{\text{bulk}} \rightarrow ZZ \rightarrow 2l + 2\text{jets}$
 - $G_{\text{RS}} \rightarrow WW/ZZ$ and $W' \rightarrow WZ$
- *Long-lived particles*
 - displaced jets
 - HSCP
- *Dark Matter*
 - monojet
 - monolepton
- *Conclusions*

DISCLAIMER

- Not an attempt to cover all Exotica results for ATLAS or CMS
 - Each collaboration has 60-70 analyses going at any given moment
 - I selected some highlights and a few to discuss in more detail
- Extra apologies to ATLAS... I probably show more CMS results
- Thanks to H. Bachacou, C. Issever, E. Thompson from which I stole a few ideas (and slides)



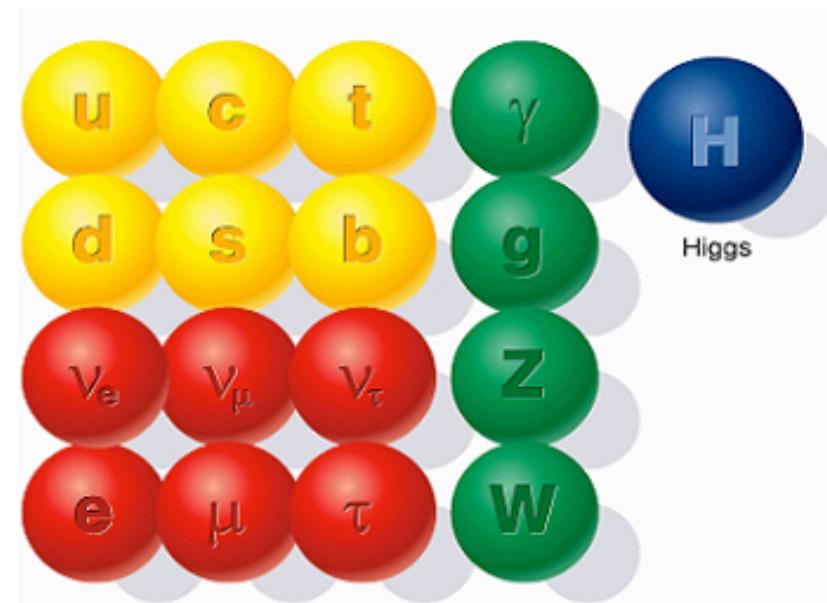
SOME BASICS...

THE STANDARD MODEL

From experimental evidence and theory insight, a simple picture emerged:

The Standard “Ingredients”:

- Handful of fundamental particles
- Particles constructed by 2 or 3 quarks and only a few rules
- A few forces mediated by bosons
- Add one particle, the Higgs give the particles mass



Standard Model Particles

The Standard Model has been incredibly successful in explaining all data...

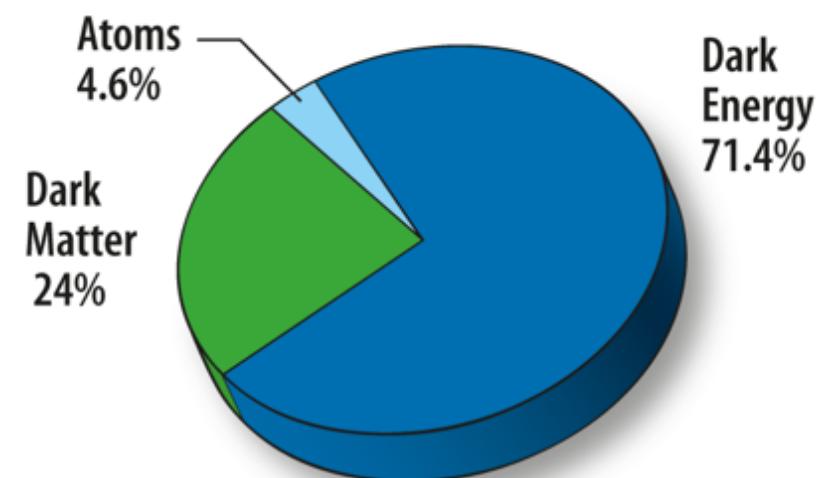
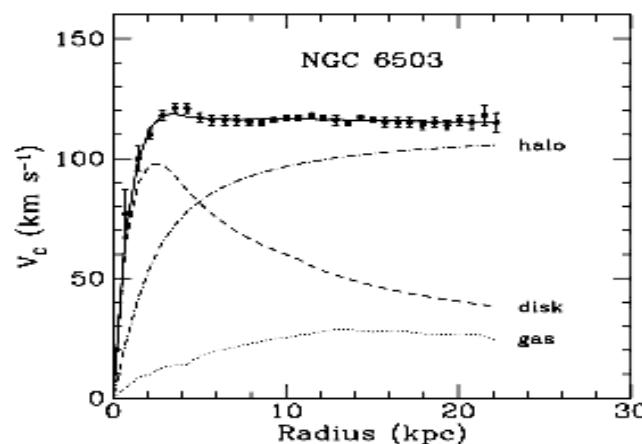
...but there are problems too

STANDARD MODEL, STANDARD PROBLEMS

- Antimatter: *What happened to all the antimatter?*



- Dark Matter: *We don't know what most of the matter in our world is made of!*



STANDARD MODEL, STANDARD PROBLEMS

“Cracks” have started to appear in the Standard Model...

Many problems identified over time

- No explanation of masses, coupling constants
- Why three families?
- Gravity not included
- The “hierarchy” problem, fine tuning...
- What is the Dark Energy?

...and yet it explains the data



The Standard Model isn't so much wrong as it is incomplete

EFFECTIVE THEORIES... AND SCALE

- Fundamental theory may be hiding at shorter distances (higher energies)
- ~1900 we reached the atomic scale
 - $10^{-8} \text{ cm} \approx \hbar^2/e^2 m_e$
 - Quantum Mechanics
 - Quantum Electrodynamics
- ~1950 we reached strong interaction scale
 - $10^{-13} \text{ cm} \approx M_{\text{exp}}[-8\pi^2 g_s^2(M)b_0]$
 - QCD
 - Quarks, Gluons
- 2010 we reach (and exceeded) the EWK scale
 - 10^{-17} cm , the TeV scale
 - EWK phase transition is happening
 - W, Z, electron...etc. acquire mass
 - $v = (\sqrt{2}G_F)^{-1/2} \approx 246 \text{ GeV} \leftarrow \text{Higgs VEV}$

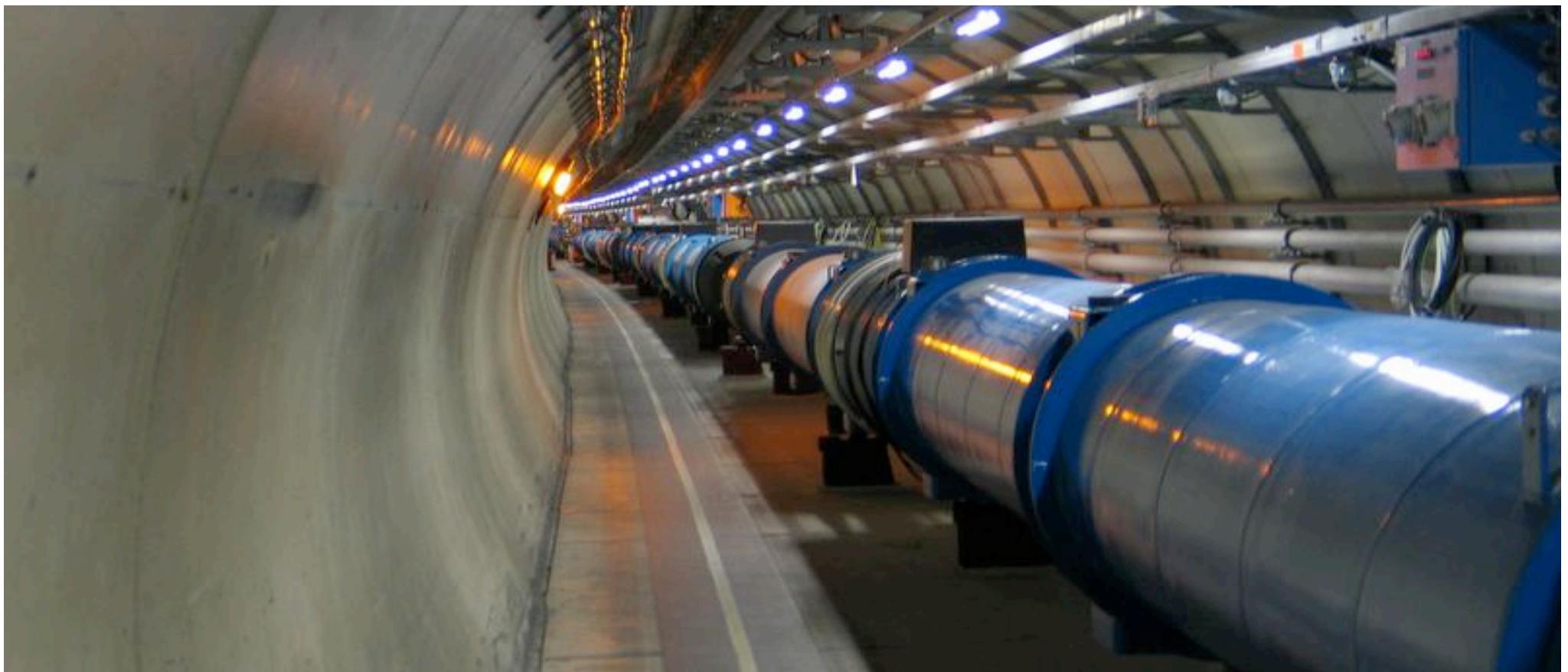
Expect lots of interesting stuff at 14 TeV!

PERIODIC TABLE OF THE ELEMENTS																			
http://www.periodictable.com																			
1	H	2	He	3	Li	4	Be	5	B	6	C	7	N	8	O	9	F	10	Ne
Alkali metal	Alkali metal	Alkaline earth metal	Alkaline earth metal	Chalcogen element	Chalcogen element	Transition metals	Transition metals	Halogen element	Halogen element	Carbon	Nitrogen	Phosphorus	Sulfur	Chlorine	Fluorine	Neon	Argon	Krypton	Xenon
Period I	Period II	Period III	Period IV	Period V	Period VI	Period VIIA	Period VIIIB	Period VIIIC	Period VIIID	Period VIIIE	Period VIIIF	Period VIIIG	Period VIIIA	Period VIIIB	Period VIIIC	Period VIIID	Period VIIIE	Period VIIIF	Period VIIIA
1.0079	4.0026	6.941	9.0122	10.811	12.0111	14.0067	15.9994	18.9984	20.9984	22.9898	24.9858	26.9868	28.9864	30.9866	31.9863	34.9888	35.9875	39.9884	40.9884
Hydrogen	Helium	Lithium	Boron	Carbon	Nitrogen	Phosphorus	Sulfur	Chlorine	Fluorine	Neon									
Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol	Symbol
Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number	Atomic Number
Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass	Rel. Atomic Mass
1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53
Li	Be	B	C	N	O	F	Ne	Na	Mg	Al	Si	P	S	Cl	Ar	Ne	Na	Mg	Al
35	36																		

ACCELERATOR ADVANCES

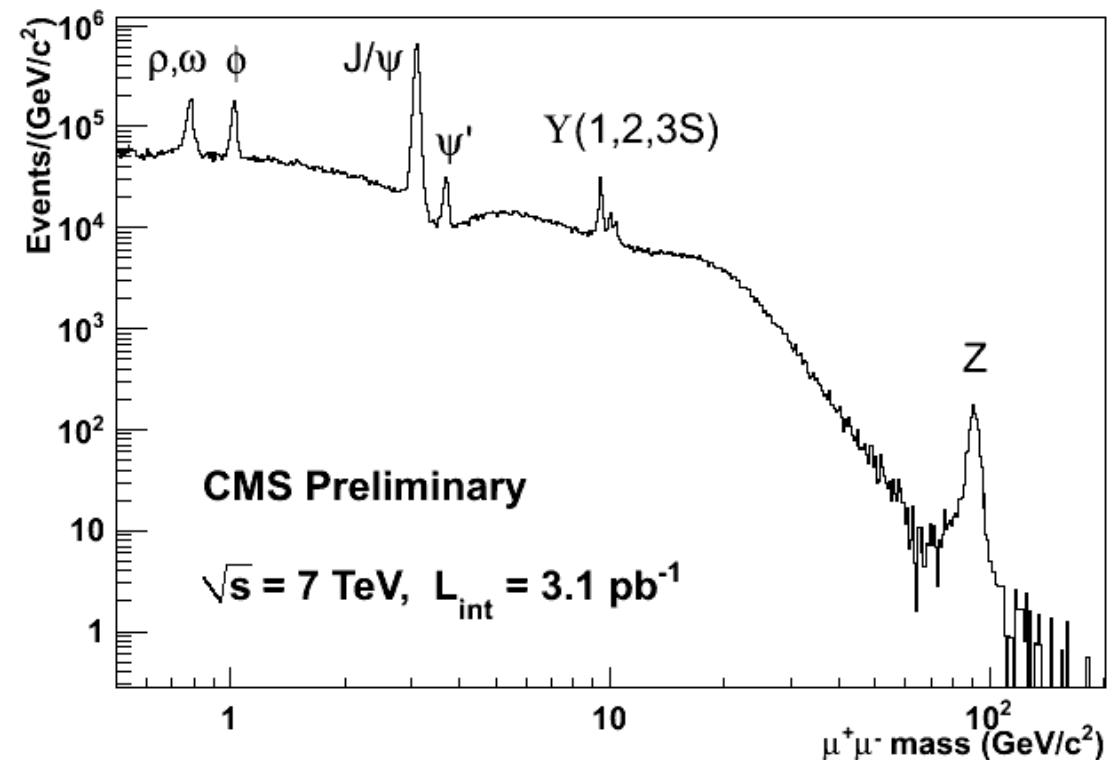
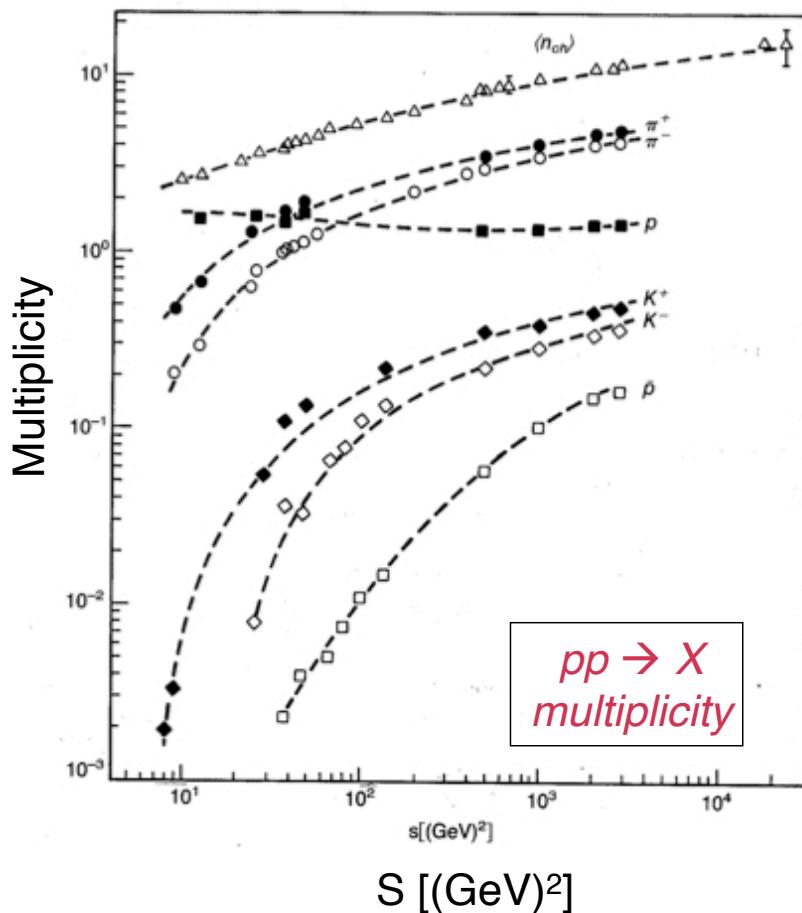
- Each advance is a revolution... but sadly only once or twice per generation
 - Previous energy record-holder (Tevatron) first started in 1983 - 30 years ago
 - LEP at CERN stopped in 2000 - 13 years ago

The jump to 14 TeV will be a huge advance for BSM searches



NEW ENERGY REGIMES

- Each advance in territory makes new discovery possible
- Many historic examples...
- Hunting for “bumps” in the mass spectra is the first step



SOME EXPERIMENTAL BASICS...

$$\text{Number of events} = \mathcal{L} \sigma \text{Br} A \varepsilon$$

- Integrated Luminosity – how much beam (protons passing each other)
- Cross section – likelihood of interaction between particles
- Branching ratio – how often we get the interaction/decay we want
- Acceptance – did the event enter the detector?
- Efficiency – was the event detected? (trigger, detector, selection...)

EXAMPLE: TOP SEARCH CIRCA 1993

- Integrated luminosity: 25 pb^{-1}
- Top production cross section: 7 pb
- Branching Fraction (not to taus):
 - Lepton+lepton (plus b jets) = 5%
 - Lepton+jets = 30%
 - Multi-jet = 44%
- Acceptance*: ~30% (varies by decay, m_T)
- Efficiencies*:
 - Trigger efficiency: ~90%
 - Lepton recon. eff.: ~90% (applied twice)
 - Lepton isolation, other cuts: ~30%
 - b-tag efficiency: ~50-60% (applied once)
- Background:
 - “noise” in dilepton channel ~0.5 event
- *(Warning! Numbers are approximate)

$$\begin{aligned}\text{Events} &= \mathcal{L} \sigma \text{Br} A \varepsilon \\ &= (25)(7)(0.05)(0.3)2(0.9)3(0.6) \\ &= \sim 2\end{aligned}$$

*Did we find it?
What can be done better?*

EXAMPLE: TOP SEARCH CIRCA 1993

- Did not find the top... can set limits on mass/cross section
- Roughly equivalent to saying “We looked at low mass and didn’t find it”

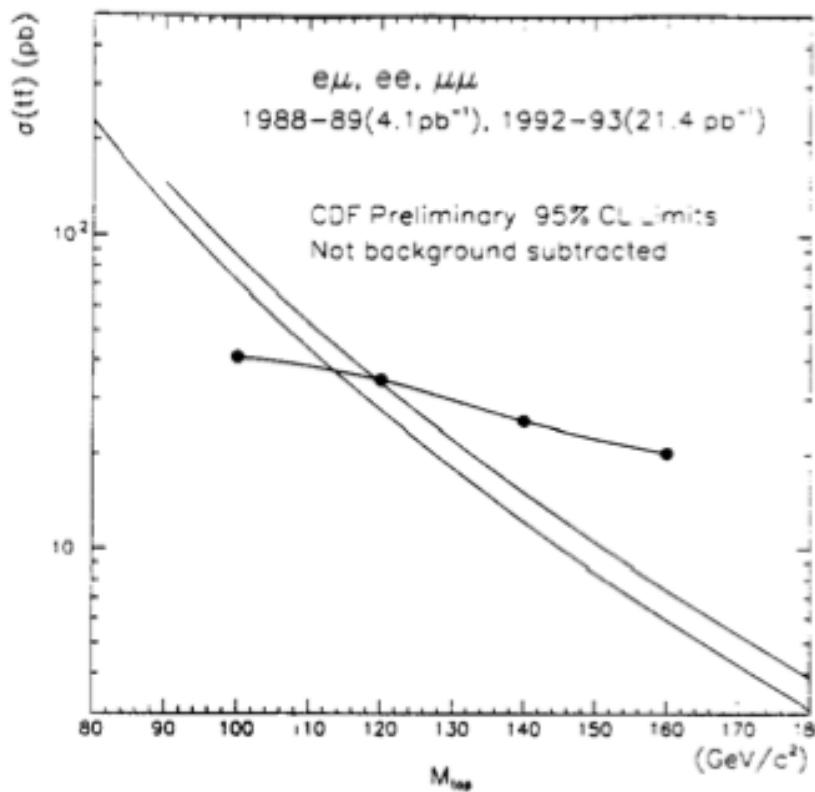
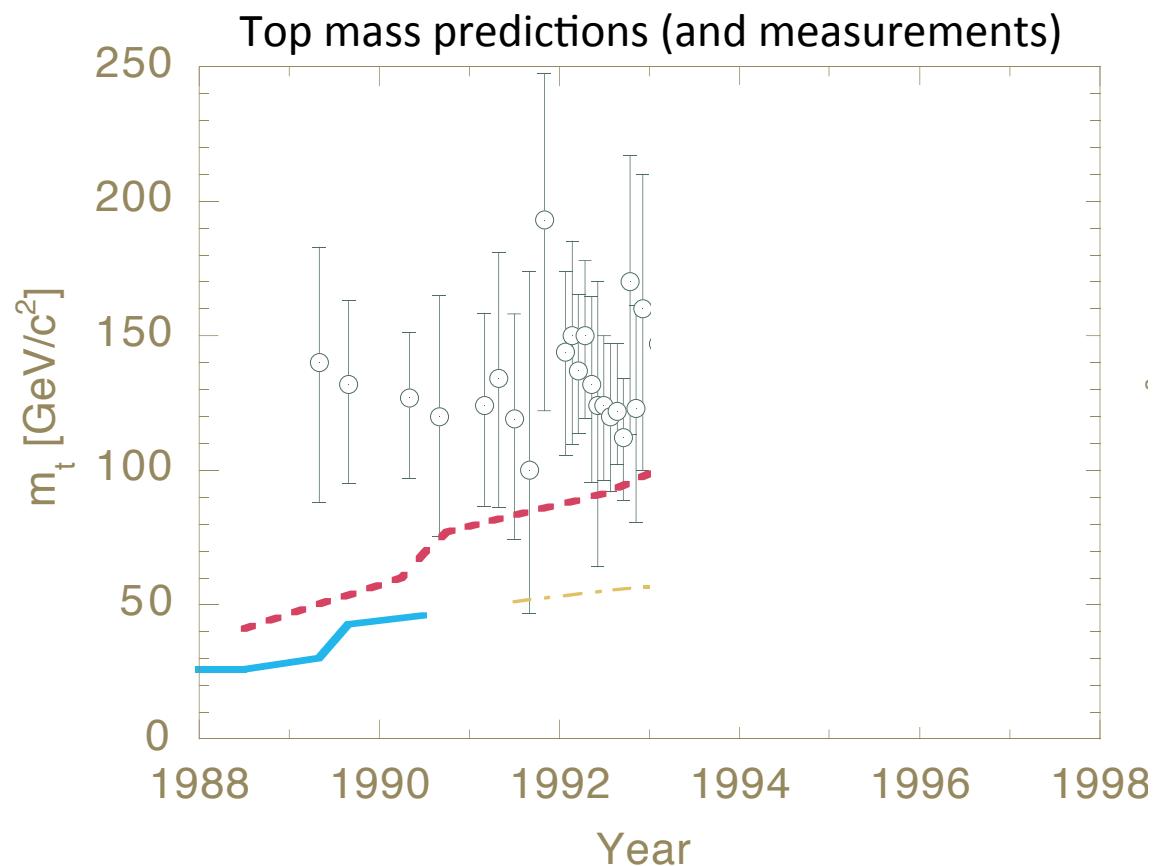


Figure 2: The experimental 95% C.L. limit on σ_{tt} as measured in the dilepton channel (points) and two theoretical predictions.

- *Knowing where it isn't... is still useful!*

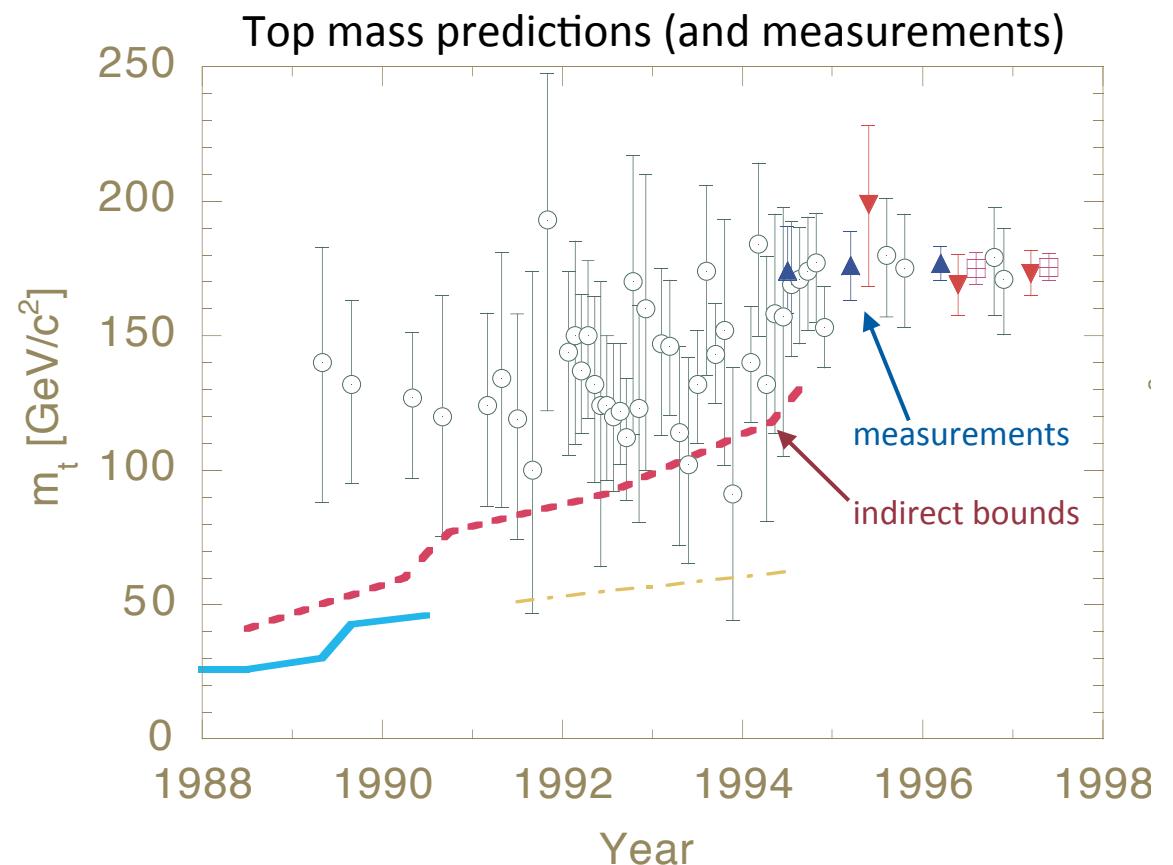
EXAMPLE: TOP SEARCH CIRCA 1993

- What can improve the sensitivity?
 - More data
 - Search more channels
 - Improve b-tagging
 - Improve selection algorithms
 - Combine many experiments
 - Go to more powerful accelerator
 - Ask a different theorist!
(look in a different place)



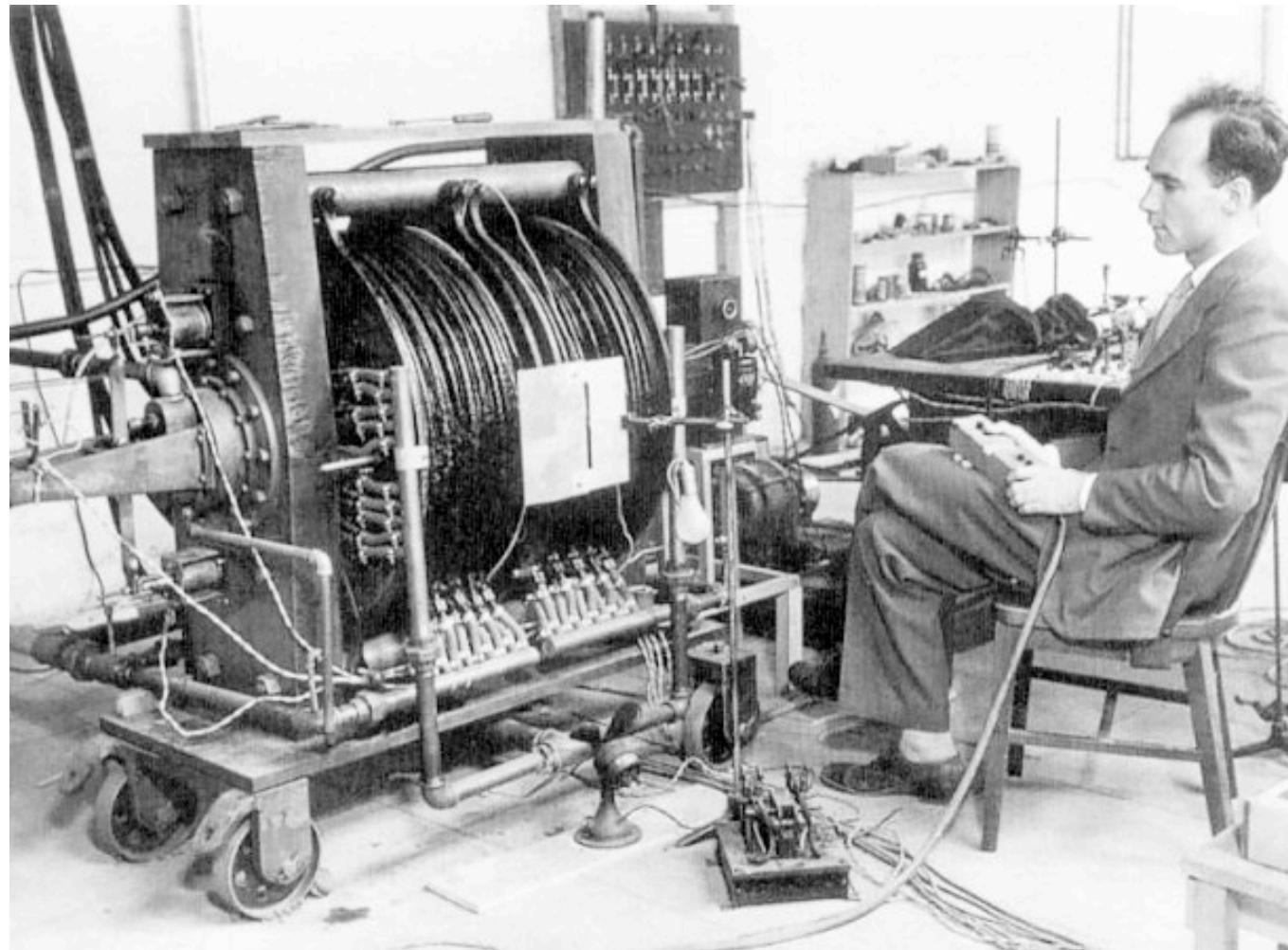
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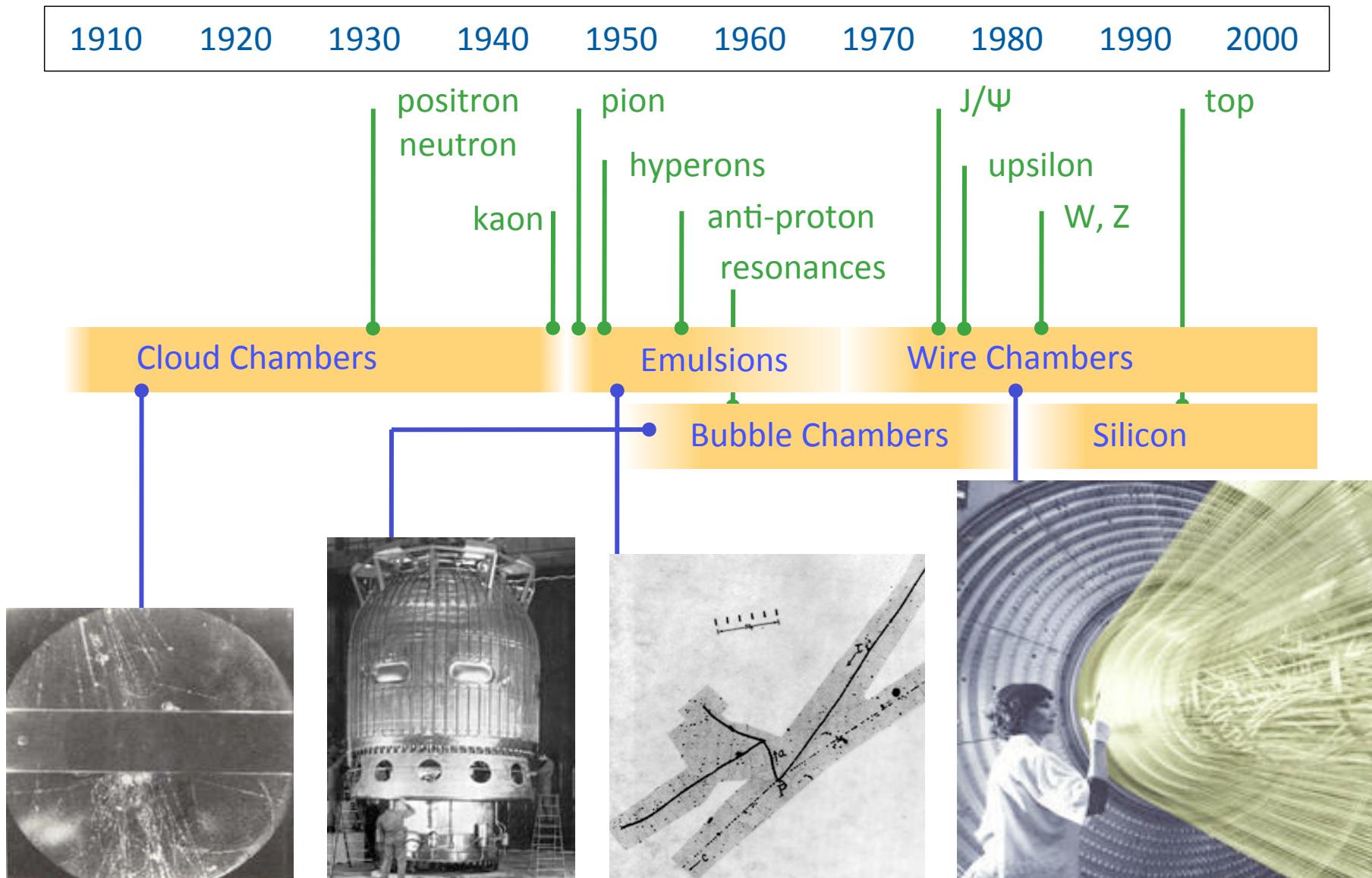
Top found in ~1995 with more data... role of detector was crucial

SEARCHES AND (OLD FASHIONED) DETECTORS



New/better detector → new physics found → Nobel prize (simple, isn't it?)

INTERPLAY OF DETECTORS AND DISCOVERIES



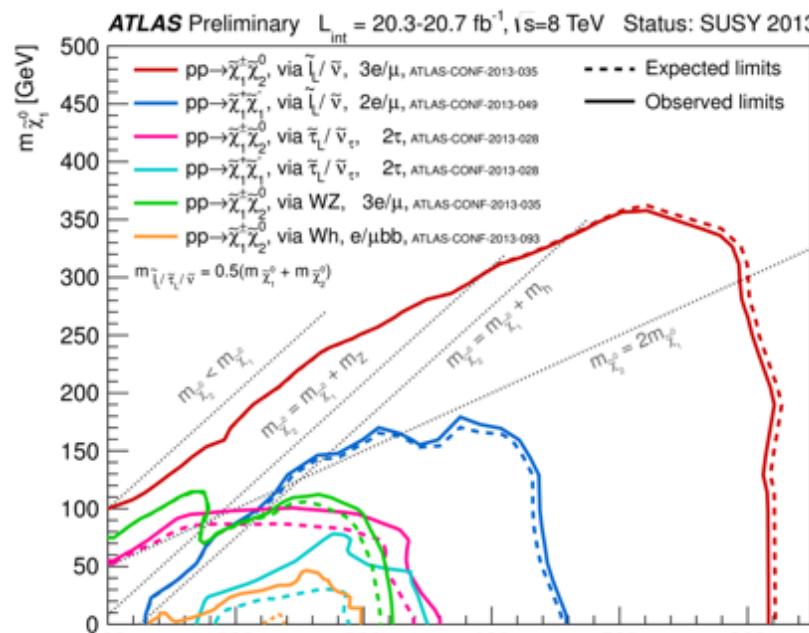
Don't underestimate the role of new technology, techniques, triggers...

WHAT IS “EXOTIC”?

BSM AND EXOTICA: WHAT IS “EXOTICA”?

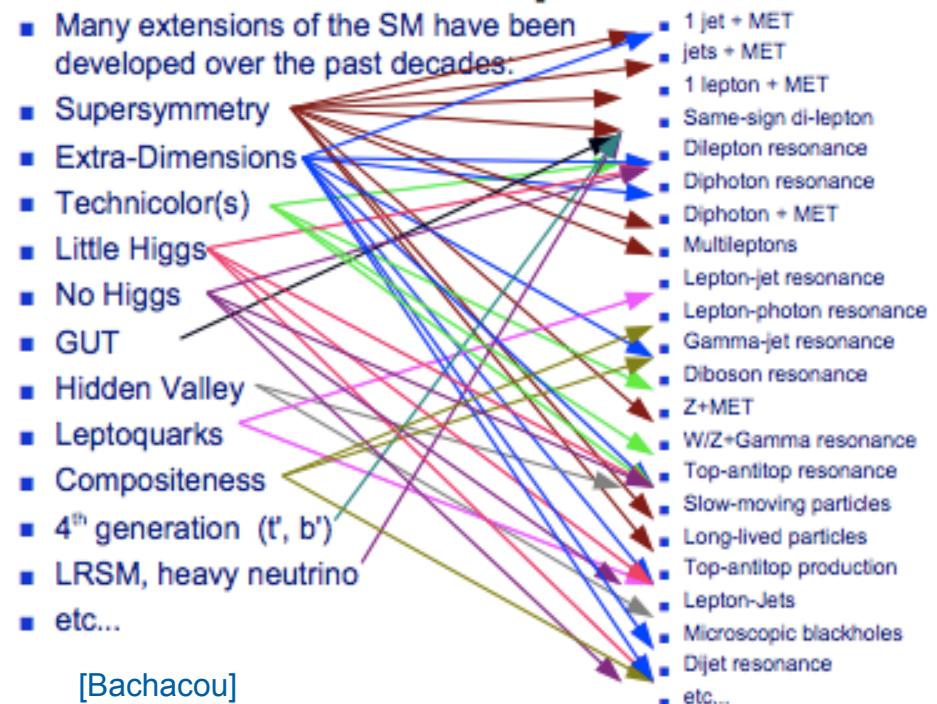
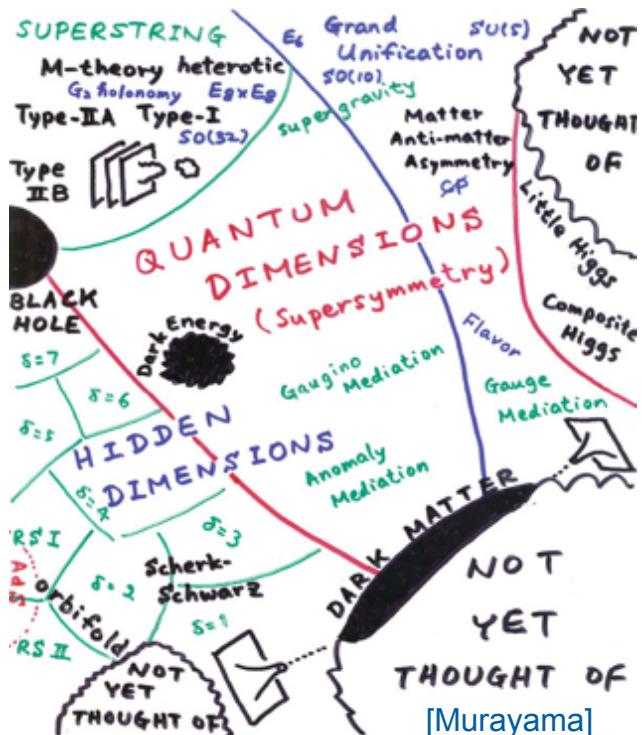
Comprehensive search of the landscape of $\sqrt{s} = 8$ TeV proton collisions

- *SUSY* provides a framework in which to search and compare results: *a map*
- *Higgs* phenomenology fleshed-out many years ago: *a guide book*
- *Exotica* searches have to cover a wide range... no set map or guide
 - with new data, race to cover as much as possible in search for BSM physics
 - now working on depth: more “general” searches, more complete coverage

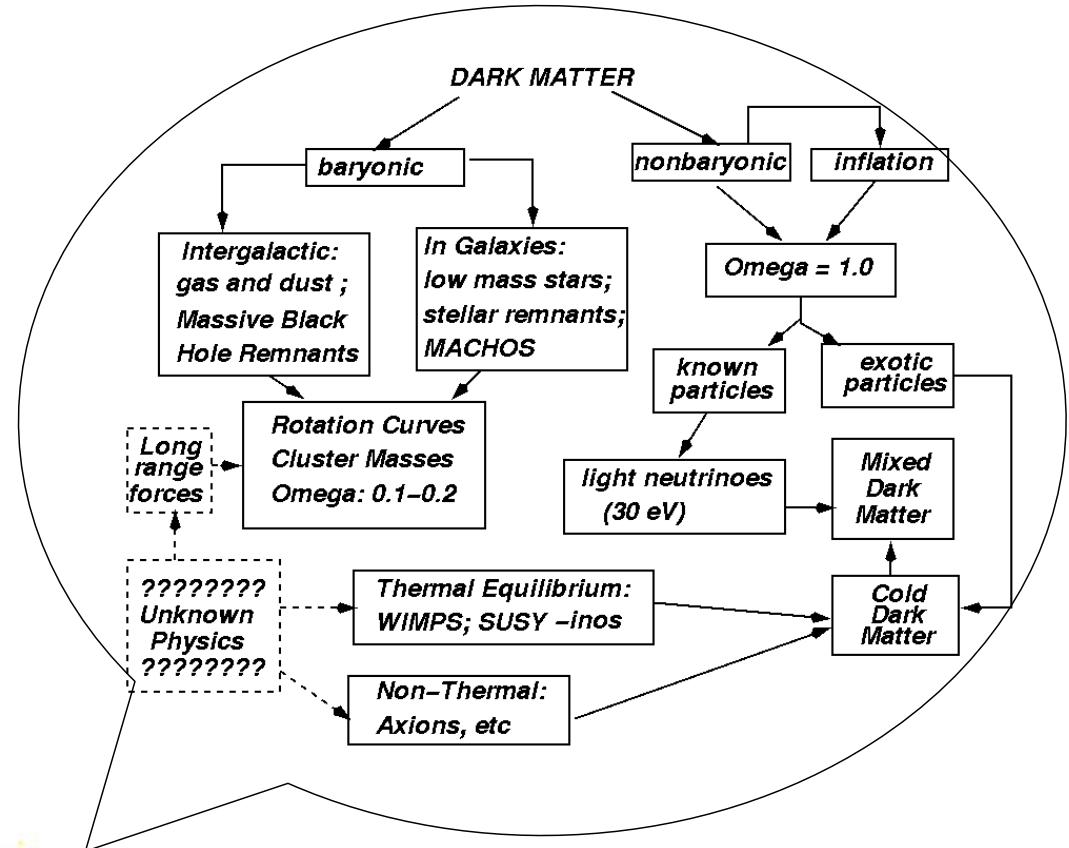


BSM AND EXOTICA: WHAT IS “EXOTICA”?

- So how to cover so much parameter space? How to conduct searches, to categorize? How do we plot out progress with few models?
- Wide variety of search strategies used
 - look for interesting features in the data – new resonant states e.g. Z' , W'
 - look at all possible channels for disagreements with expectation – leptons, photons, jets
 - follow-up interesting new BSM models



CAN THEORY HELP?



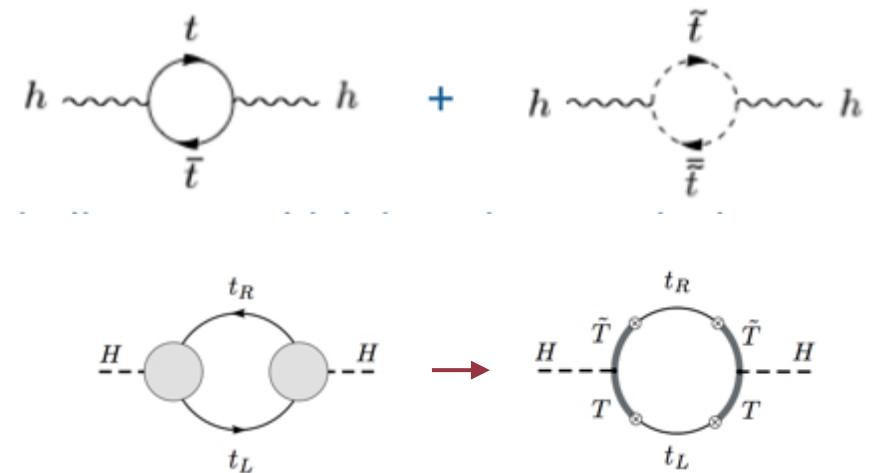
*Yes, but they also need to see something new in the data
(no shortage of good ideas)*



THE HIERARCHY PROBLEM

Low-mass Higgs causes tensions, but also motivates new physics at $\sim\text{TeV}$

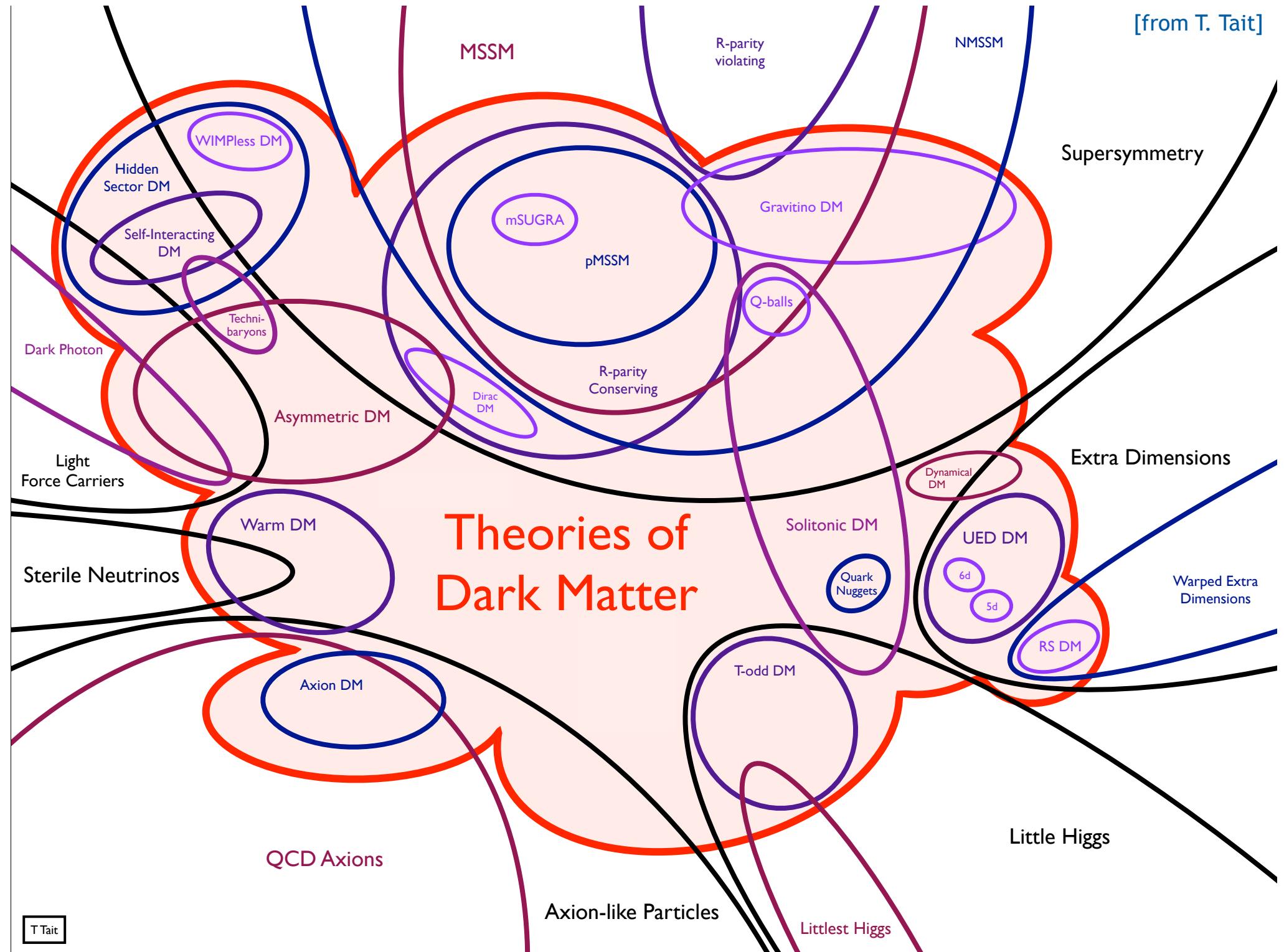
- Implications of Higgs as a scalar
 - radiative corrections to self-energy are divergent
 - maybe we are missing a new (fermion/scalar/vector) degree of freedom?
- Supersymmetry:
 - sparticles cancel particle contributions
 - well-studied but so far undetected
- Higgs is a composite:
 - strongly-coupled BSM, yielding vector-like q
 - multi-jet resonances from heavy gluons
- Extra Dimensions:
 - Higgs is a vector in 5D
 - motivates LED searches, KK excitations
- Or maybe there is no fine tuning problem!
 - See for example [arXiv:1306.5647](https://arxiv.org/abs/1306.5647)



$$M_{Pl}^2 \sim M_D^{2+n} R^n$$

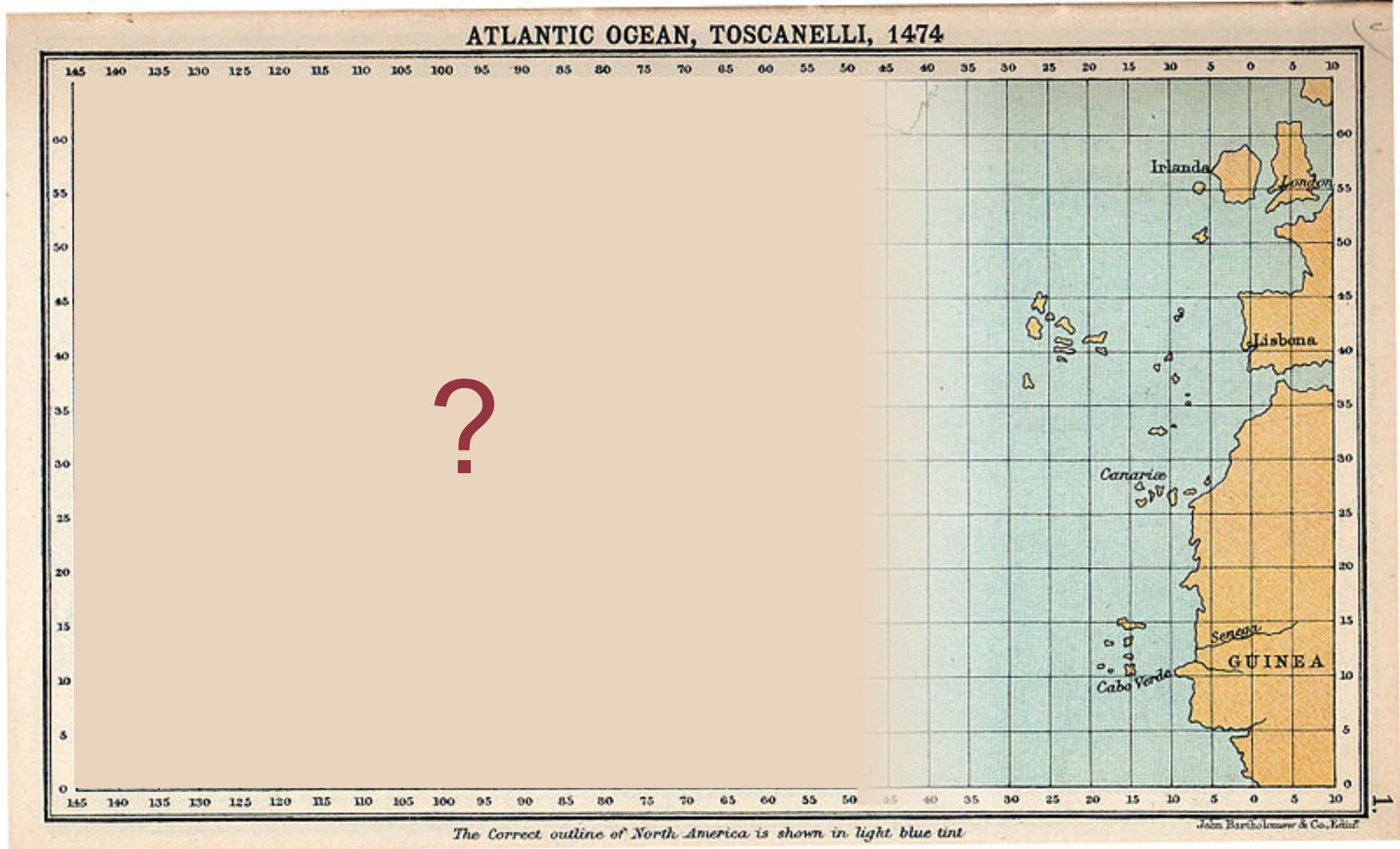
[from T. Tait]

Theories of Dark Matter



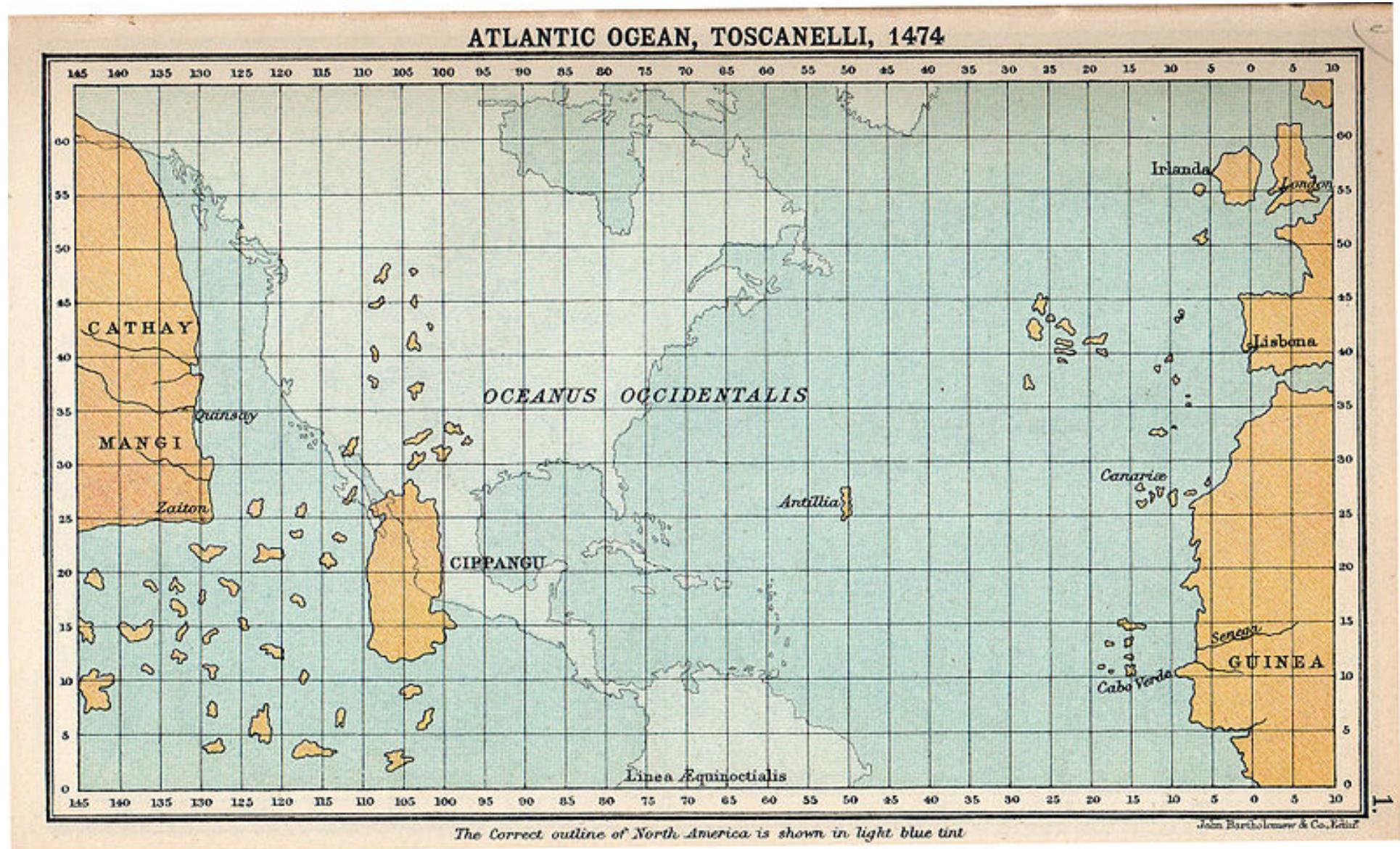
ROLE OF MODELS IN EXOTICA

[borrowed from C. Issever]

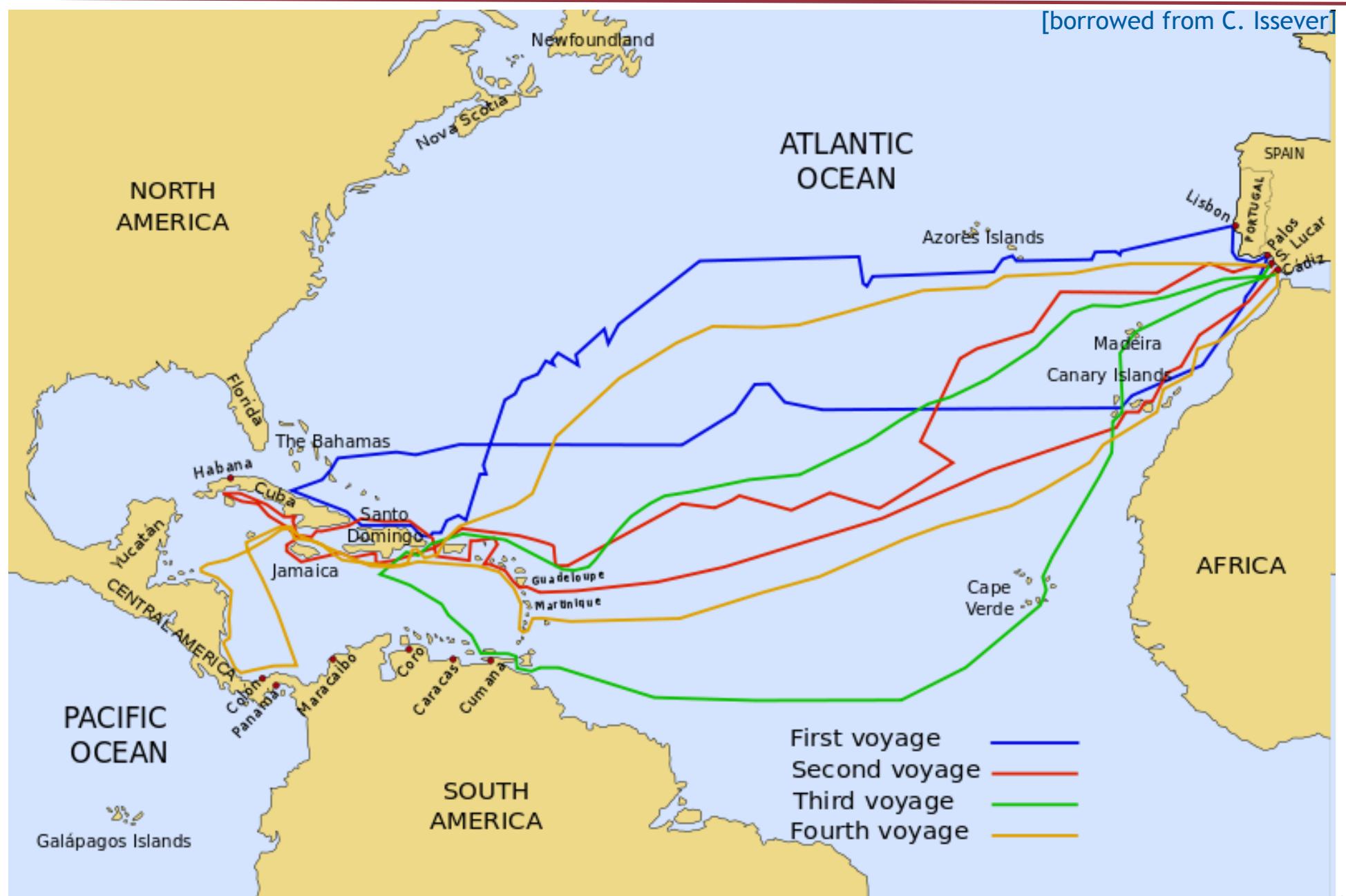


ROLE OF MODELS IN EXOTICA

[borrowed from C. Issever]

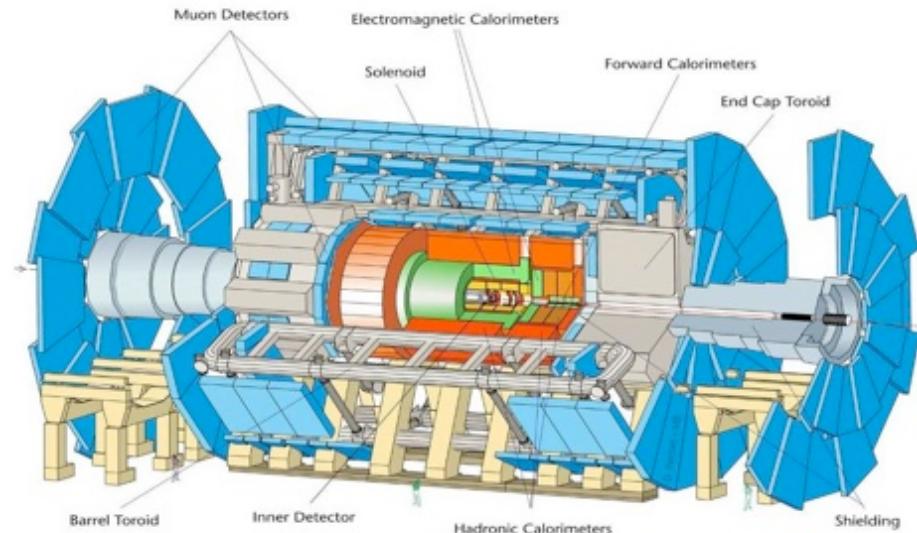


ROLE OF MODELS IN EXOTICA

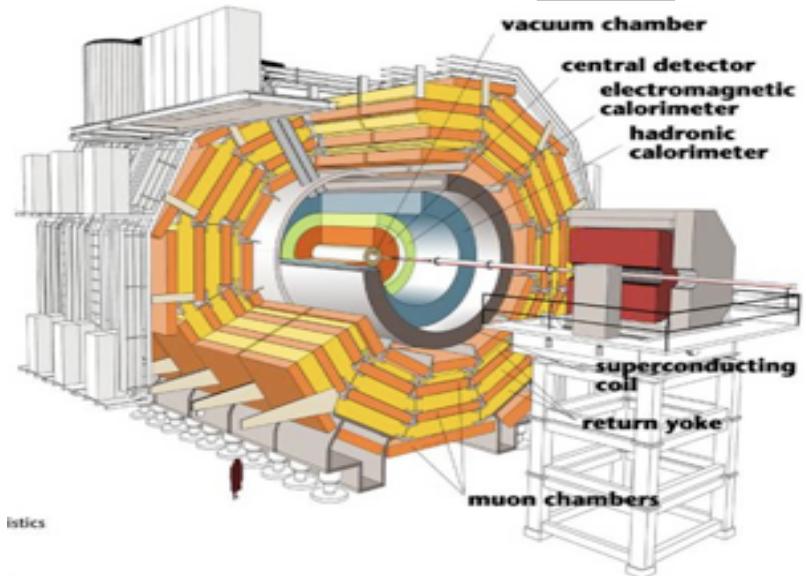


THE FOUR MAIN LHC EXPERIMENTS

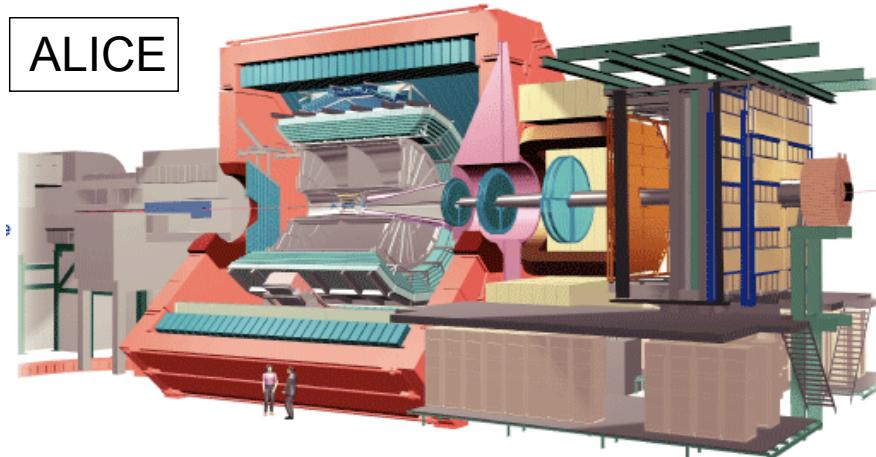
ATLAS



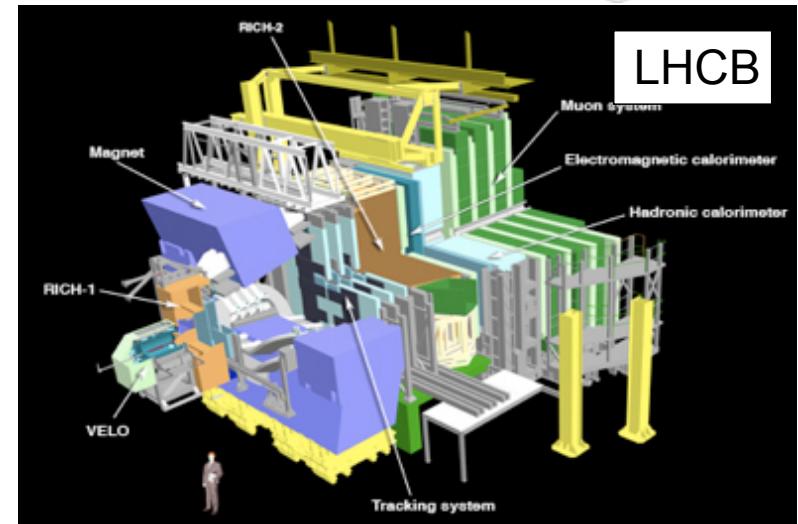
CMS



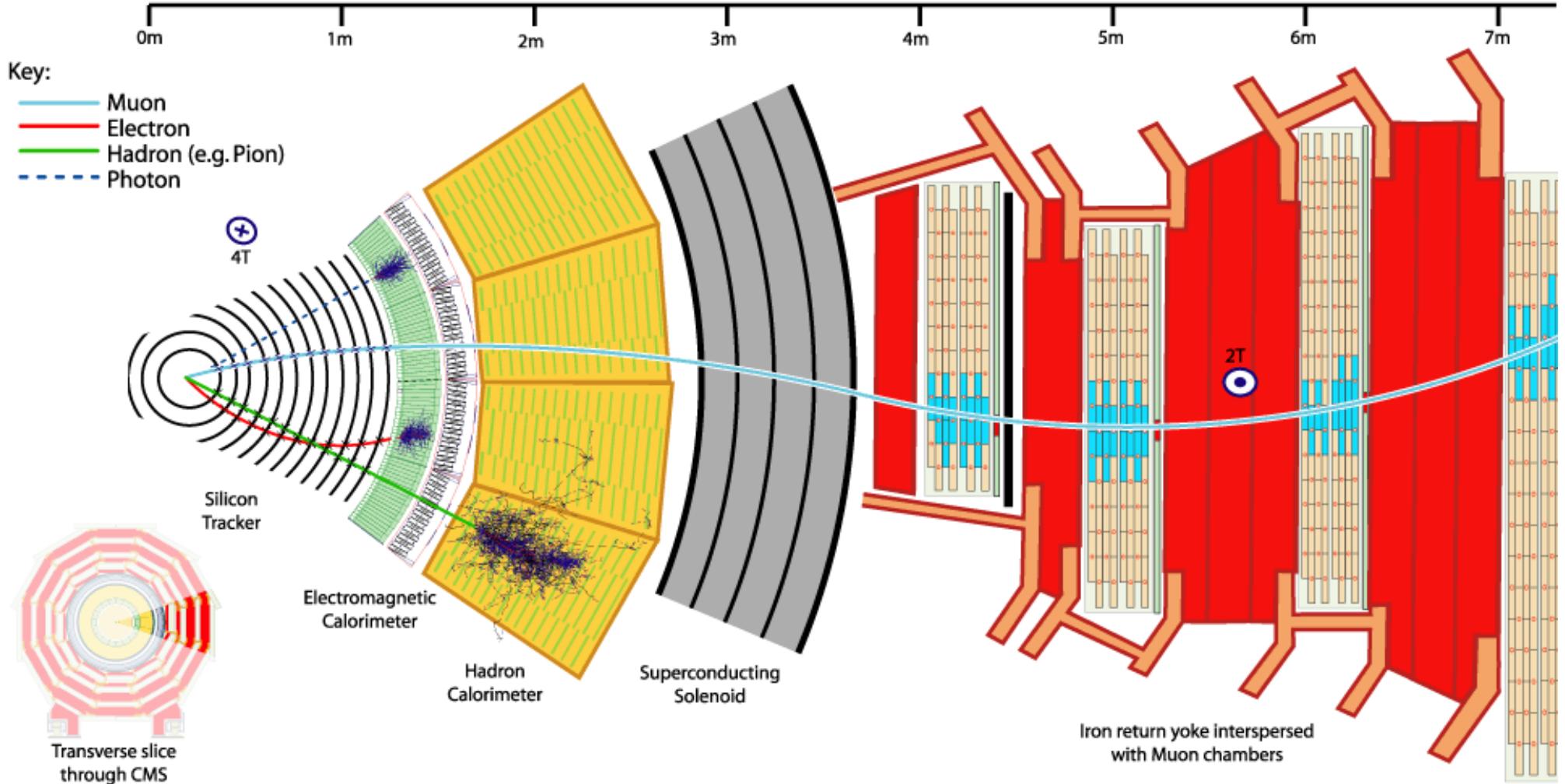
ALICE



LHCb

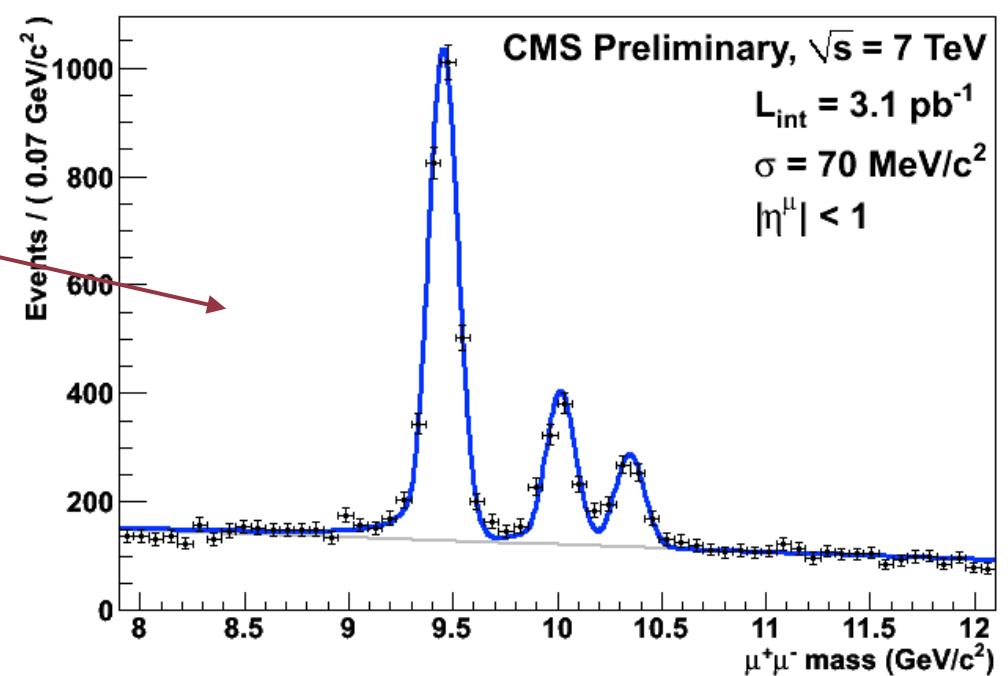
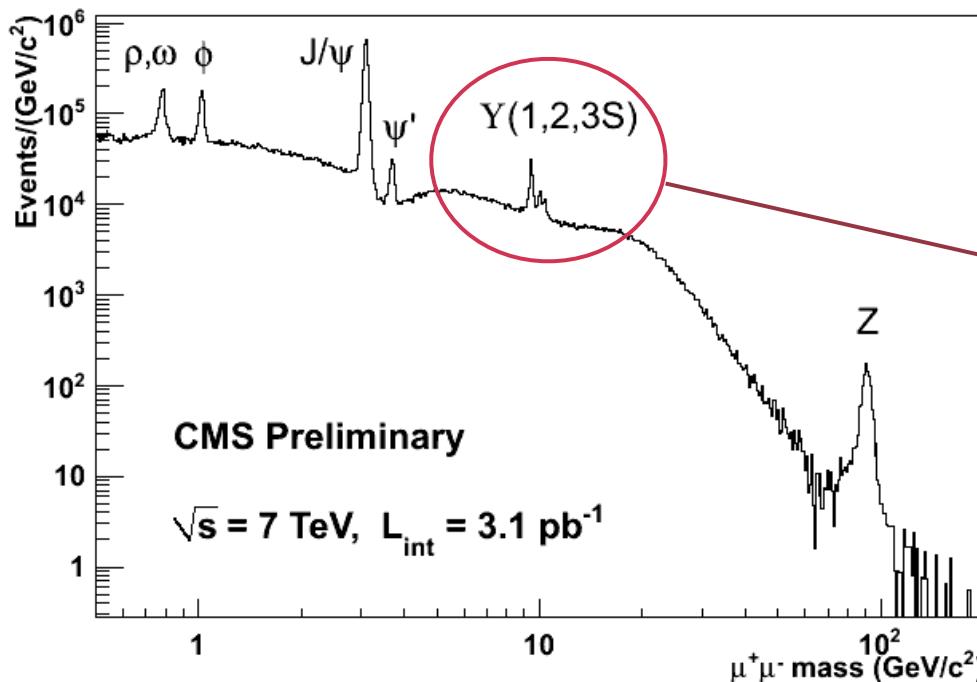


(NORMAL) PARTICLES IN THE DETECTOR



UPSILONS

- CMS: clear calibration peaks after only a few weeks
 - Width related to precision with which we can reconstruct the particles
 - Careful alignment will improve things, but only a little: already extremely good at 3 pb^{-1}
- ATLAS, Alice, LHCb similar: no major problems at startup



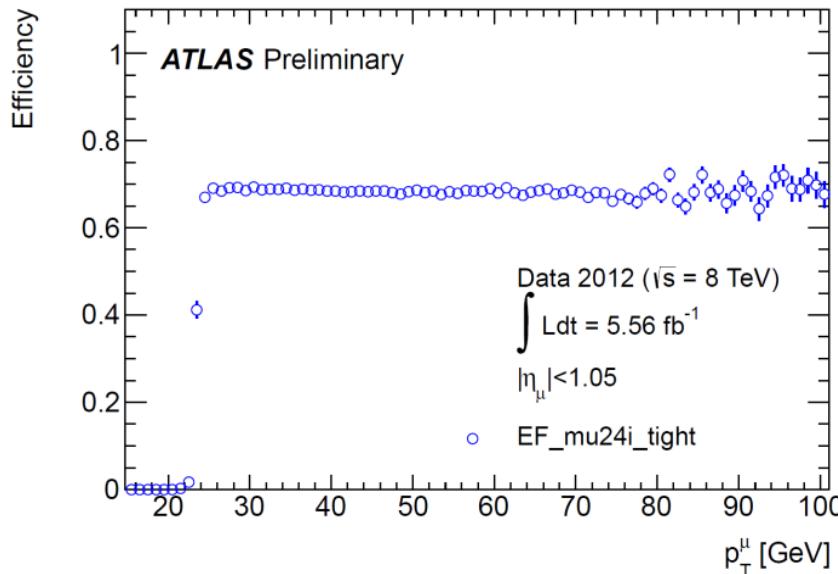
Now on to searches for higher-mass resonances...

RESONANCES AND OTHER “USUAL SUSPECTS”: Z' , W' , DIJETS, ...

DILEPTON RESONANCE SEARCH: TRIGGER

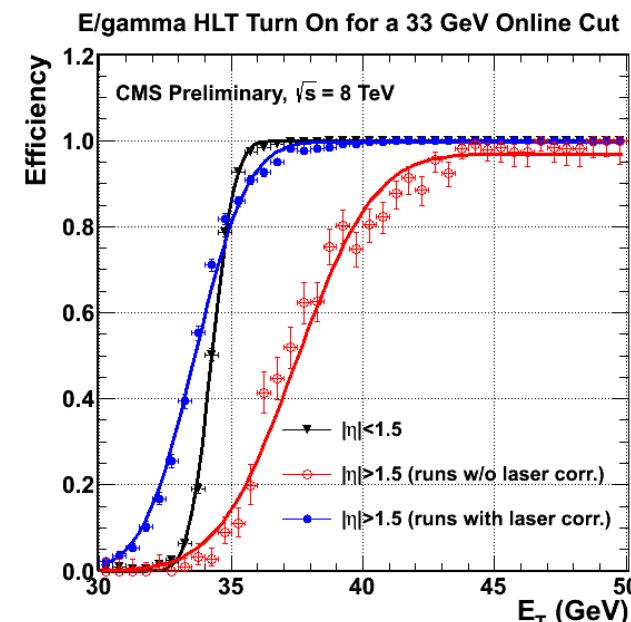
ATLAS

- ee channel
 - Diphoton trigger
 - $E_T > 35 \text{ GeV}$ and $E_T > 25 \text{ GeV}$
- $\mu\mu$ channel
 - Single muon triggers
 - $E_T > 24 \text{ GeV}$ or $E_T > 36 \text{ GeV}$



CMS

- ee channel
 - Dielectron trigger
 - Both clusters w $E_T > 33 \text{ GeV}$
- $\mu\mu$ channel
 - single muon trigger
 - $E_T > 40 \text{ GeV}$



EM CALORIMETRY (ATLAS)



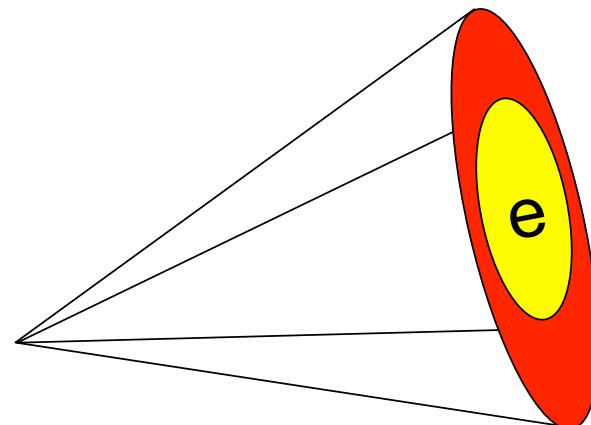
Barrel Liquid Argon Calorimeter



Accordion Sampling Layers

DILEPTON RESONANCE SEARCH: ELECTRONS

- Electron p_T thresholds similar
- Jets fake electrons: use isolation to reduce fakes



ATLAS	CMS
$E_T^1 > 40 \text{ GeV}$	$E_T^1 > 35 \text{ GeV}$
$E_T^2 > 30 \text{ GeV}$	$E_T^2 > 35 \text{ GeV}$

	ATLAS	CMS
leading	$ l_{0.2}^{\text{calo}} < 0.7\% \cdot E_T + 5 \text{ GeV}$	$ l_{0.3}^{\text{tracker}} < 5 \text{ GeV}$
subleading	$ l_{0.2}^{\text{calo}} < 2.2\% \cdot E_T + 6 \text{ GeV}$	$ l_{0.3}^{\text{Calo}} < 3\% \cdot E_T$

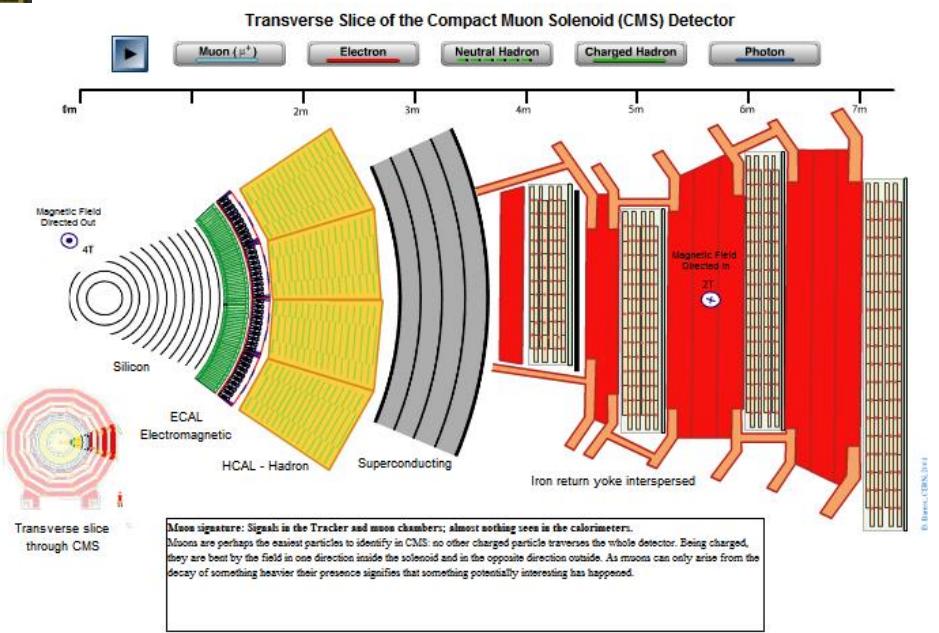
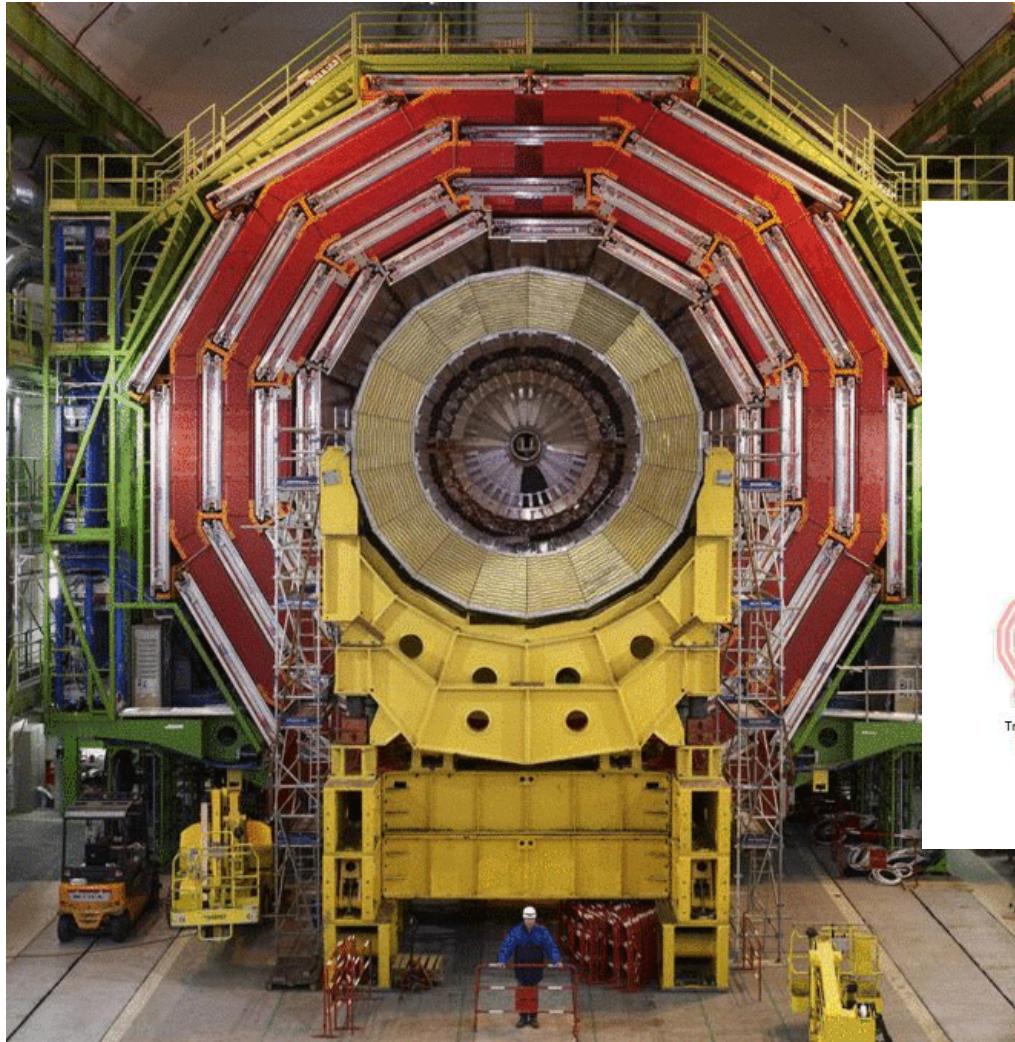
ATLAS

- $A \times \varepsilon = 73\%$ ($m = 2 \text{ TeV}$)

CMS

- $A \times \varepsilon = 67\%$ ($m = 2.5 \text{ TeV}$)

MUON DETECTORS (CMS)



DILEPTON RESONANCE SEARCH: MUONS

ATLAS

- Single muon triggers
 - $pT > 25 \text{ GeV}$
 - $|\eta| < 2.4$
- Suppress cosmic rays
 - $|d_0| < 0.2 \text{ mm}$
 - $|z_{0-\text{vertex}}| < 1 \text{ mm}$
- Suppress jets faking μ 's
 - $\sum pT(\Delta R < 0.3) < 5\% \cdot pT$
- Require opposite charge
- $A \times \varepsilon = 46\%$ ($m = 2 \text{ TeV}$)

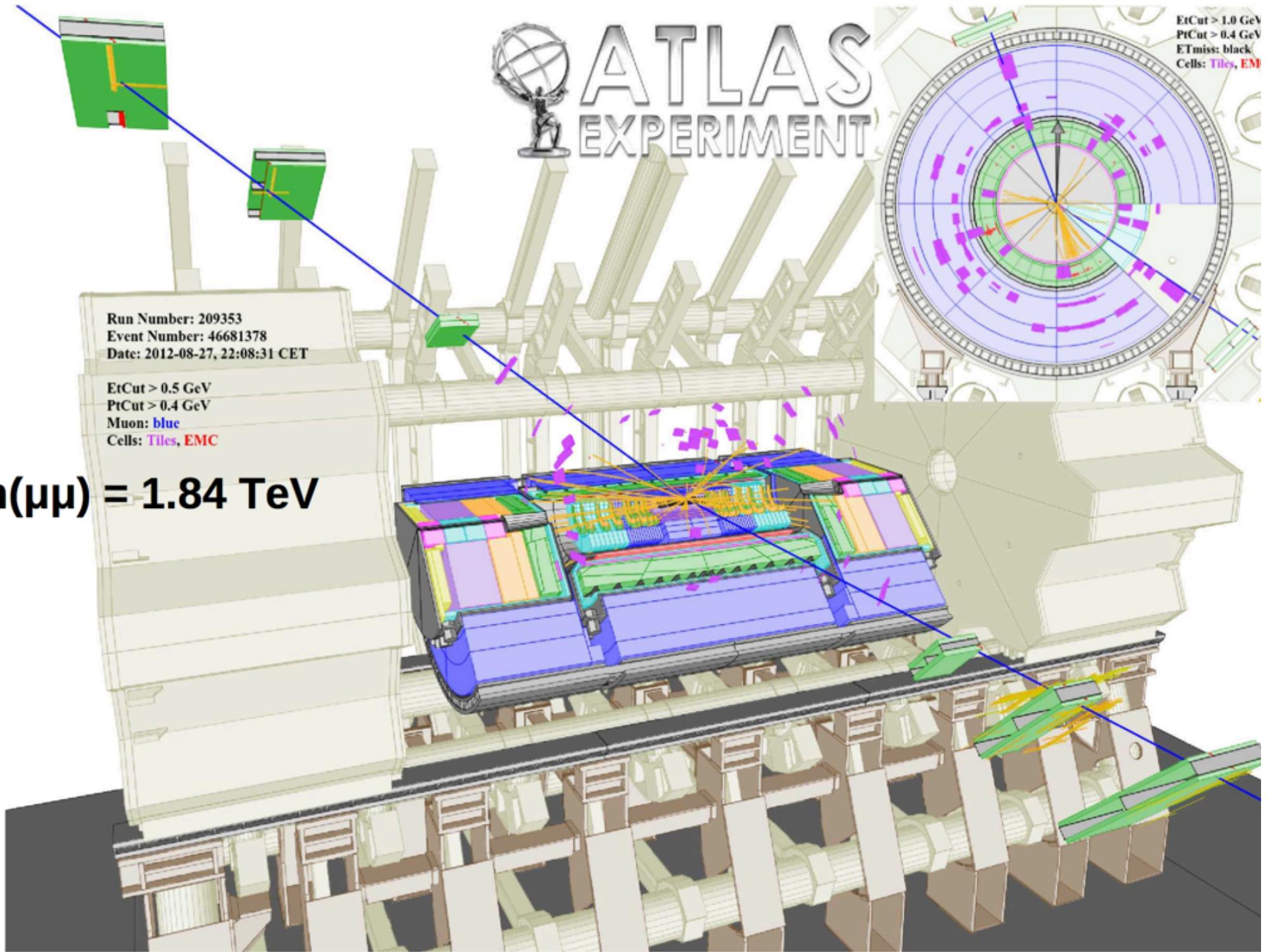
CMS

- Single muon trigger
 - $pT > 45 \text{ GeV}$
 - $|\eta| < 2.4$
- Suppress cosmic rays
 - $|d_0| < 0.2 \text{ mm}$
 - $|z_{0-\text{vertex}}| < 24 \text{ cm}$
- Suppress jets faking μ 's
 - $\sum pT(\Delta R < 0.3) < 10\% \cdot pT$
 - $|z_{0-\text{vertex}}| < 0.2 \text{ mm}$
- Require opposite charge
- $A \times \varepsilon = 80\%$ ($m = 2.5 \text{ TeV}$)

DILEPTON RESONANCE SEARCH: BACKGROUNDS

- SM Drell-Yan: $\gamma^*/Z \rightarrow l+l-$
 - shape taken from Monte Carlo
 - normalisation taken from Z peak in data
- t-tbar:
 - where tt goes to e+e-, $\mu^+\mu^-$
 - est. from MC, cross-checked in data
 - also includes $Z \rightarrow \tau\tau$, WW, WZ
- Jet Background (for ee):
 - di-jet, W+jet events where the jets are misidentified as electrons/muons
- Cosmic Ray Background (for $\mu\mu$):
 - muons from cosmic rays
 - estimated <0.1 event after vertex and angular difference requirements

...Start taking data and what do we see?

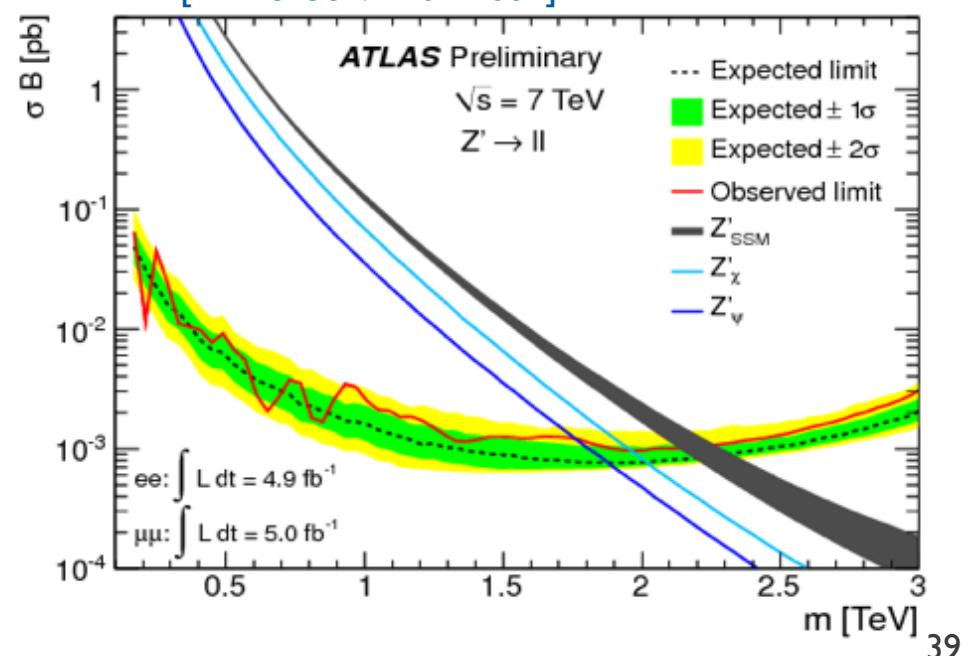
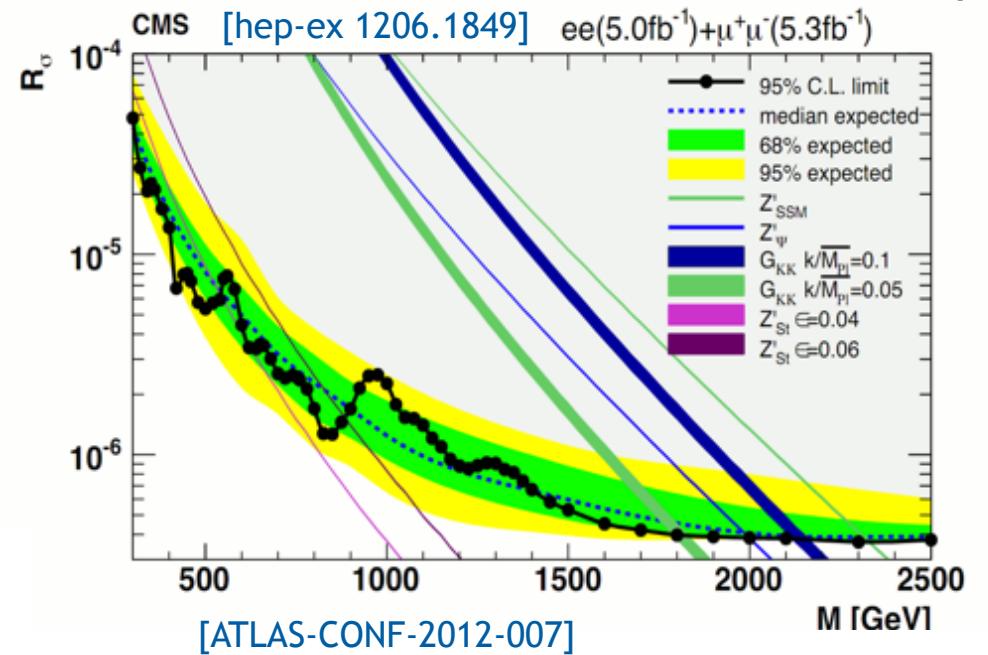
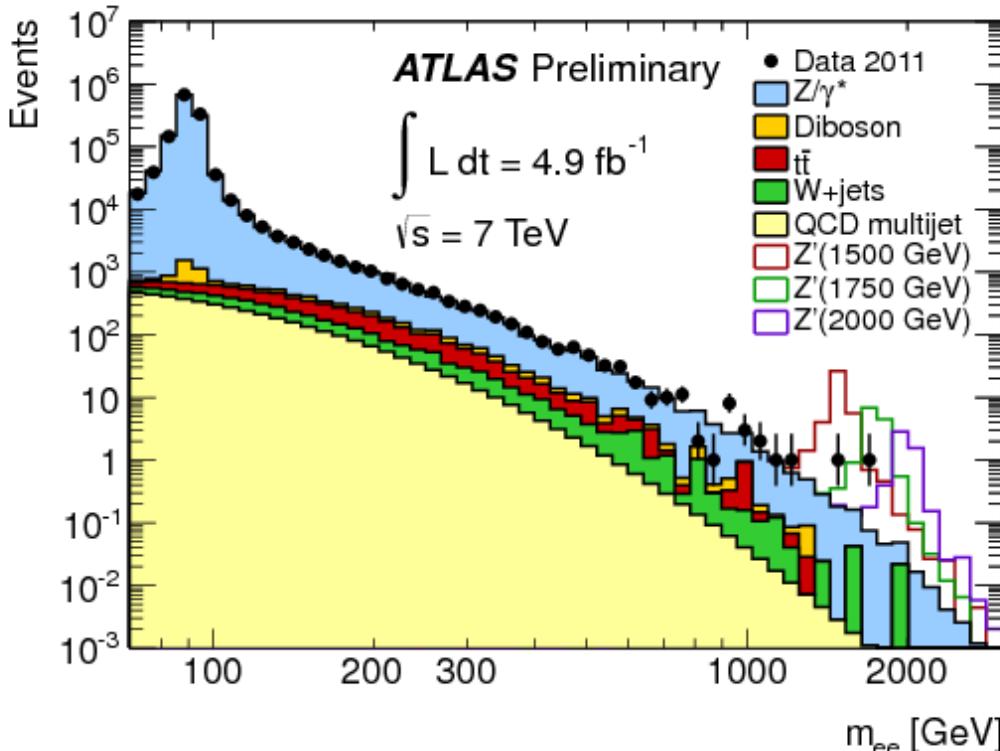


slide circa
March 2012

Z' IN 2011 DATA?

- Many new models have Z-like narrow resonances decaying to dileptons
- Interesting features in dilepton spectra
 - around 2σ each for CMS & ATLAS in $e+\mu$
 - similar in scale to 2011 Higgs excess

Worth watching in 2012's 8 TeV data...

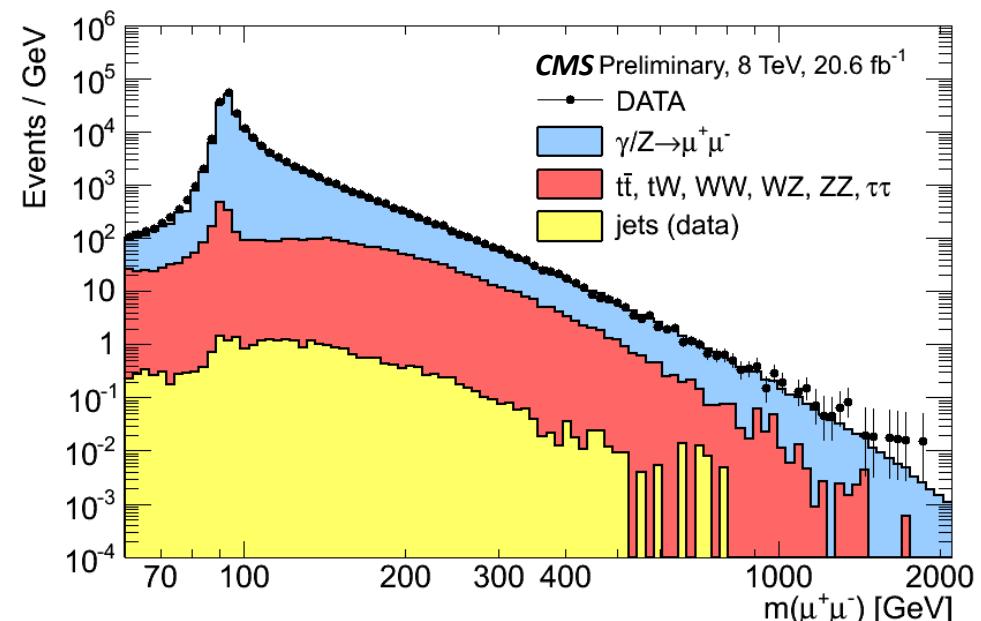
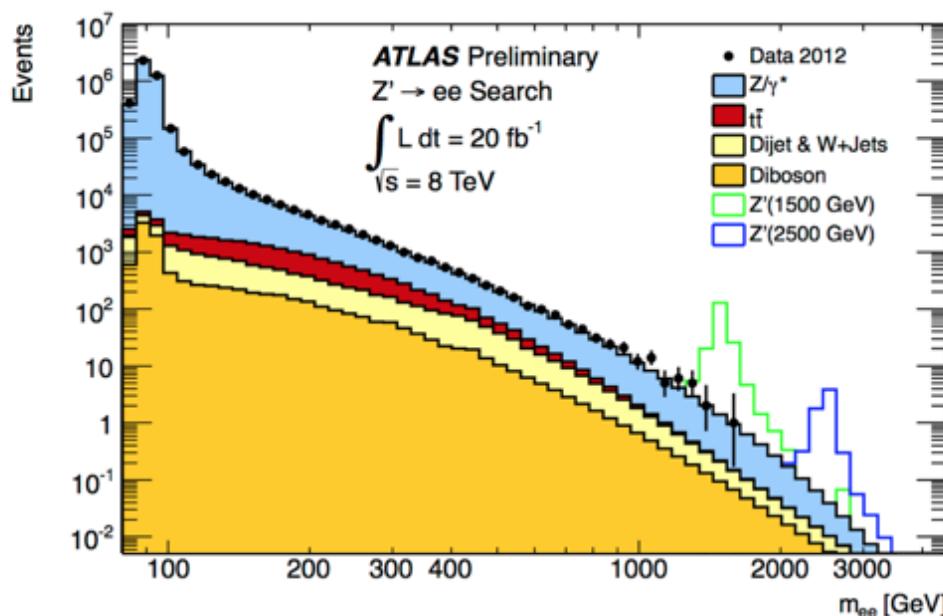


DILEPTON RESONANCE SEARCH

[ATLAS-CONF-2013-017, CMS EXO-12-061]

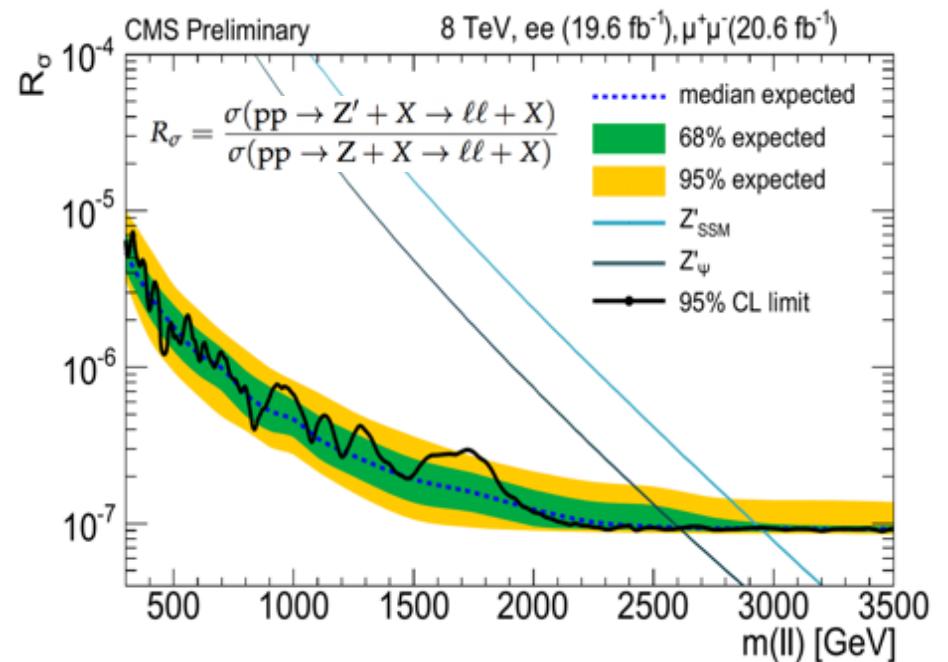
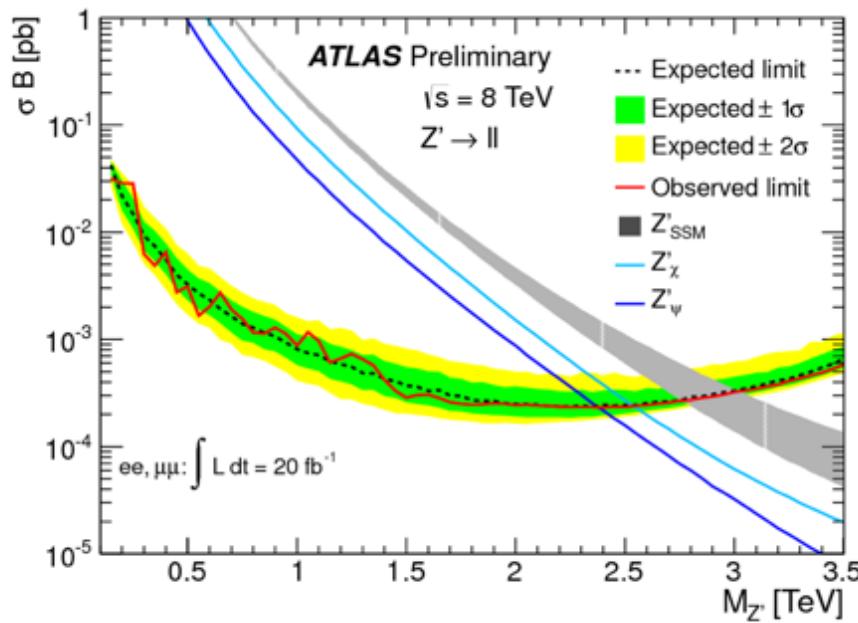
- Event selection
 - CMS: $E_T(e_1, e_2) > 35 \text{ GeV}$, $p_T(\mu_1, \mu_2) > 45 \text{ GeV}$, plus isolation criteria
 - ATLAS: $E_T(e_1, e_2) > (40, 30) \text{ GeV}$, $p_T(\mu_1, \mu_2) > 25 \text{ GeV}$, plus isolation criteria
- Backgrounds
 - Z/γ^* , $t\bar{t}$, tW , VV , $Z \rightarrow \tau\tau$, multijets with ≥ 1 jet reconstructed as lepton
 - estimated by functional fit

No obvious excess observed in 2012 data



SEARCH FOR Z' (DILEPTON RESONANCE)

- Both experiments analysed full 8 TeV datasets, combined ee and $\mu\mu$ channels
- No excess; limits set for a variety of narrow resonances (Z'_{SSM} , Z'_ψ , etc.)

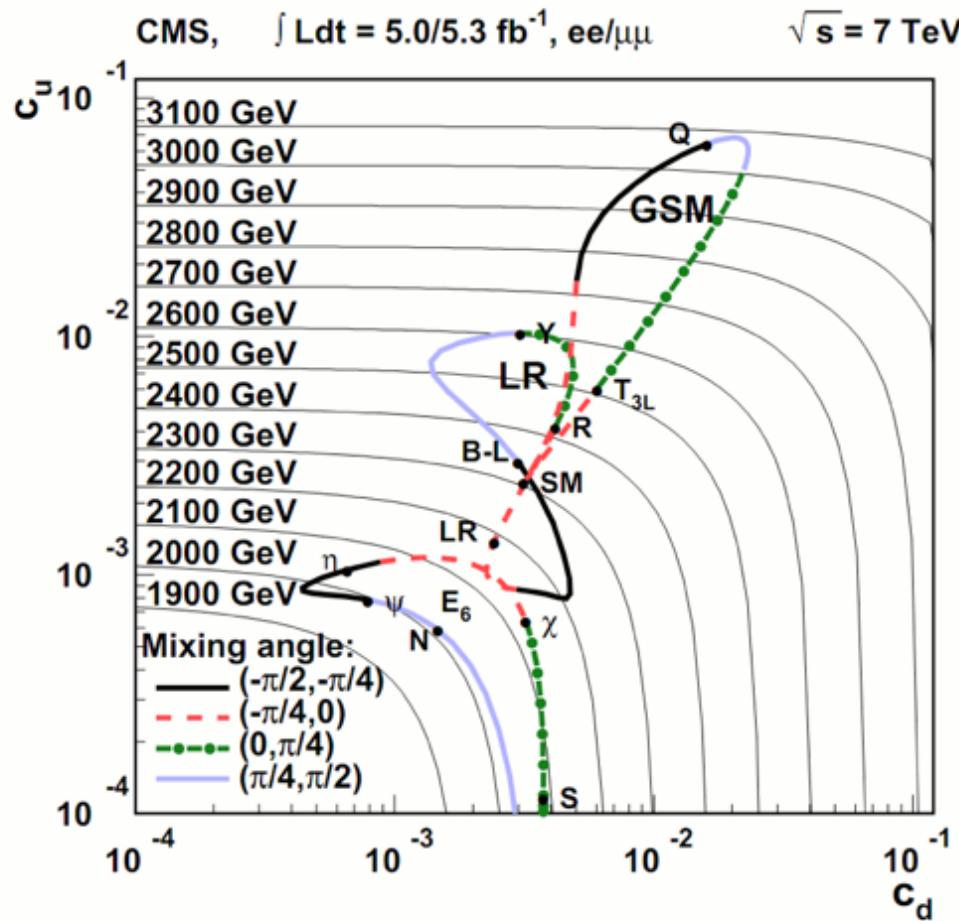


$M(Z'_{SSM})$	expected	observed
CMS	$> 2.96 \text{ TeV}$	$> 2.96 \text{ TeV}$
ATLAS	$> 2.85 \text{ TeV}$	$> 2.86 \text{ TeV}$

Excess in 2011 data just below 1 TeV all but gone in 2012

SEARCH FOR Z' (DILEPTON RESONANCE)

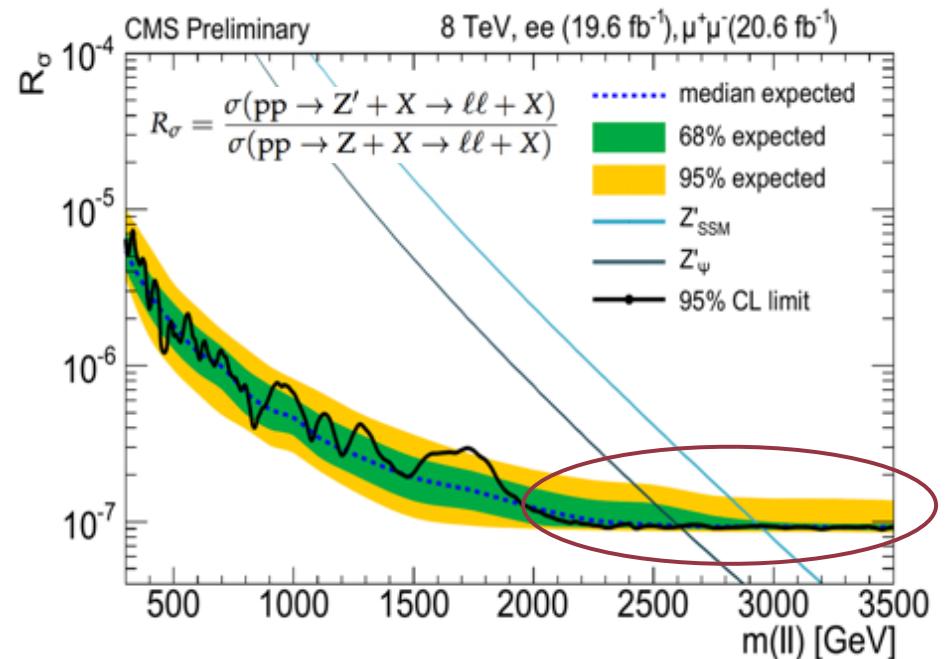
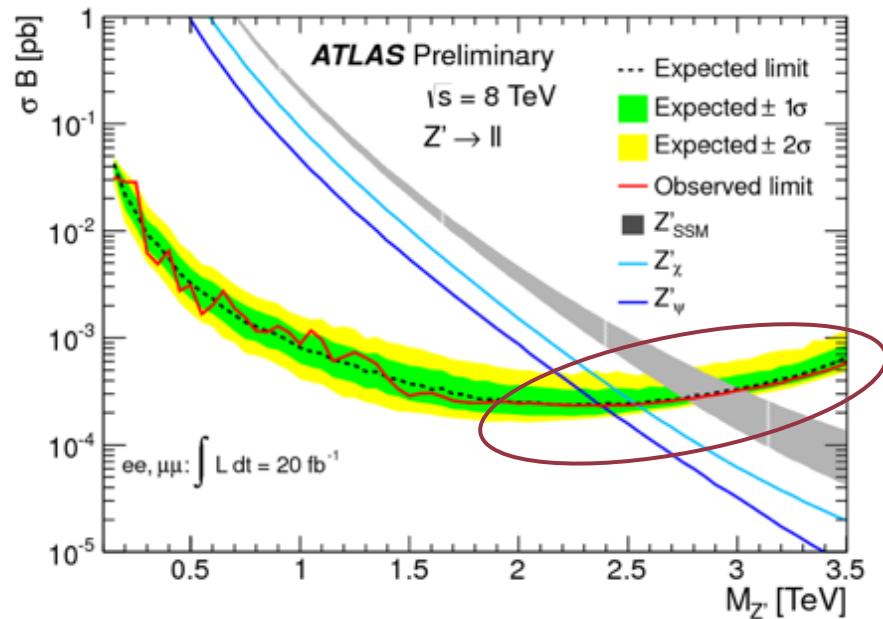
- Working to extend usefulness of search by (for example) plotting vs coupling
- Can plot many Z'-like models in one, unfortunately not a simple mass-vs-coupling



For theorists: if our searches are almost what you want, but not quite, then contact us

SEARCH FOR Z' (DILEPTON RESONANCE)

- Both experiments analysed full 8 TeV datasets, combined ee and $\mu\mu$ channels
- No excess; limits set for a variety of narrow resonances (Z'_{SSM} , Z'_ψ , etc.)

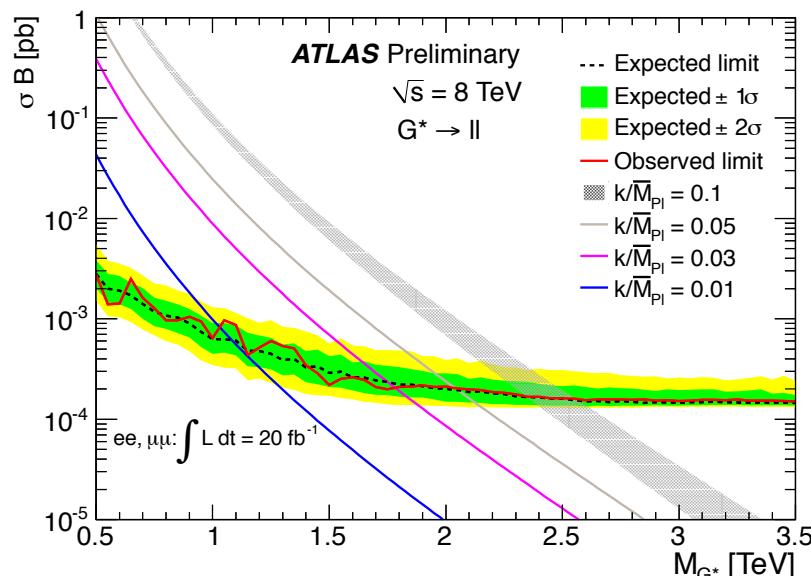


Interesting also to look at the differences in the searches...

ATLAS AND CMS DIFFERENCES?

ATLAS

- Uses signal templates for Z' limits
- Loss of sensitivity at high masses (parton luminosities)
- Cross section limits model-specific
- narrow resonance G^* has no rise:



CMS

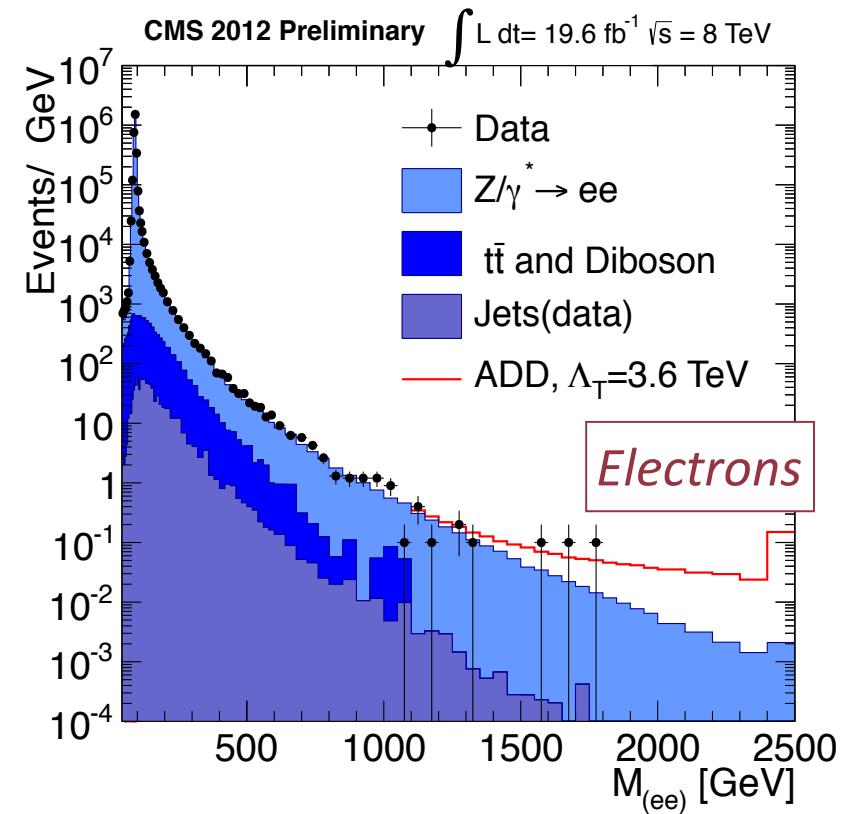
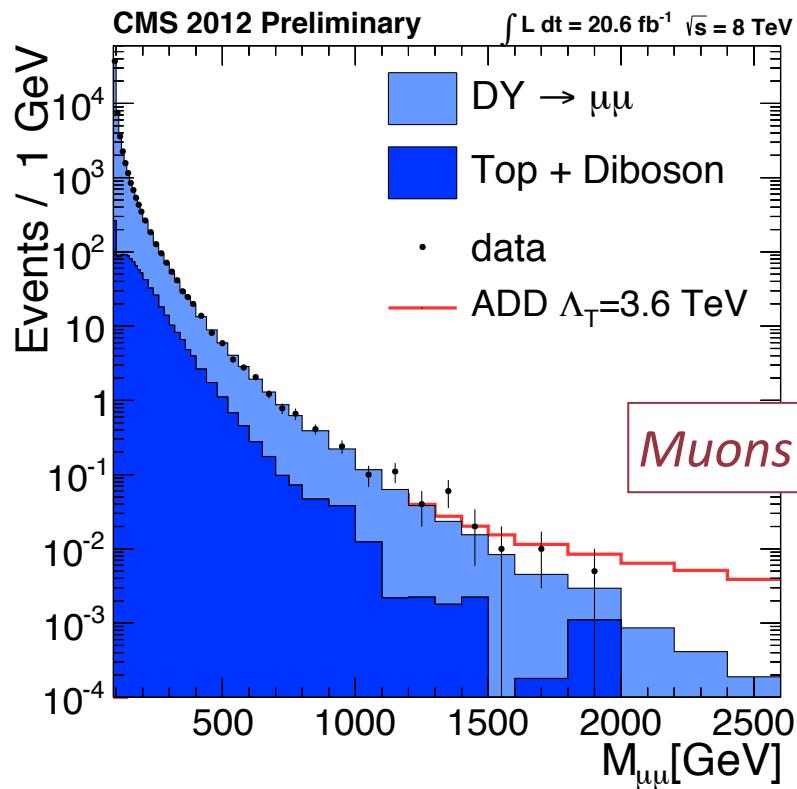
- Use narrow resonance window
 - Cross section upper limits less model dependent
 - Give outside world description of what was done
- Take signal shapes within $\pm 40\%$ of the mass peak into account to compute theory curves
- Not sensitive to parton luminosities
- Generic resonance search

Big difference in efficiency and approach... small difference in limit

SEARCH FOR EXTRA DIMENSIONS IN DILEPTONS

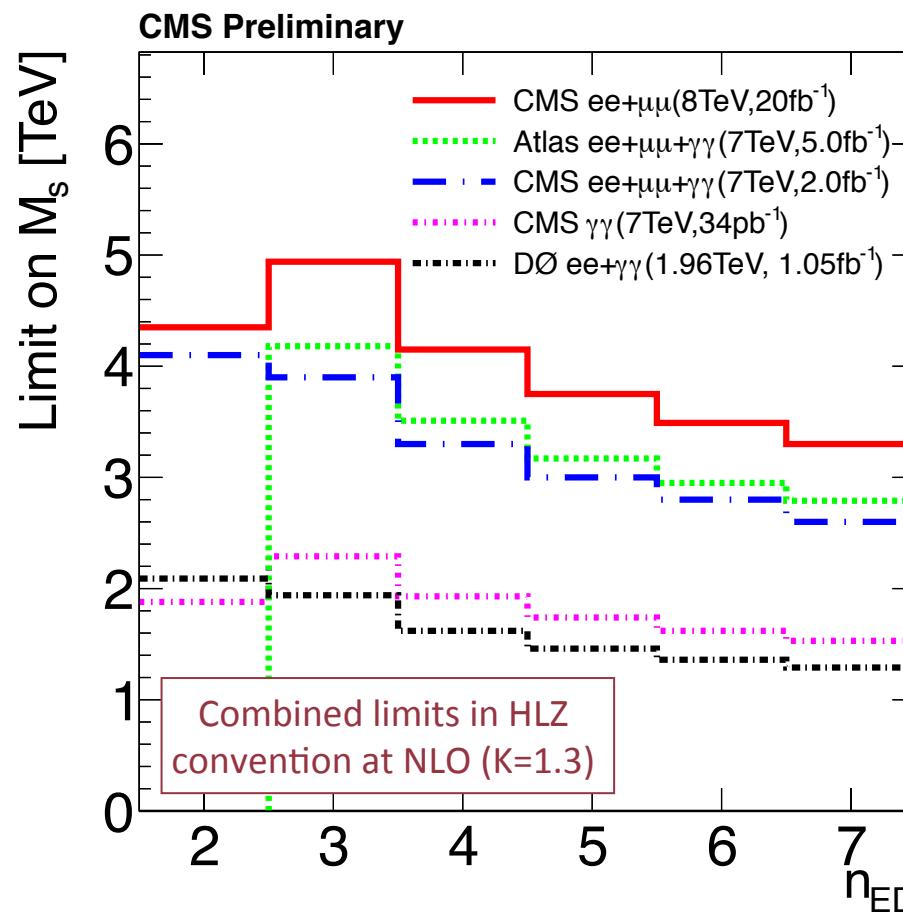
[CMS EXO-12-027, CMS EXO-12-031]

- Extra Dimension search in dilepton invariant mass spectra (same as Z')
- Simple counting experiment on integral above a mass threshold (Bayesian)
- Leading systematics from momentum scale (muons) and PDF (electrons)



EXTRA DIMENSIONS IN DILEPTONS

[CMS EXO-12-027, CMS EXO-12-031]



M_s (ADD) at LO	Lumi.	$\delta=3$	$\delta=3$	$\delta=6$	$\delta=6$	Λ_T (GRW)
95% CL limits	[fb^{-1}]	Exp.	Obs.	Exp.	Obs.	[TeV]
CMS dimuon	20.6	4.34	4.33	3.07	3.06	3.64
CMS dielectron	19.6	4.62	4.64	3.27	3.28	3.90
Combined:	20.6+19.6	4.76	4.77	3.37	3.37	4.01

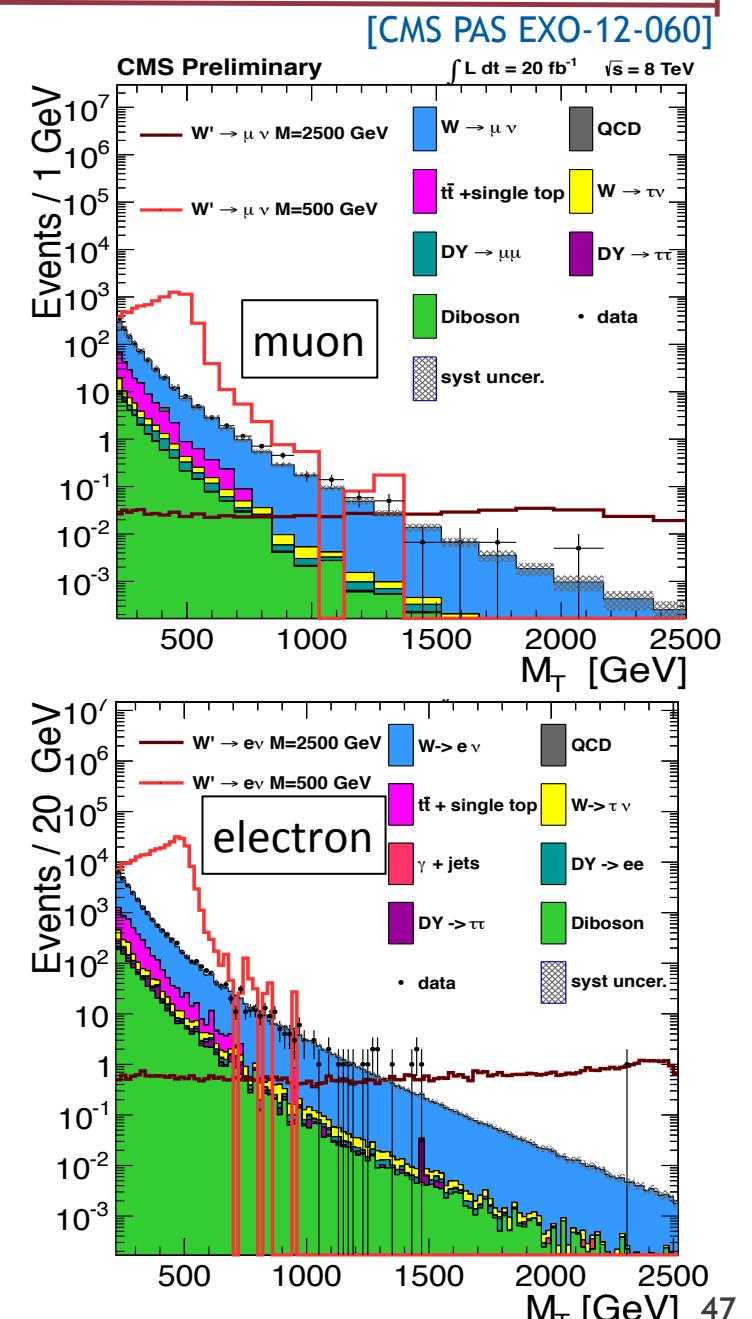
$W' \rightarrow l\nu$ IN 8 TeV DATA

- Search for a new heavy gauge boson W' decaying to a charged lepton (μ or e) and ν

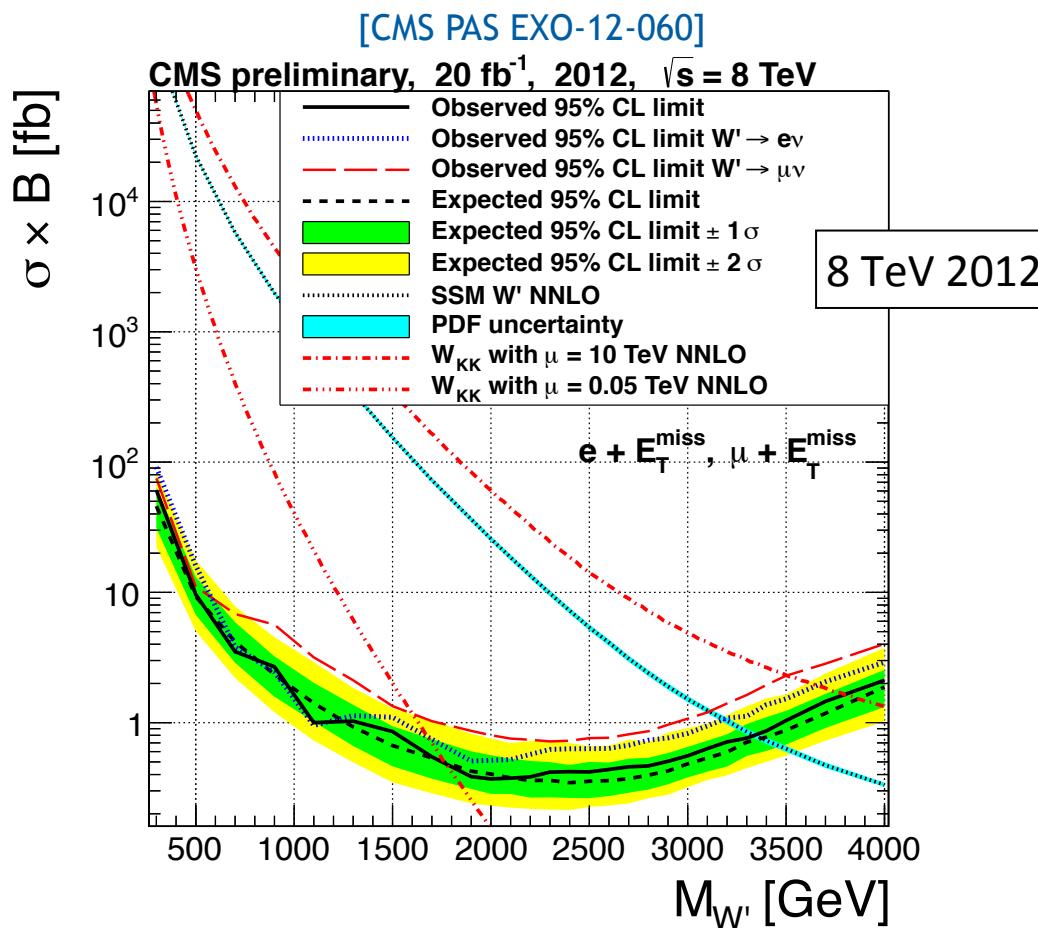
$$M_T = \sqrt{2 \cdot p_T^\ell \cdot E_T^{\text{miss}} \cdot (1 - \cos \Delta\phi_{\ell,\nu})}$$

- Many models possible
 - right-handed W' bosons with standard-model couplings
 - left-handed W' bosons including interference
 - Kaluza-Klein W' -states in split-UED
 - Excited chiral boson (W^*)
- Event Selection and Backgrounds
 - back-to-back isolated lepton and E_T^{miss}
 - Plot transverse mass of $l\nu$ system
 - backgrounds from W , QCD, $t\bar{t}$ +single t , DY, VV from data

No significant excess observed



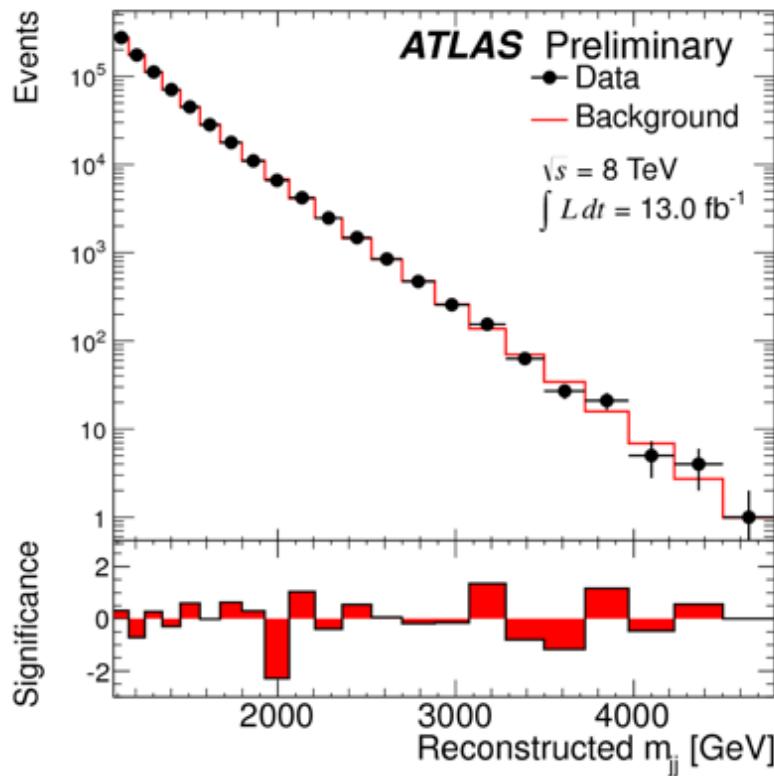
$W' \rightarrow l\nu$ IN 7 AND 8 TeV



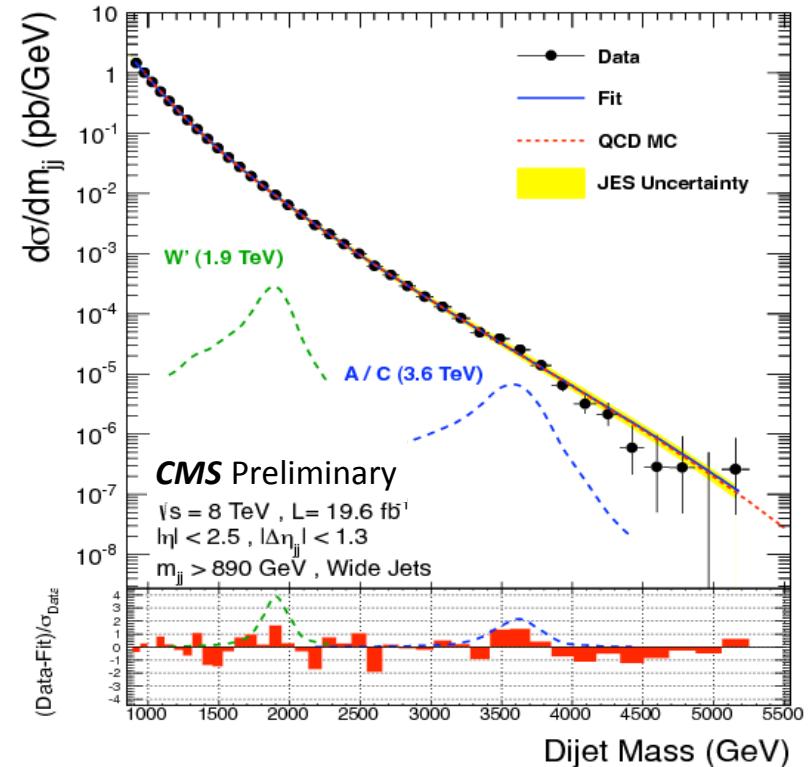
$M(W'_{\text{SSM}}) \text{ 95% CL}$	Luminosity	Expected	Observed
ATLAS $e+\mu$, 2011	4.7	$> 2.55 \text{ TeV}$	$> 2.55 \text{ TeV}$
CMS $e+\mu$, 2012	3.7	$> 2.80 \text{ TeV}$	$> 2.85 \text{ TeV}$
CMS $e+\mu$, 2012	20	$> 3.37 \text{ TeV}$	$> 3.35 \text{ TeV}$

DIJETS IN 8 TeV DATA

[ATLAS-CONF-2012-148]



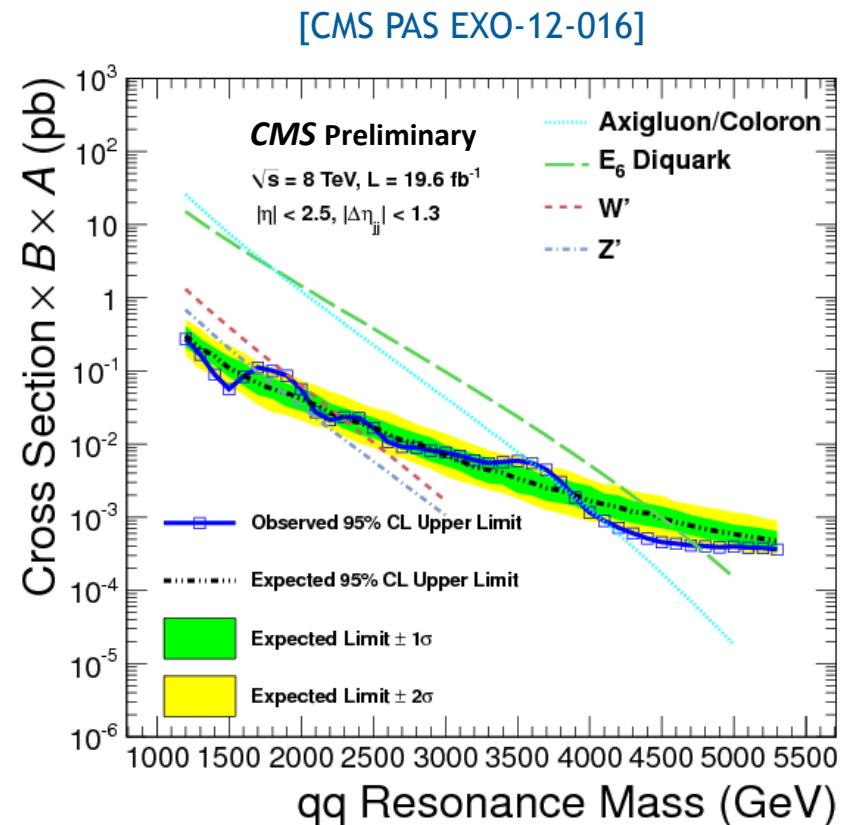
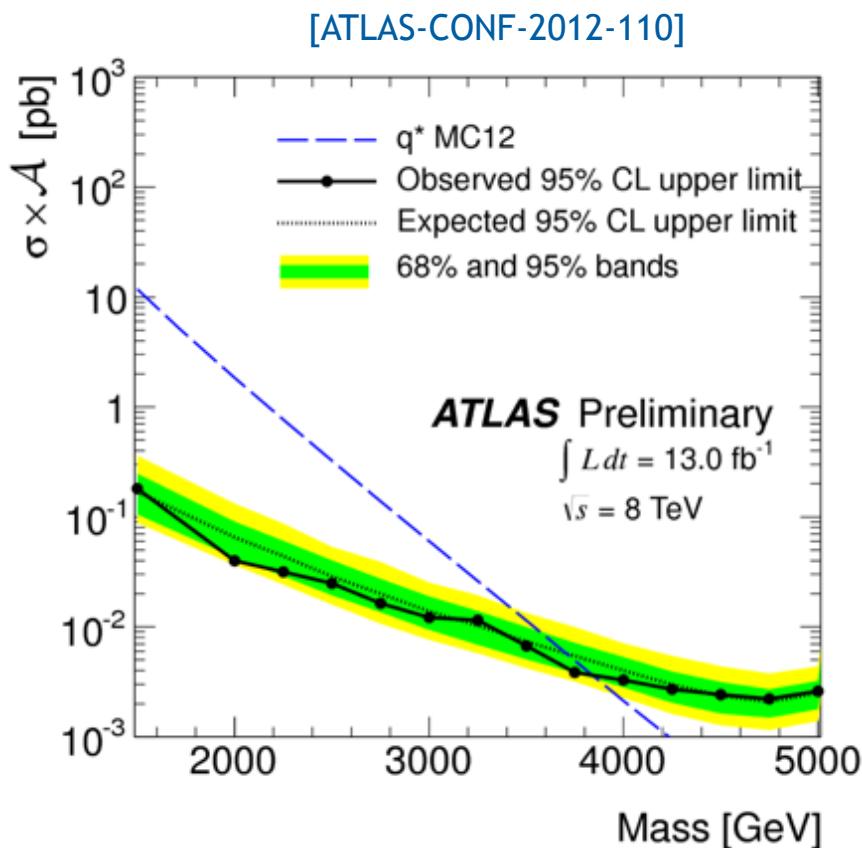
[CMS PAS EXO-12-059]



- Search for dijet resonance in smoothly falling mass spectrum
 - leading jet mass $m_{jj} > 0.9\text{-}1 \text{ TeV}$ from trigger and other constraints
 - Background estimated from smooth functional fit

$$\frac{d\sigma}{dm_{jj}} = \frac{P_0(1-x)^{P_1}}{x^{P_2+P_3 \ln(x)}}$$

DIJETS IN 8 TeV DATA



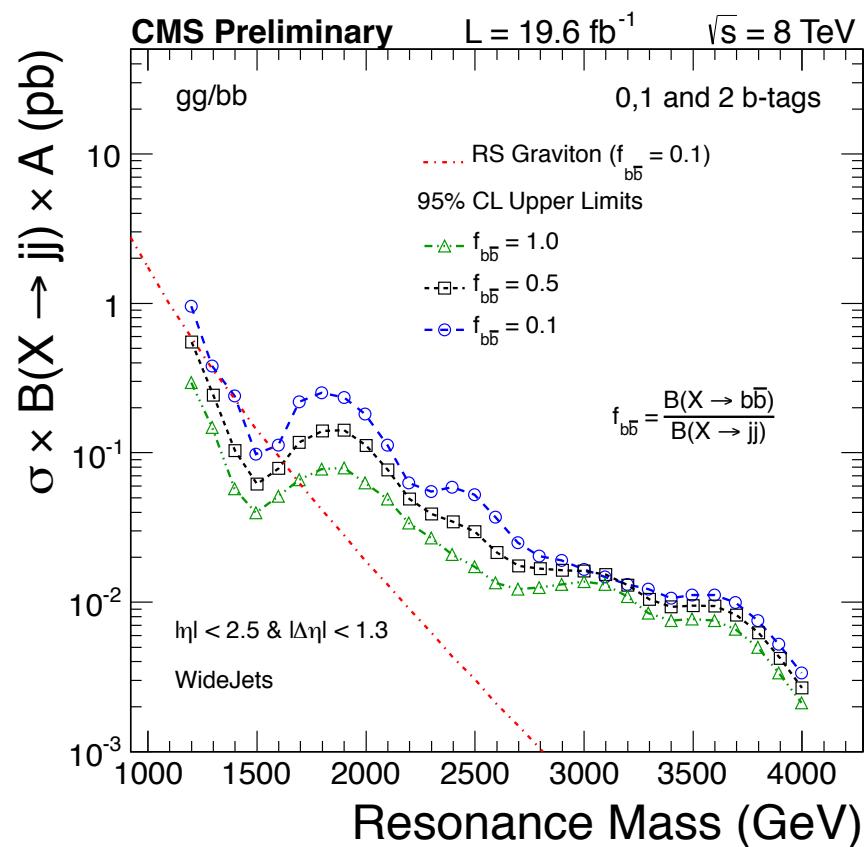
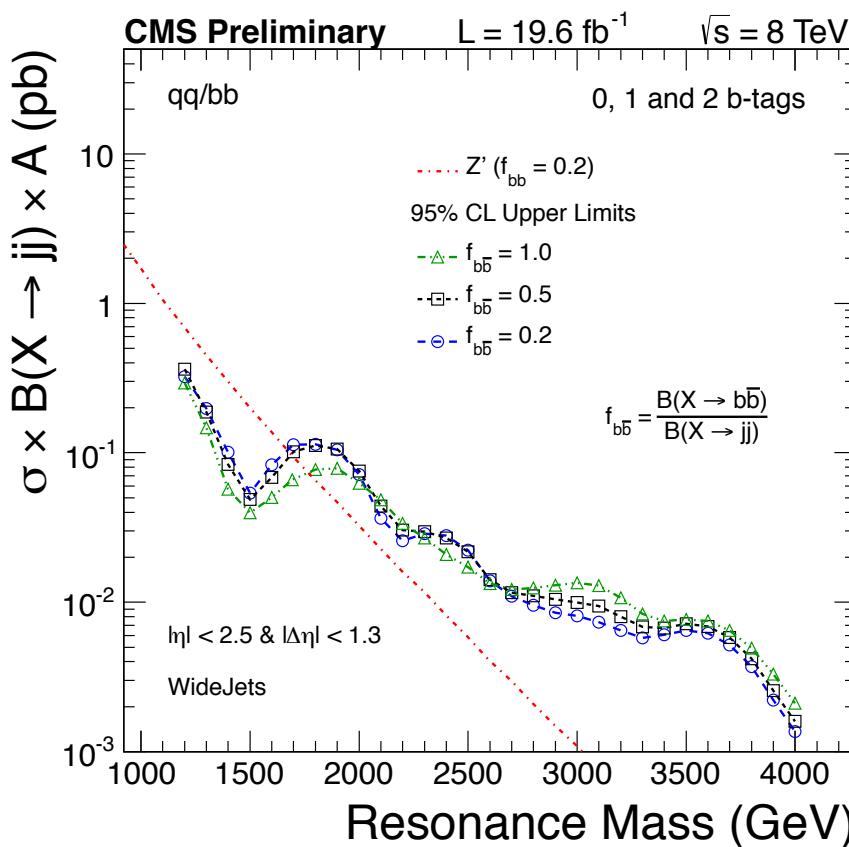
$M(q^*)$ 95% CL	Luminosity	Expected	Observed
ATLAS 2011	4.8	> 3.09 TeV	> 3.55 TeV
CMS 2011	5.0	> 3.27 TeV	> 3.05 TeV
ATLAS 2012	13.0	> 3.70 TeV	> 3.84 TeV
CMS 2012	19.6	> 3.75 TeV	> 3.50 TeV

DIJET WITH b-TAG

[CMS EXO-12-023]

- Dijet with 0, 1, 2 b-tags
 - model-independent limits vs. BR
 - Simultaneous search in 0, 1 and 2 b-tags
 - Limits set on qq, gg and bg (Z' , G_{RS} and b^* models)

$$f_{b\bar{b}} = \frac{\text{BR}(X \rightarrow b\bar{b})}{\text{BR}(X \rightarrow jj)}$$

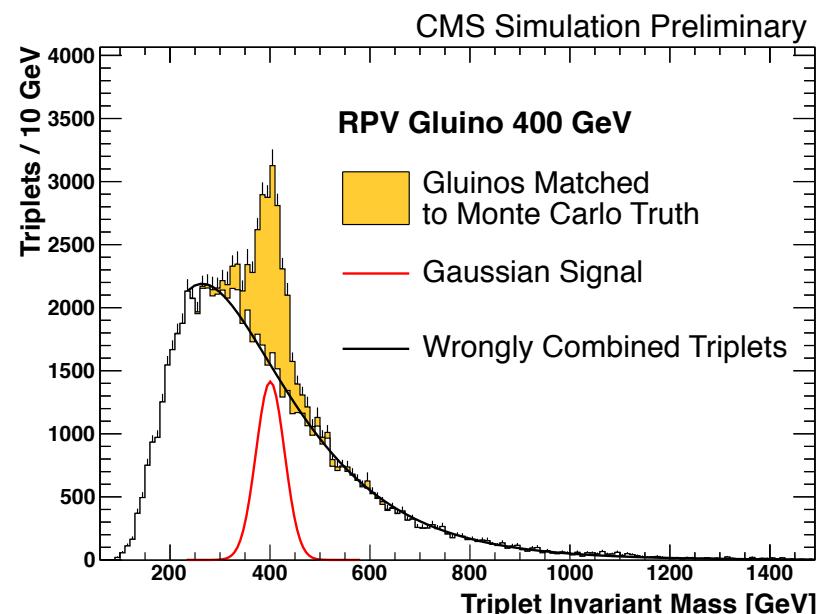
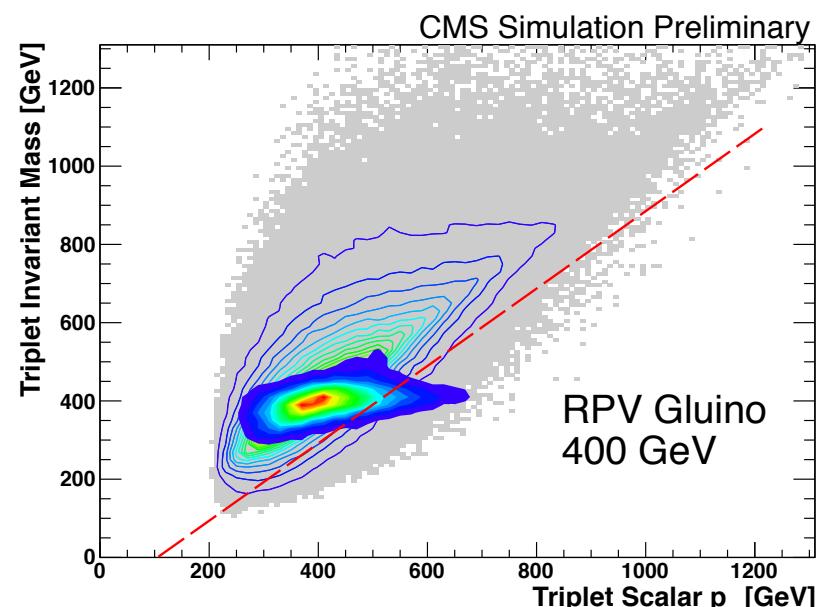


PAIR-PRODUCED THREE JET RESONANCES

- Search for strongly coupled resonances decaying to three jets
- Benchmark: pair produced gluinos to three jets through UDD RPV coupling
- Event Selection
 - ≥ 6 jets > 60 GeV ($1^{\text{st}}\text{--}4^{\text{th}}$ jet > 80 GeV)
 - Use sphericity to suppress backgrounds at high mass
 - Apply b-tagging for gluino \rightarrow udb/csb scenario
- Combine the six highest p_{T} jets into 20 unique triplet combinations
- To suppress wrong combinations and QCD, only accept triplets that satisfy

$$M_{jjj} < \sum_{i=1}^3 |p_{\text{T}}|_i - \Delta$$

[CMS EXO-12-049]

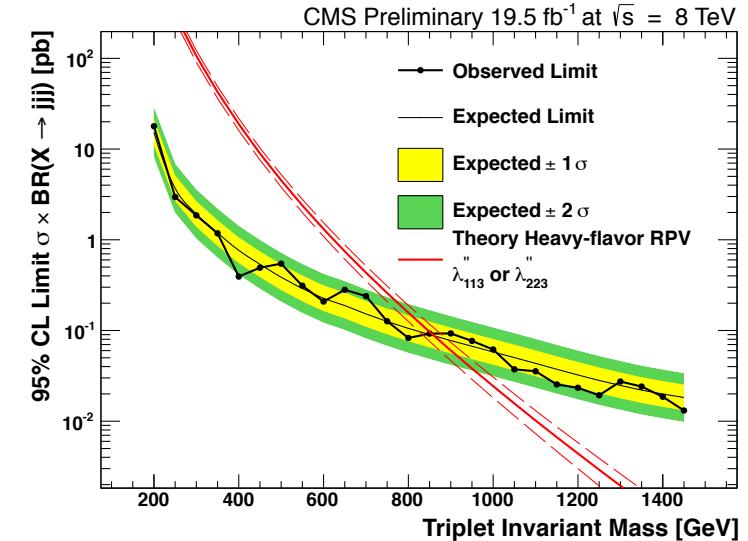
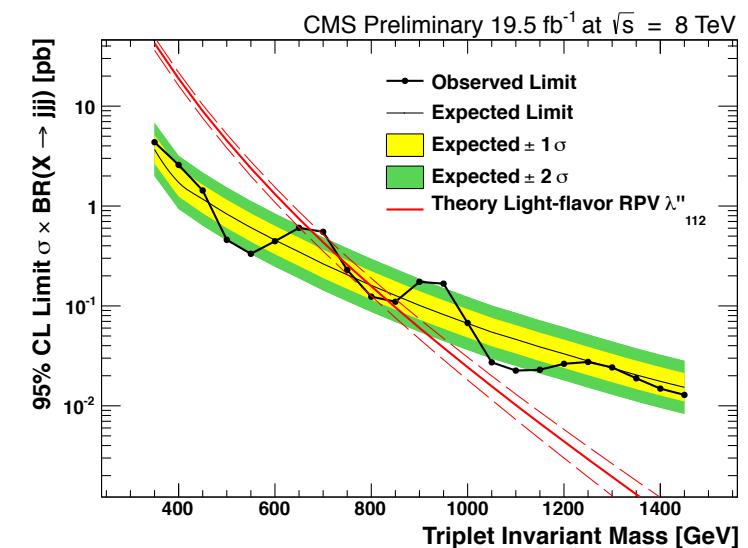
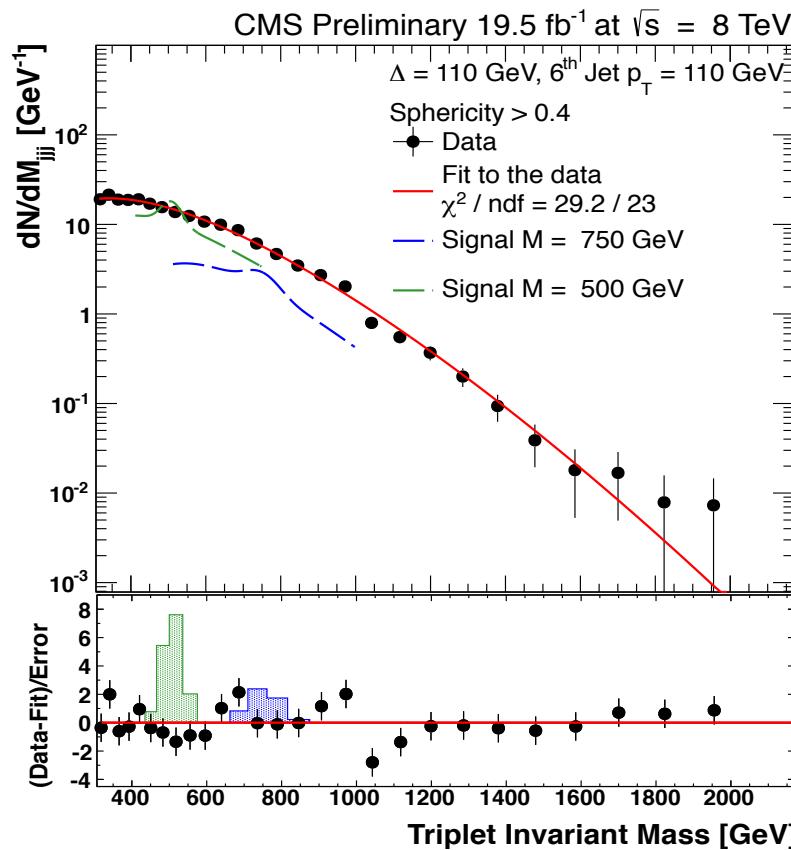


THREE JET RESONANCES

[CMS EXO-12-049]

- Look for bump in falling spectrum
 - For b-tagged result, QCD shape from b-vetoed data
 - All-hadronic top bkgnd consistent w/ SM rate

*Limits on RPV gluinos < 650 GeV (light-flavor)
and between 200 and 835 GeV (heavy-flavor)*



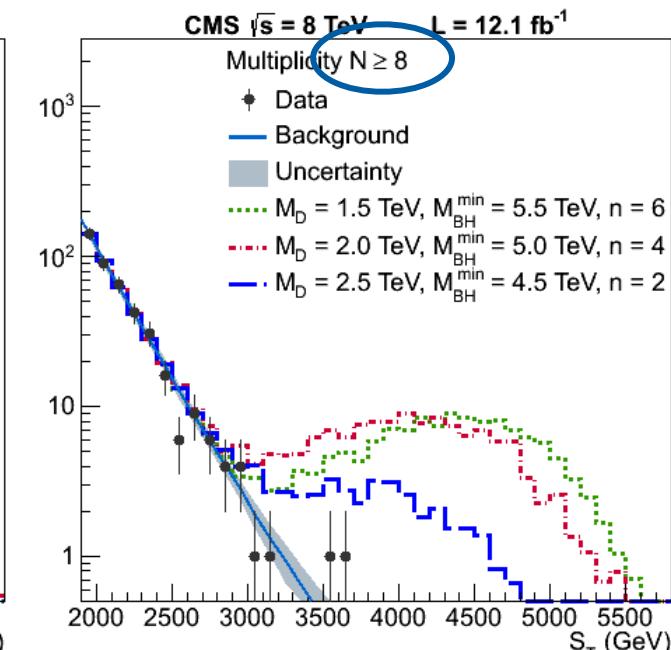
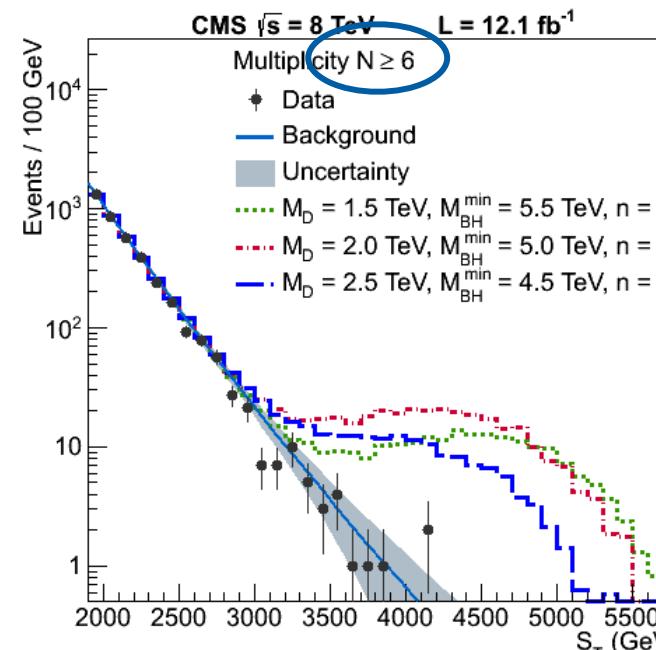
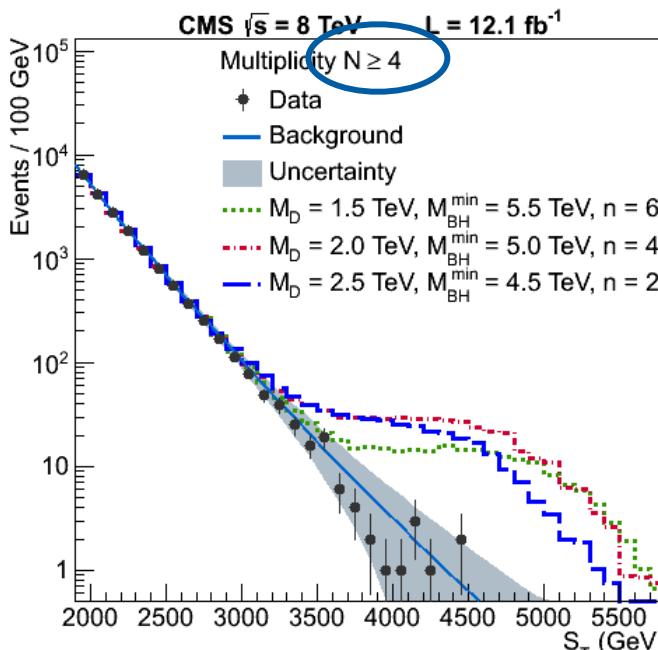
SEARCH FOR BLACK HOLES

[arXiv:1303:5338, EXO-12-009]

- Search for microscopic Black Holes in 12 fb^{-1} of 8 TeV data
 - Hypothetical BH would evaporate into many high- p_T objects
 - Estimate by S_T , the p_T sum of physics objects with $p_T > 50 \text{ GeV}$
- Main background of QCD estimated by fit to $n=2$ distribution
 - Normalised for each multiplicity bin separately at $S_T = 1.8\text{--}2.2 \text{ TeV}$
 - Model-independent limits vs S_T and multiplicity

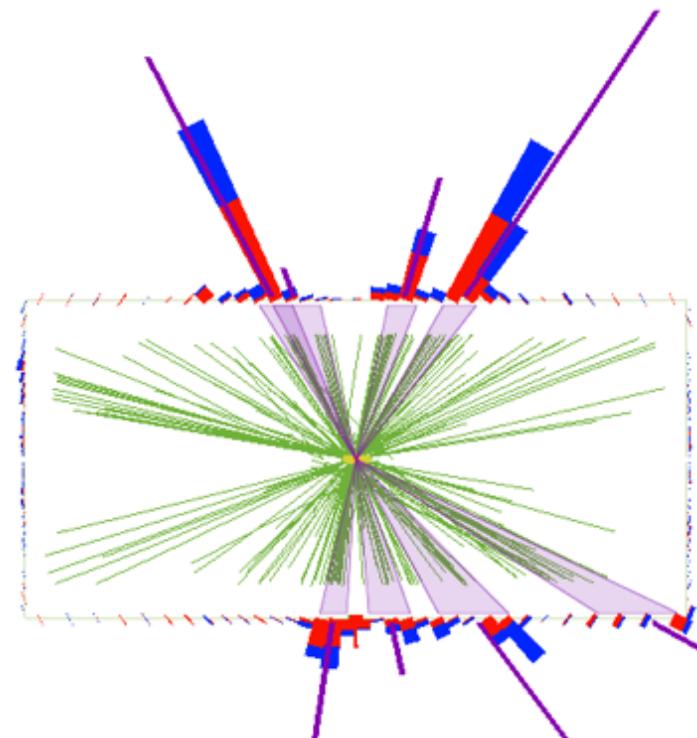
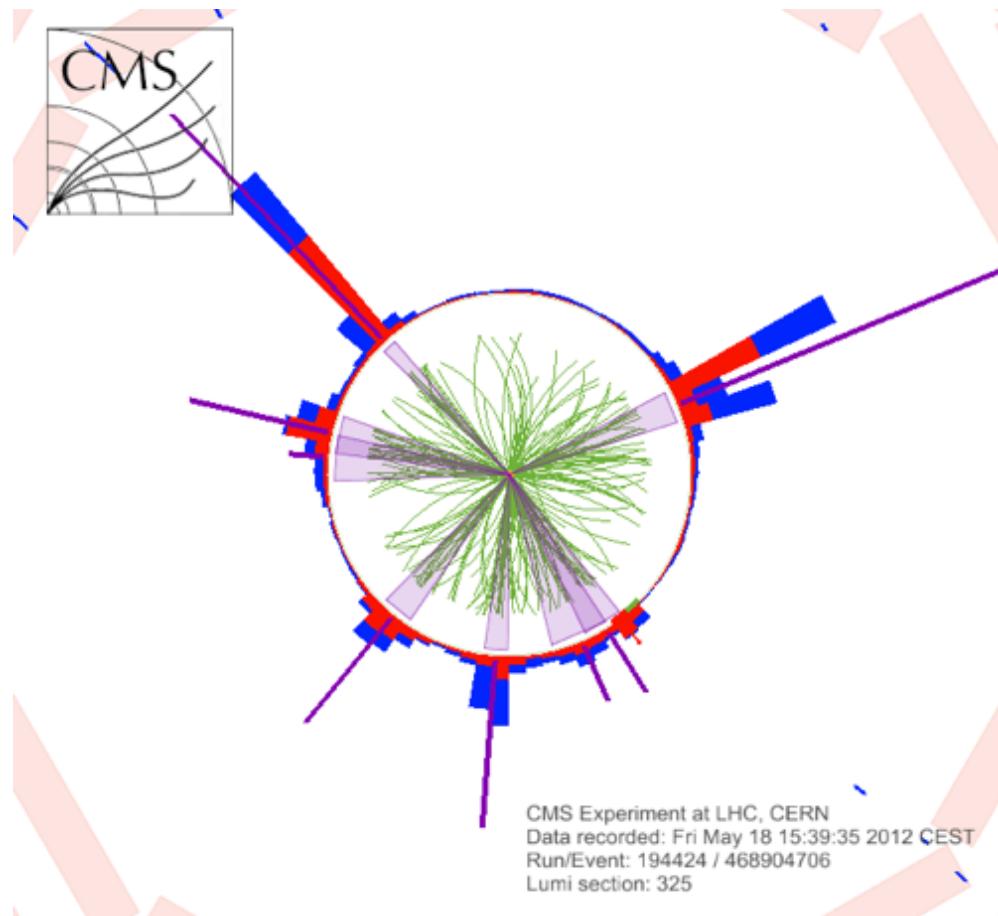
$$S_T = \sum_{j,e,\mu,\gamma,MET}^N p_T$$

Significant improvement in sensitivity ($\sim 15\text{--}20\%$) with respect to 7 TeV data



8-JET EVENT, $S_T = 3$ TeV

[arXiv:1303:5338, EXO-12-009]



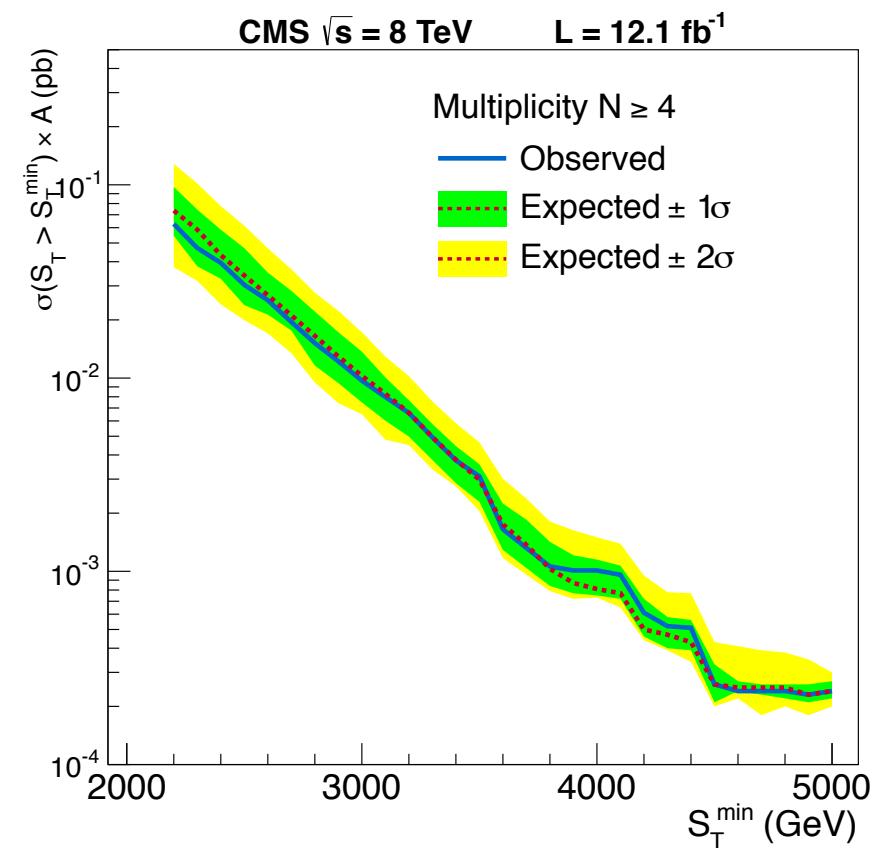
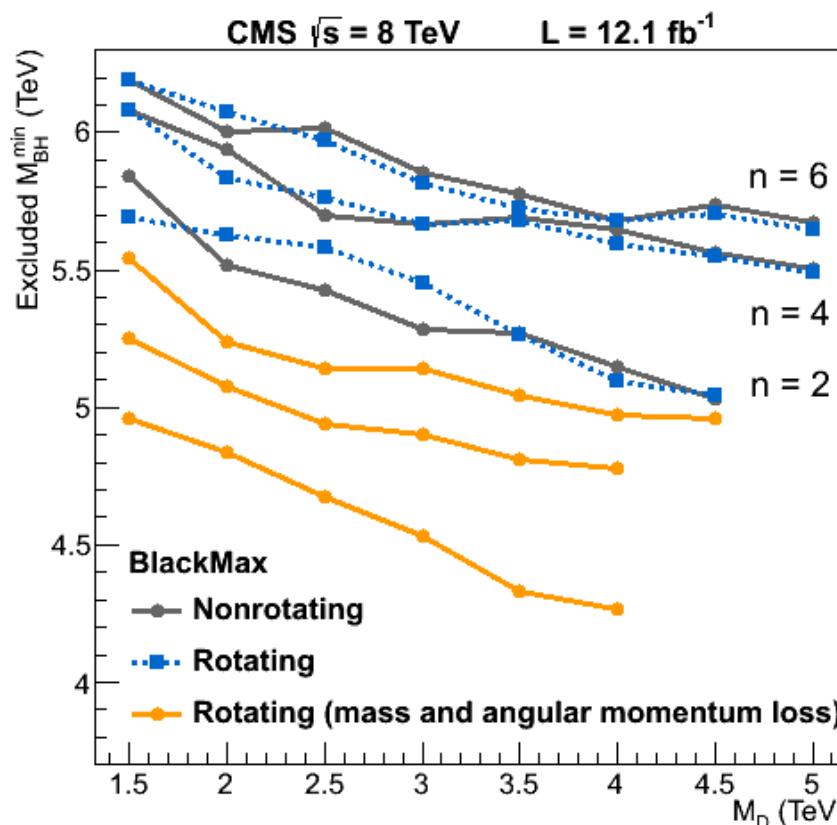
Many interesting events found!

SEARCH FOR BLACK HOLES - HIGH S_T

[arXiv:1303:5338, EXO-12-009]

- No excess of events above expected backgrounds observed
 - Limits on ADD parameter M_D assuming specific BH models (Charybdis, BlackMax, ...)
 - Model-specific limits on semiclassical BH masses in the 4.3 – 6.2 TeV range

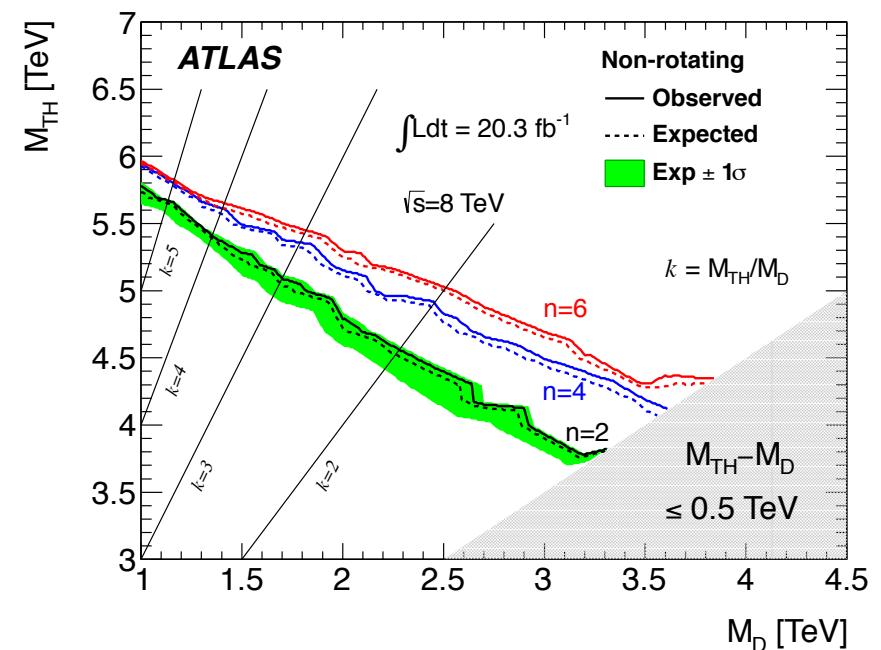
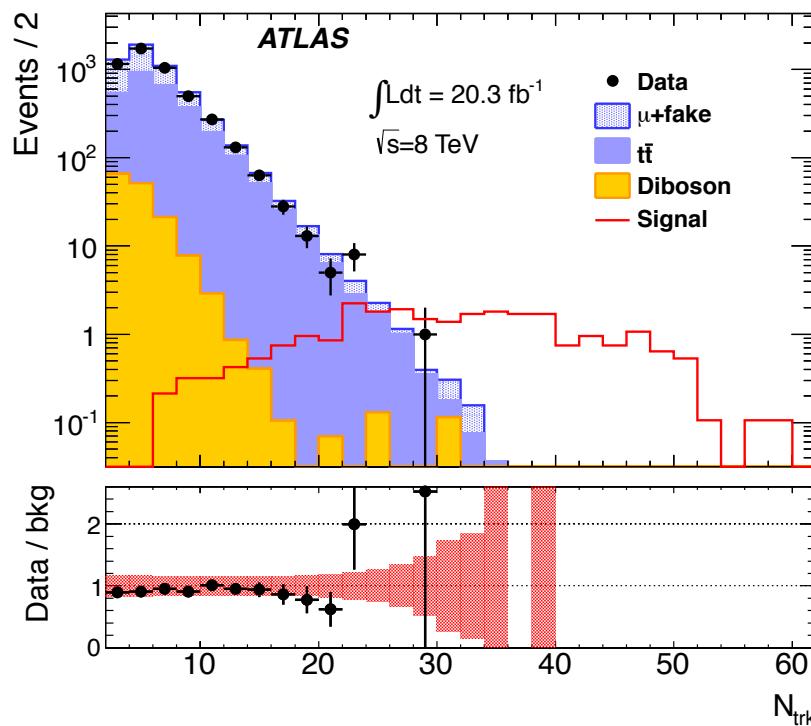
Also interesting as a model-independent search vs S_T and multiplicity



SEARCH FOR BLACK HOLES - SAME-SIGN $\mu\mu$

[ATLAS arXiv:1308:4075]

- If black holes lead to high multiplicity, can just count tracks for same-sign $\mu\mu$
- Use track multiplicity for ($p_T\{\text{trk}\} > 10 \text{ GeV}$) for events with $p_T(\mu) > 100 \text{ GeV}$
- 95% CL exclusion contours for non-rotating BH, rotating BH, and stringball models with $n = 2, 4$, and 6 : up to $\sim 6 \text{ TeV}$ for $M_D = \sim 1 \text{ TeV}$



COFFEE BREAK?

THE Sun

STEPHEN HAWKING
EVEREST CLIMB
LATEST UPDATE
Page 14

Thursday July 5 2012 30p

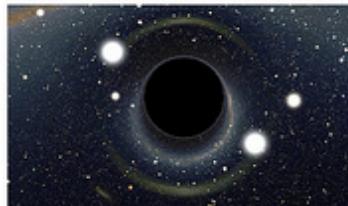
BLACK HOLE EATS CANADA

**COSMIC GIANT PASSES EARTH AND
SNACKS ON NORTH AMERICA**

Yesterday, the entire country of Canada became victim of a hungry Black Hole from outer space. The passing space muncher, who NASA scientists believe is a gravitational mass that will 'pig out' on anything - had decided to pass planet Earth because it felt a tad peckish. Why the galactic 'lard arse' chose Canada is a mystery, but some experts believe that to a Black Hole - the shape of Canada, with all its lakes and mountains, had just simply looked - TASTY!!

(continued on Page 2)

SO WHAT WILL IT HAVE FOR DESSERT? - READ EXPERT OPINIONS - Pages 7 & 8



The Black Hole - looking a bit stuffed after munching up Canada.