

# Data Analysis I

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# Introduction

- ▶ 4 + 4 lectures to present observation systems and the way we analyse their data
- ▶ Observation systems:
  - LIGO / Virgo
  - LISA
  - Pulsar Timing Array
- ▶ Analysis
  - Bayesian / Bayesian
  - Sources: binaries, stochastic background

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# Overview of Data Analysis I

## Statistic basis of DA

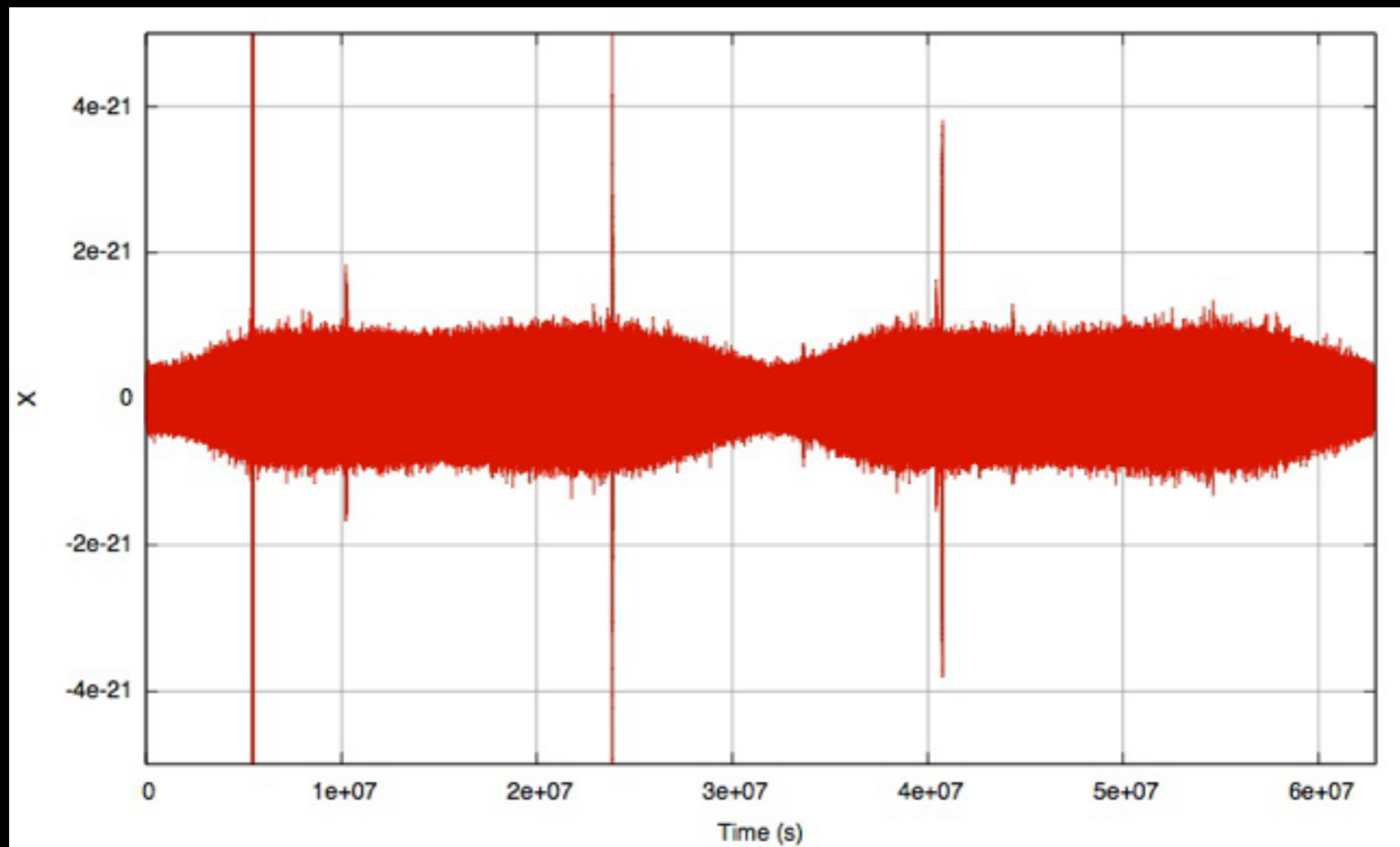
- ▶ Extracting signals from noisy data
  - the likelihood function
- ▶ Noise weighted inner products, Match-filter, Signal to Noise Ratio and
- ▶ Two approaches
  - Frequentist approach
  - Bayesian approach
- ▶ Detection statistics and model evidence

# References

- ▶ Romano & Cornish Liv. Rev. Relativ. (2017) 20:2 “Detection methods for stochastic GW backgrounds: a unified treatment”
- ▶ Janarowski & Krolak LRR (2012) 15, 4 “Gravitational-Wave Data Analysis. Formalism and Sample Applications: The Gaussian Case”
- ▶ Rover et al.
- ▶ Design document of the LISAPathfinder parameter estimation pipeline

# Extracting signal from data

- ▶ Example of eLISA simulated data (LISACode):
  - about 100 SMBHs,
  - Galactic binaries



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- ▶ First « simple data analysis » :



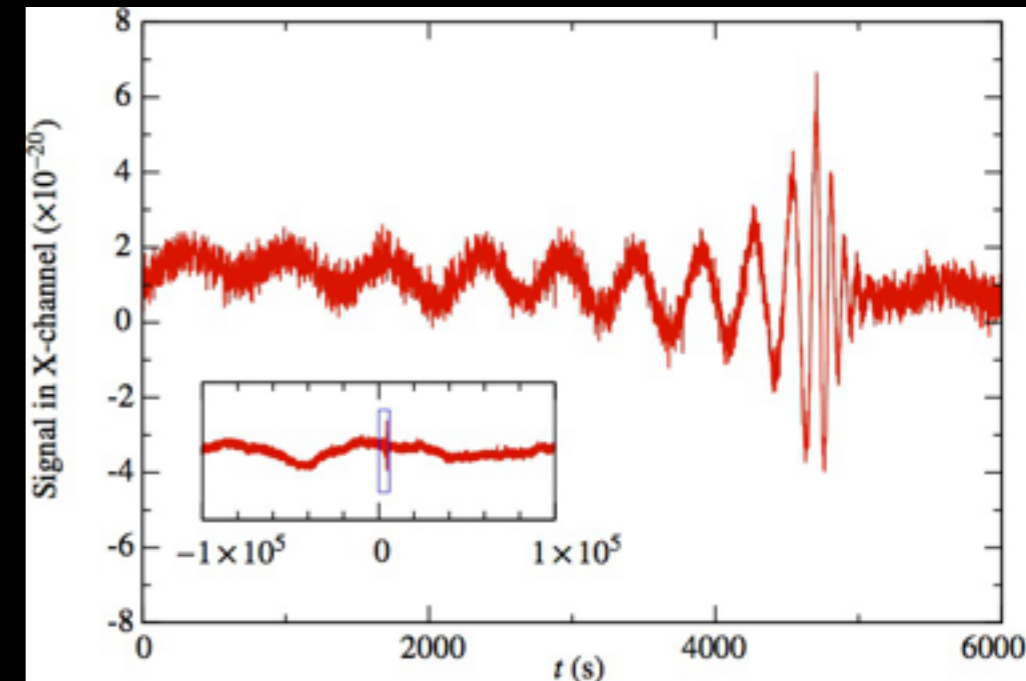
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## ▶ First « simple data analysis » :

- low pass + zoom : we can « see » end of waveform for powerful sources



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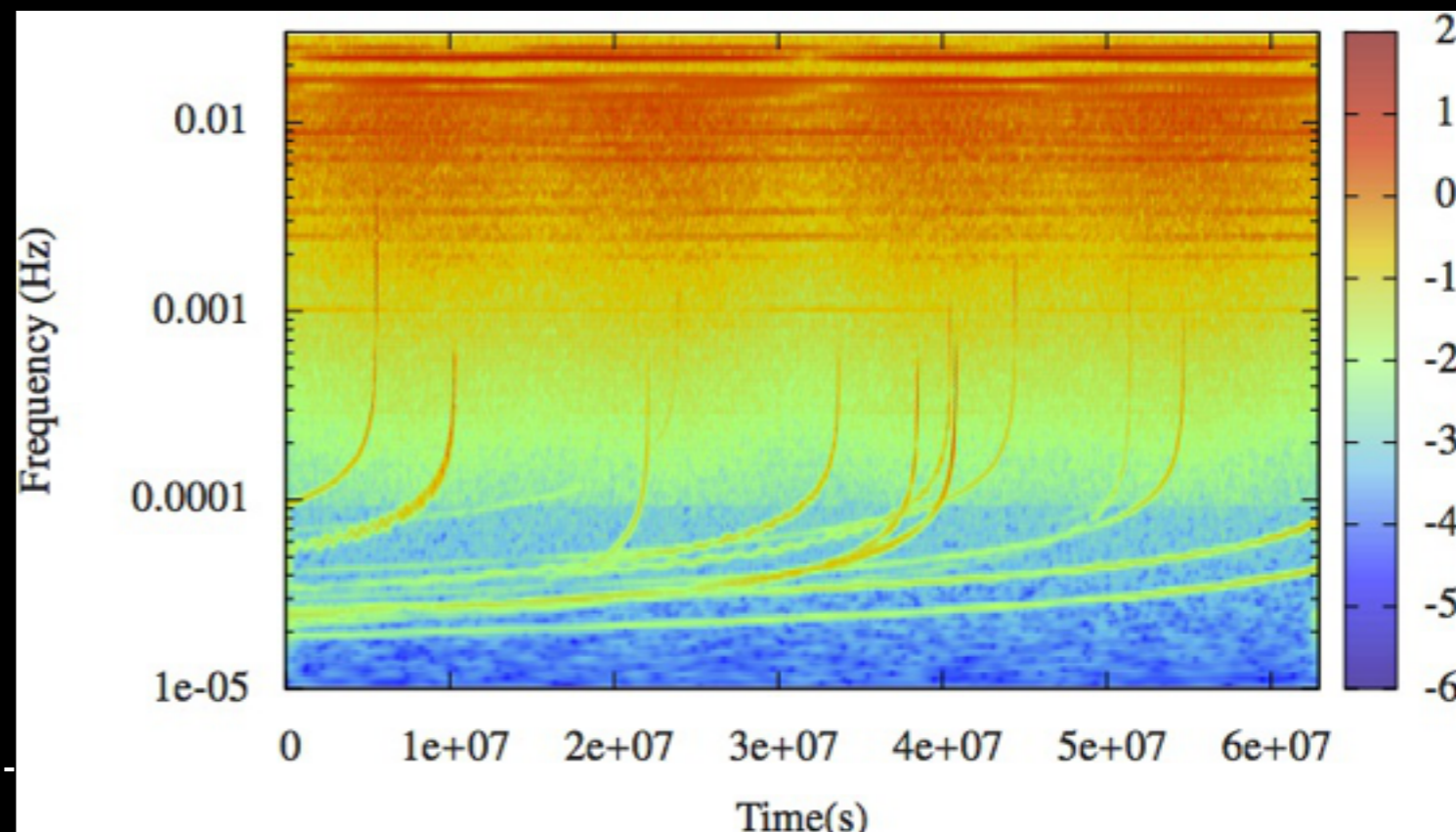
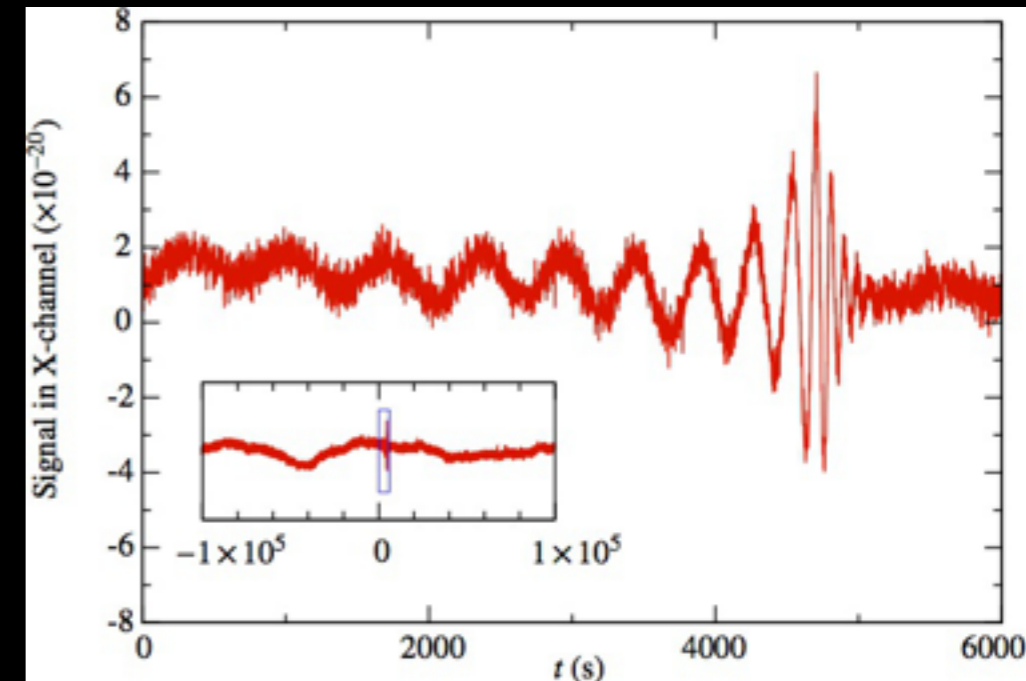
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- wavelet transform :
  - chirps
  - Galactic binaries



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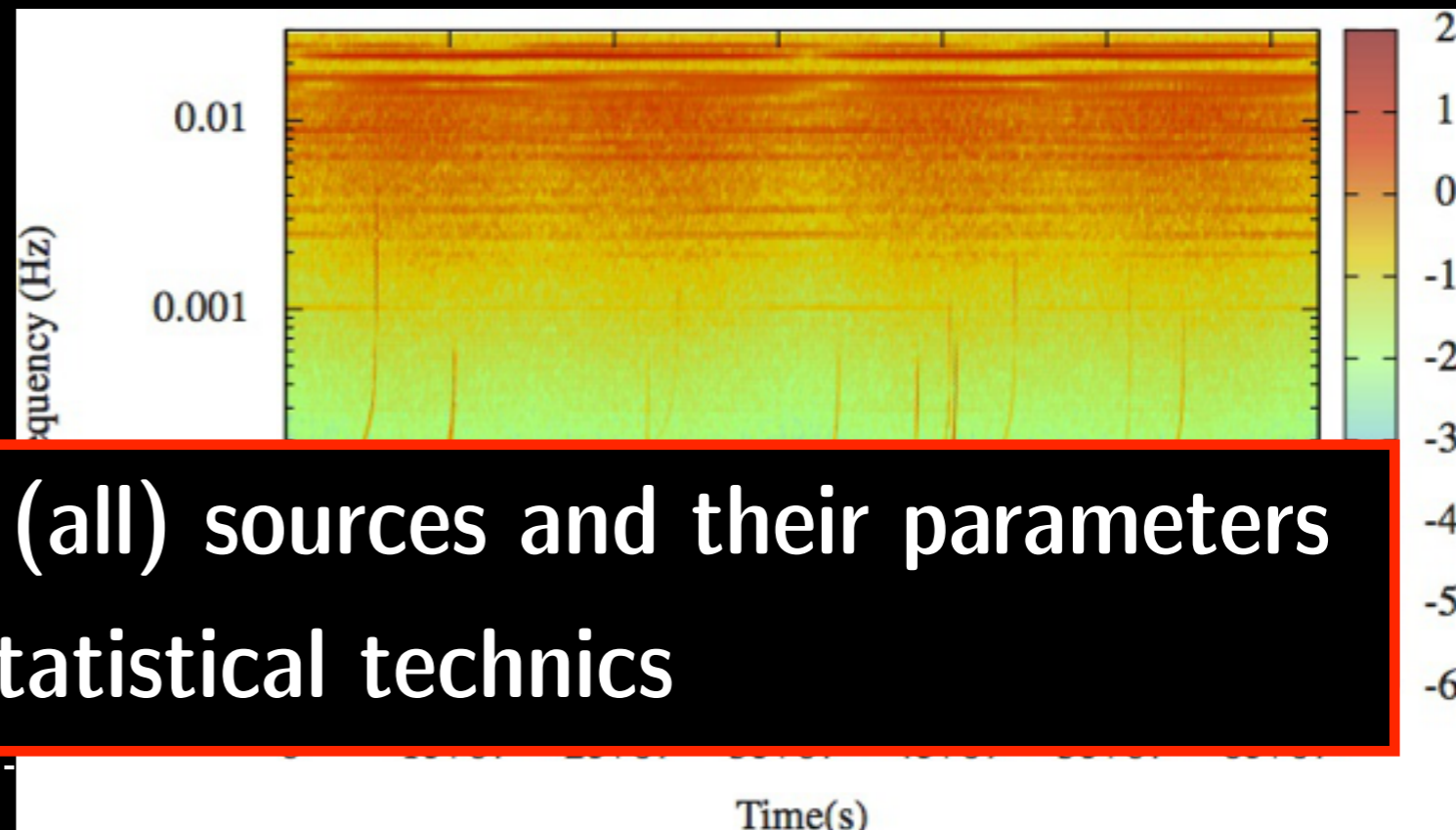
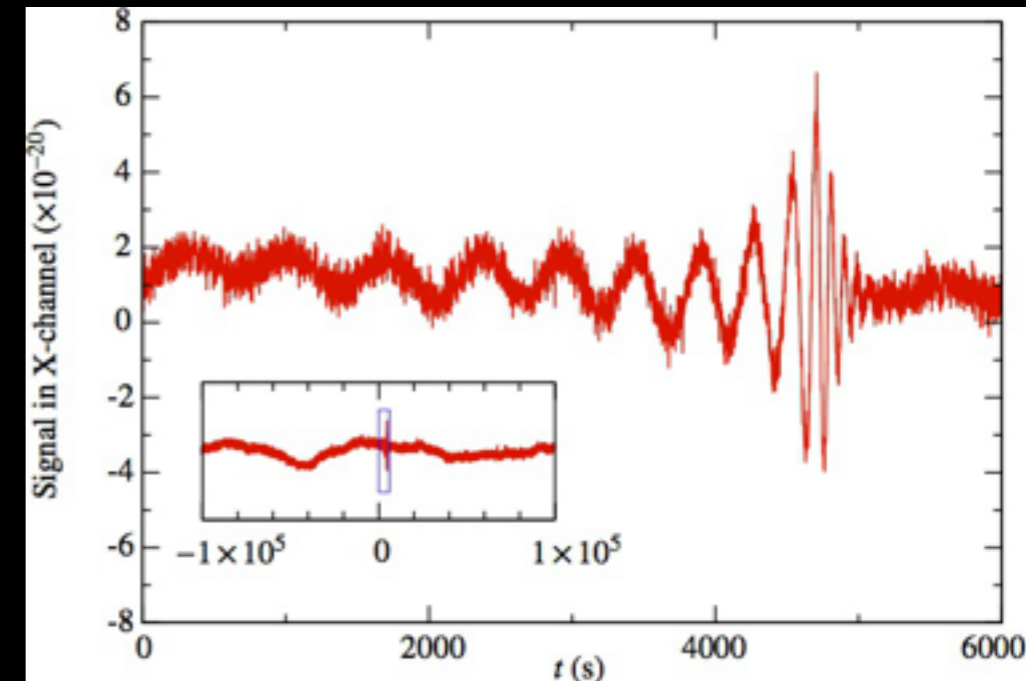
## ▶ Example of eLISA simulated data (LISACode):

- about 100 SMBHs,
- Galactic binaries

## ▶ First « simple data analysis » :

- low pass + zoom : we can « see » end of waveform for powerful sources

- wavelet transform :
  - chirps
  - Galactic binaries



▶ BUT for identifying more (all) sources and their parameters we need more advanced statistical techniques

# Data, signal and noise

## ▶ The time data $d(t)$ contains:

- $h(t)$  : signals that can be characterized by a sets of parameters
  - deterministic / stochastic
  - resolvable or not
- $n(t)$  : noises from
  - instrument itself
  - other sources

## ▶ Assumption 1: GW and noise are linearly independent:

$$d(t) = h_{real}(t) + n(t)$$

- $h(t)$  : GW perturbation  $h_{ab}(t, \vec{x})$  convolved with instrument response

# Likelihood

- ▶ **Goal:** find the  $h_{model} = h_{real}$
- ▶ **Likelihood:** found by demanding residual compatible with noise distribution  $p_n(x)$ :
  - The likelihood of observing  $d \equiv \{d_1, d_2, \dots, d_N\}$  where  $d_i = d(t_i)$ , is given by:

$$p(d(t)/h_{real}(t)) = p_n(r(t)) = p_n(d(t) - h_{real}(t))$$

- ▶ **So if  $p(d(t)/h_{model}(t))$  is compatible with the noise distribution:**  $h_{model}(t) = h_{real}(t)$



# Likelihood

- ▶ Usual case: noise is a multi-variate gaussian distribution:

$$p(d|h) = p_n(r) = \frac{1}{\sqrt{\det(2\pi C_n)}} e^{-\frac{1}{2} \sum_{i,j} r_i (C_n^{-1})_{ij} r_j}$$

where the correlation matrix is :  $C_n = \langle n_i n_j \rangle - \langle n_i \rangle \langle n_j \rangle$

- ▶ Generalization for a network of detectors:

$$p(d|h) = \frac{1}{\sqrt{\det(2\pi C_n)}} e^{-\frac{1}{2} \sum_{Ii, Jj} r_{Ii} (C_n^{-1})_{Ii, Jj} r_{Jj}}$$

where  $I, J$  labels the detector and  $i, j$  the discrete time or frequency sample

# Likelihood

► Inner product:

$$\langle x|y \rangle = \sum_{Ii, Jj} x_{Ii} (C_n^{-1})_{Ii, Jj} y_{Jj}$$

► Likelihood:

$$\mathcal{L} = p(d|h) = \frac{1}{\sqrt{\det(2\pi C_n)}} e^{-\frac{1}{2} \langle d-h|d-h \rangle}$$

► If  $C_n^{-1}$  is diagonal with  $1/\sigma_i^2$  the inner product is similar to

$$\chi^2 = \sum_i \left( \frac{d_i - h_i}{\sigma_i} \right)^2 \Rightarrow \mathcal{L} = C e^{-\frac{1}{2} \langle d-h|d-h \rangle} = C e^{-\frac{1}{2} \chi^2}$$

# Likelihood

## ► If stationary noise

→  $C_n$  only depend to  $|t_i - t_j|$

→  $C_n \sim$  diagonal in the Fourier domain (Discrete Fourier Transform) with on the diagonal  $S_{n,k} T/2$

→ Inner product:  $\langle \tilde{x} | \tilde{y} \rangle = 2 \sum_{j=0}^{N/2-1} \Delta f \frac{\tilde{x}_j^* \tilde{h}_j + \tilde{x}_j \tilde{h}_j^*}{S_{n,j}}$

→ Continuous limit:  $\langle \tilde{x} | \tilde{y} \rangle = 2 \int_0^\infty df \frac{\tilde{x}^*(f) \tilde{h}(f) + \tilde{x}(f) \tilde{h}^*(f)}{S_n(f)}$

$$= 4 \Re \int_0^\infty df \frac{\tilde{x}^*(f) \tilde{h}(f)}{S_n(f)}$$

→ Match-filtering



# Likelihood

- ▶ If noise  $C_n$  is known (i.e. known parameters) and stationary, the factor in front is neglected and we only consider the logarithm of likelihood:

$$\begin{aligned}\log \mathcal{L} &= -\frac{1}{2} \langle d - h | d - h \rangle \\ &= \langle d | h \rangle - \frac{1}{2} \langle h | h \rangle - \frac{1}{2} \langle d | d \rangle\end{aligned}$$

- ▶  $\langle d | d \rangle$  is fixed so the relevant term that is usually used is the reduced likelihood:

$$\log \mathcal{L}' = \langle d | h \rangle - \frac{1}{2} \langle h | h \rangle$$

# Maximum likelihood

- ▶ Considering the signal:  $h = A h_A$

$$\log \mathcal{L}' = A \langle d|h_A \rangle - \frac{A^2}{2} \langle h_A|h_A \rangle$$

- ▶ If the maximum likelihood corresponds to  $A_{ML}$ :

$$\left. \frac{\partial \log \mathcal{L}'}{\partial A} \right|_{A_{ML}} = 0 \quad \Rightarrow \quad A_{ML} = \frac{\langle d|h_A \rangle}{\langle h_A|h_A \rangle}$$

- ▶ then the maximum likelihood is :

$$\max(\log \mathcal{L}') = \frac{\langle d|h_A \rangle^2}{2 \langle h_A|h_A \rangle} = \frac{\langle d|h \rangle^2}{2 \langle h|h \rangle}$$

# Signal to Noise Ratio

- ▶ We can define the SNR using the power ratio

$$SNR^2 = \frac{P_{signal}}{P_{noise}}$$

- ▶ Average noise power:  $P_{noise} = \int_0^{\infty} df S_n(f)$

- ▶ Average signal power:  $P_{signal} = 2 \int_0^{\infty} df |h(f)|^2$

- ▶ Optimal SNR:

$$SNR_{opt}^2 = 2 \int_0^{\infty} df \frac{|h(f)|^2}{S_n(f)} = \langle h|h \rangle$$

# SNR vs likelihood

▶ “ Usual ” SNR:  $SNR = \frac{\langle d|h \rangle}{\sqrt{\langle h|h \rangle}}$

▶ Maximum likelihood:  $max(\log \mathcal{L}') = \frac{\langle d|h \rangle^2}{2 \langle h|h \rangle}$

▶ The relation between SNR and maximum likelihood is simply:

$$SNR = \sqrt{2 \max(\log \mathcal{L}')}$$

# Statistical inference

- ▶ Our core tool: the likelihood:  $\mathcal{L} = p(d/h)$
- ▶ **Likelihood** measures the **probability** of having **data  $d$**  given **the hypothesis/model  $h$** .
- ▶ How to use it to infer about :
  - detectability of a signal in data ?
  - value of parameters of a signal ?
- ▶ Depending where the uncertainty is put, 2 approaches:
  - the frequentist inference
  - the bayesian inference

# Frequentist inference

## ▶ General ideas:

- Uncertainties in the data
- Parameters of the system we want to observe are fixed
- “long-run relative occurrence of an event in a set of identical experiments”
- Probability related to the frequencies of events

## ▶ Probability of observing the data $d$ given the hypothesis $H$

## ▶ Measured data drawn from an underlying probability distribution $p(d/H)$ .

# Frequentist inference

- ▶ **Statistic: function of the data**
  - Likelihood or something else
- ▶ **Required:**
  - knowledge of the probability distribution of the statistic (analytic or simulation).
- ▶ **Problem:**
  - Distribution is constructed on non observed data

# Frequentist: hypothesis testing

- ▶ Hypothesis:
  - $H_0$  : no signal
  - $H_1$  : signal: composite hypothesis  $\Rightarrow$  several parameter's values
- ▶ Argue for  $H_1$  by arguing against  $H_0$
- ▶ Construct a statistic  $\Lambda$ , called a test or detection statistic
- ▶ Calculate  $p(\Lambda/H_0)$ : the sampling distribution of  $\Lambda$  under assumption of null hypothesis
- ▶ If data distribution different  $\Rightarrow$  reject  $H_0$  and accept  $H_1$  at  $p \times 100\%$  level

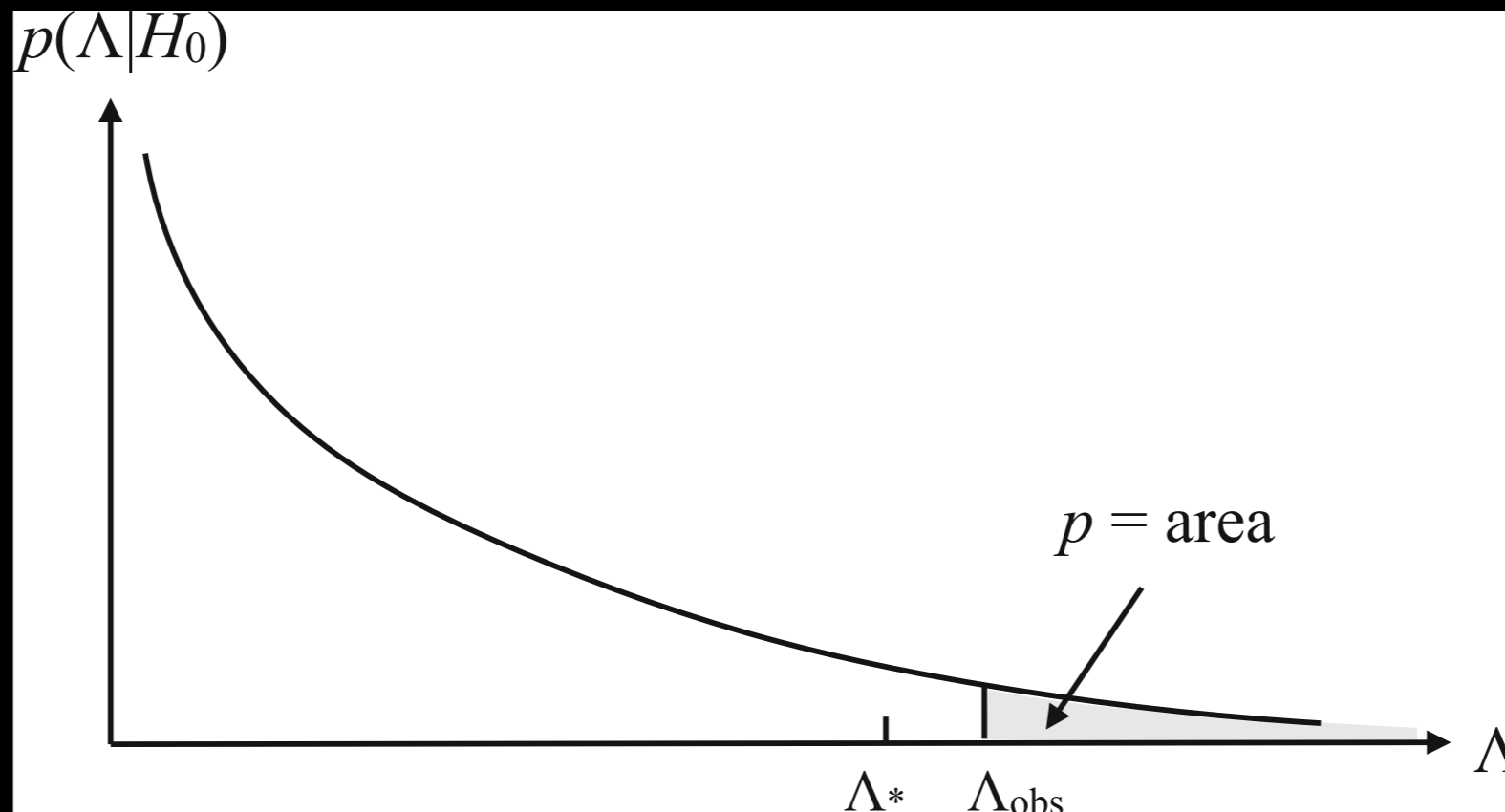


# Frequentist: hypothesis testing

- ▶ If data distribution  $\neq p(\Lambda|H_0) \Rightarrow$  reject  $H_0$  and accept  $H_1$  at  $p \times 100\%$  level

$$p \equiv \text{prob}(\Lambda > \Lambda_{obs} | H_0) \equiv \int_{\Lambda_{obs}}^{\infty} d\Lambda p(\Lambda | H_0)$$

- ▶ p-value required to reject the null hypothesis determines a threshold  $\Lambda^*$ .



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- ▶ p-value required to reject the null hypothesis determines a threshold  $\Lambda^*$ .

- ▶ Errors: 2 types:

- FAP: false alarm error:  $\Lambda_{obs} > \Lambda^*$  : reject  $H_0$  but no signal:

$$FAP \equiv \alpha = \text{prob}(\Lambda > \Lambda^* | H_0)$$

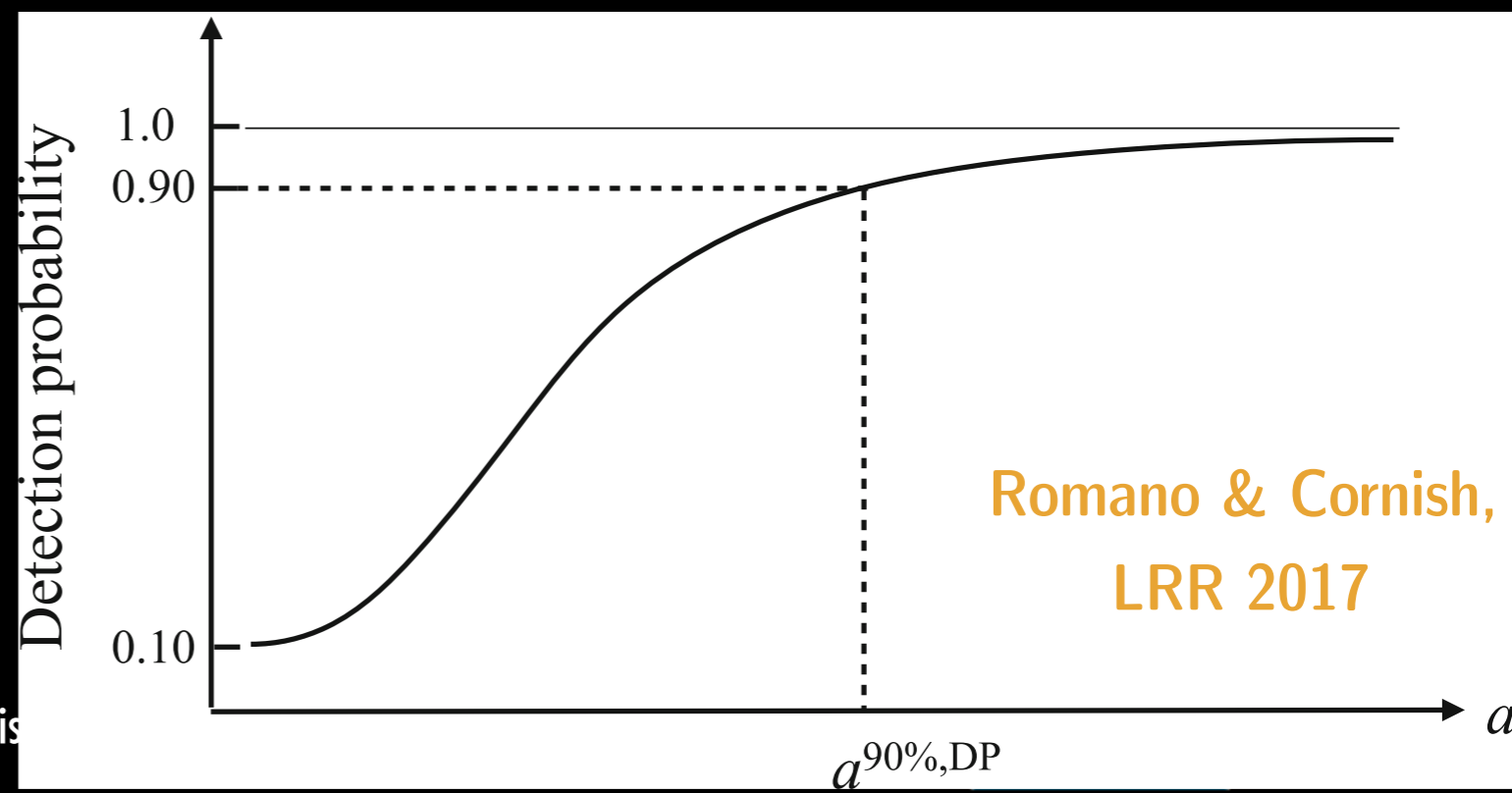
- FDP: false dismissal error:  $\Lambda_{obs} < \Lambda^*$  : accept  $H_0$  but signal:

$$FDP \equiv \beta(a) = \text{prob}(\Lambda < \Lambda^* | H_a)$$

# Frequentist: hypothesis testing

- ▶ Ideally: FAP and FDP as small as possible ... but compete
  - GW: FAP very small
  - Medical: FDP very small
- ▶ **Newman-Pearson** criterion: for fixed FAP, the best statistic is the one minimizing FDP:  $FAP \Rightarrow \Lambda^*$ .
- ▶ **Detection probability**:  $1 - \beta(a) = 1 - \text{prob}(\Lambda < \Lambda^* / H_0)$  :

- Independent from data
- Depends only on
  - sampling distribution
  - FAP



# Frequentist: upper limit

▶ No detection  $\Rightarrow$  set an upper limit (amplitude) based on:

- $\Lambda_{\text{obs}}$ : observed detection statistic
- Confidence level

▶ Example:

- 90% confidence-level upper-limit  $a^{90\%, UL}$

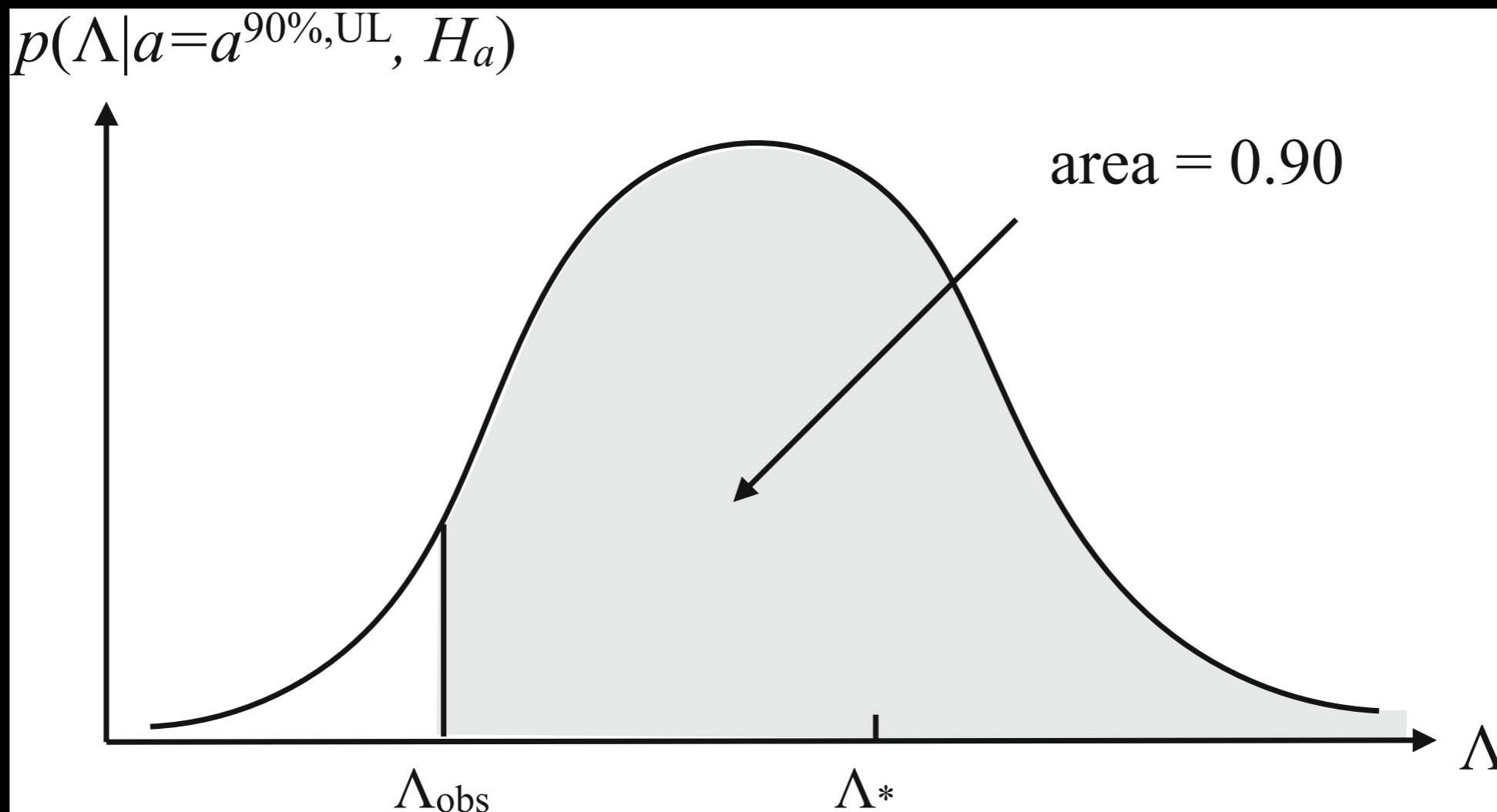
= minimal value of  $a$  for which  $\Lambda > \Lambda_{\text{obs}}$  at least 90% of the time

$$\text{prob}(\Lambda > \Lambda_{\text{obs}} | a \geq a^{90\%, UL}, H_a) \geq 0.9$$

# Frequentist: upper limit

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= minimal  $a$  for which  $\Lambda > \Lambda_{obs}$  at least 90% of the time

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# Frequentist: upper limit

- ▶ In practice, you can use injection in your data
- ▶ Example: GW signal from binary described by amplitude  $a$  + other parameters

## 1. Test amplitude $a$

1.1. Produce fake data by injecting a signal in your data signal with a given  $a$  randomizing other parameters

1.2. Calculate  $\Lambda$  for the fake data

1.3. Repeat 1.1 large number of time

2. Compute the ratio  $N(\Lambda > \Lambda_{obs}) / N_{total}$

3. If adjust  $a$  and restart from 1. until you get

$$N(\Lambda > \Lambda_{obs}) / N_{total} = \text{confidence level}$$

# Frequentist: parameter estimation

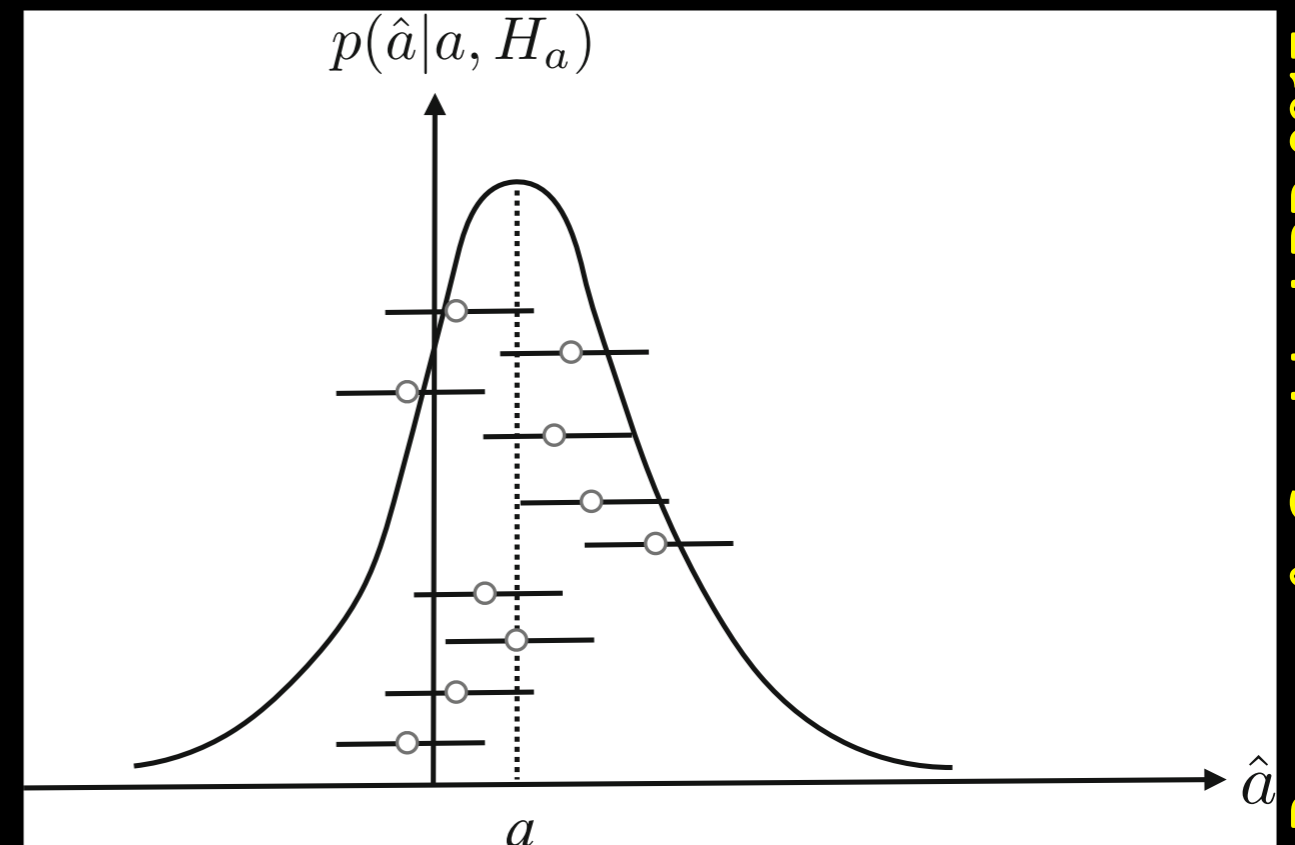
- ▶ Construct the estimator  $\hat{a}$  of the parameter  $a$  : it's a statistic that can be maximum likelihood or others estimators.
- ▶ Calculate the sampling distribution  $p(\hat{a}/a, H_a)$
- ▶ Using  $p(\hat{a}/a, H_a)$  + a confidence level of 95%, construct the frequentist confidence interval  $[\hat{a}-\Delta, \hat{a}+\Delta]$ , such as
$$\text{prob}(\hat{a} - \Delta < a < \hat{a} + \Delta) = 0.95$$
- ▶ Interpretation:
  - in a set of many repeated experiments, in 95% of the case the true value of  $a$  is in the intervals
  - $a$  not a random variable so its not a probability on  $a$ .

# Frequentist: parameter estimation

- ▶ Using  $p(\hat{a}/a, H_a)$  + a frequentist confidence interval of 95%, construct interval  $[\hat{a}-\Delta, \hat{a}+\Delta]$ , such as

$$\text{prob}(\hat{a} - \Delta < a < \hat{a} + \Delta) = 0.95$$

- ▶ Not physical value are allowed.



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# Frequentist: summary

▶ **Uncertainty in the data**  $\Rightarrow$  define a **statistic  $\Lambda$**  for the data and define its sampling distribution

▶ p-value required to reject  $H_0 \Rightarrow \Lambda^*$  :

$$FAP \equiv \alpha = \text{prob}(\Lambda > \Lambda^* | H_0)$$

▶ **Detection: compare  $\Lambda_{\text{obs}}$  and  $\Lambda^*$**

- $\Lambda_{\text{obs}} > \Lambda^* \Rightarrow H_0$  rejected  $\Rightarrow$  detection

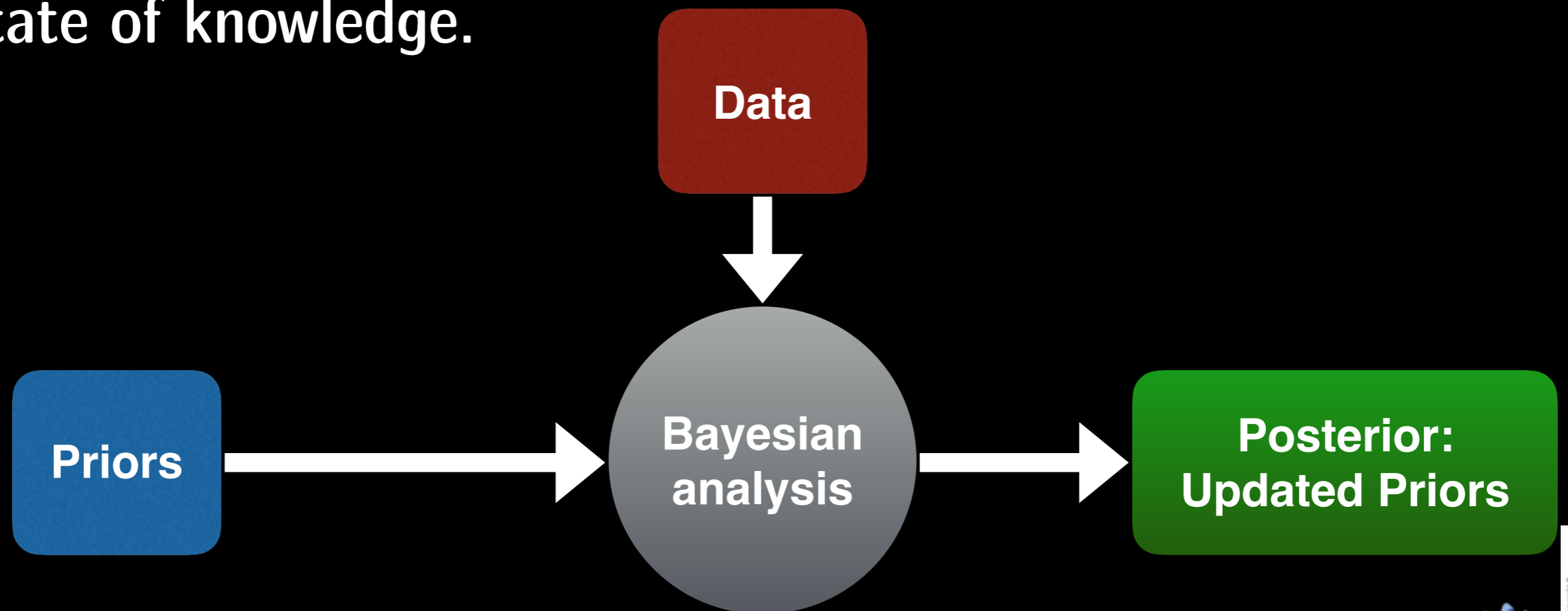
- **Parameter estimation**: estimator  $\hat{a} \rightarrow$  distribution  $p(\hat{a}/a, H_a) +$  confidence level  $\rightarrow$  frequentist confidence interval

- $\Lambda_{\text{obs}} < \Lambda^* \Rightarrow H_0$  accepted  $\Rightarrow$  no detection

- **Upper limit**: minimal value of  $a$  for which  $\Lambda > \Lambda_{\text{obs}}$  at least CL% of the time, with CL the confidence interval

# Bayesian inference

- ▶ Data are given
- ▶ The uncertainties are on the model / parameters
- ▶ Our prior knowledge is updated by what we learn from the data, as measured by the likelihood to give our posterior state of knowledge.



# Bayesian inference

- ▶ Bayes theorem:

$$p(a|d) = \frac{p(a) p(d|a)}{p(d)}$$

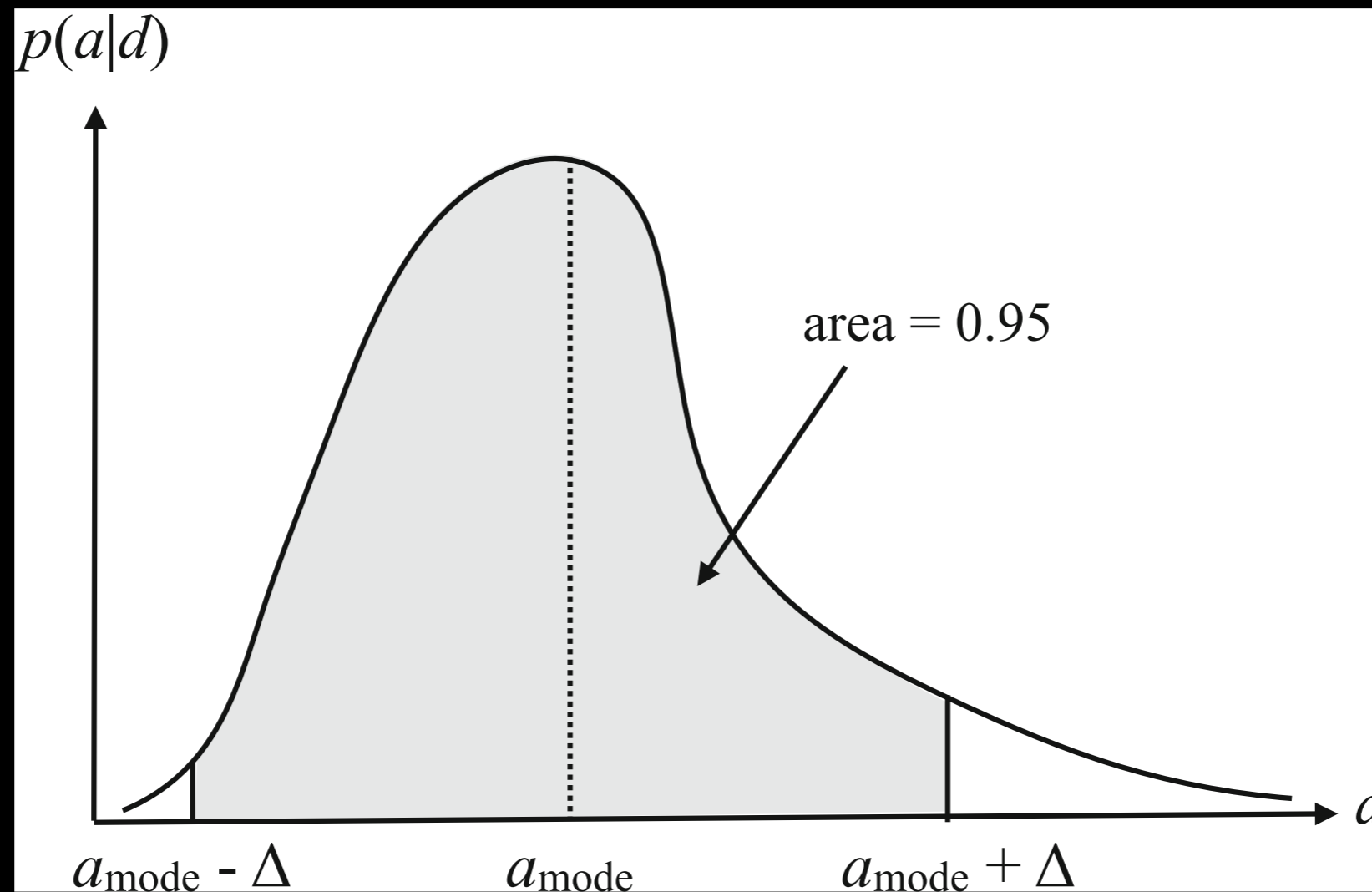
Diagram illustrating Bayes' theorem with labels and arrows:

- prior** (arrow pointing to  $p(a)$ )
- likelihood** (arrow pointing to  $p(d|a)$ )
- evidence** (arrow pointing to  $p(d)$ )
- posterior distribution** (arrow pointing to  $p(a|d)$ )

- ▶ “Everything” about the parameters is in the posterior distribution

# Bayesian inference

- ▶ Confidence interval = credible interval (degree of belief): area under the posterior between one parameter value and another



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# Bayesian: marginalization

▶ More than one parameter in your model but there are parameters we don't care about => marginalized over them

▶ Example with 2 parameters a,b:

- care only about a => marginalized over b

- $p(a|d) = \int db p(a, b|d)$

- relation between joint probabilities and conditional probabilities

$$p(a, b) = p(a|b) p(b)$$

- => the marginalization over b is simply

$$p(a|d) = \int db p(a|b, d) p(b)$$

# Bayesian: information

- ▶ Information gain:

$$I = \int da p(a|d) \log \left( \frac{p(a|d)}{p(a)} \right)$$

- ▶ If there no gain of information from the data, likelihood

$\mathcal{L}(a)$  is constant

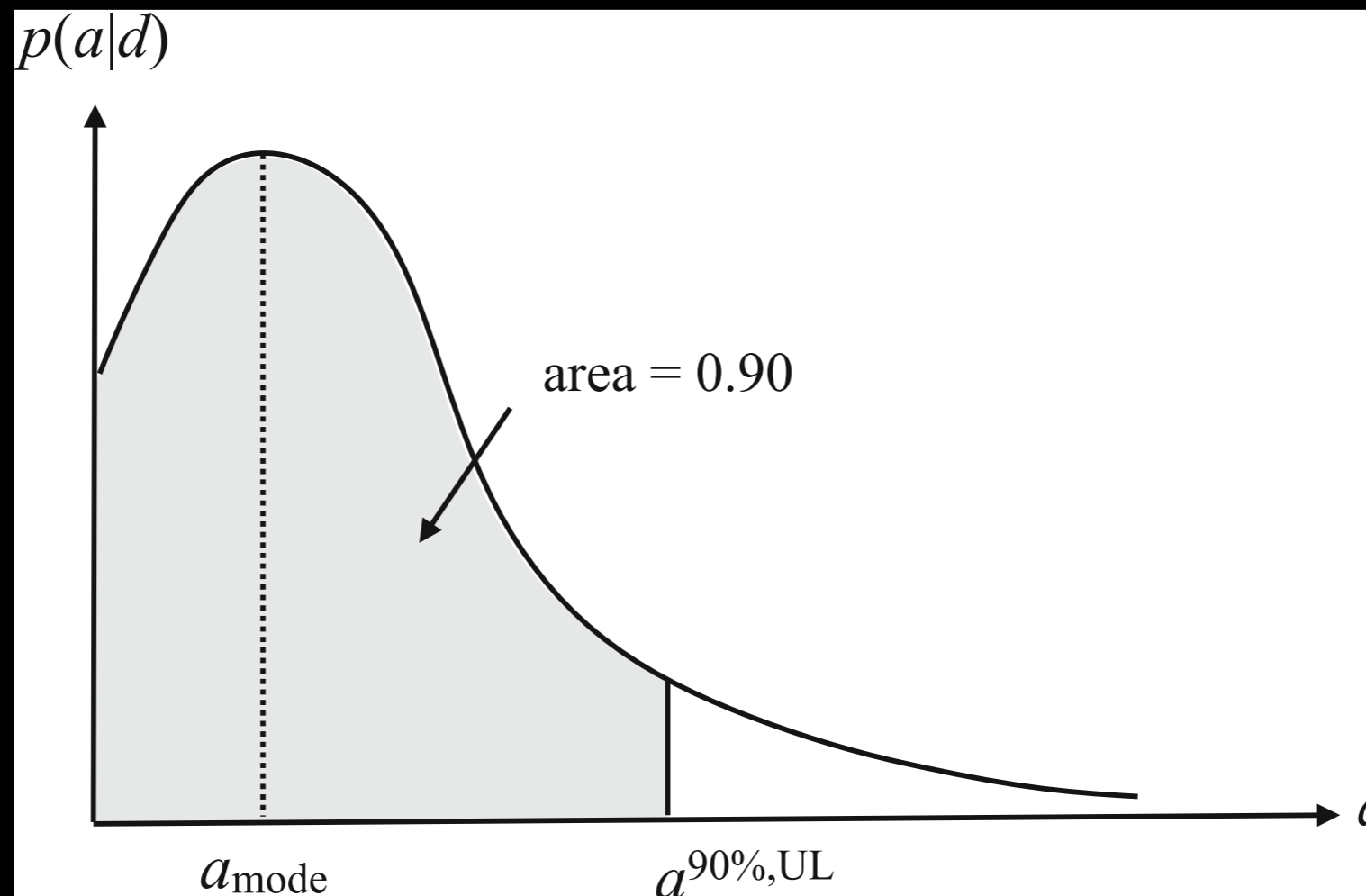
$$\Rightarrow p(a/d) = p(a)$$

$$\Rightarrow I = 0$$

# Bayesian: upper limit

- ▶ If the Bayesian credible interval is compatible with the minimum value for the parameter, we can set an upper limit for a “confidence level”:

$$prob(0 < a < a^{UL,90\%} | d) = 0.9$$



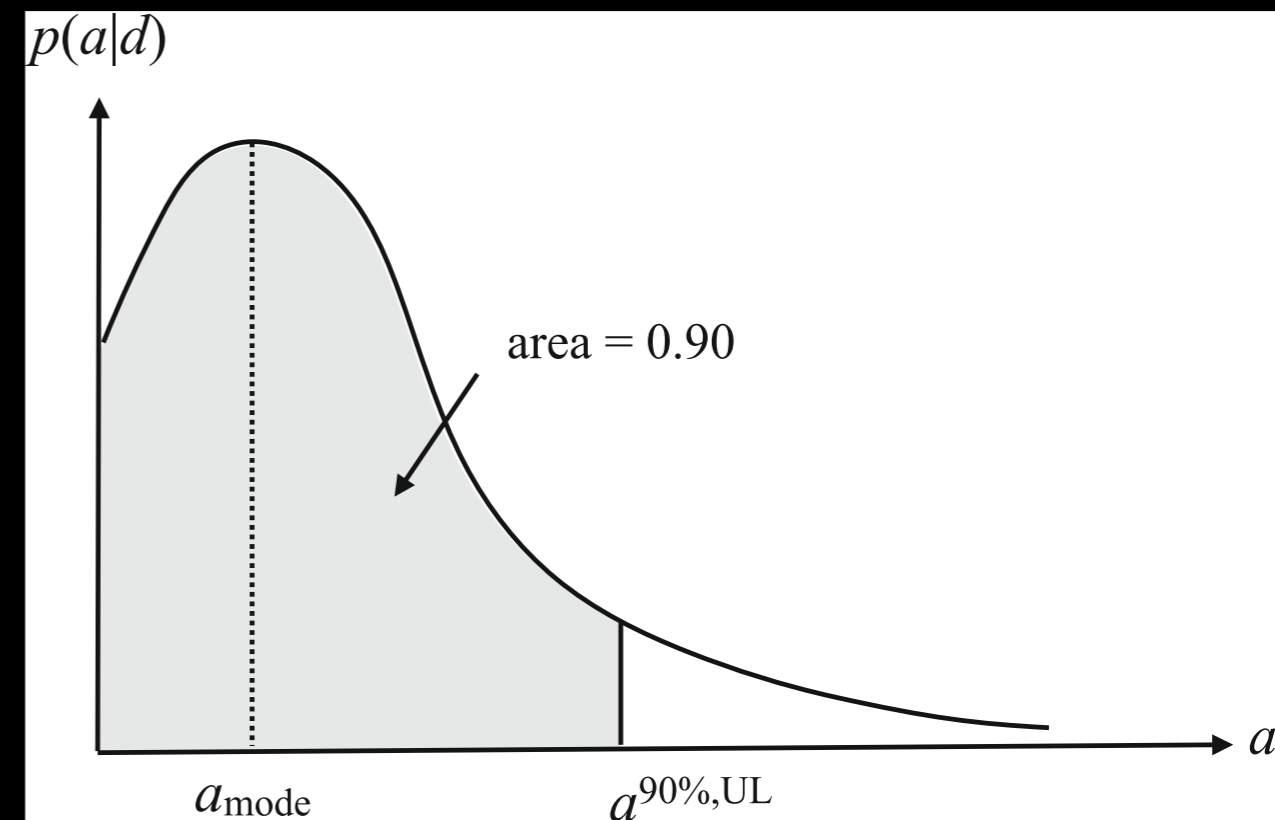
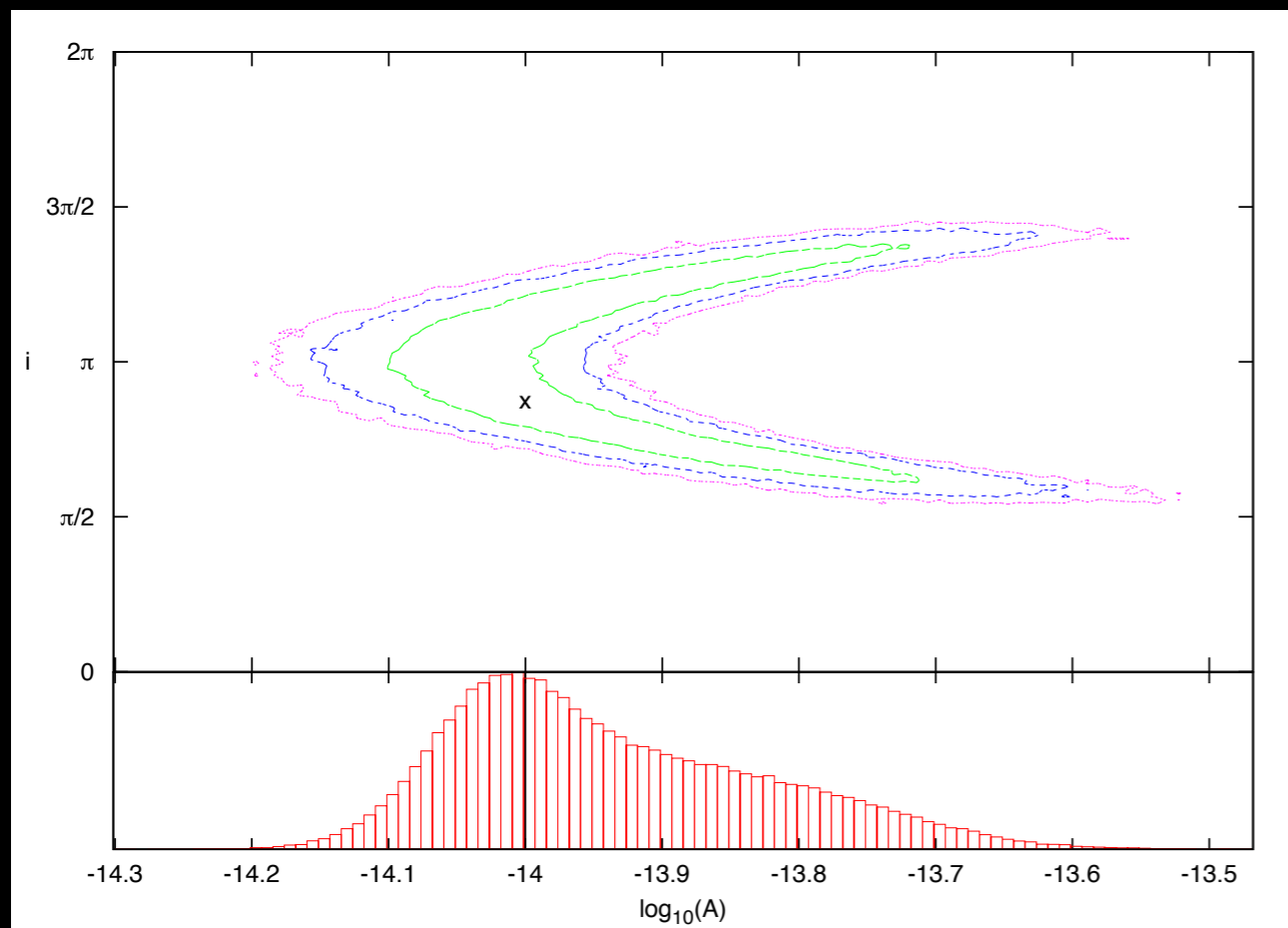
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# Bayesian: upper limit

## ► If the parameter is the amplitude

- Confidence interval exclude 0  $\Rightarrow$  potential detection ...
- Confidence interval include 0  $\Rightarrow$  result compatible with no detection  $\Rightarrow$  upper limit

$$prob(0 < a < a^{UL,90\%} | d) = 0.9$$



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# Bayesian: model selection

Goal: use Bayes theorem to compare models

▶  $\mathcal{M}_\alpha$  : models                       $\theta_\alpha$  : parameters

▶ Posterior distribution for given the model :

$$p(\theta_\alpha | d, \mathcal{M}_\alpha) = \frac{p(d | \theta_\alpha, \mathcal{M}_\alpha) p(\theta_\alpha | \mathcal{M}_\alpha)}{p(d | \mathcal{M}_\alpha)}$$

▶ Evidence given a model:

$$p(d | \mathcal{M}_\alpha) = \int d\theta_\alpha p(d | \theta_\alpha, \mathcal{M}_\alpha) p(\theta_\alpha | \mathcal{M}_\alpha)$$

# Bayesian: model selection

- ▶ Posterior probability of models  $\mathcal{M}_\alpha$  :

$$p(\mathcal{M}_\alpha|d) = \frac{p(d|\mathcal{M}_\alpha) p(\mathcal{M}_\alpha)}{p(d)}$$

- ▶ Evidence: sum of all possible model ... but total number unknown  $\Rightarrow$  use a subset

$$p(d) = \sum_{\alpha} p(d|\mathcal{M}_\alpha) p(\mathcal{M}_\alpha)$$

- ▶ Odds ratio between 2 models:

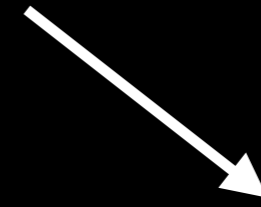
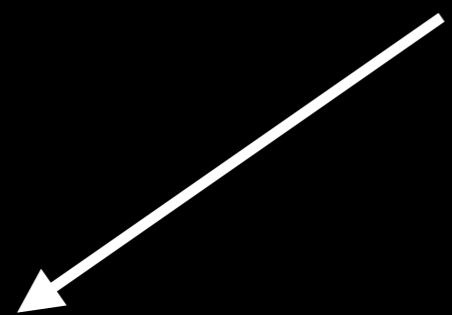
$$\mathcal{O}_{\alpha\beta} = \frac{p(\mathcal{M}_\alpha|d)}{p(\mathcal{M}_\beta|d)}$$

# Bayesian: model selection

- ▶ Odds ratio between 2 models:

$$\mathcal{O}_{\alpha\beta} = \frac{p(\mathcal{M}_\alpha|d)}{p(\mathcal{M}_\beta|d)}$$

$$\mathcal{O}_{\alpha\beta} = \frac{p(d|\mathcal{M}_\alpha)}{p(d|\mathcal{M}_\beta)} \frac{p(\mathcal{M}_\alpha)}{p(\mathcal{M}_\beta)}$$



prior odds ratio

Bayes factor evidence ratio

$$\mathcal{B}_{\alpha\beta} = \frac{p(d|\mathcal{M}_\alpha)}{p(d|\mathcal{M}_\beta)}$$

# Bayesian: Bayes factor

- ▶ Bayes factor: usual tool to compare model, in particular signal versus no signal  $\mathcal{M}$
- ▶ Problem: interpretation of Bayes factor

$\mathcal{B}_{\alpha\beta}$	$2 \ln \mathcal{B}_{\alpha\beta}(d)$	Evidence for model $\mathcal{M}_\alpha$ relative to $\mathcal{M}_\beta$
$< 1$	$< 0$	Negative (supports model $\mathcal{M}_\beta$ )
1 - 3	0 - 2	Not worth more than a bare mention
3 - 20	2 - 6	Positive
2 - 150	6 - 10	Strong
$>150$	$>10$	Very strong

- ▶ Need proper calibration (simulations, ...)

# Bayesian in practice

- ▶ In practice, we need to sample the parameters space computing likelihood to construct the posterior distribution of parameters.
- ▶ Several methods:
  - Monte-Carlo Markov Chain,
  - Metropolis Hasting Markov Chain,
  - Multi-Nest,
  - EMCEE,
  - ...

# Bayesian vs frequentist

	Frequentist	Bayesian
Probabilities	Probabilities assigned only to propositions about <b>outcomes of repeatable experiments</b> , not to hypotheses or <b>parameters</b> which have <b>fixed but unknown values</b>	<b>Probabilities</b> can be assigned to <b>hypotheses</b> and <b>parameters</b> since probability is degree of belief in any proposition
Data	Assumes <b>measured data</b> are drawn from an underlying <b>probability distribution</b> , which assumes the truth of a particular hypothesis or model ( <b>likelihood</b> function)	Same
Input	Constructs a <b>statistic</b> to estimate a parameter <b>or</b> to decide whether or not to claim a detection	Needs to specify <b>prior</b> degree of belief in a particular hypothesis or parameter
Methods	Calculates the <b>probability distribution</b> of the statistic (sampling distribution)	Uses <b>Bayes' theorem</b> to update the prior degree of belief in light of new data
Results	Constructs <b>confidence intervals</b> and <b>p-values</b> for parameter estimation and hypothesis testing	Constructs <b>posteriors</b> and odds ratios for parameter estimation and hypothesis testing/model comparison

# Bayesian vs frequentist: GW obs

- ▶ In the past, almost only frequentists
- ▶ Now, Bayesian methods become more and more popular
- ▶ For all GW observatories, we used the two approaches and hybrid approaches mixing the two.
  - LIGO:
    - methods based on Freq. or Bayesian
  - LISA:
    - mainly Bayesian methods
  - PTA:
    - methods based on Freq. or Bayesian

# Bayesian vs frequentist

- ▶ But at the end, what we need is computing a large number of likelihood, or equivalent estimators
  - Main computing cost
- ▶ Joke about Bayesian inference from a colleague:
  - “That’s the beauty of Bayesian inference:
    - likelihood\*prior
    - realize that you have no idea how to pick the prior
    - assume flat prior
    - realize is a likelihood computation
  - Now you just computed a likelihood, but you are cool because you did it in a Bayesian way.”



# Likelihood / noise knowledge

- ▶ Depending on the level of knowledge of the noise, different flavors of likelihood. Some examples:
  - Perfectly known ( $S_n$ )  $\Rightarrow$  reduced likelihood
  - Known shape components
    - $\Rightarrow C_n$  described using parameters included in the search with model parameters
  - Partially known noise levels, and taking into account heavier tail distribution effects
    - $\Rightarrow$  Student-t [Rover 2011]: each frequency bin follows a multi-variate distribution with  $\nu_j$  degrees of freedom.

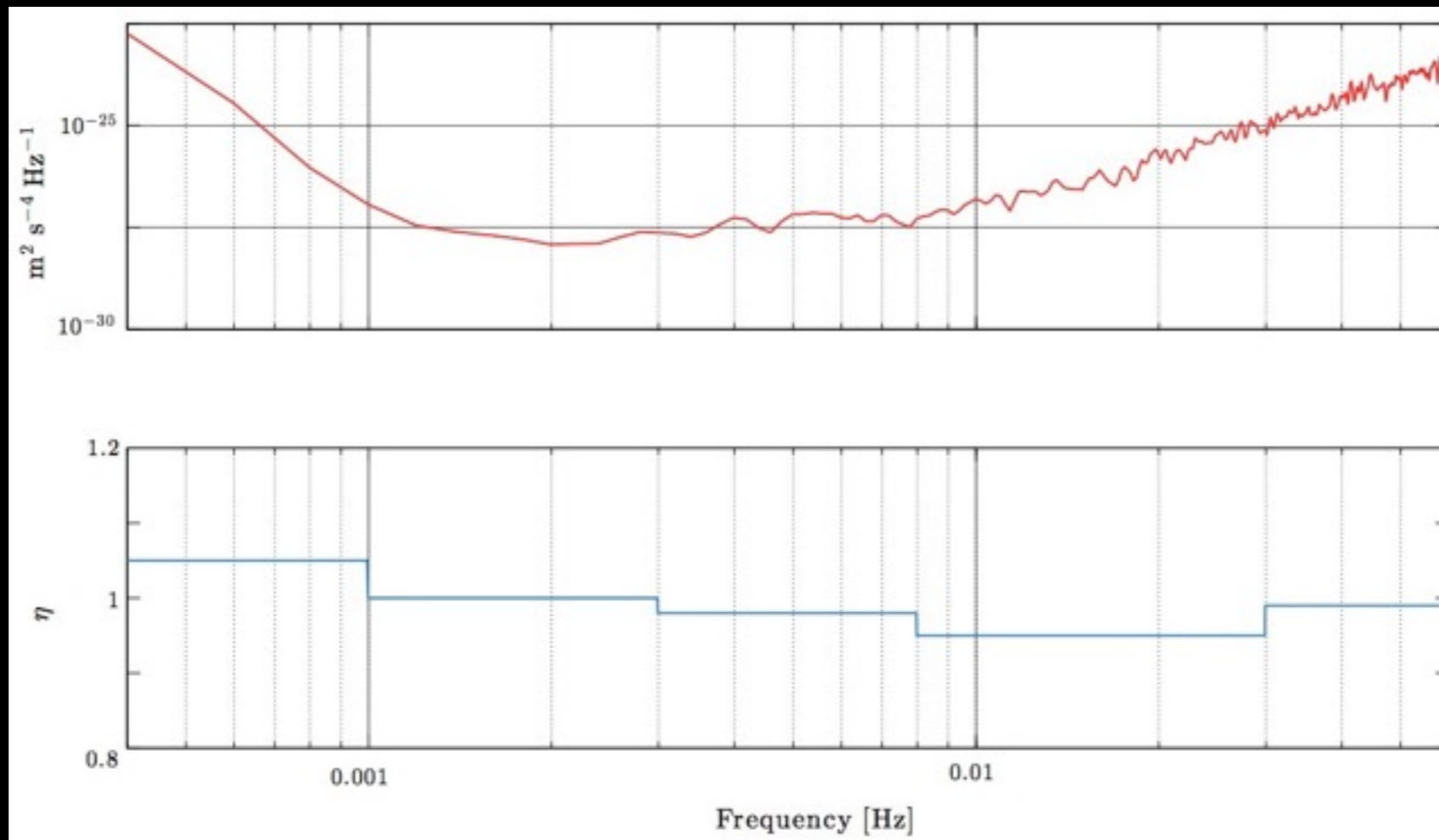
$$\log \mathcal{L} = - \sum_j \frac{\nu_j + 2}{2} \log \left( 1 + \frac{1}{\nu_j} \chi^2 \right)$$

# Likelihood / noise knowledge

- Partially known noise levels but fluctuations of  $S_n$  by segment  
=> one parameter per segment

$$S_{n,i} \rightarrow \eta_j S_{n,i}, \quad i_j < i \leq i_{j+1}$$

$$\log \mathcal{L} = -\frac{1}{2} \left( \chi^2 + N_{j,bins} \sum_j \log \eta_j \right)$$



# Thank you