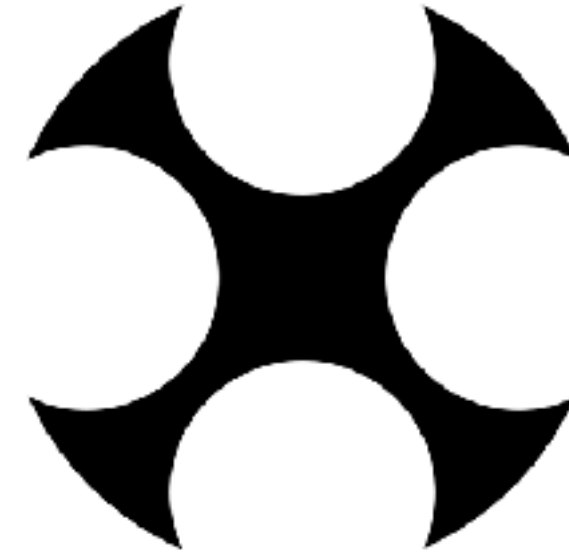


Tests of Gravity with LISA and 3rd Generation Detectors

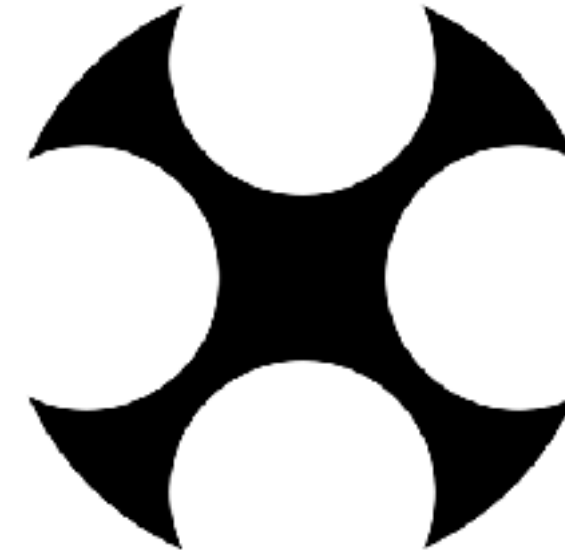
Nicolas Yunes
eXtreme Gravity Institute
Montana State University

Numerical Relativity Beyond GR,
Benasque, Spain
June 8th, 2018





eXtreme Gravity Institute



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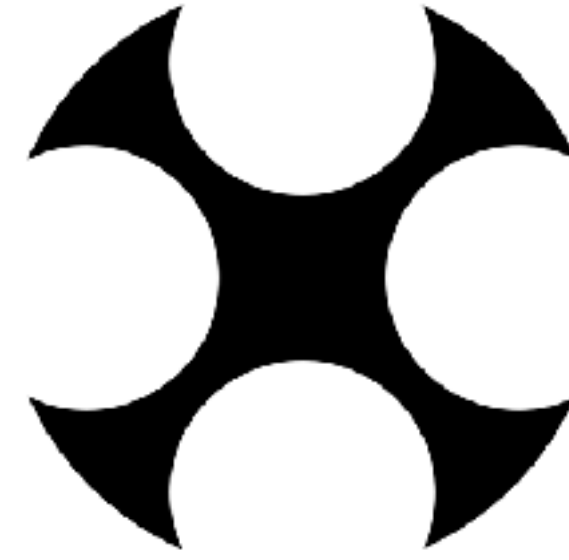
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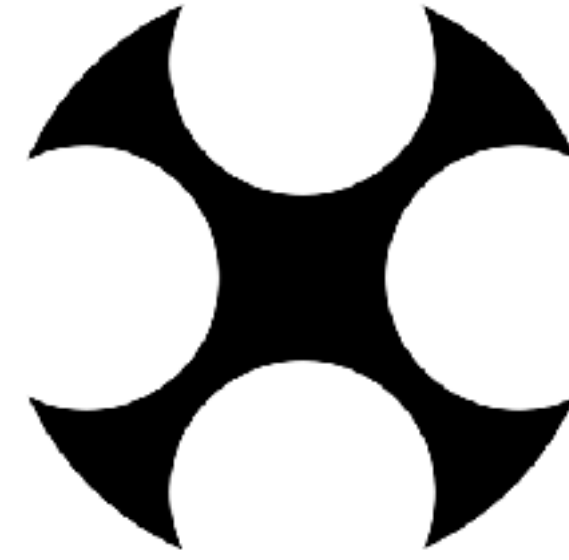
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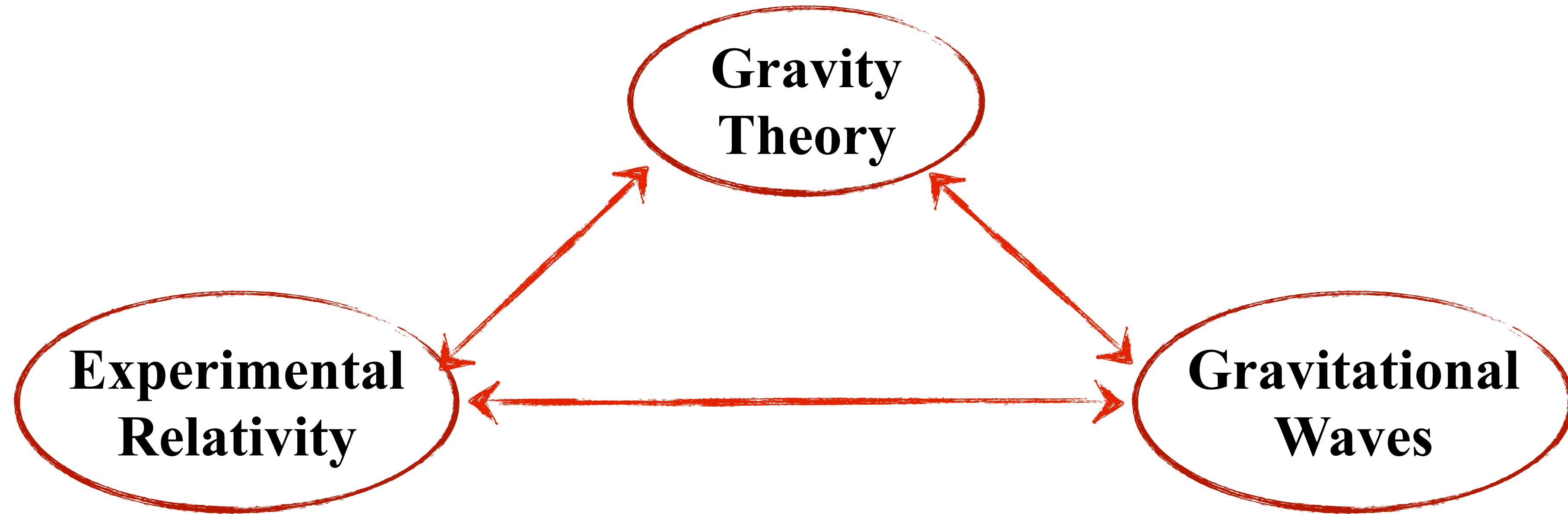
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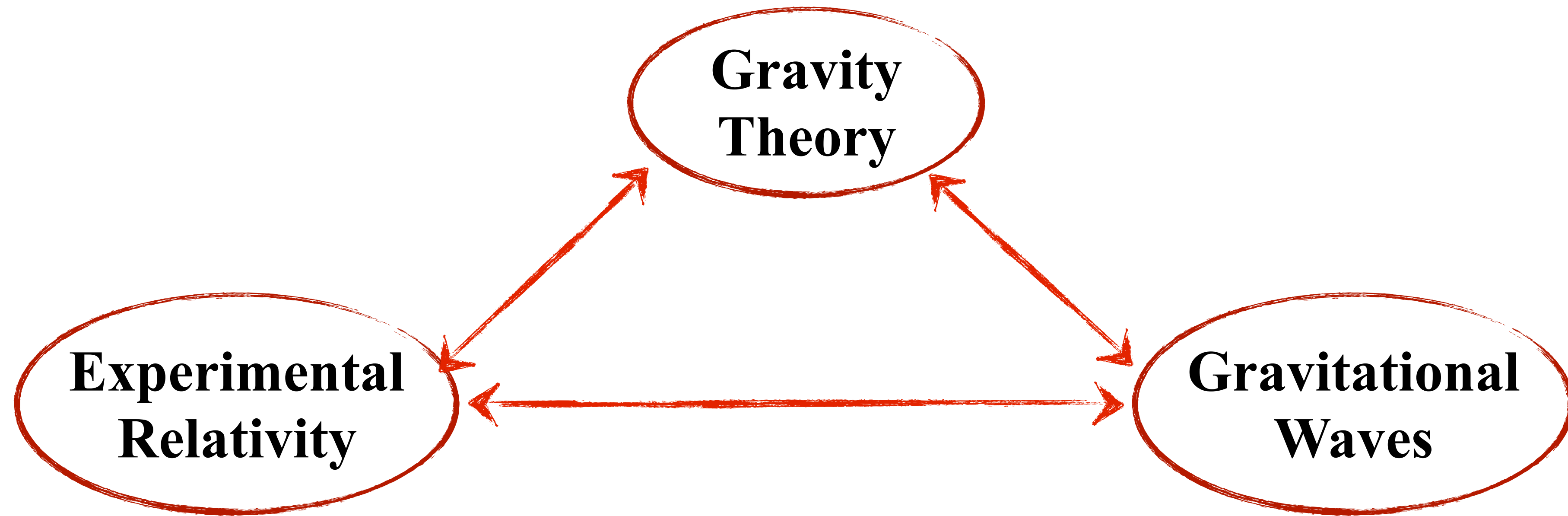
**Co-Chair of NASA PCOS GWSIG &
Fundamental Physics in LISA Consortium**



Who are you again?



Who are you again?



**What can we learn about theoretical physics
from future GW observations?**

A Fantastical Timeline

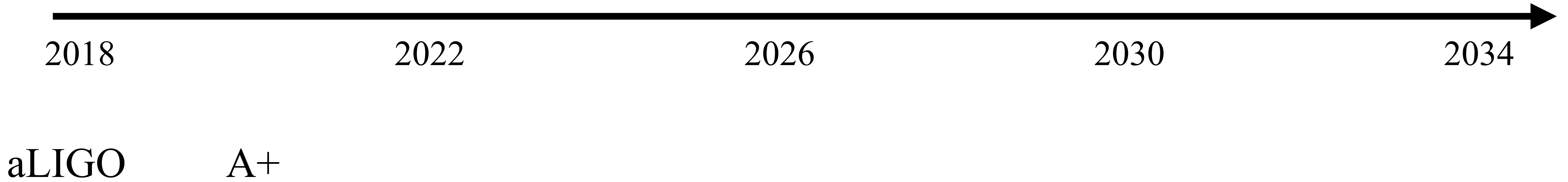


A Fantastical Timeline

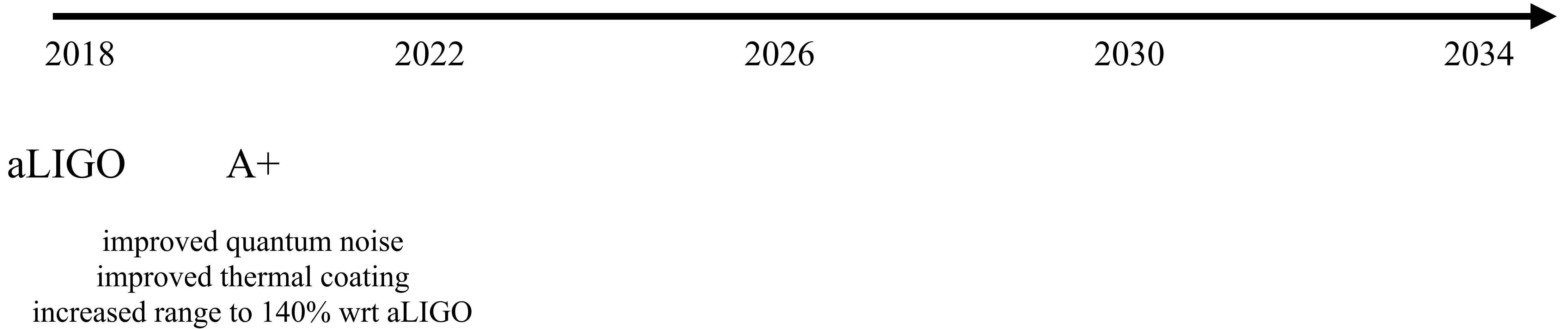


aLIGO

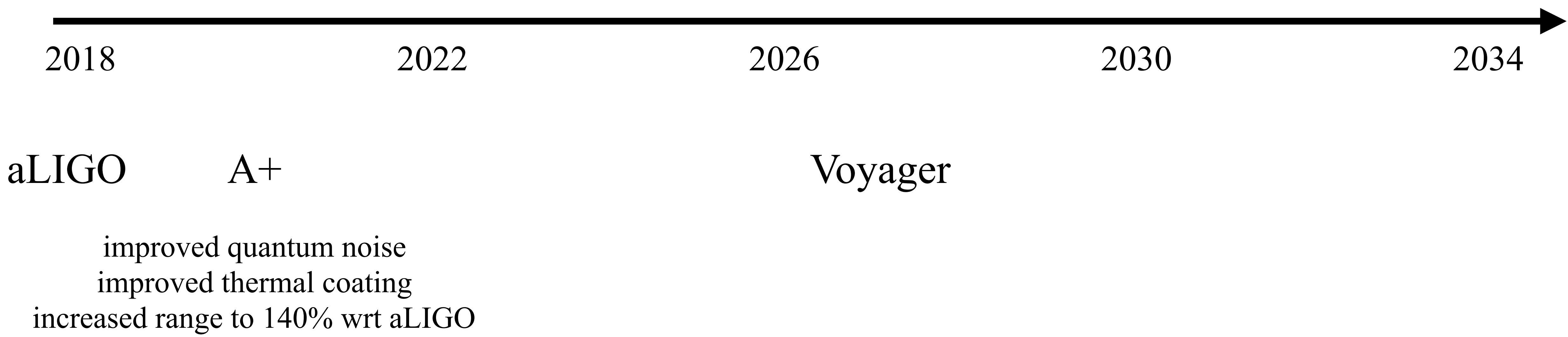
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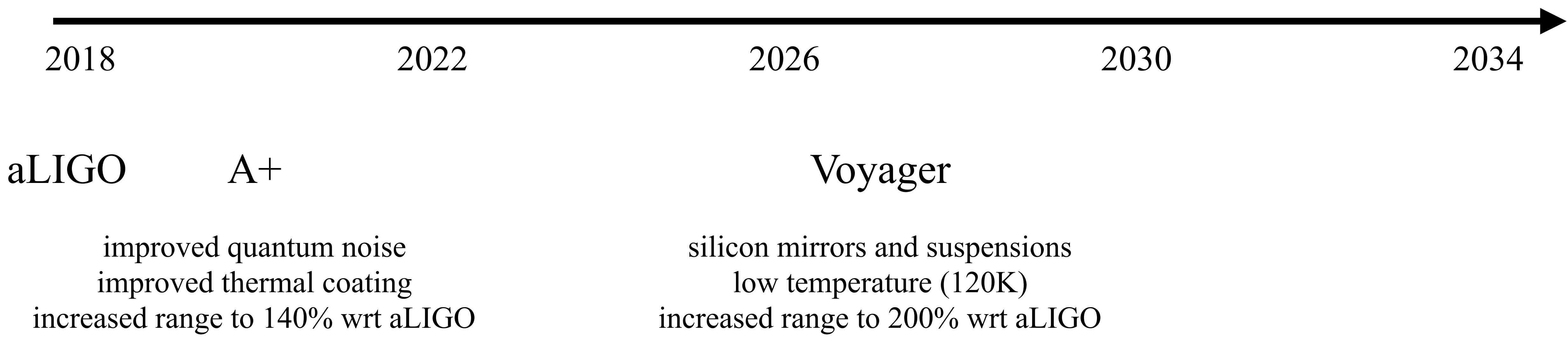
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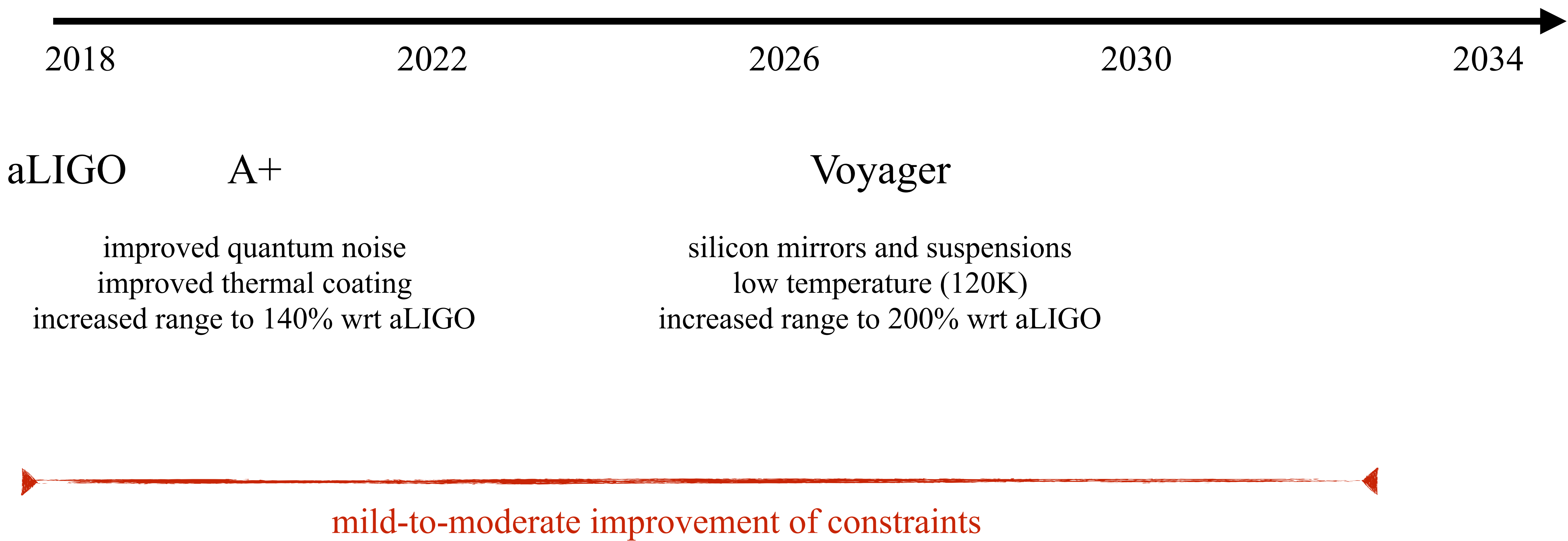
A Fantastical Timeline



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A Fantastical Timeline



Example: Ground-based constraints on Graviton

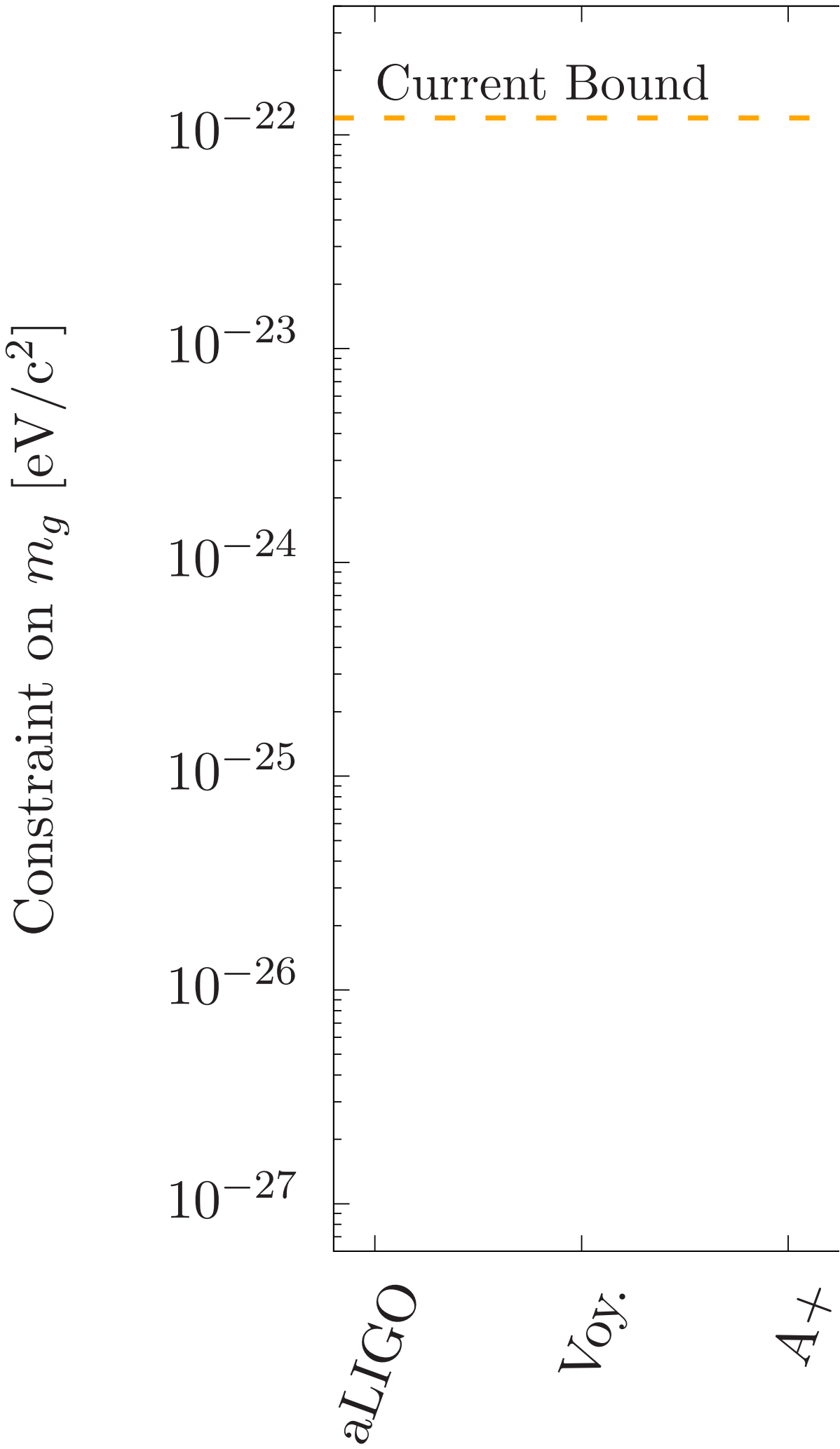
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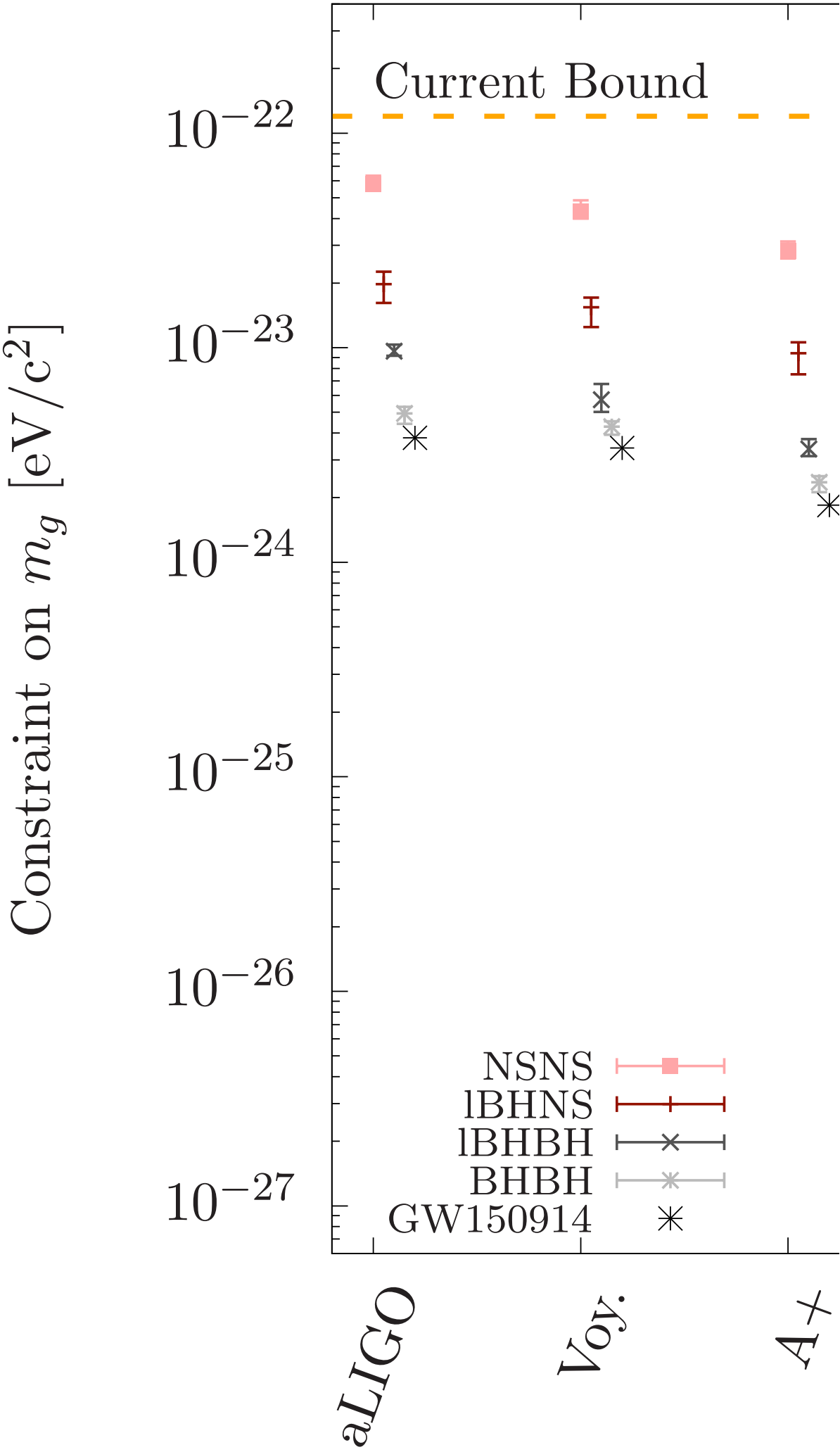
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[Chamberlain & Yunes, PRD '17]

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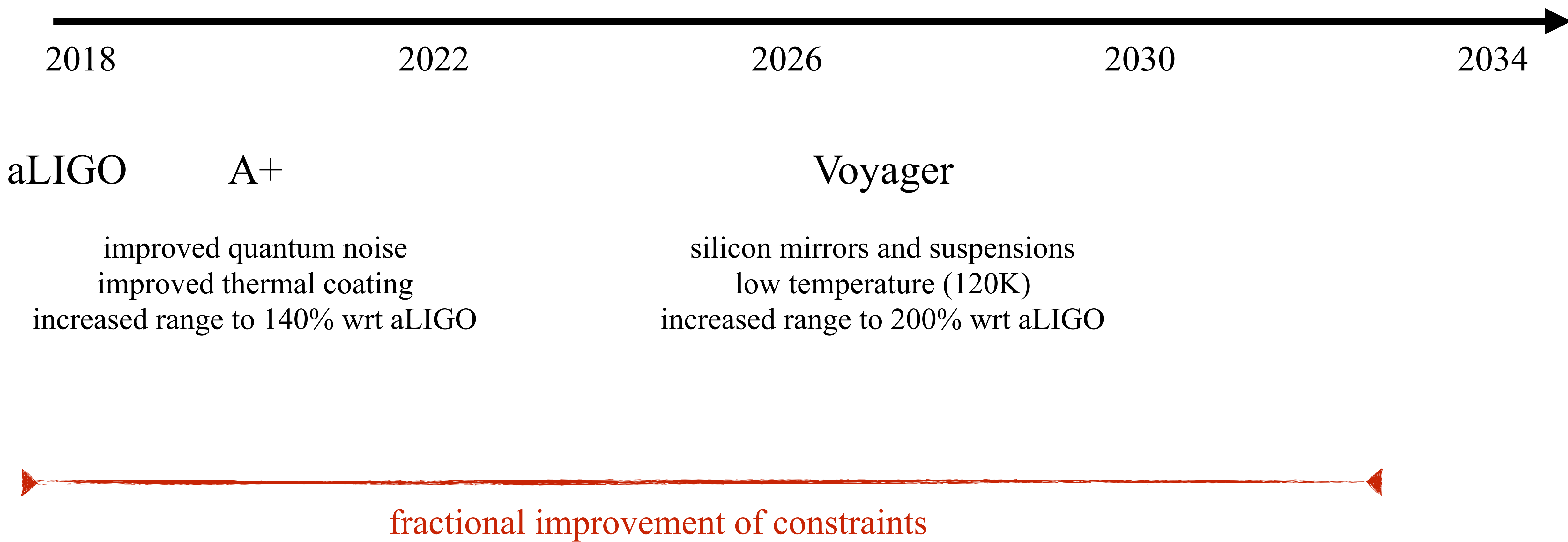
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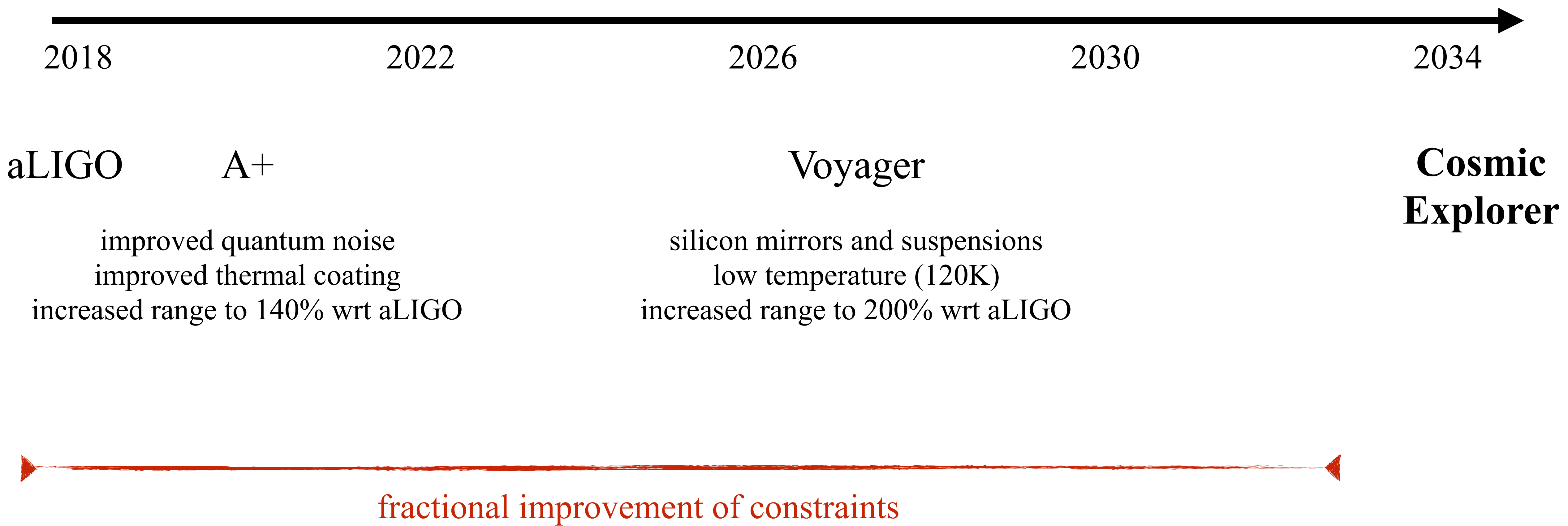
Instrument

[Chamberlain & Yunes, PRD '17]

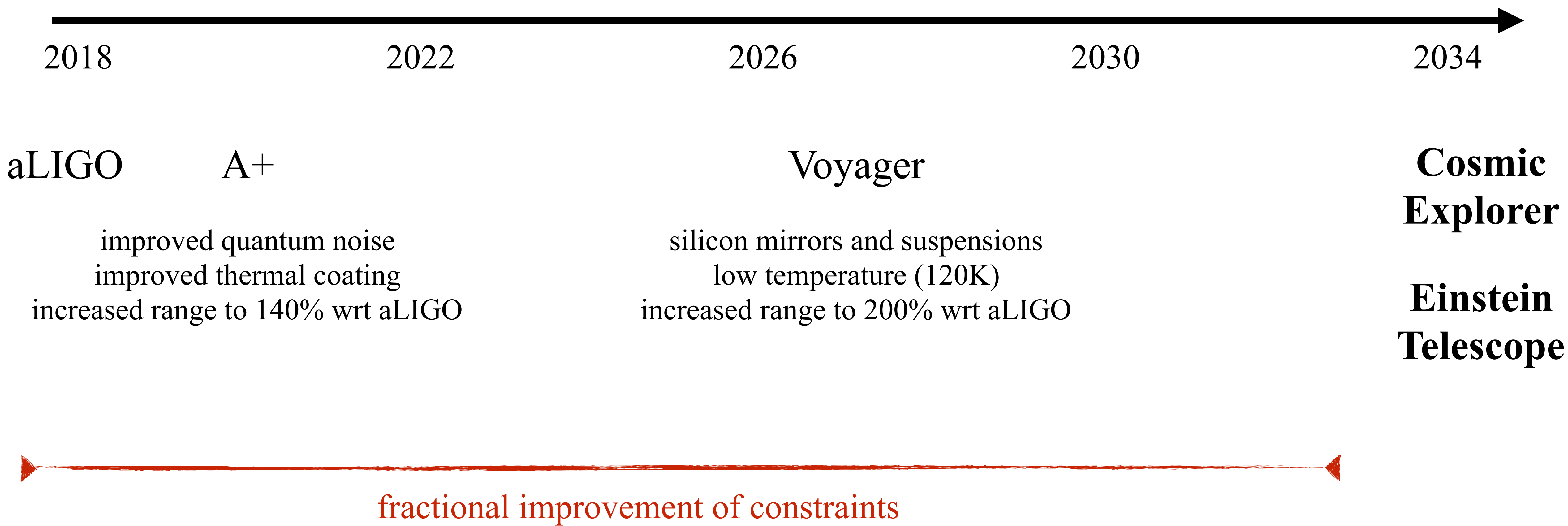
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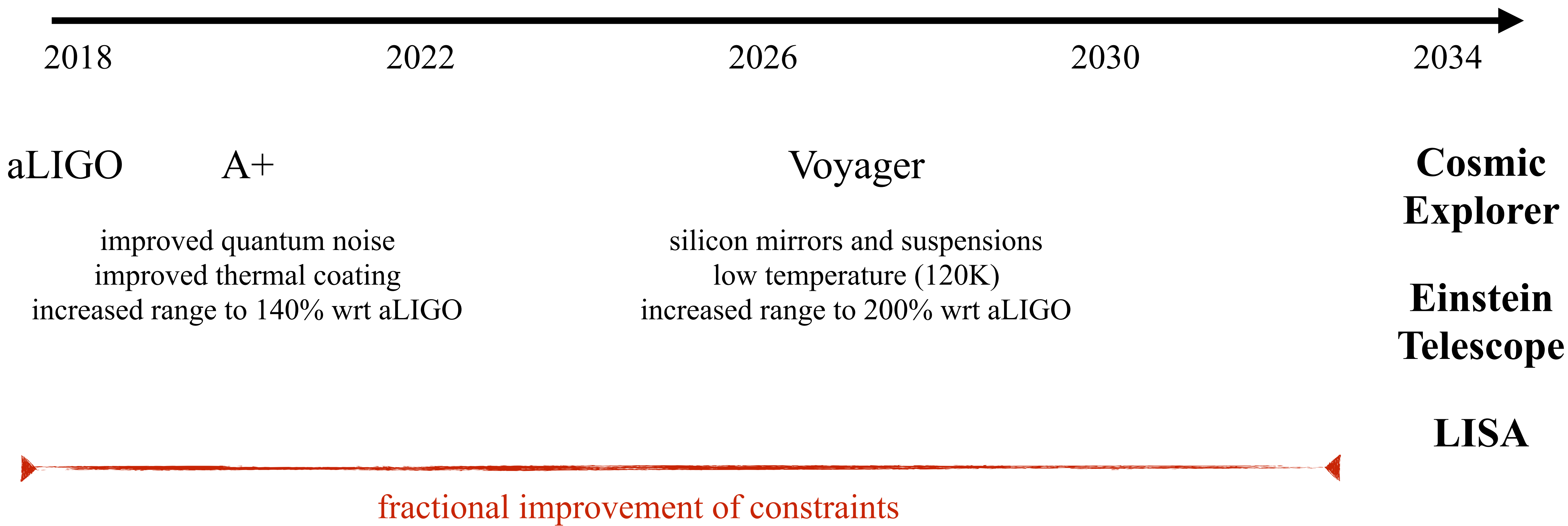
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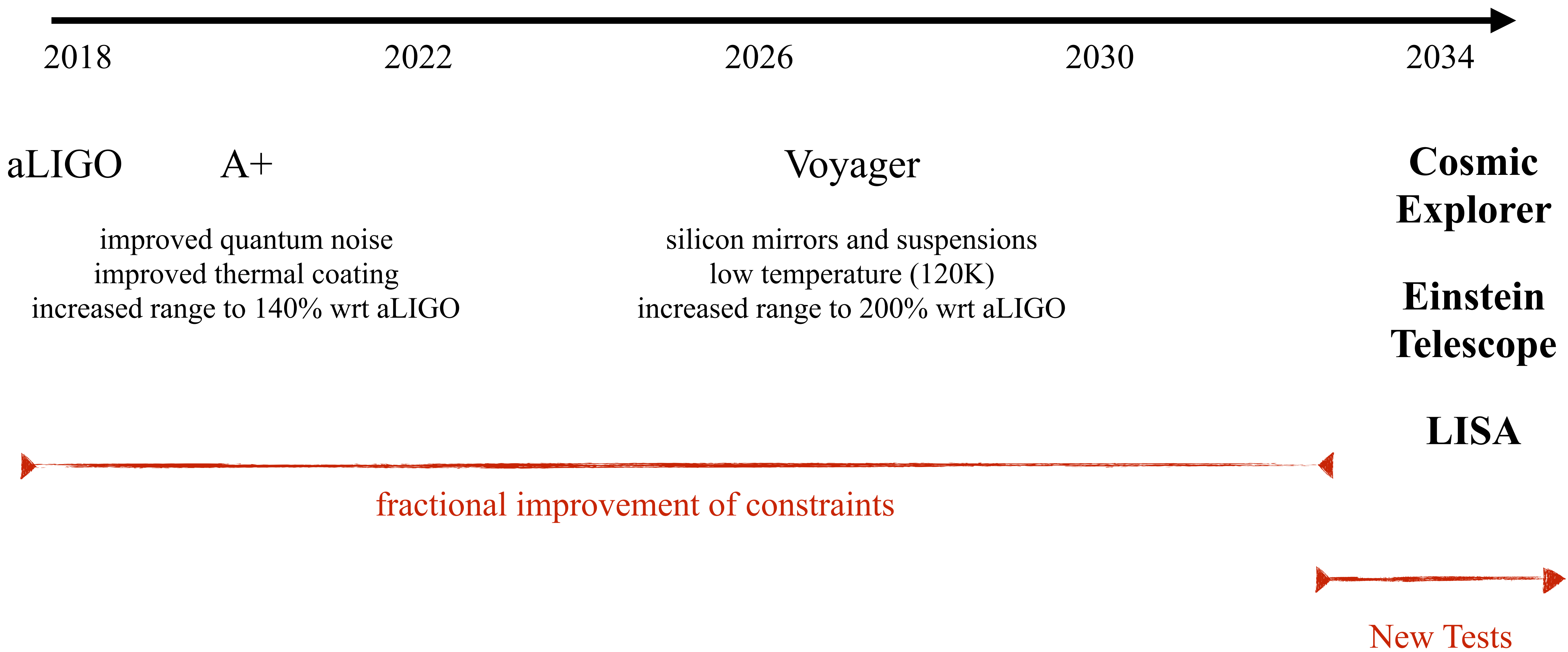
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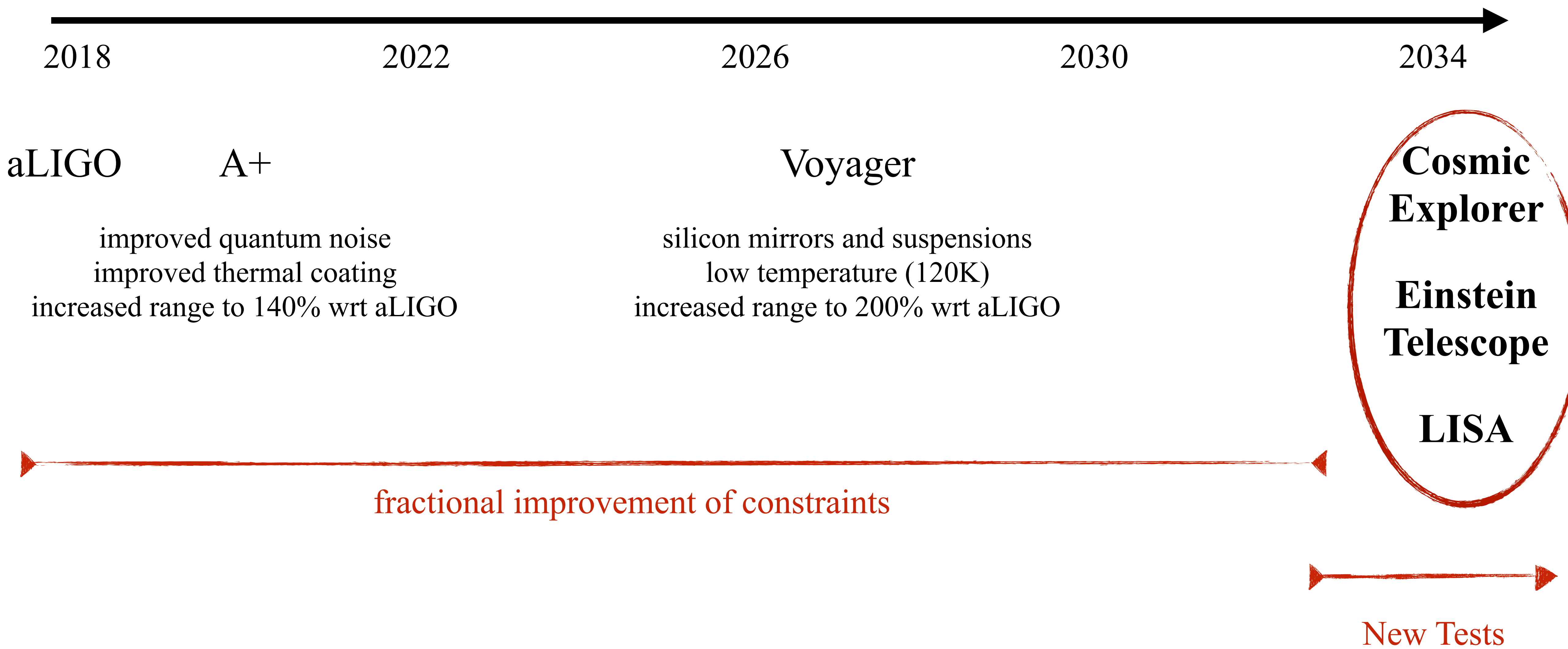
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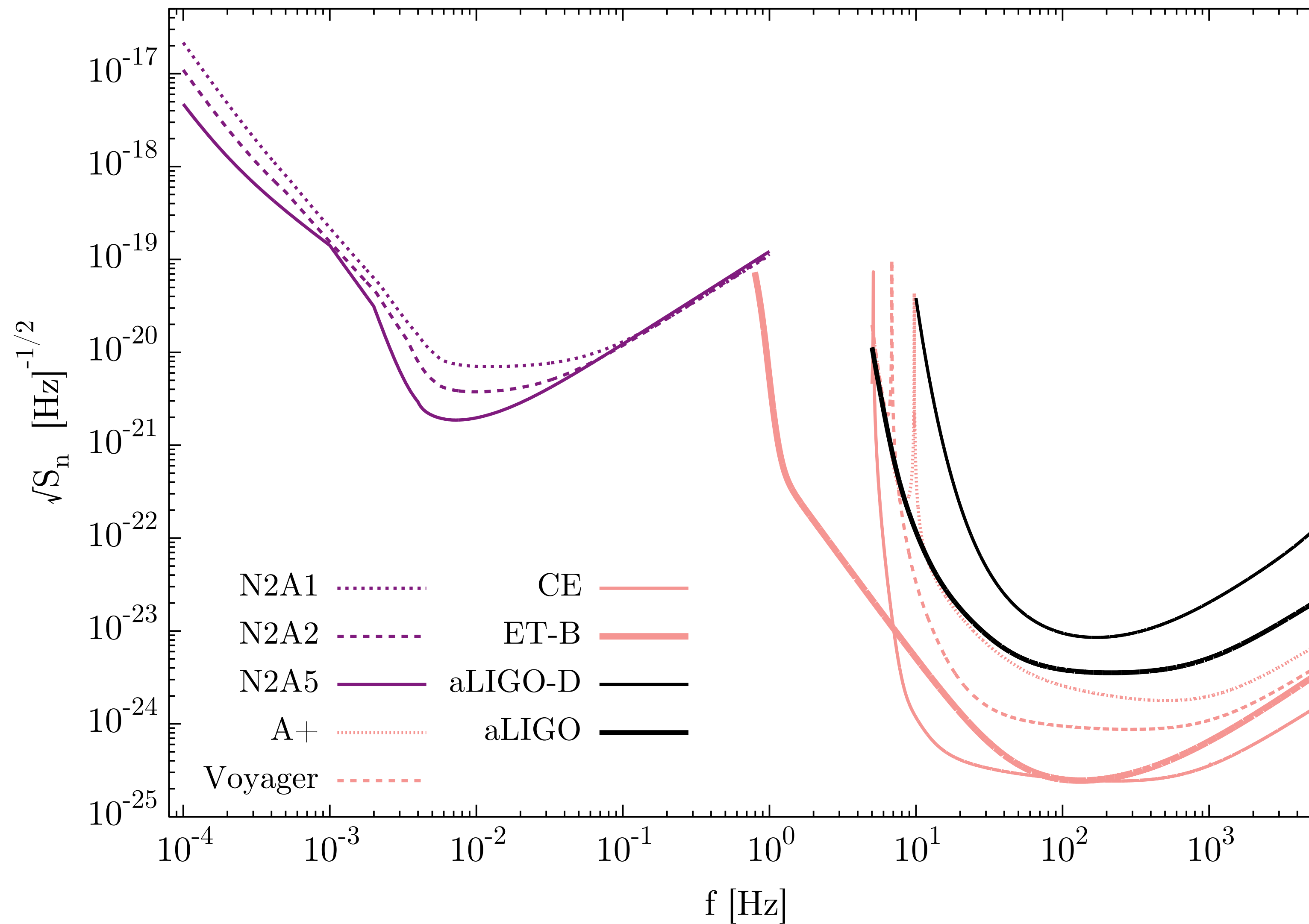
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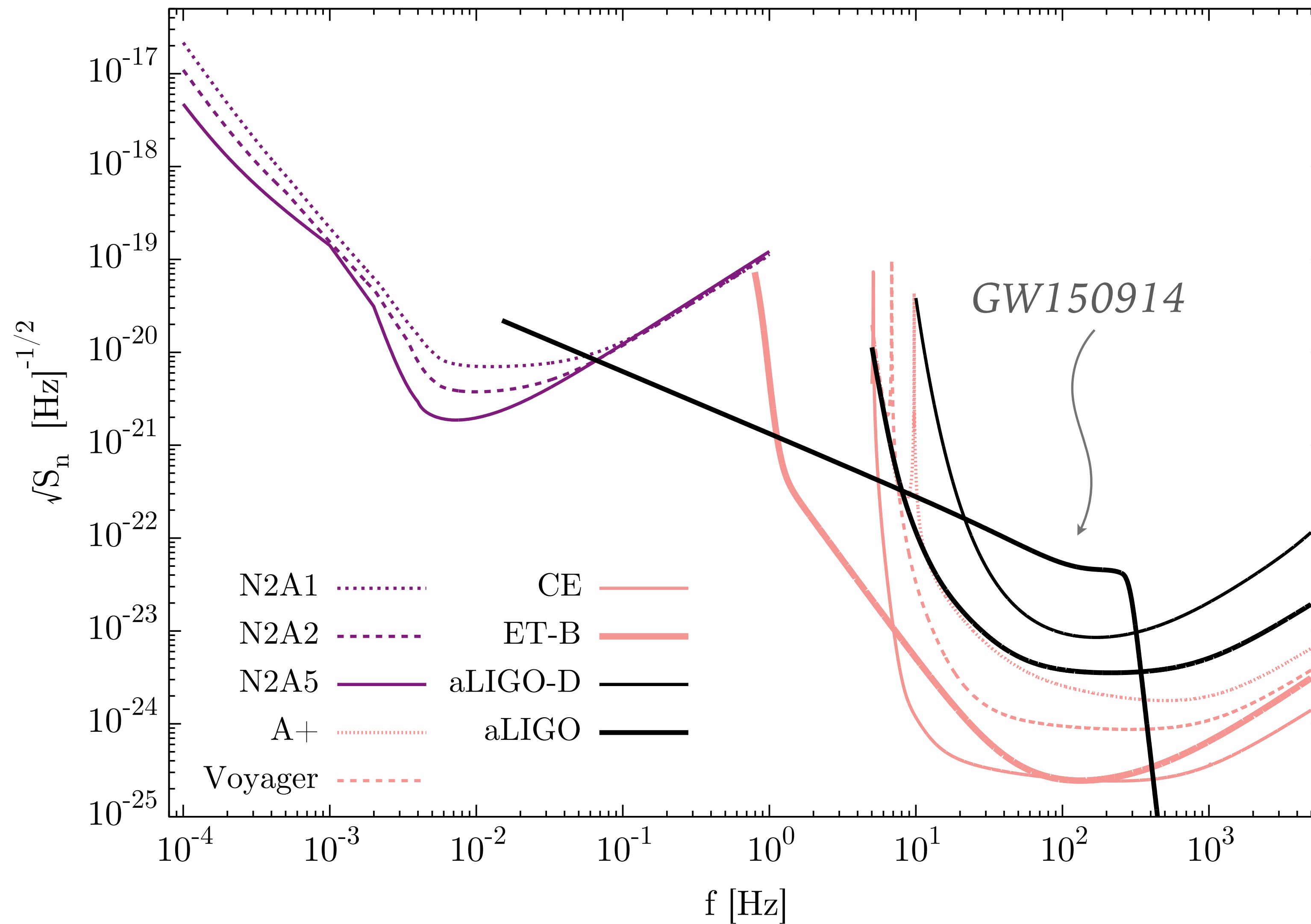
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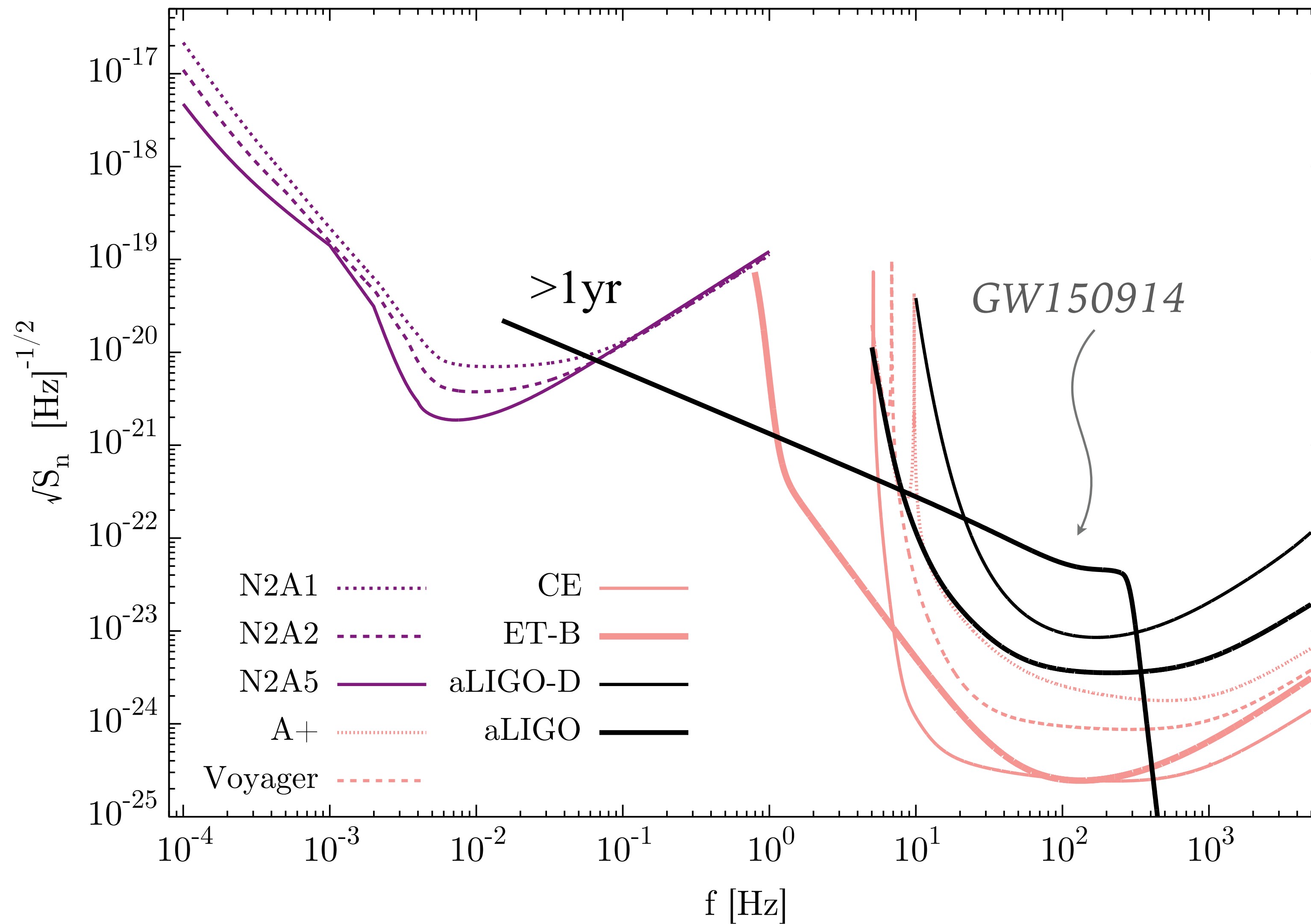
New how? Example: Multi-Band Events



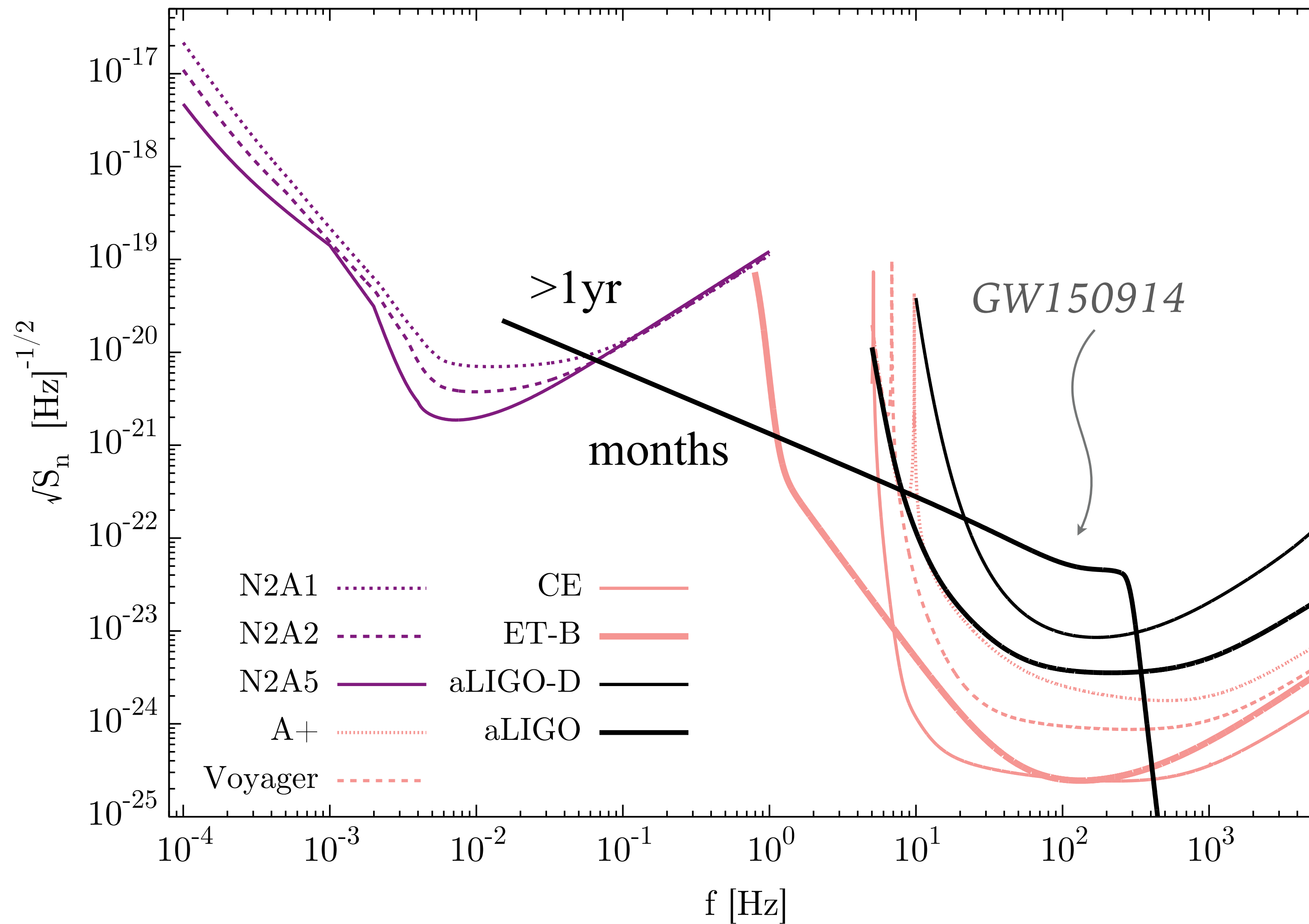
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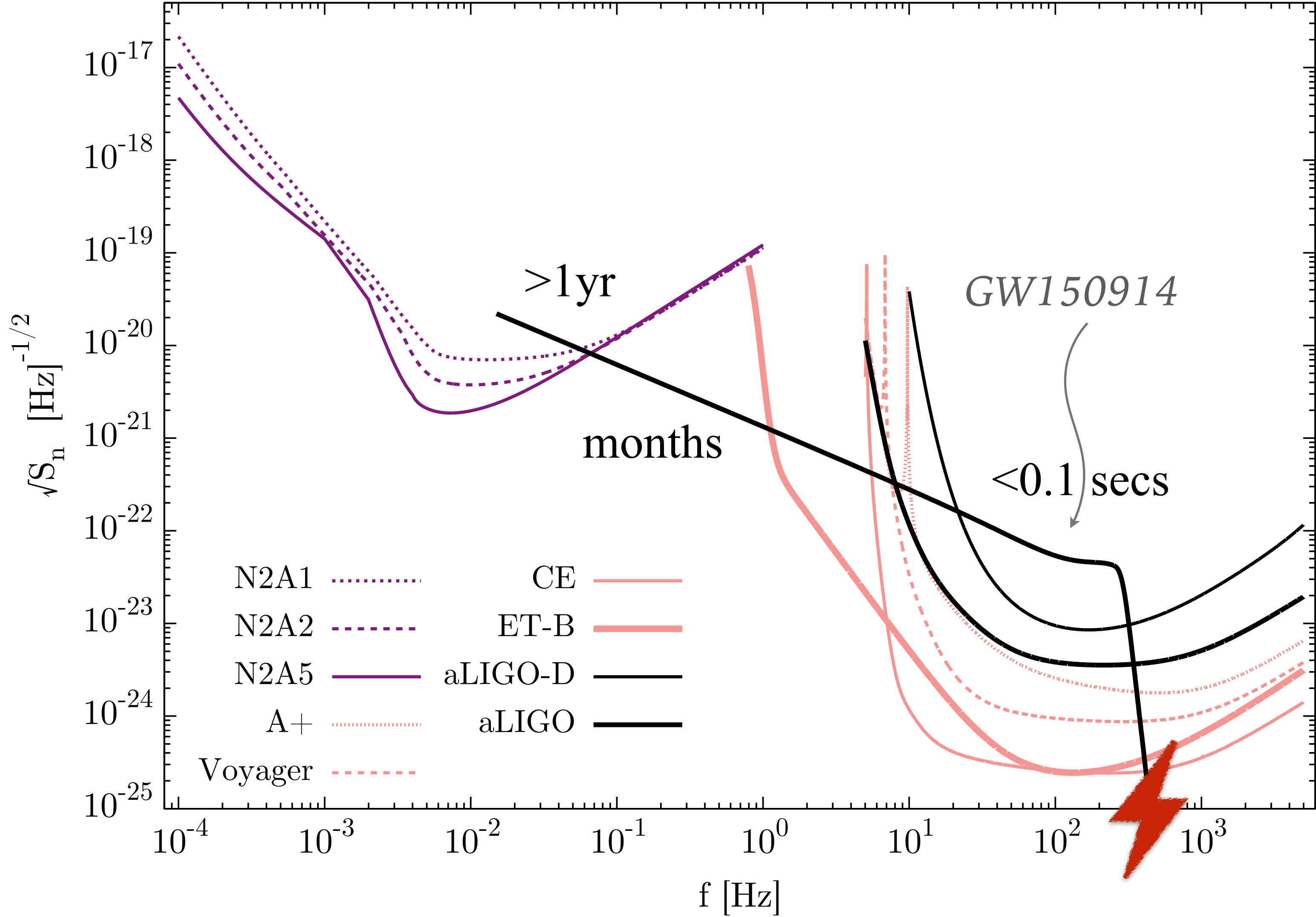
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New how? Example: Multi-Band Events



New how? Example: Multi-Band Events



New how? Example: Multi-Band Constraints

Case Study: Dipole Radiation

$$\dot{E}_{\text{GW}} = \dot{E}_{\text{GR}} \left[1 + B \left(\frac{Gm}{r_{12}c^2} \right)^{-1} \right]$$

$$\tilde{h}(f) = \tilde{h}_{\text{GR}}(f) (1 + \alpha f^a) e^{i\beta f^b}$$

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[Barausse, Yunes, Chamberlain, PRL '16]

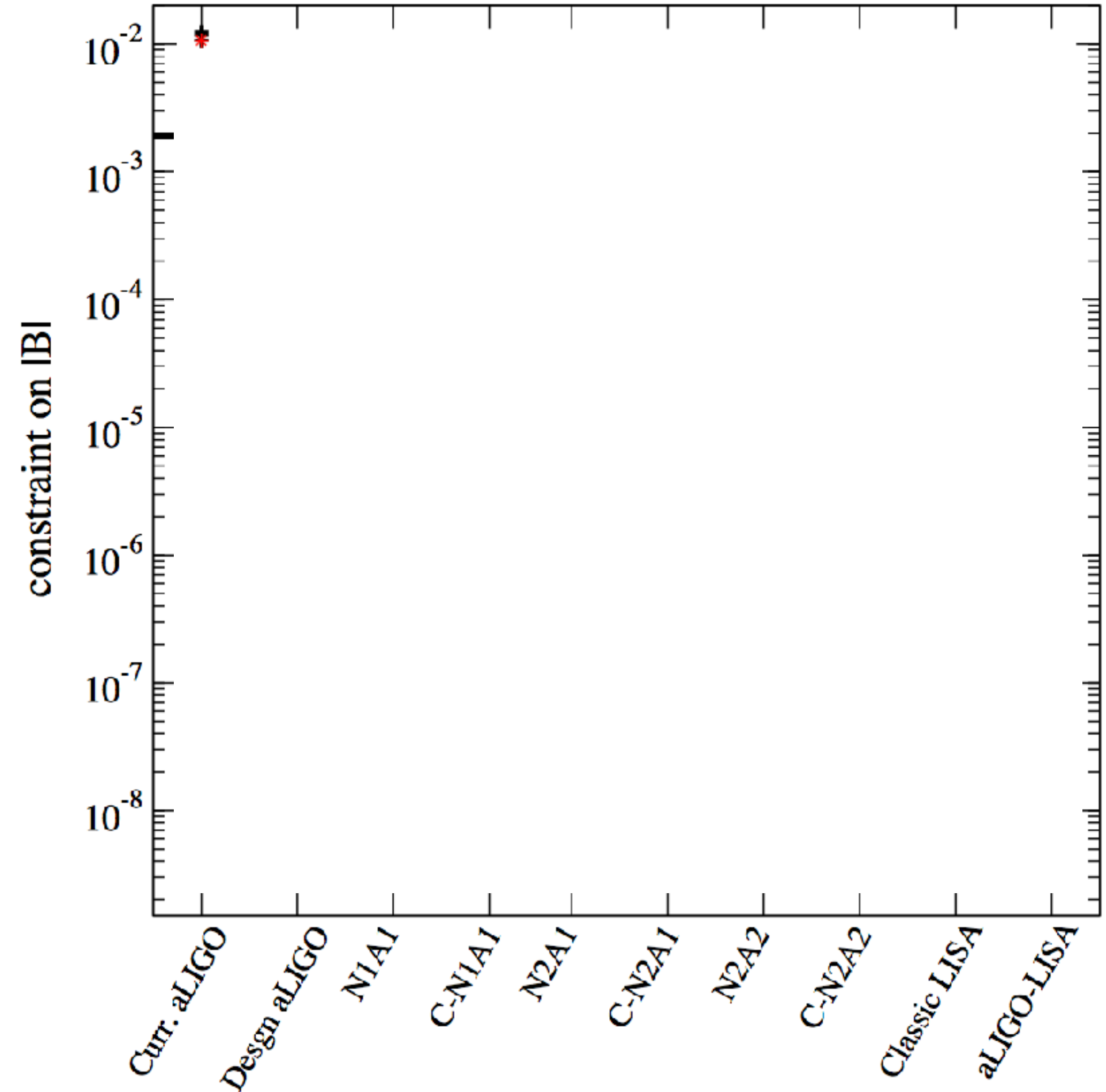
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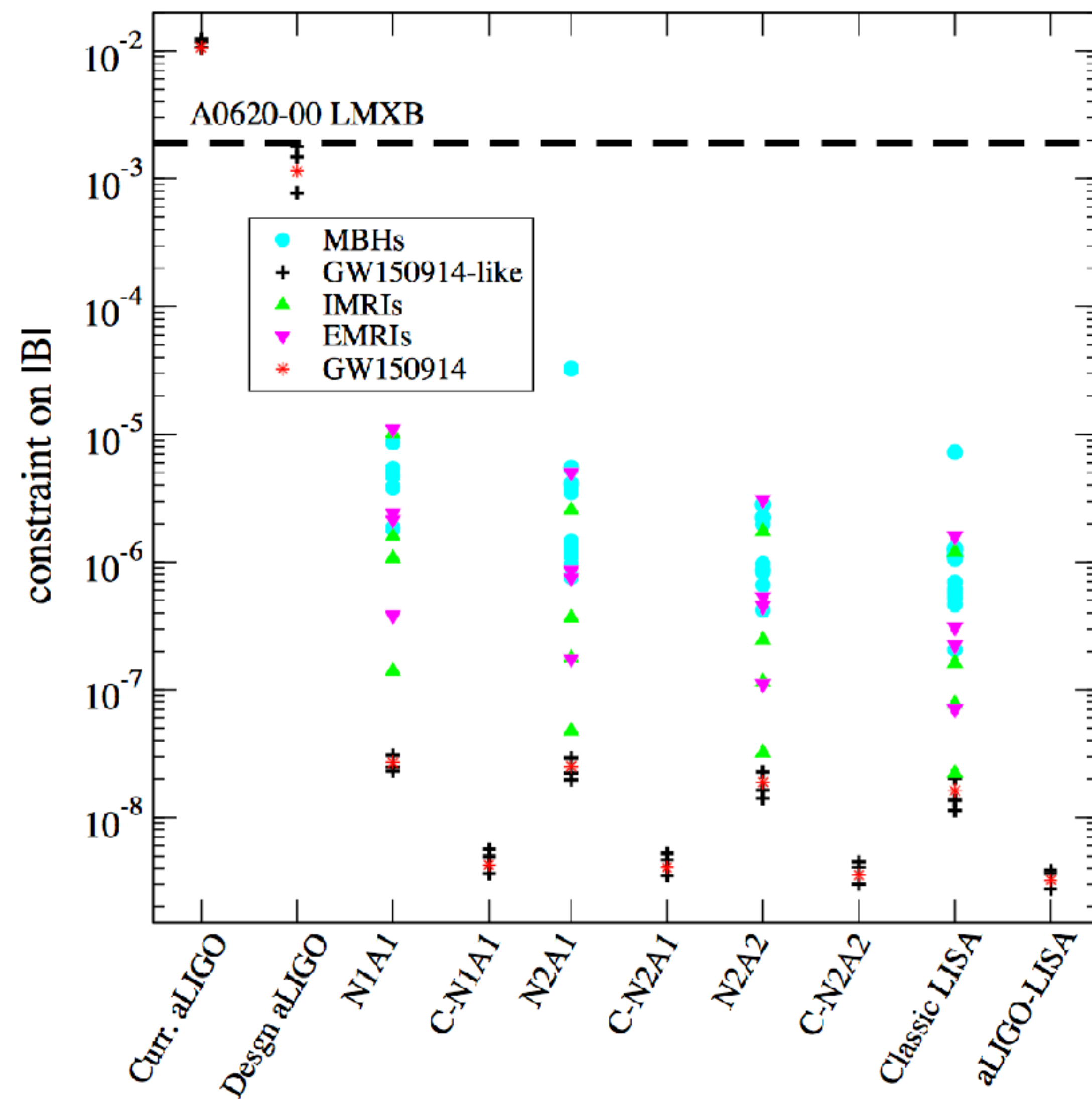
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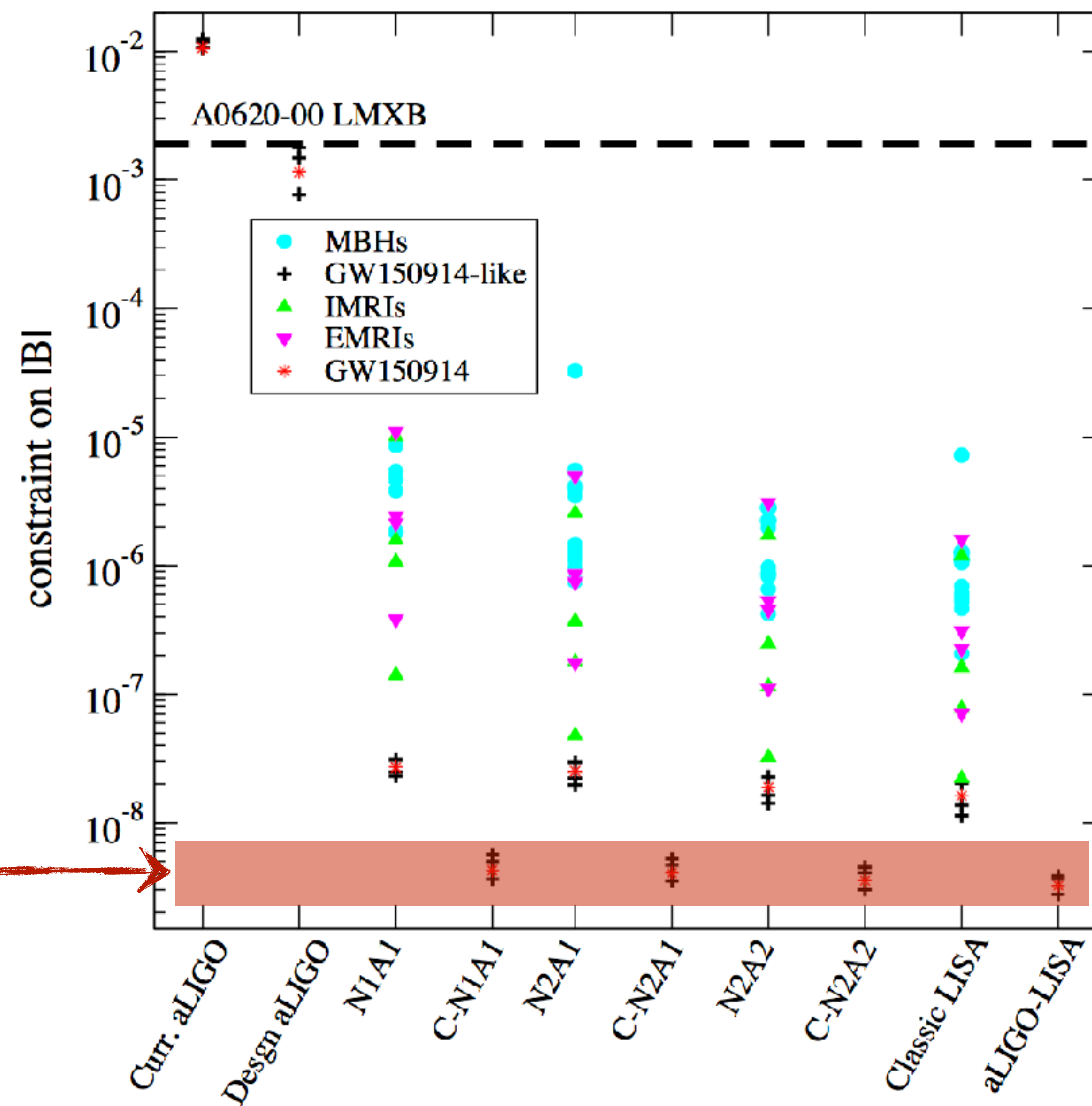
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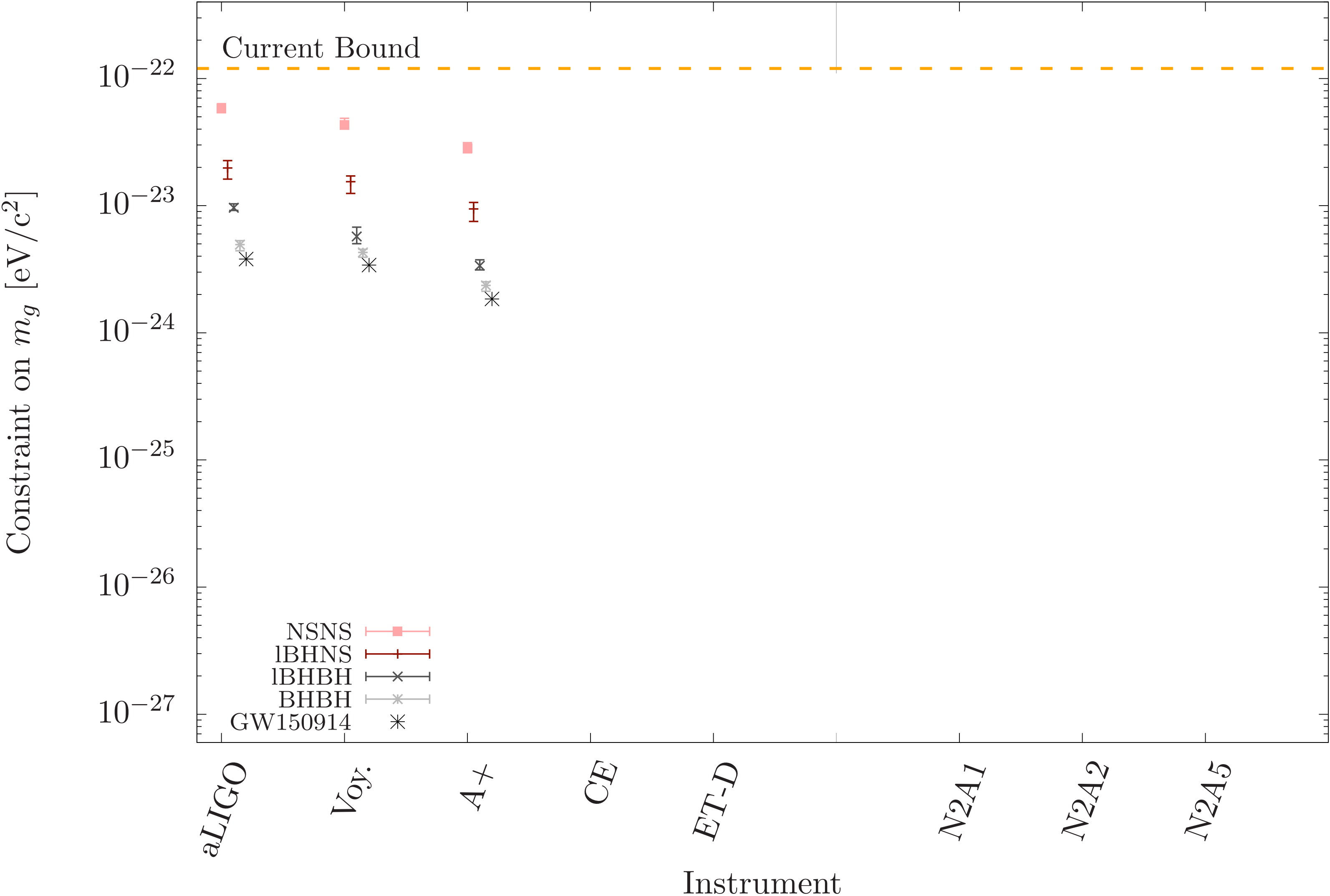
**10⁶ times better than
current bounds!!**



[Barausse, Yunes, Chamberlain, PRL '16]

New how? Example: Precision Tests

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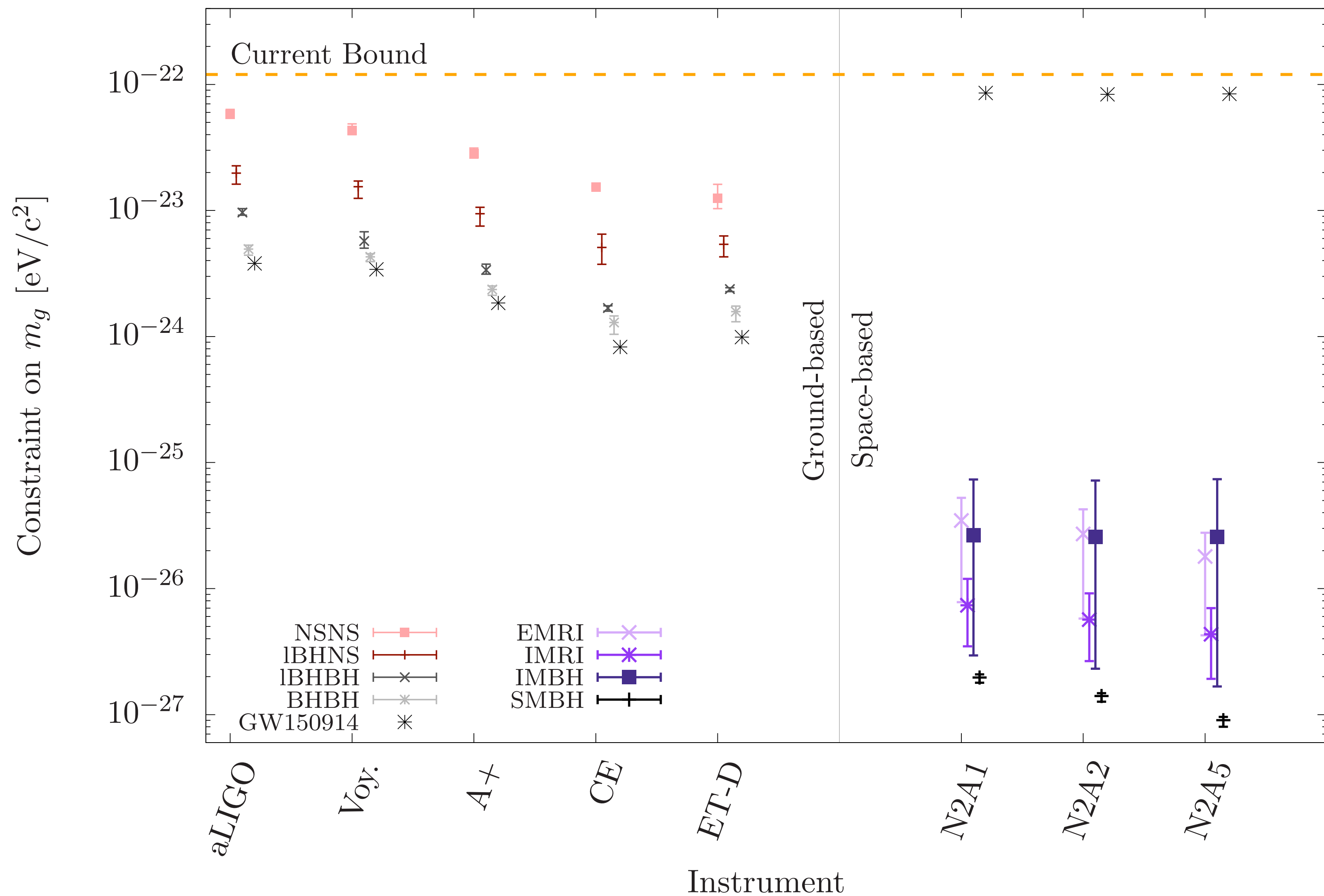
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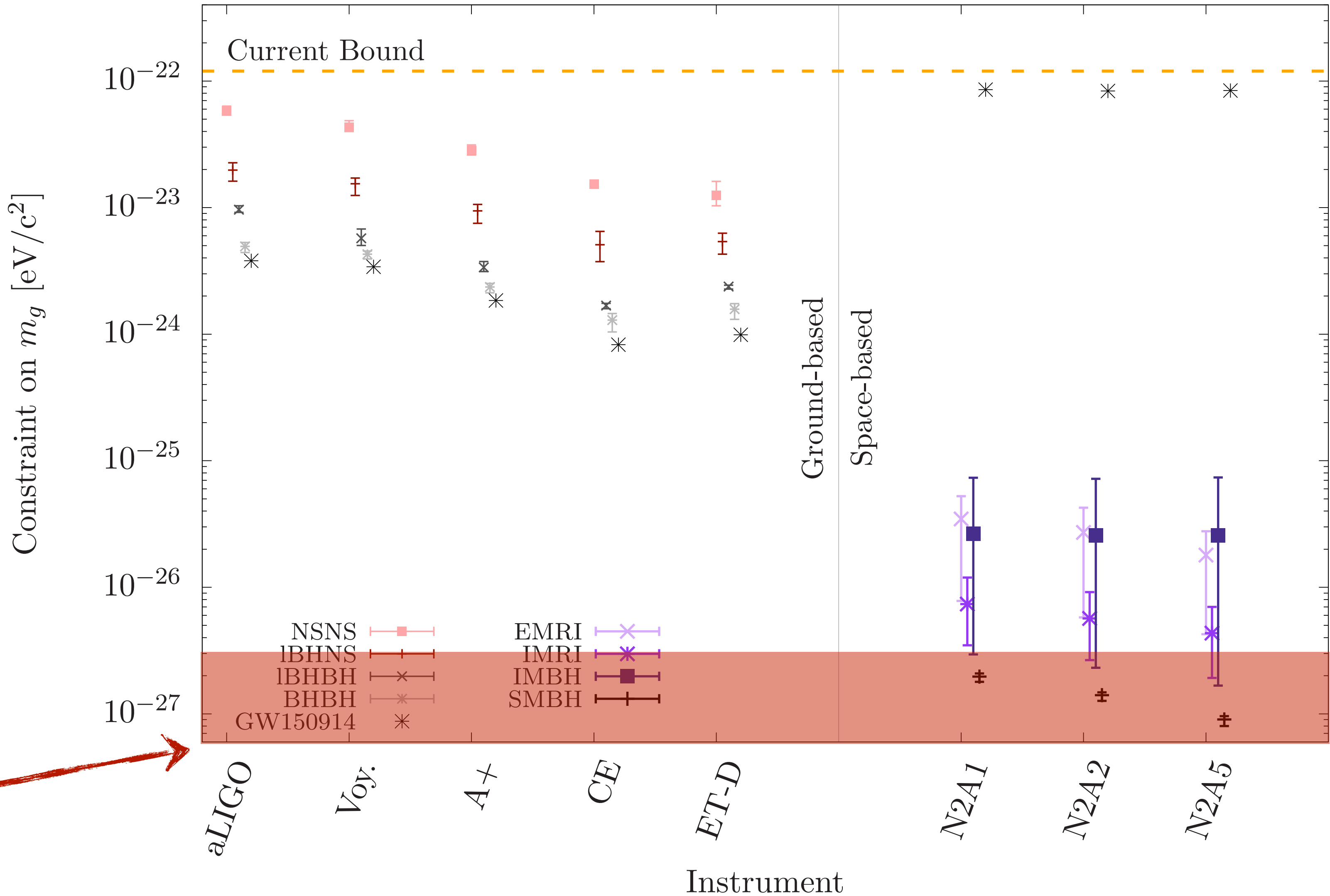
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A Brave New World: A Classification

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New Sources

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SMBH Mergers

Final State Conjecture through QNMs, graviton mass, modified dispersion.

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EMRIs Kerr Hypothesis via GW geodesy, strong equivalence principle, chaos.

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Stochastic Backgrounds

Cosmological modified gravity, cosmic strings ?

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Connection to cosmological modified theories?

You ready for this? Necessary Simulations

New Sources

BH Mergers

EMRIs

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New Data

Much Higher SNR

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Multi-Band Sources

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The whole NR enchilada in MG (spin-precessing + eccentric + higher modes)

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Multi-Band Sources

Two stage-analysis (PN for early LISA inspiral → NR for LIGO merger)

What are the NR challenges?

New Sources

SMBH Enchilada

EMRIs

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High SNR/High Precision

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Multi-Band Sources

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Intermediate mass-ratio inspiral modeling...

3G Detectors Will Allow for New Tests of GR
(precision tests, new sources, new effects)

Conclusions

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Control over systematic errors in parameter estimation? etc...

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If it bleeds,
can we kill it?



Thank You