

The background of the slide is a composite image. On the left, the curved horizon of Earth is visible from space, showing clouds and the dark surface. The rest of the background is a deep space scene featuring a large, bright, irregularly shaped galaxy in the center. The galaxy is surrounded by a complex network of purple and blue filaments, representing the cosmic web. In the upper right corner, there is a circular, multi-colored (green, blue, red) pattern with a bright white center, resembling a galaxy cluster or a specific astronomical data visualization. The overall scene is set against a dark, star-filled sky.

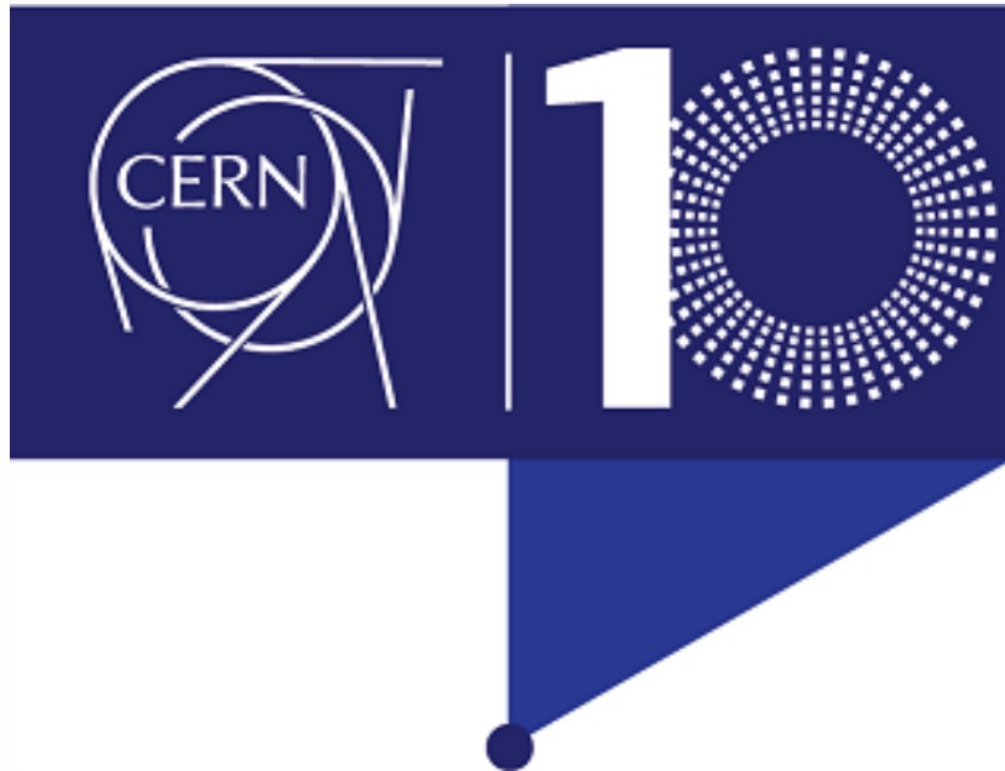
Precision Cosmology  
IMFP Winter Meeting 2022  
Benasque, 6<sup>th</sup> September 2022

Juan García-Bellido  
IFT-UAM/CSIC

# Overview

From Fundamental Physics to Precision Cosmology

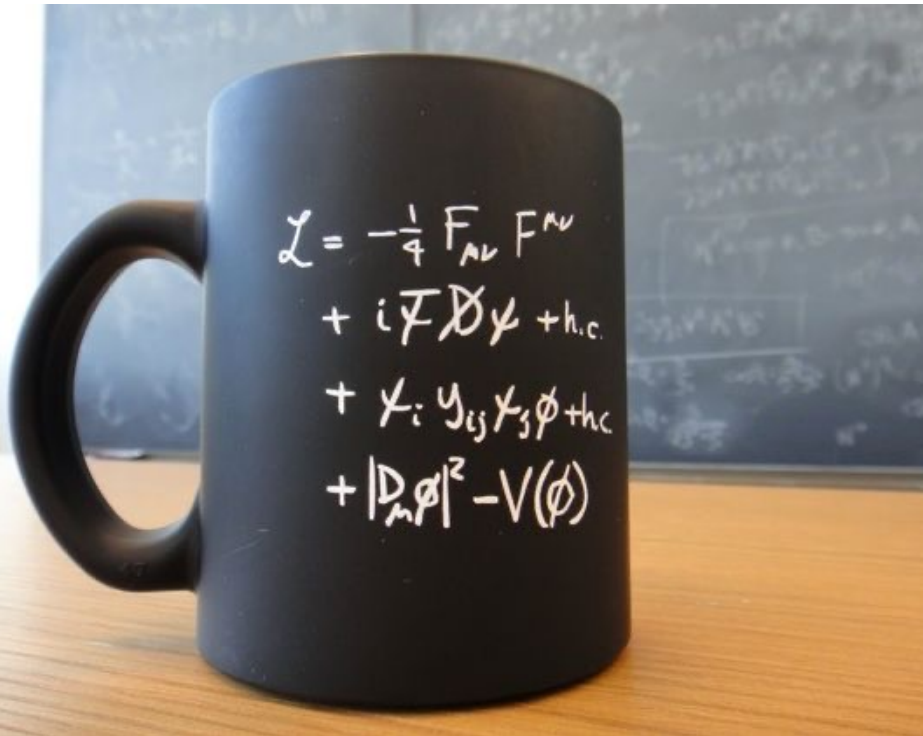
- From non-minimal coupling  $\xi$  to Higgs Inflation
- Renormalization Group Equations ( $\lambda(\mu)$ ,  $\xi(\mu)$ )
- Critical Higgs Inflation and PBH = DM
- Inevitable Quantum Diffusion = PNG
- Signatures of QDiff in CMB and LSS
- Signatures of PBH in microlensing & GW
- Hubble tension and future surveys.
- Entropic Forces and Structure Formation
- Cosmic Acceleration from First Principles



years  
**HIGGS boson**  
discovery

4<sup>th</sup> July 2012 - Higgs boson discovery

# Standard Model Lagrangian



$$\begin{aligned} \mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ & + i\bar{\psi} \not{D} \psi + \text{h.c.} \\ & + \bar{\psi}_i y_{ij} \psi_j \phi + \text{h.c.} \\ & + |D_\mu \phi|^2 - V(\phi) \\ & + \xi |\phi|^2 R \end{aligned}$$

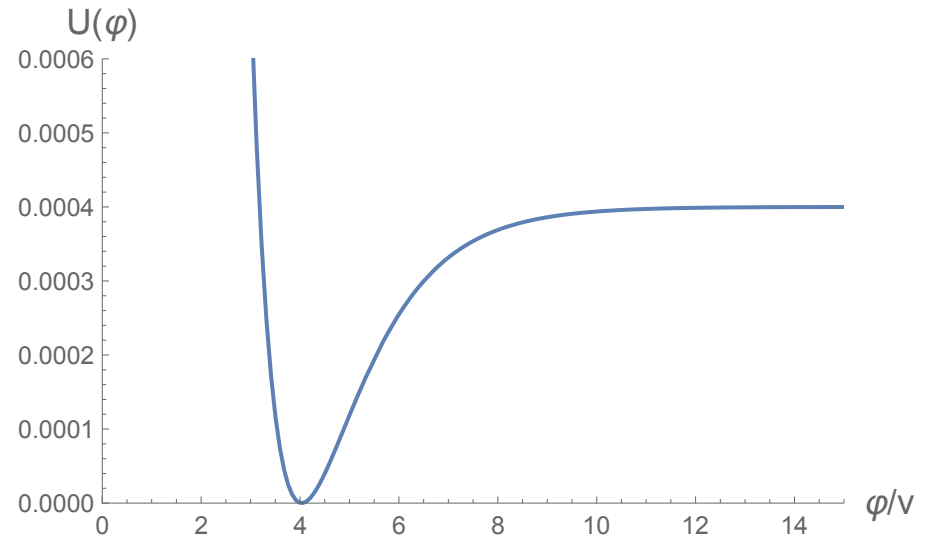
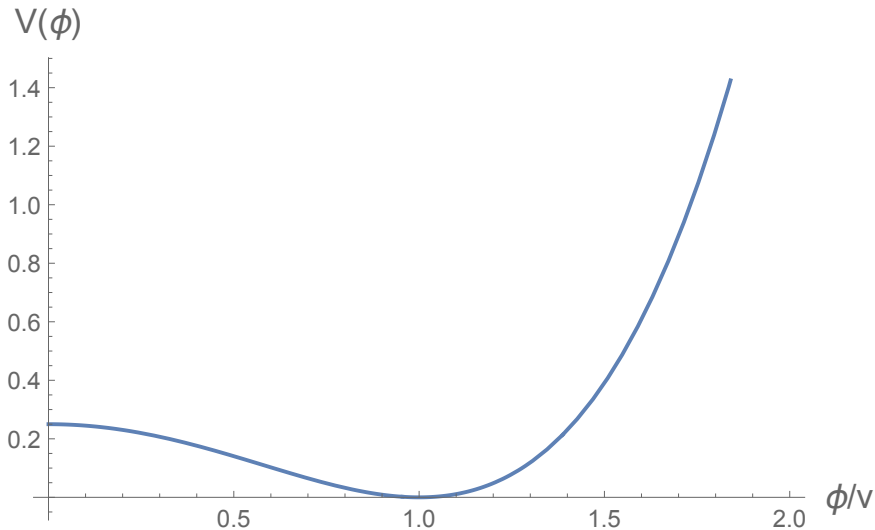
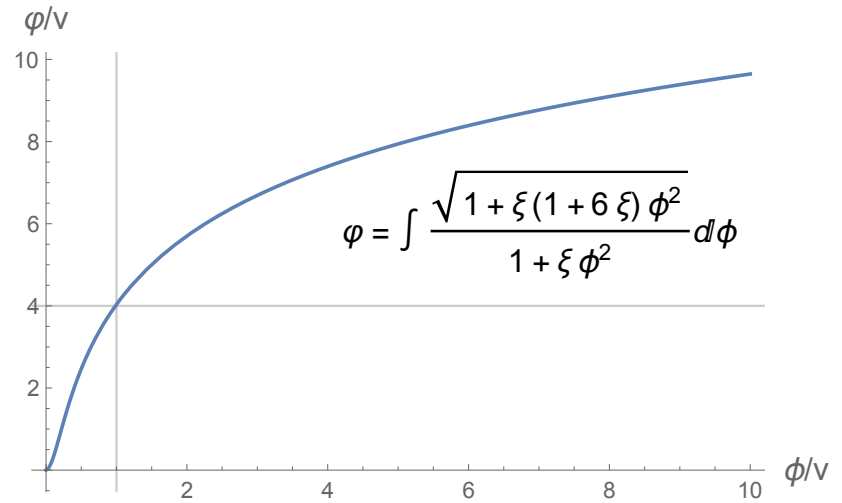
$$R = 12H^2 + 6\dot{H} \rightarrow R_0 = 9.2 H_0^2 \rightarrow m_H = \sqrt{\xi R_0} = 2 \times 10^{-32} \text{ eV}$$

# Conformal redefinition of metric and Higgs

$$g_{\mu\nu} \rightarrow (1 + \xi\phi^2)g_{\mu\nu}$$

$$\phi \rightarrow \varphi$$

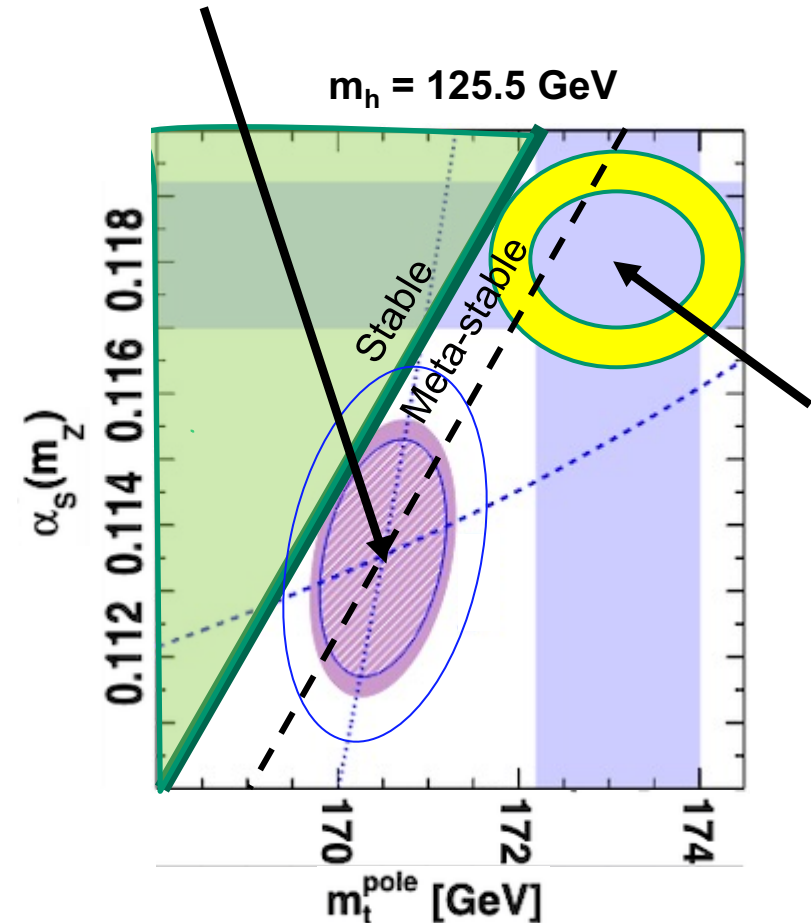
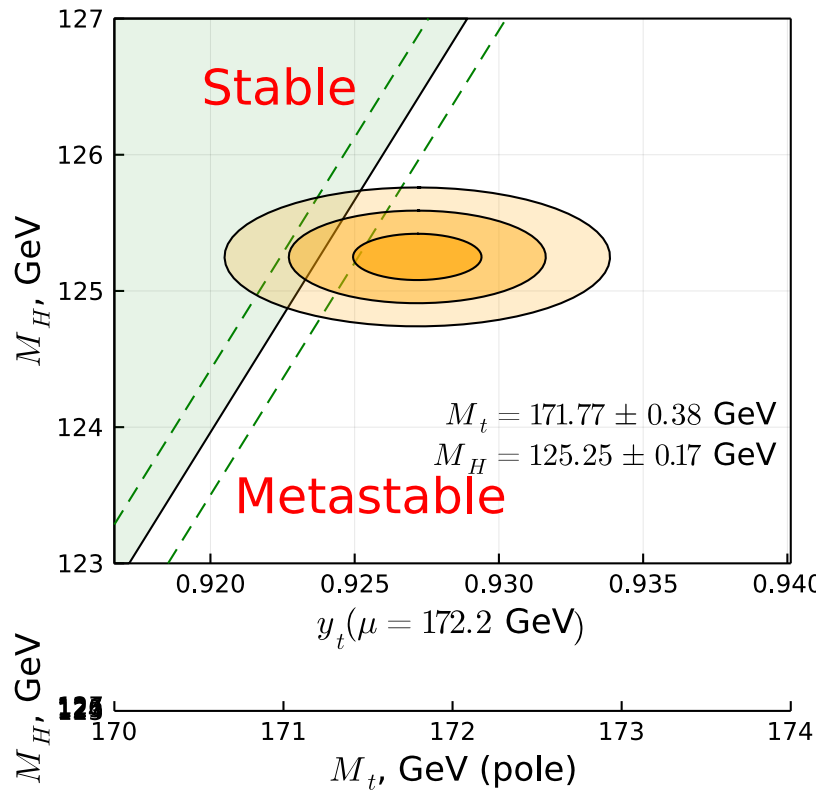
$$V(\phi) \rightarrow \frac{V(\phi)}{(1 + \xi\phi^2)^2}$$



# EW vacuum metastability

LHC-CMS Collab. (2020)

<https://arxiv.org/abs/1904.05237>



$$m_t^{\text{pole}} = 170.5 \pm 0.8 \text{ GeV}$$

$$\alpha_S(m_Z) = 0.1135^{+0.0021}_{-0.0017}$$

Buttazzo et al. (2012)

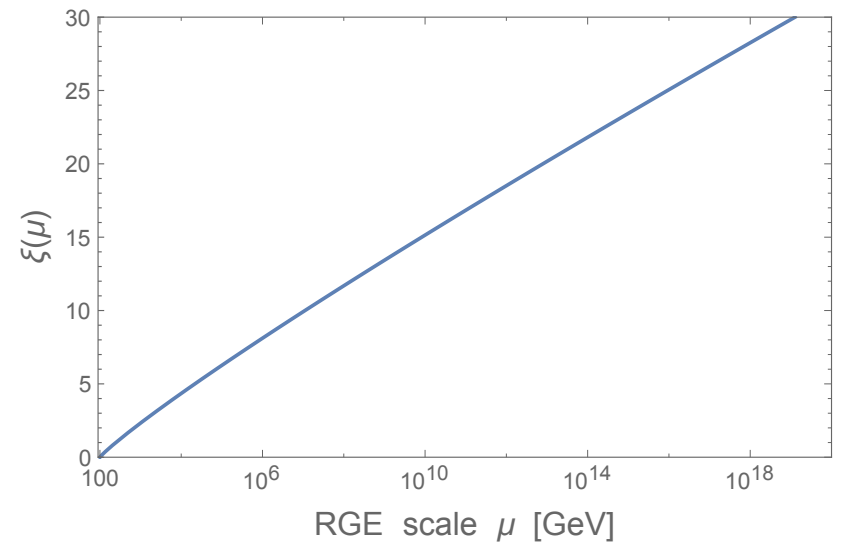
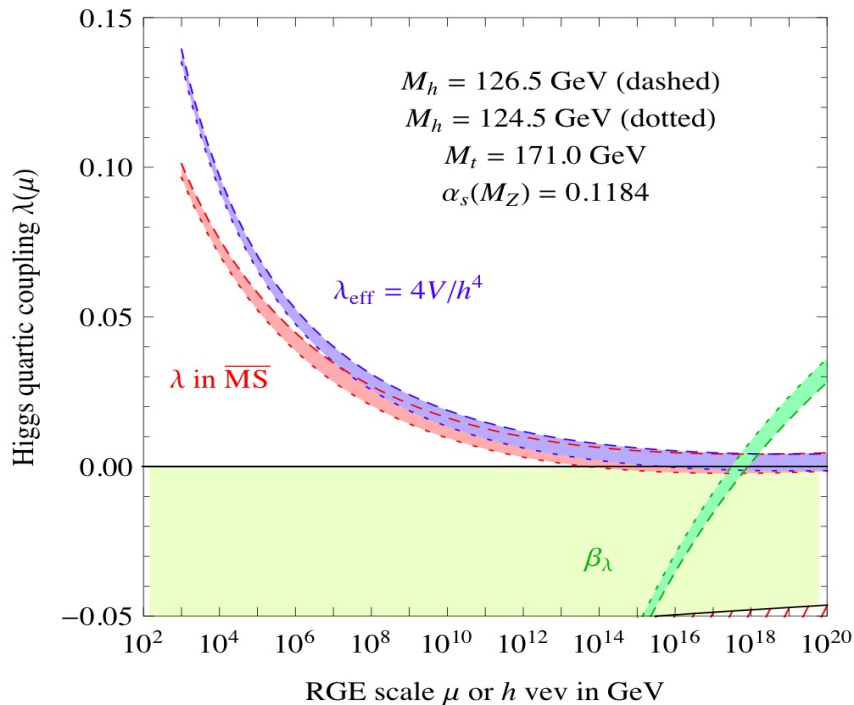
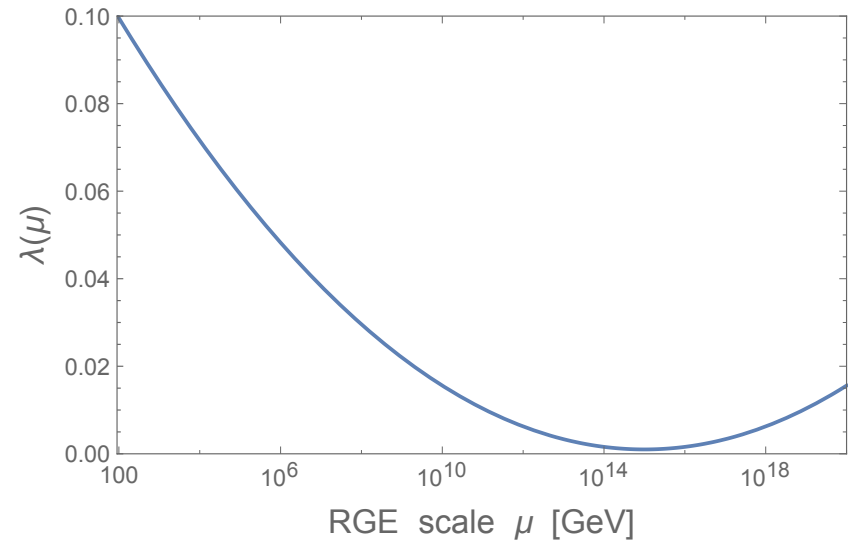
<https://arxiv.org/pdf/1112.3022.pdf>

# Renormalization of Higgs couplings

$$\lambda(\phi) = \lambda_0 + b_\lambda \ln^2(\phi/\mu),$$

$$\xi(\phi) = \xi_0 + b_\xi \ln(\phi/\mu),$$

Buttazzo et al (2014)



# Critical Higgs Inflation

Ezquiaga, JGB, Ruiz Morales (2017)

$$S = \int d^4x \sqrt{g} \left[ \left( \frac{1}{2\kappa^2} + \frac{\xi(\phi)}{2} \phi^2 \right) R - \frac{1}{2} (\partial\phi)^2 - \frac{1}{4} \lambda(\phi) \phi^4 \right]$$

$$\lambda(\phi) = \lambda_0 + b_\lambda \ln^2(\phi/\mu) ,$$

$$\xi(\phi) = \xi_0 + b_\xi \ln(\phi/\mu) ,$$

$$\frac{d\varphi}{d\phi} = \frac{\sqrt{1 + \xi(\phi) \phi^2 + 6 \phi^2 (\xi(\phi) + \phi \xi'(\phi)/2)^2}}{1 + \xi(\phi) \phi^2}$$

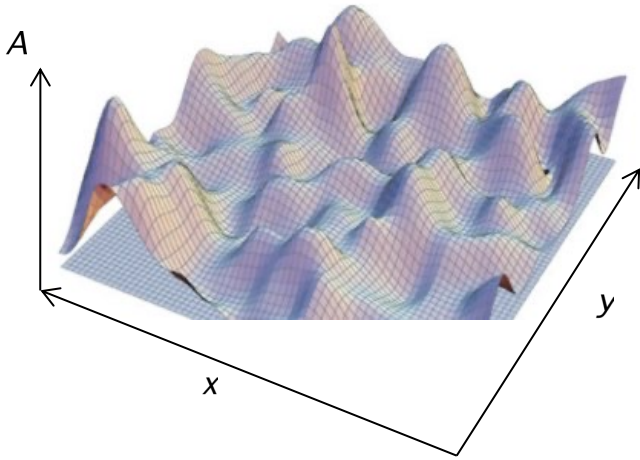
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$$V(x) = \frac{V_0 (1 + a \ln^2 x) x^4}{(1 + c (1 + b \ln x) x^2)^2} \quad x = \phi/\mu$$

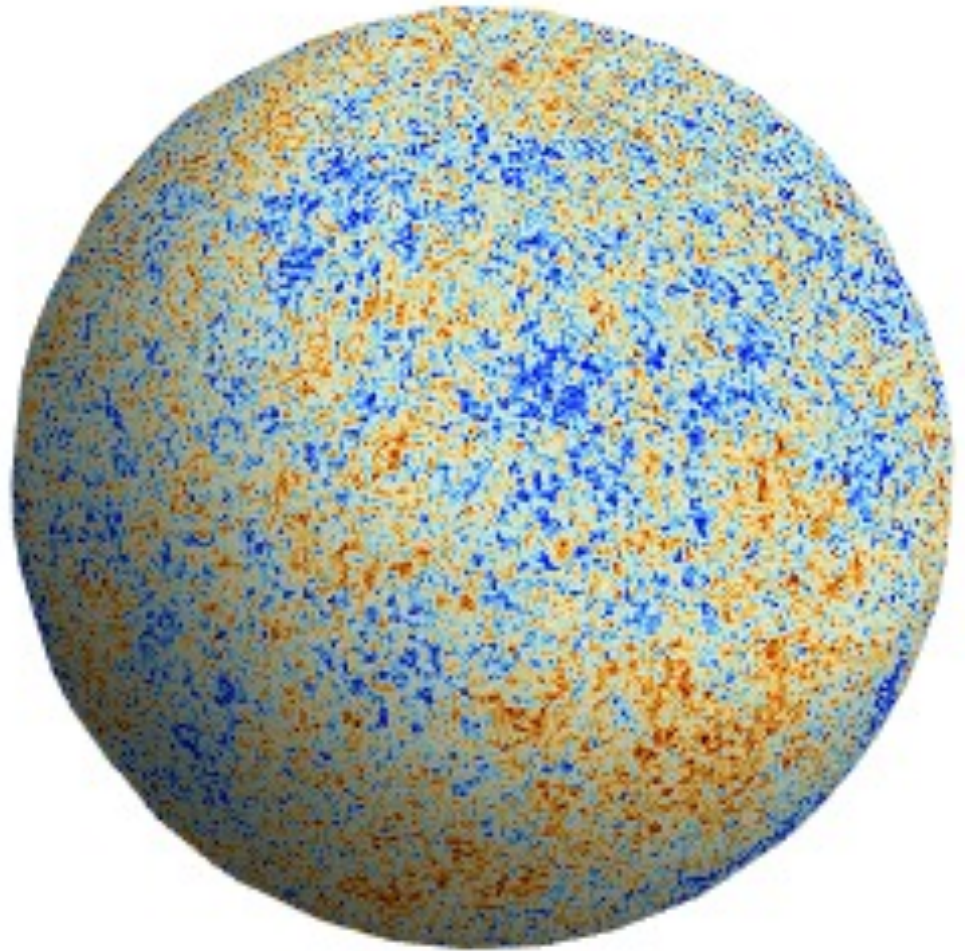
$$V_0 = \lambda_0 \mu^4 / 4, \quad a = b_\lambda / \lambda_0, \quad b = b_\xi / \xi_0 \quad \text{and} \quad c = \xi_0 \kappa^2 \mu^2$$



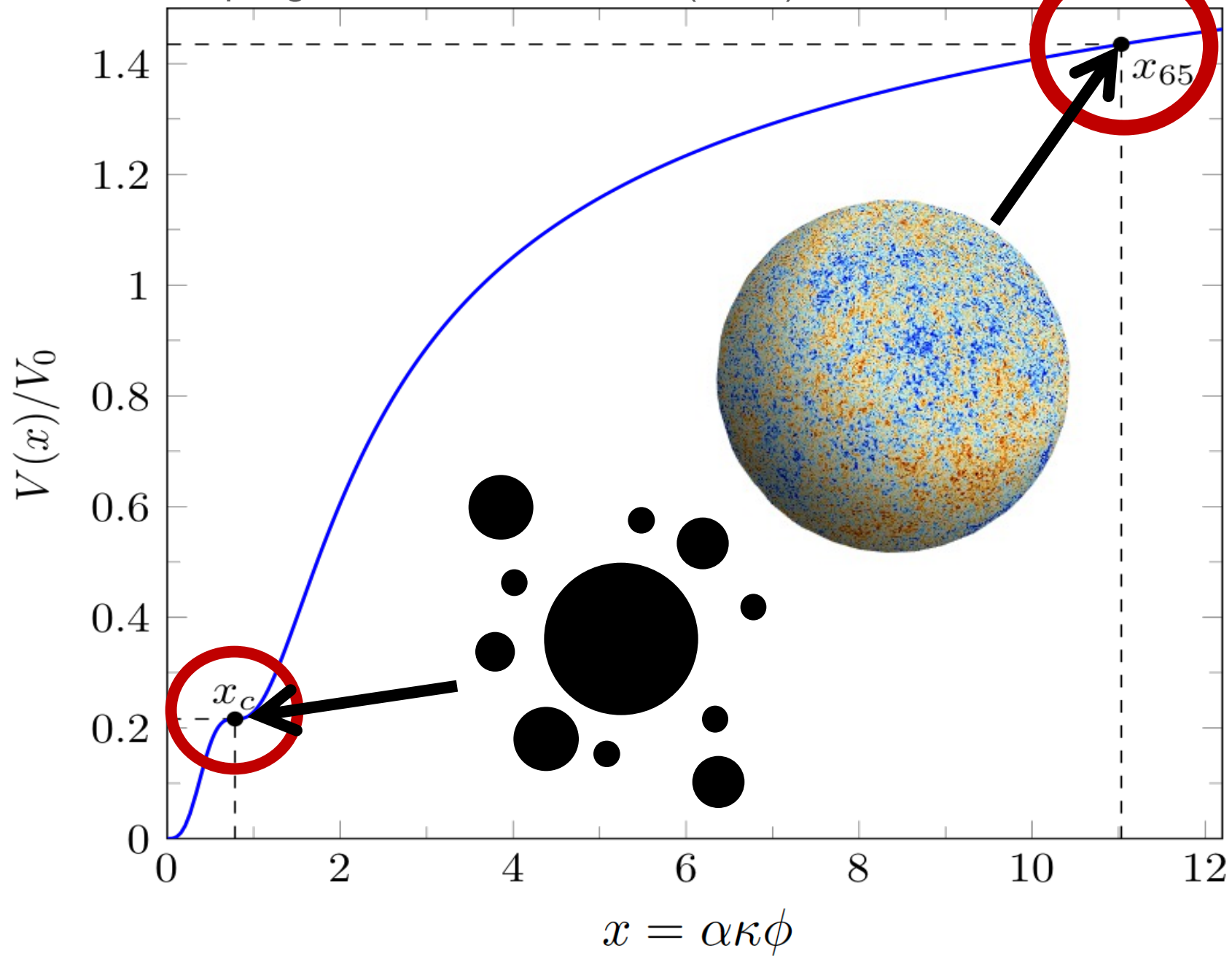
# Inflation



Quantum Fluctuations=  
Ripples in Space-Time

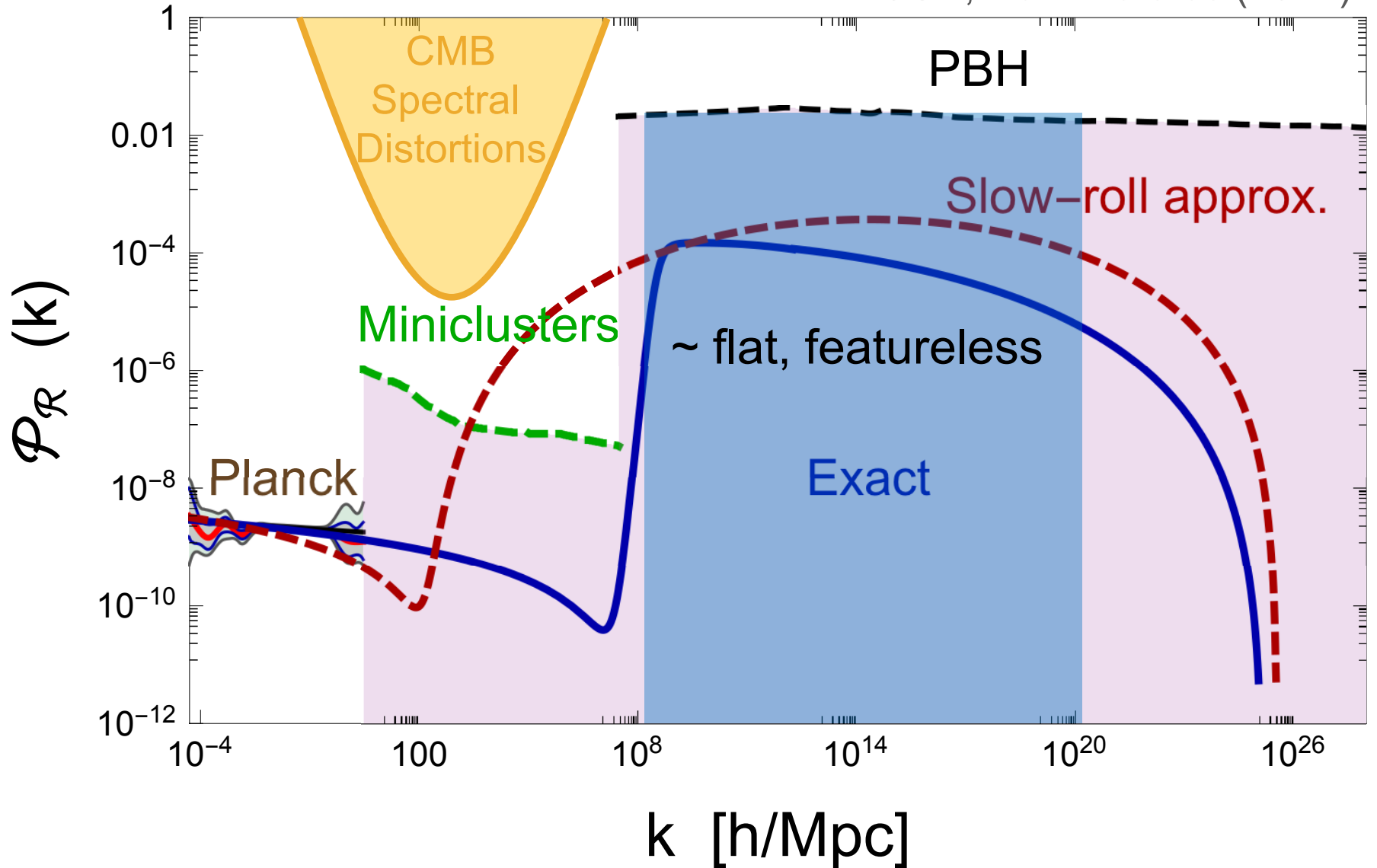


Stretched to cosmological distances

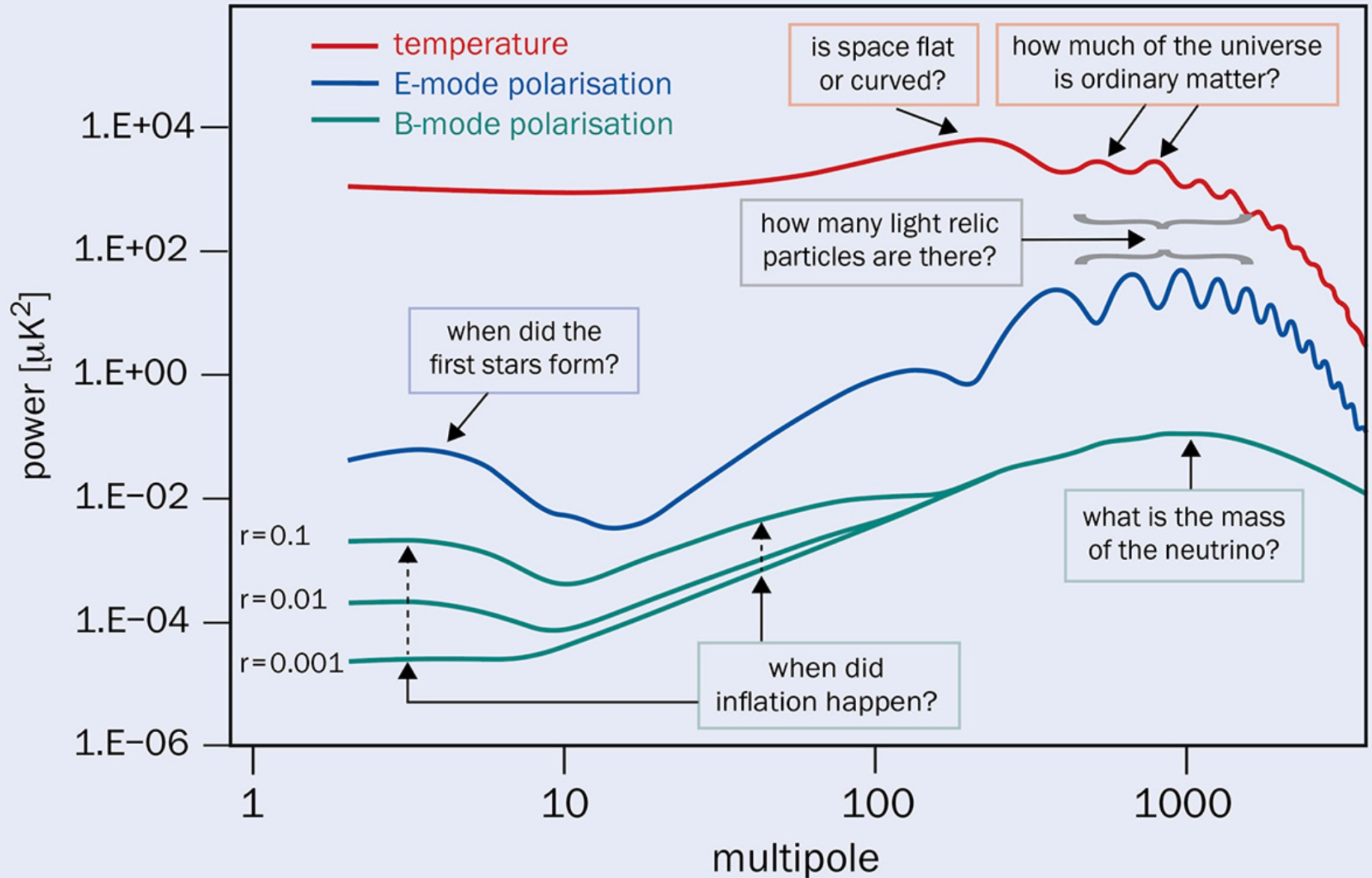


# Primordial Power Spectrum

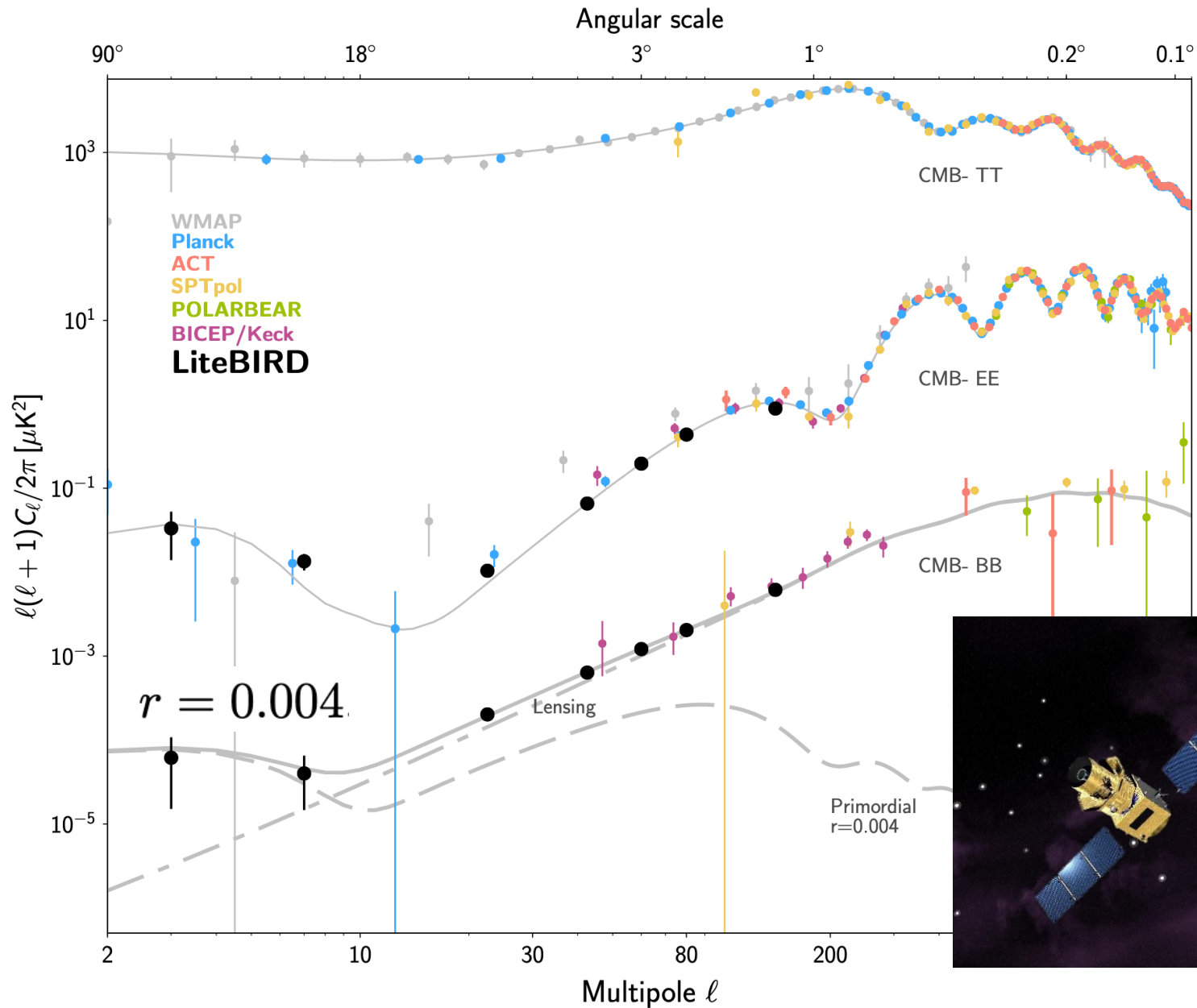
JGB, Ruiz Morales (2017)



# Future experiments will search for B modes



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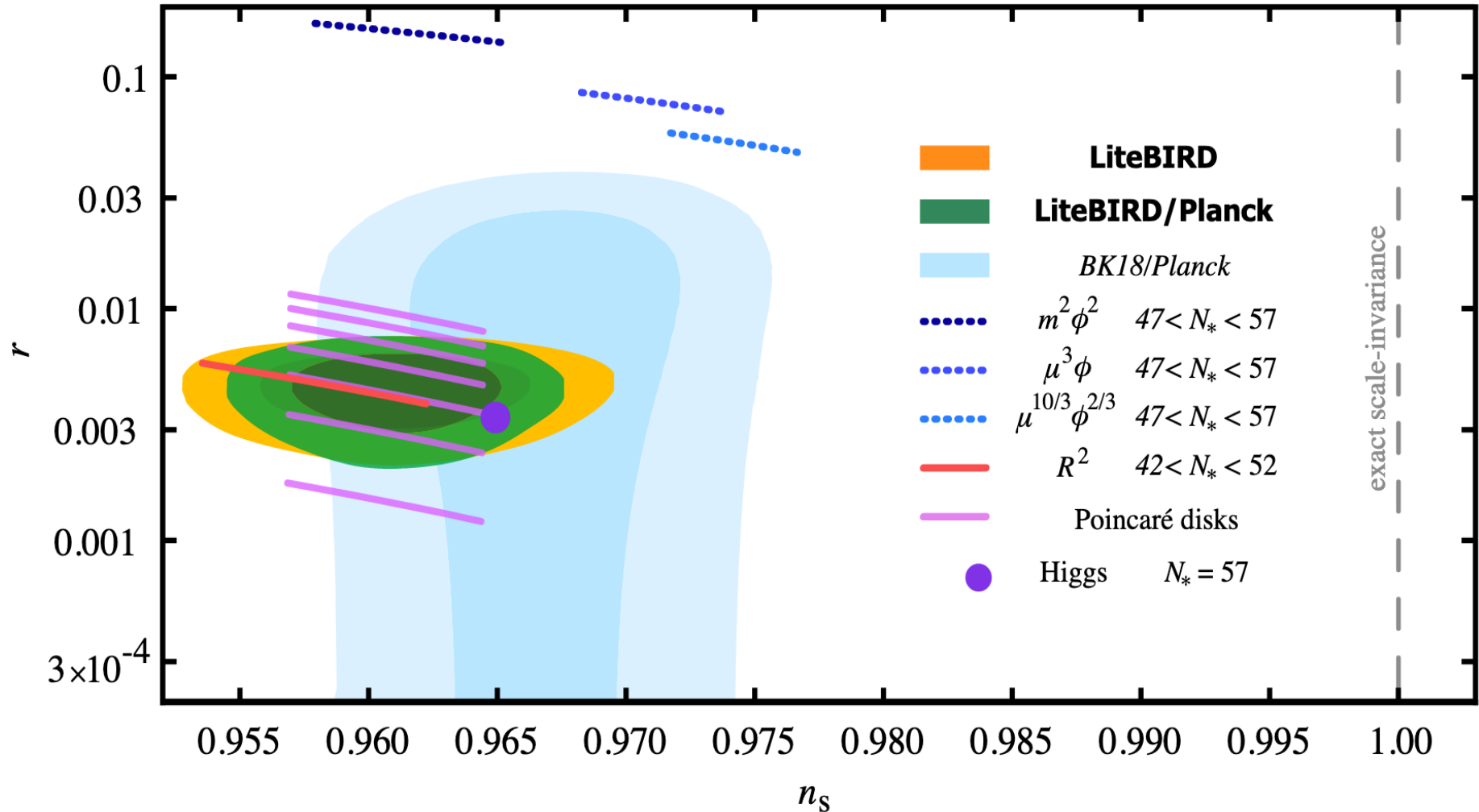


LiteBird 2202.02773



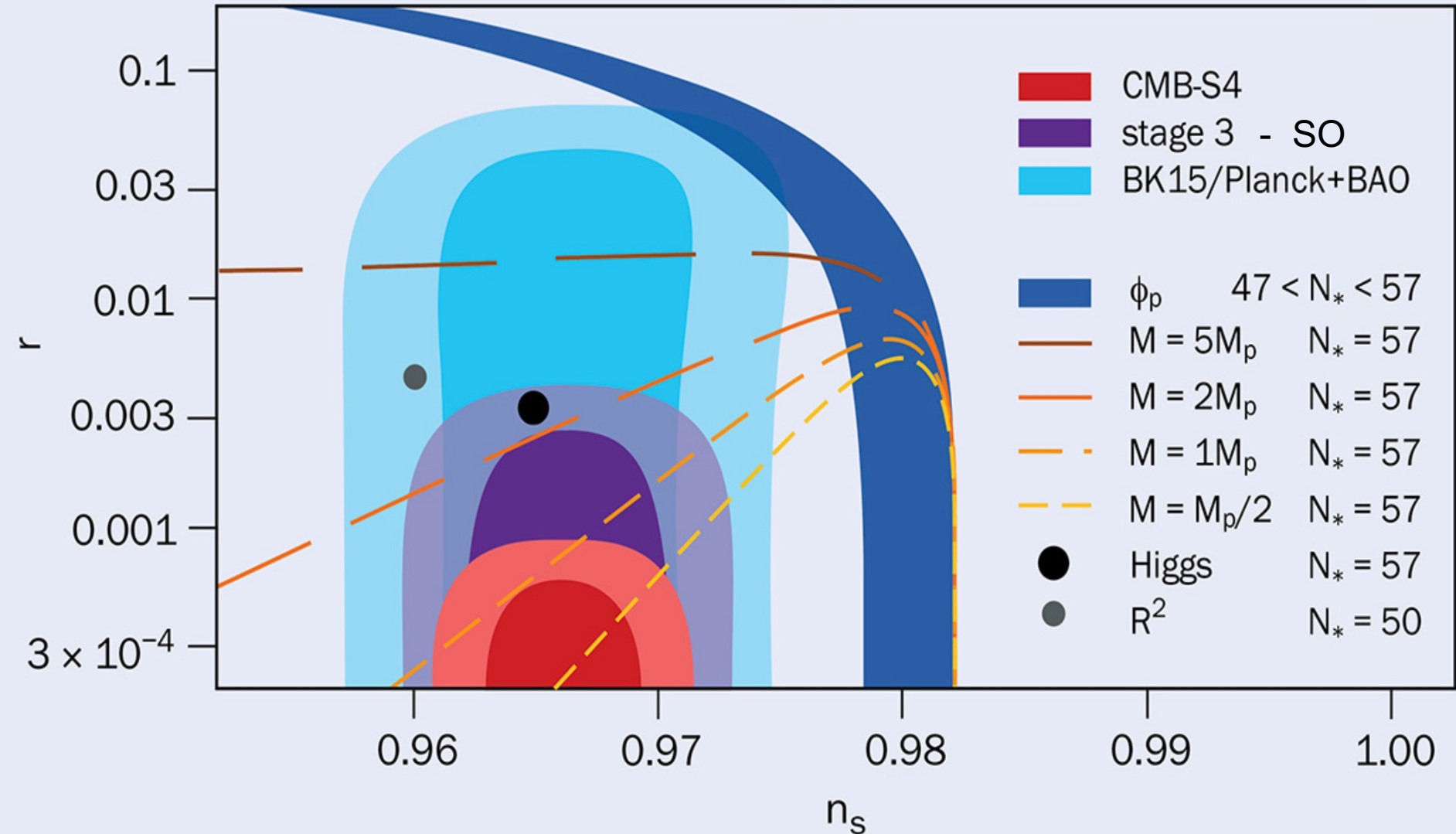
# Future experiments will search for B modes

LiteBird 2202.02773



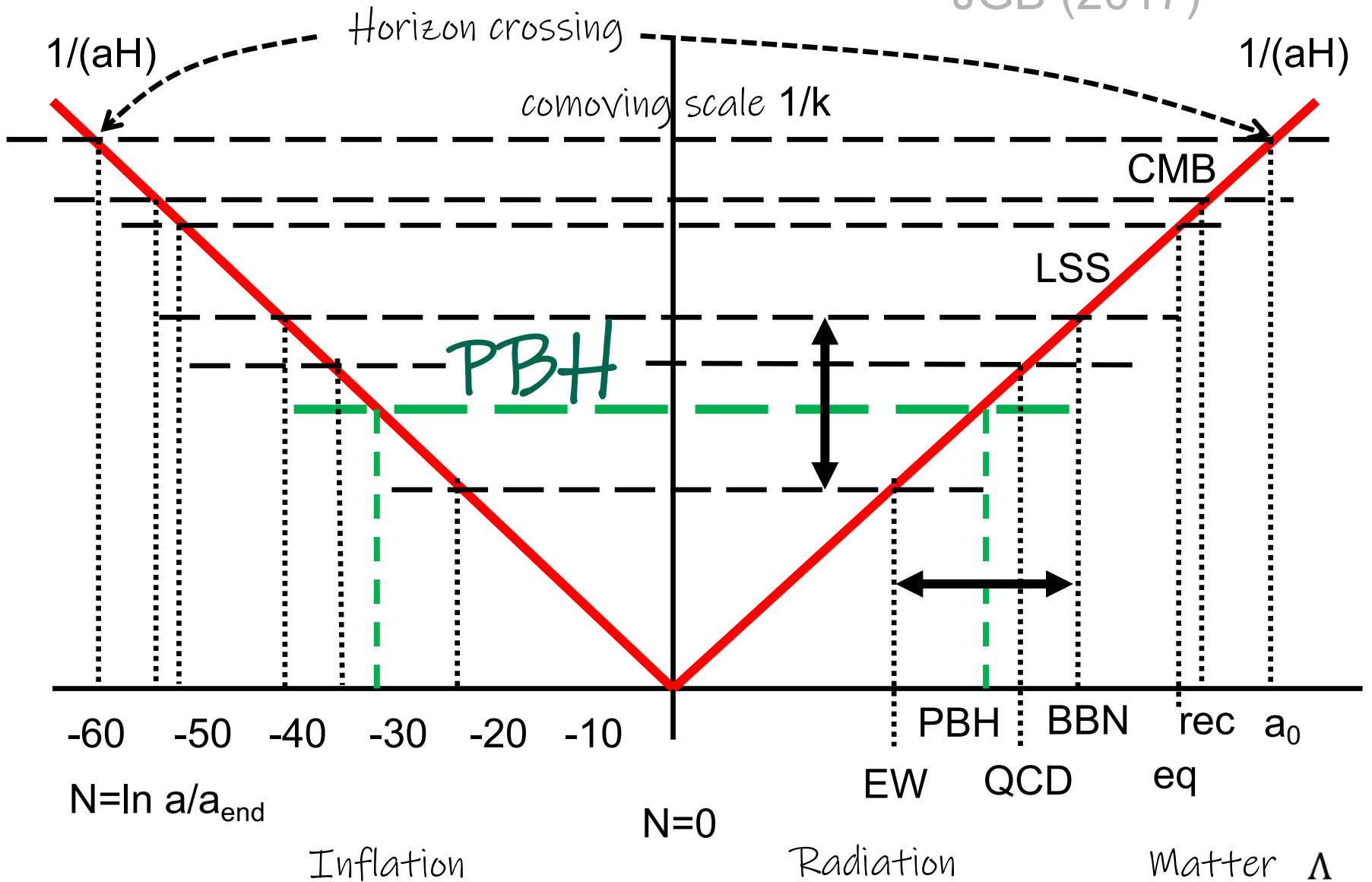
# Future experiments will search for B modes

CMB-S4 (2022)



# Inflation

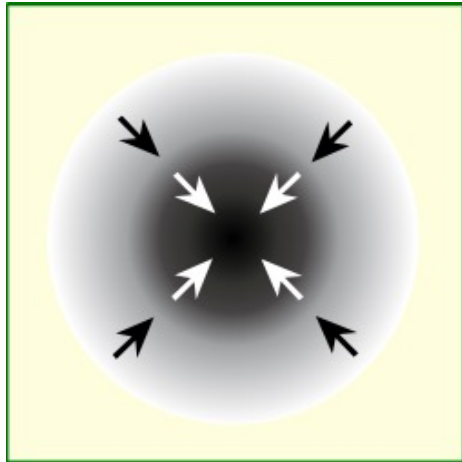
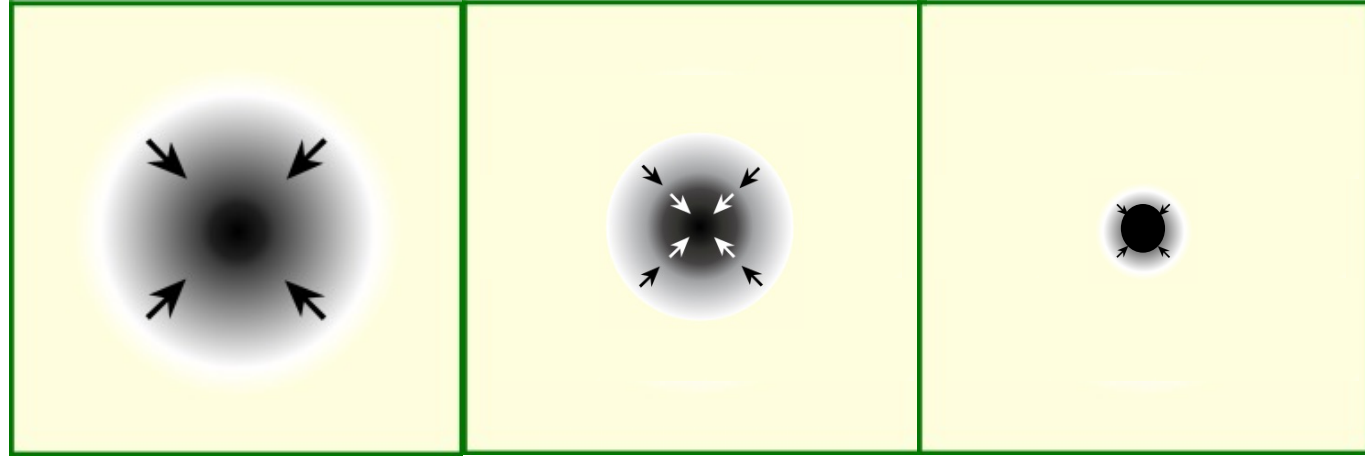
JGB (2017)



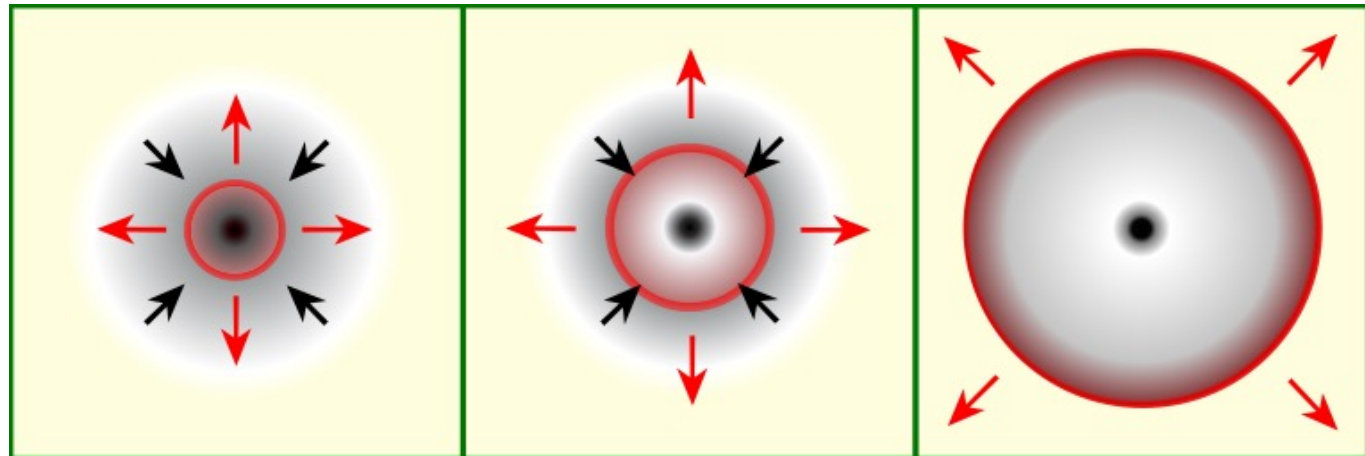


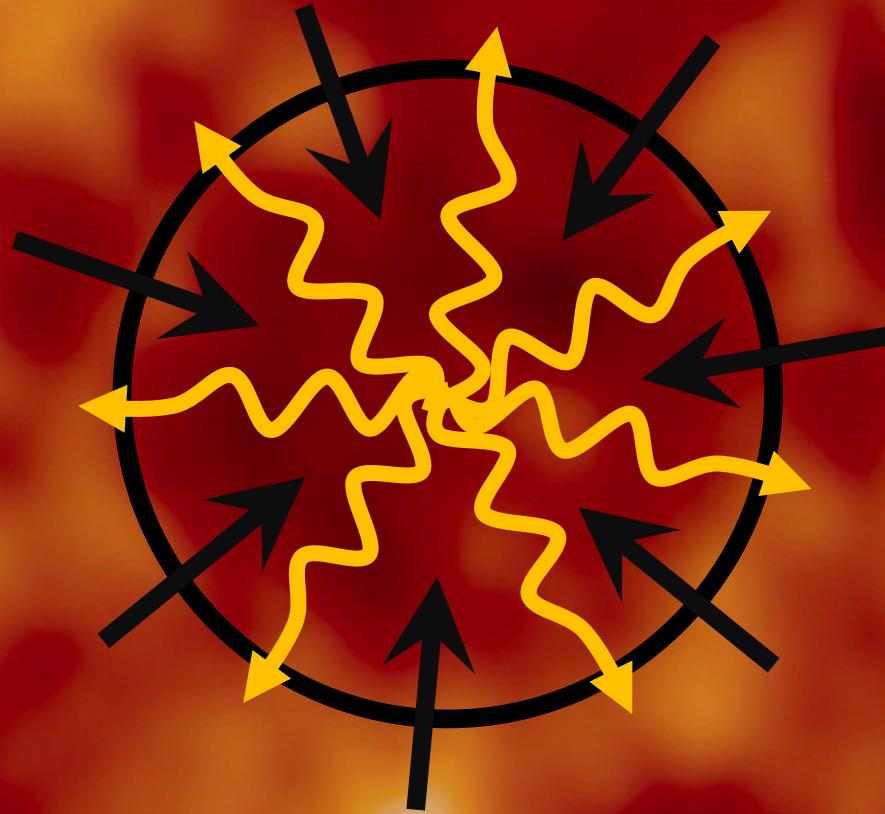
# Gravitational Collapse

Gravity wins



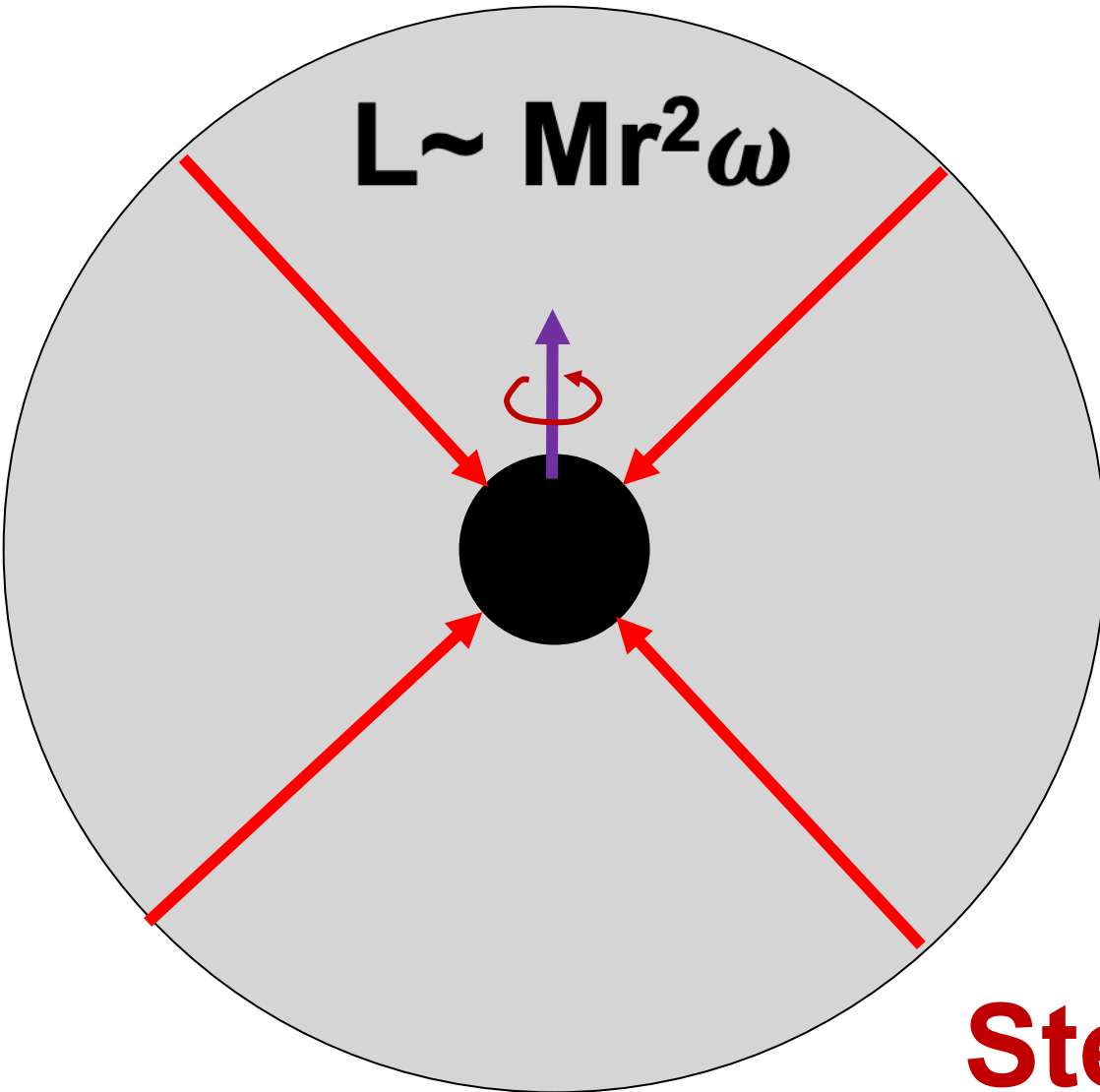
Radiation wins





Primordial plasma

PBH are  $\sim$  spinless

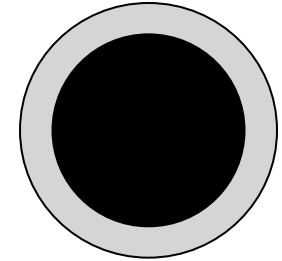


**Primordial**

**BH**

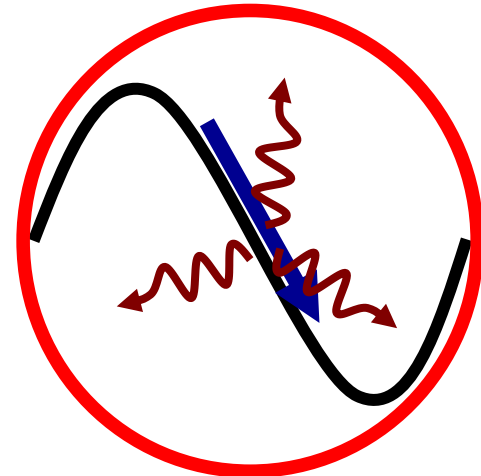
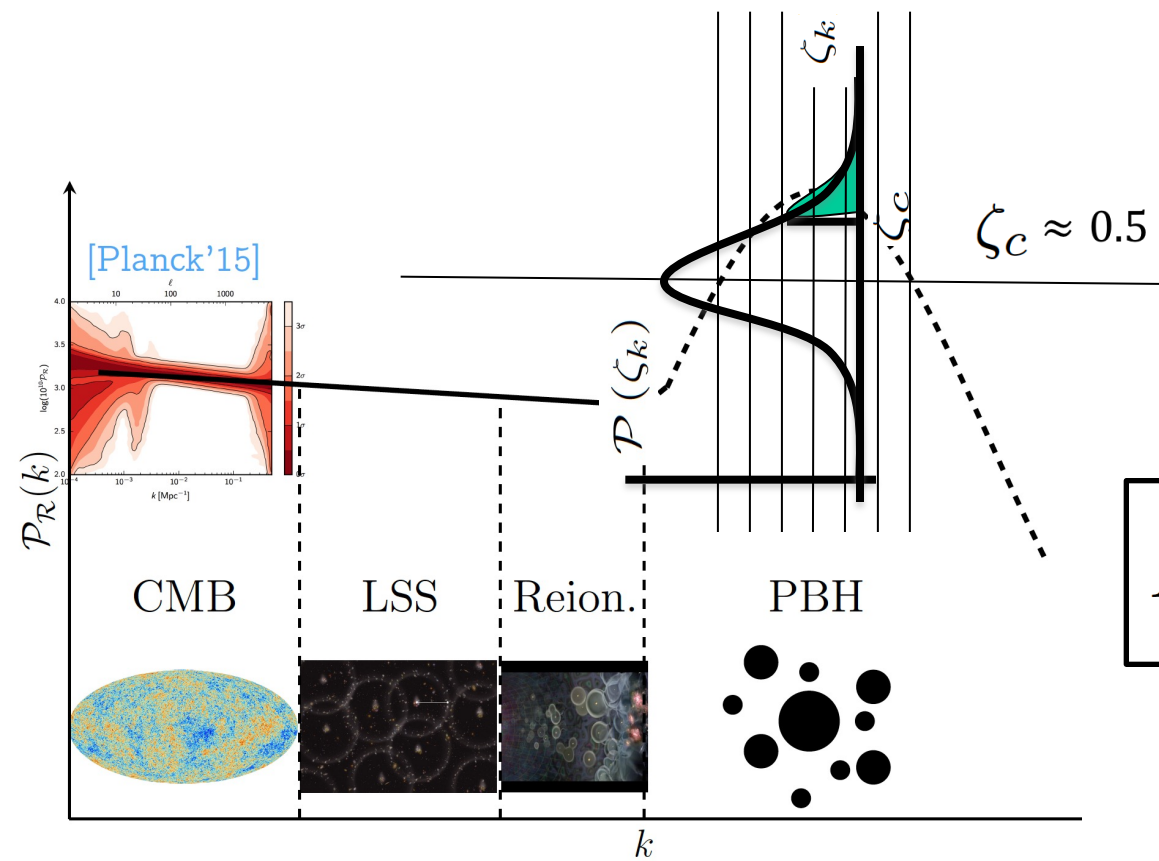
=

**Mass**



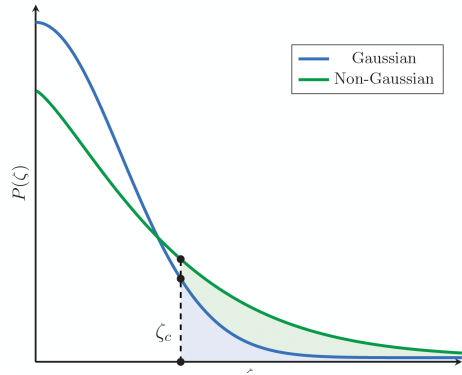
**Stellar BH**

# Gravitational Collapse of PBH



$$M_{\text{PBH}} \simeq 30 M_{\odot} e^{2(N-36)}$$

$$\beta^{\text{form}}(M_k) = \int_{\zeta_c}^{\infty} \mathcal{P}(\zeta_k) d\zeta_k$$



$$\beta(N) = \begin{cases} \text{Erfc} \left( \frac{\zeta_c}{\sqrt{2P_{\zeta}(N)}} \right), & \text{Gaussian statistics,} \\ \text{Erfc} \left( \sqrt{\frac{1}{2} + \frac{\zeta_c}{\sqrt{2P_{\zeta}(N)}}} \right), & \chi^2 \text{ statistics} \end{cases}$$

# Stochastic $\delta N$ - formalism

## Coarse-grained curvature perturbation

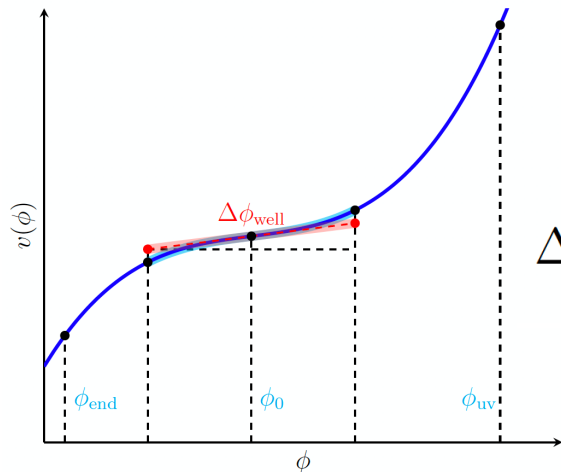
$$ds^2 = -dt^2 + a^2(t)e^{2\zeta(t, \mathbf{x})} \delta_{ij} dx^i dx^j \quad \zeta_{\text{cg}}(\mathbf{x}) = \delta N_{\text{cg}}(\mathbf{x}) = \mathcal{N}(\mathbf{x}) - \langle \mathcal{N} \rangle$$

$$\frac{1}{M_{\text{pl}}^2} \frac{d}{d\mathcal{N}} P_{\Phi}(\mathcal{N}) = \left( - \sum_i \frac{v_{\phi_i}}{v} \frac{\partial}{\partial \phi_i} + v \sum_i \frac{\partial^2}{\partial \phi_i^2} \right) \cdot P_{\Phi}(\mathcal{N}) \quad \text{Fokker-Planck Diffusion Eq.}$$

Determined by the poles of the characteristic function

$$P_{\phi}(\mathcal{N}) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{-it\mathcal{N}} \chi_{\mathcal{N}}(t, \phi) dt = \sum_n a_n(\phi) e^{-\Lambda_n \mathcal{N}}$$

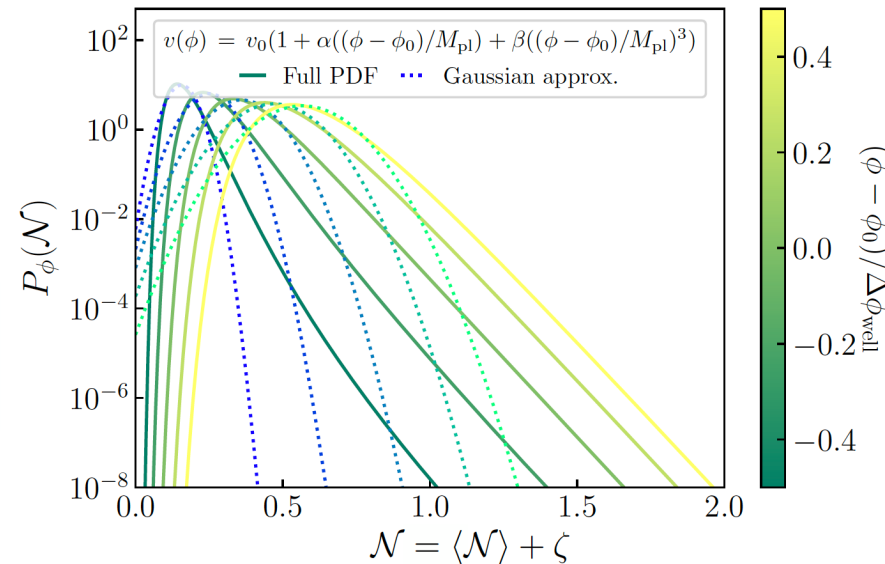
$$\chi_{\mathcal{N}}(t, \phi) = \sum_n \frac{a_n(\phi)}{\Lambda_n - it} + \text{regular func.}$$



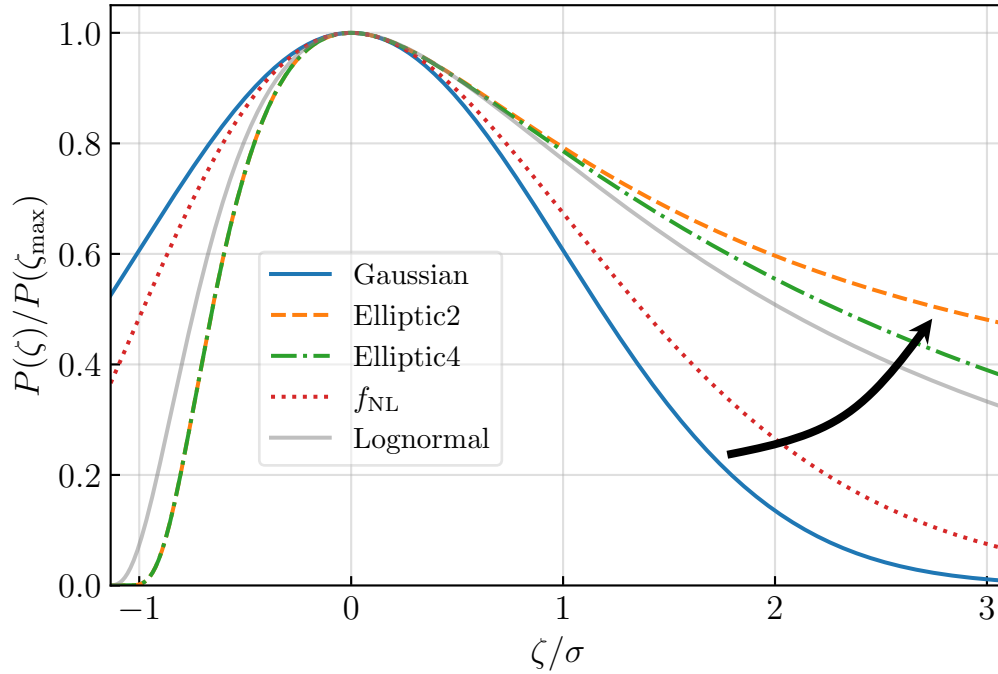
$$\alpha \gg (v_0^2 \beta)^{1/3}$$

$$\Delta\phi_{\text{well}} \simeq 2M_{\text{pl}} \sqrt{\frac{\alpha}{3\beta}}$$

Ezquiaga, JGB, Vennin (2019)



# Quantum Diffusion $\ni$ CMB & LSS



Ezquiaga, JGB, Vennin (2022)

$$P_2(\zeta_k) = -\frac{\pi}{2\mu^2} \vartheta'_2 \left( \frac{\pi\alpha_k}{2}, e^{-\frac{\pi^2}{\mu^2} \mathcal{N}_k} \right)$$

$$P_4(\zeta_k) = \frac{\pi}{2\mu^2\alpha_k} \vartheta'_4 \left( \frac{\pi\alpha_k}{2}, e^{-\frac{\pi^2}{\mu^2} \mathcal{N}_k} \right)$$

$$\zeta(x) = \zeta_G(x) + \frac{3}{5} f_{\text{NL}} \left[ \zeta_G^2(x) - \sigma_G^2(x) \right]$$

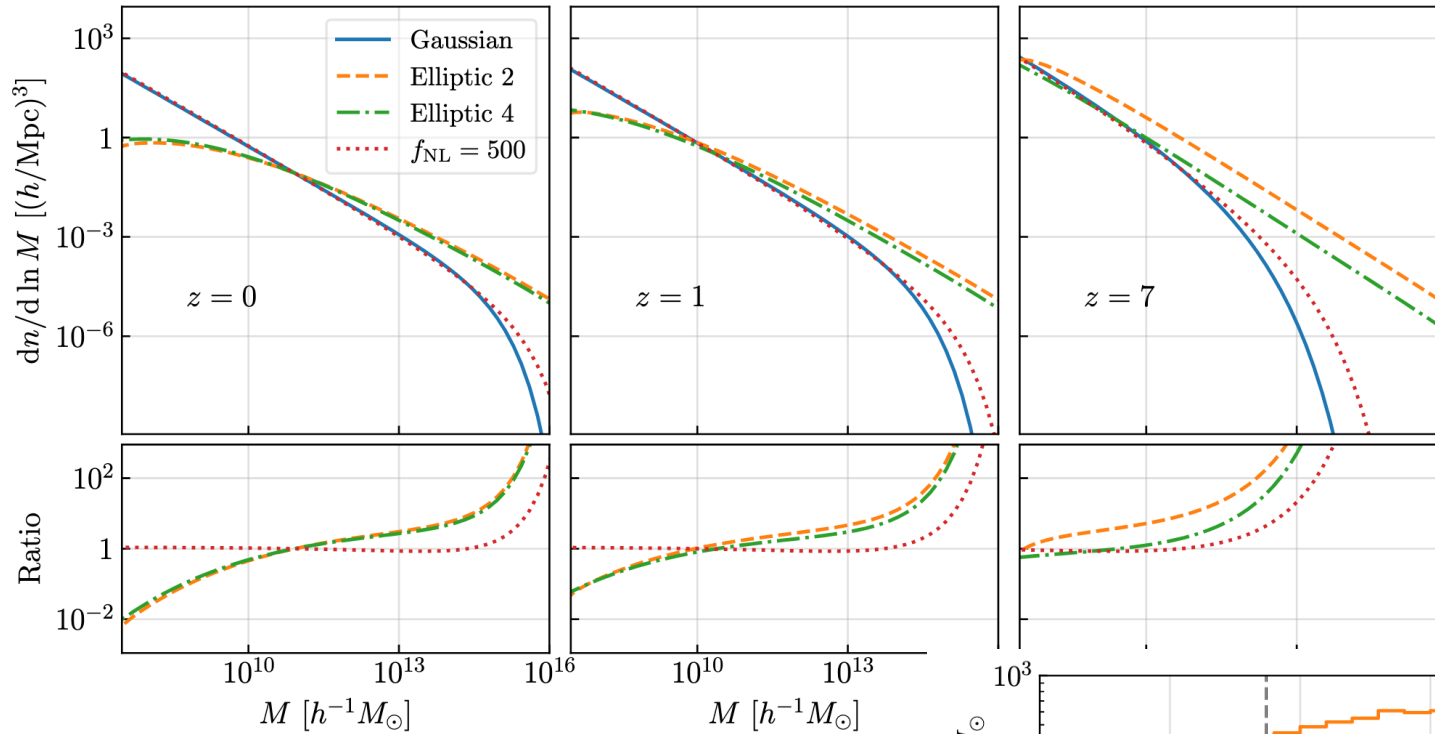
$$\text{LN}(x, \rho, \sigma) = \frac{1}{\rho\sigma\sqrt{2\pi}} \exp \left[ -\frac{\ln(x/\rho)^2}{2\sigma^2} - \frac{\sigma^2}{2} \right]$$

$$P_{\text{NL}}(\zeta) = \frac{1}{\sqrt{2\pi\sigma_G^2\Delta}} \left[ e^{-\frac{25(\sqrt{\Delta}-1)^2}{72f_{\text{NL}}^2\sigma_G^2}} + e^{-\frac{25(\sqrt{\Delta}+1)^2}{72f_{\text{NL}}^2\sigma_G^2}} \right]$$

$$\text{G}(x, \rho, \sigma_G) = \frac{1}{\sigma_G\sqrt{2\pi}} \exp \left[ -\frac{(x-\rho)^2}{2\sigma_G^2} \right]$$

where  $\Delta(\zeta) = 1 + \frac{12}{5} f_{\text{NL}}\zeta + \frac{36}{25} f_{\text{NL}}^2\sigma_G^2$ .

# Quantum Diffusion $\ni$ CMB & LSS

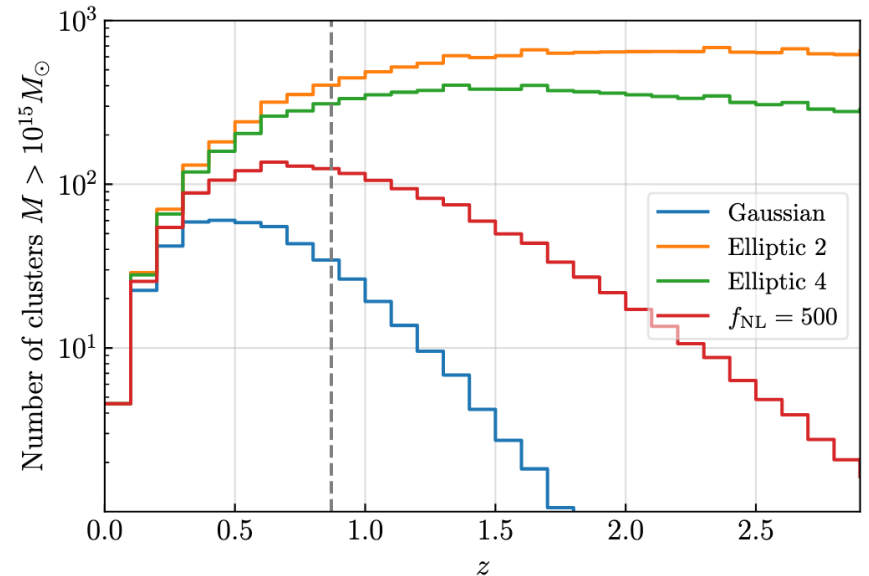


Halo  
Mass  
Function

Ezquiaga, JGB, Vennin (2022)

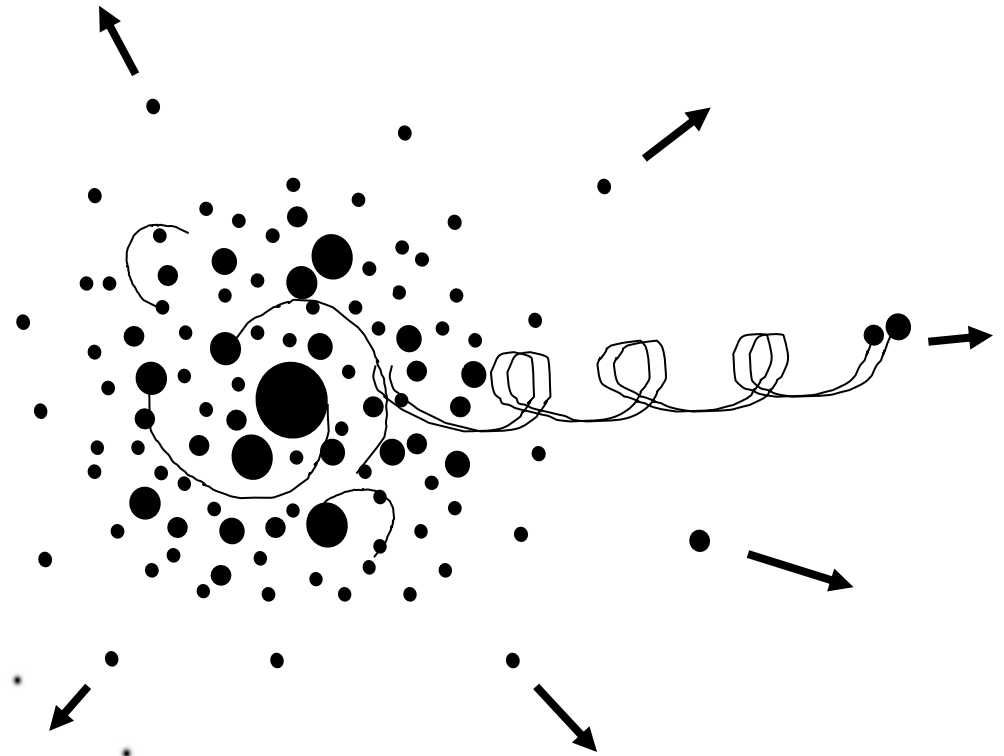
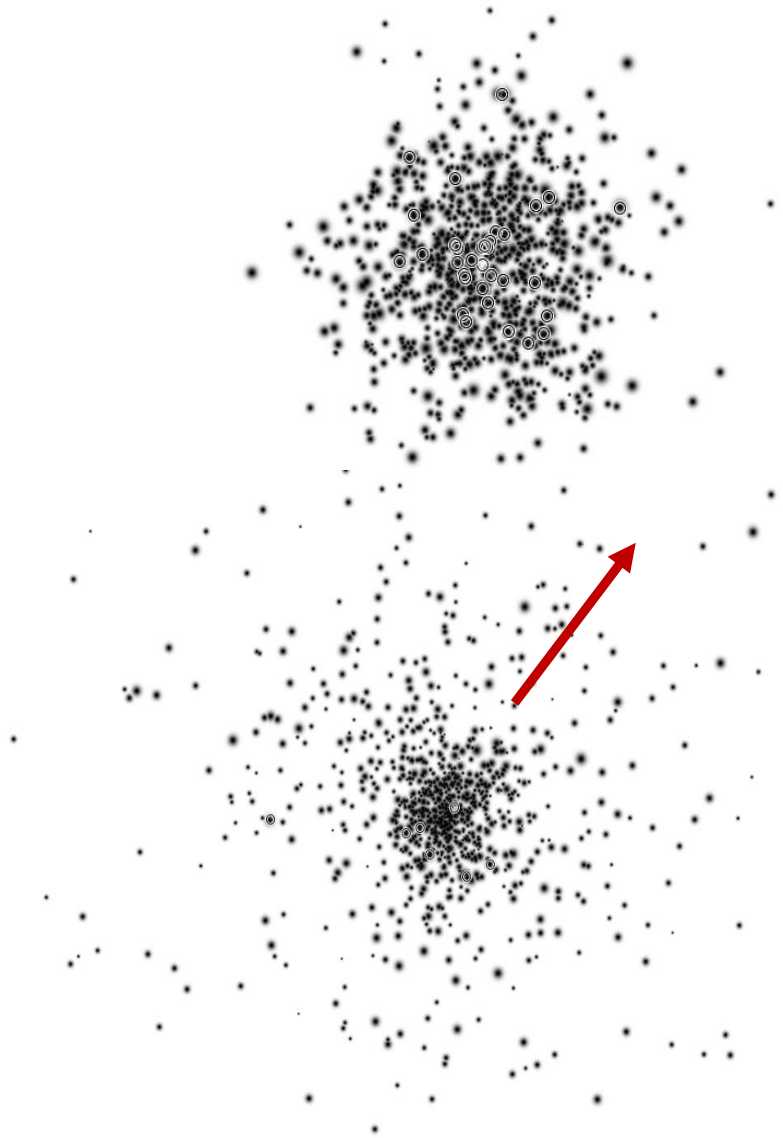
El Gordo

$M \sim 3 \cdot 10^{15} M_\odot$  at  $z = 0.87$



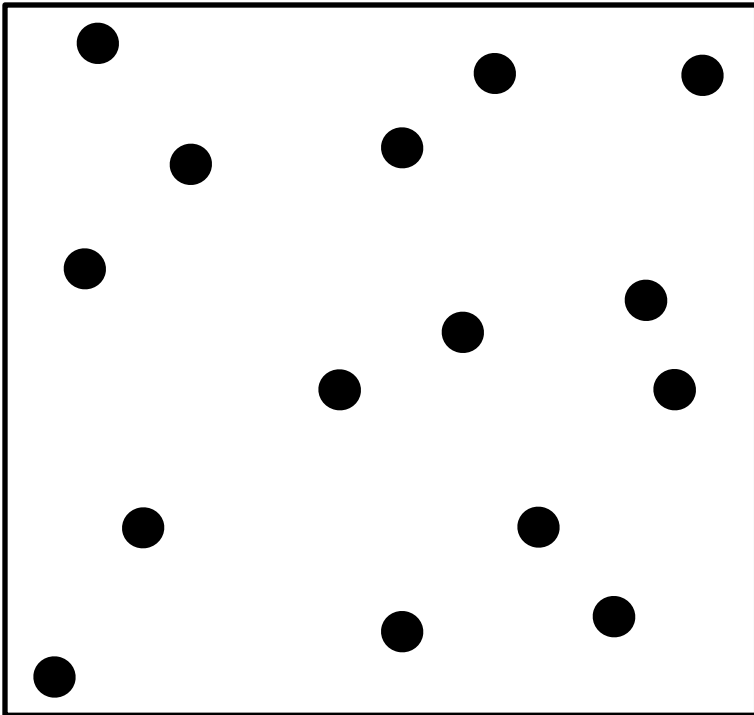
# PBH clusters

Trashorras , JGB, Nesseris (2020)

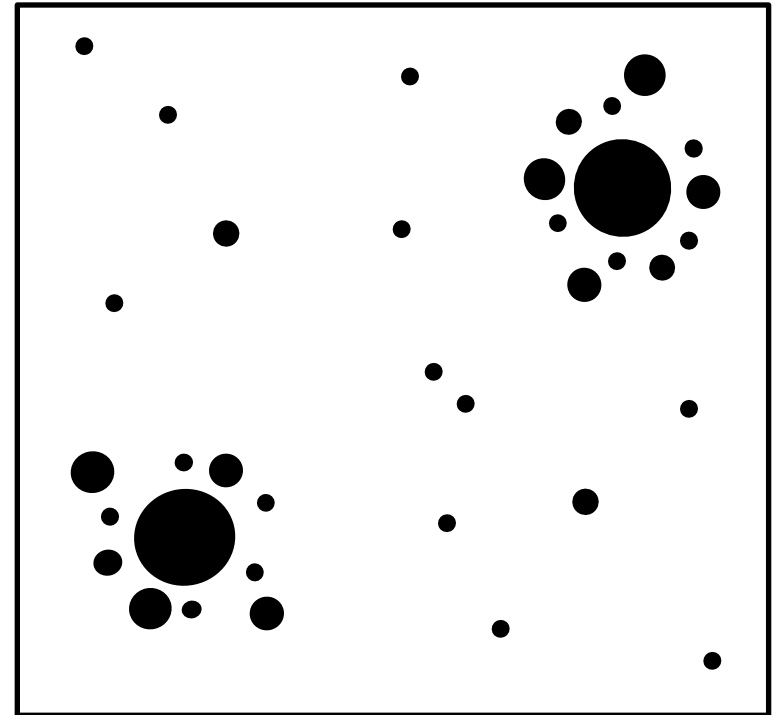




# Spatial Distribution PBH



- Monochromatic
- Uniformly distributed



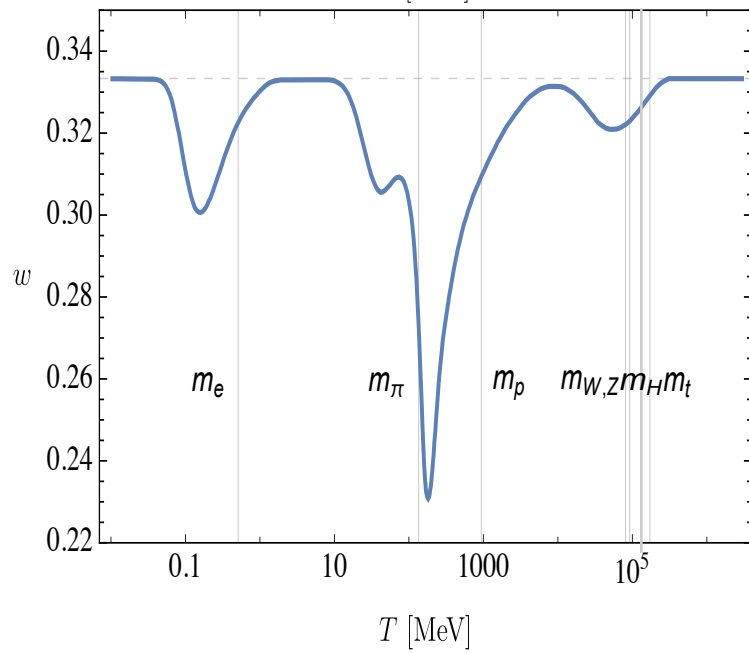
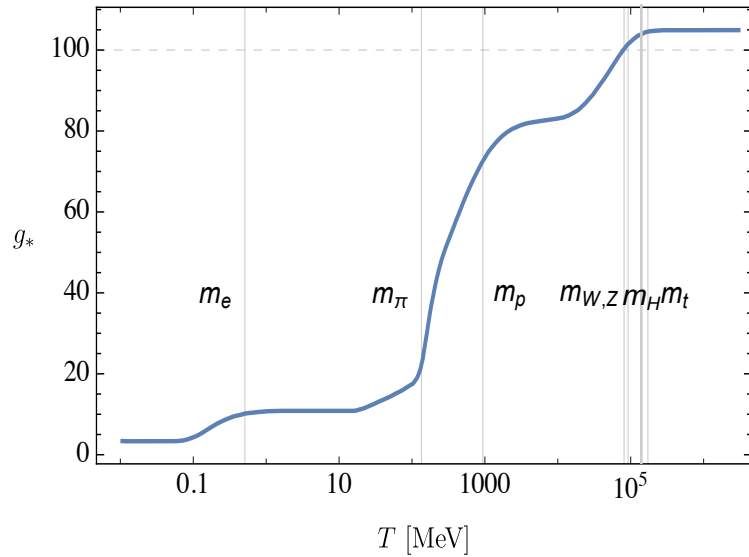
- Broad range of masses
- PBH in clusters



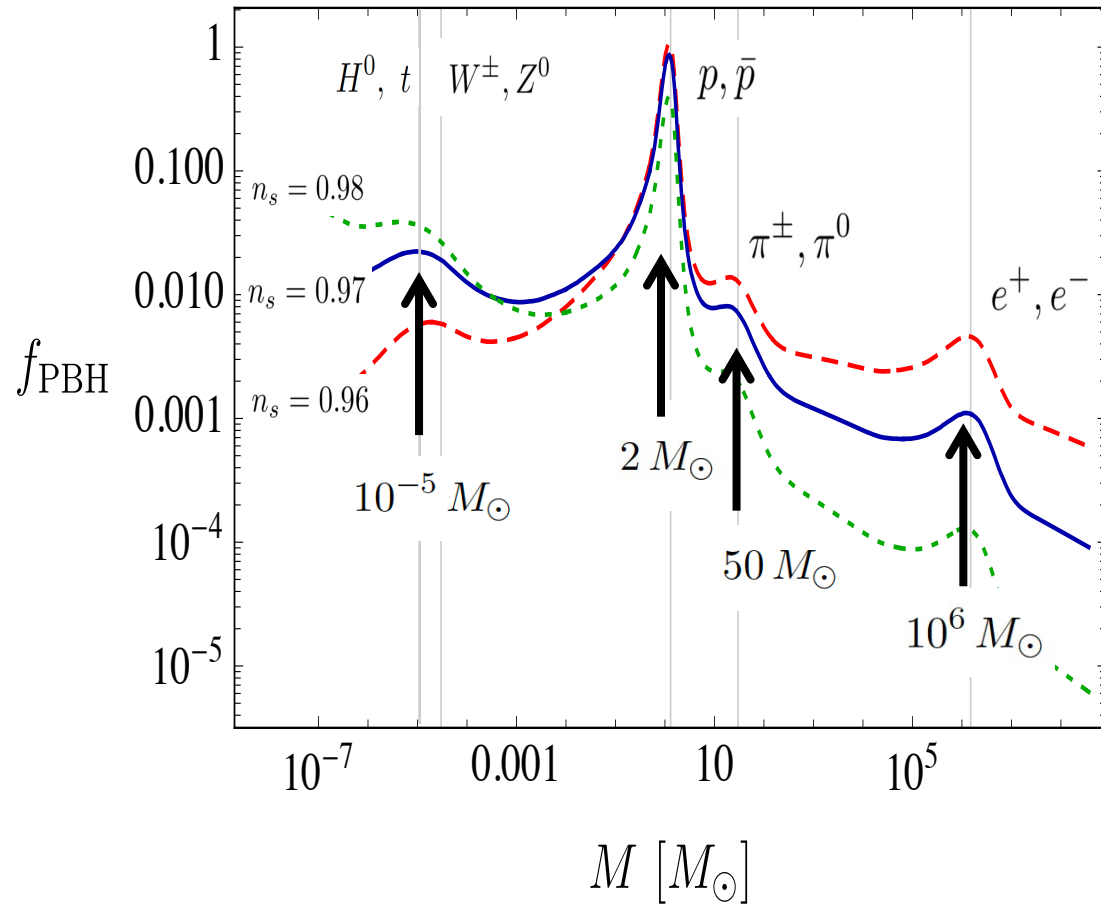
JGB (2017)

# Thermal history of the universe

Carr, Clesse, JGB, Kühnel (2019)

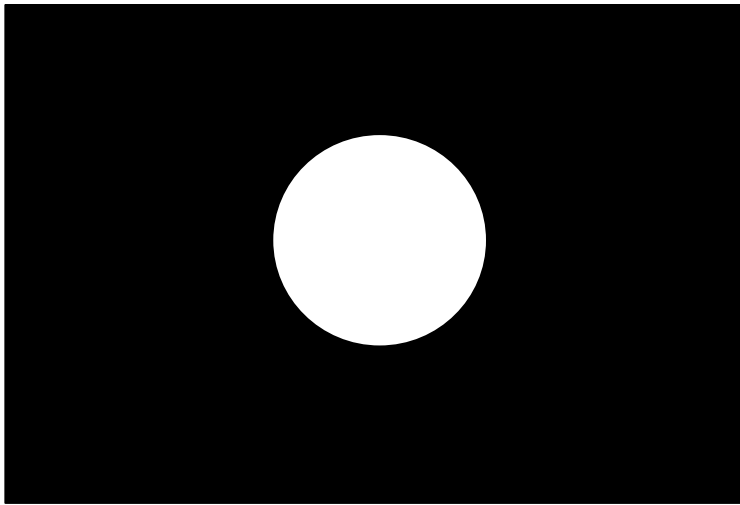


## PBH mass spectrum

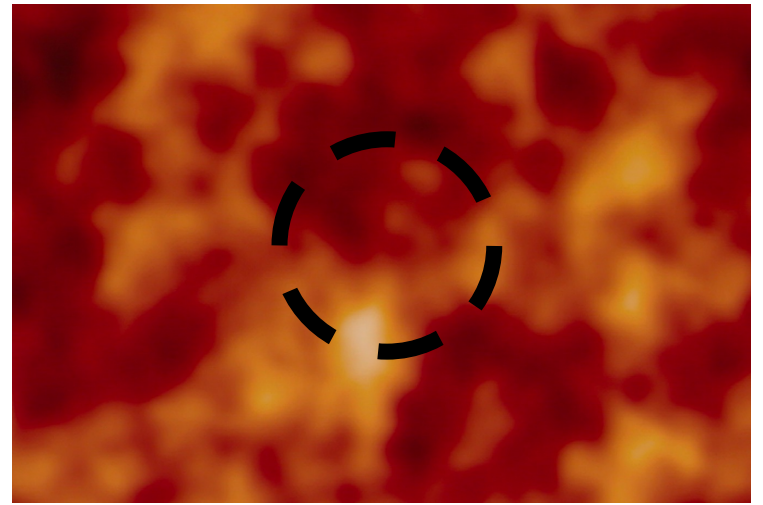


# Origin of stellar & PBH masses

Chandrasekhar mass  
(Pauli Excl.)



Mass within horizon  
 $\propto$  QCD (Causality)



$$M_{\text{Chandra}} \simeq \frac{2}{\mu^2} \sqrt{\frac{3\pi}{4}} \frac{M_{\text{P}}^3}{m_{\text{p}}^2} \simeq 1.4 M_{\odot}$$

$$M_{\text{PBH}} \simeq \frac{4\pi}{3} \rho_{\text{rad}} d_{\text{H}}^3 \simeq \frac{3\sqrt{5}}{4\pi^{3/2} \sqrt{g_*}} \frac{M_{\text{P}}^3}{\Lambda_{\text{QCD}}^2} \simeq 2 M_{\odot}$$

# Electroweak baryogenesis $\simeq$ QCD

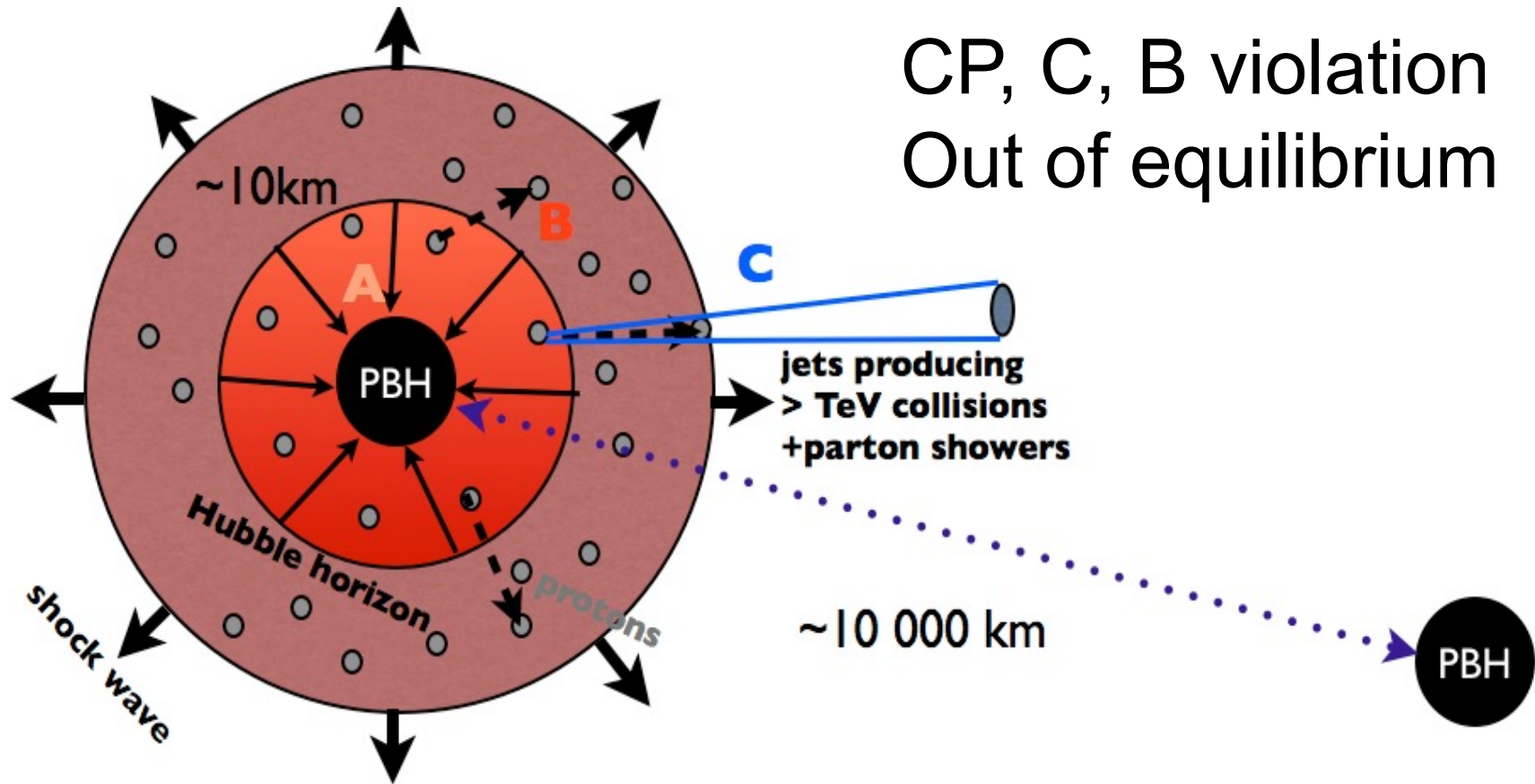
## “Primordial supernova”

JGB, Carr, Clesse (2019)

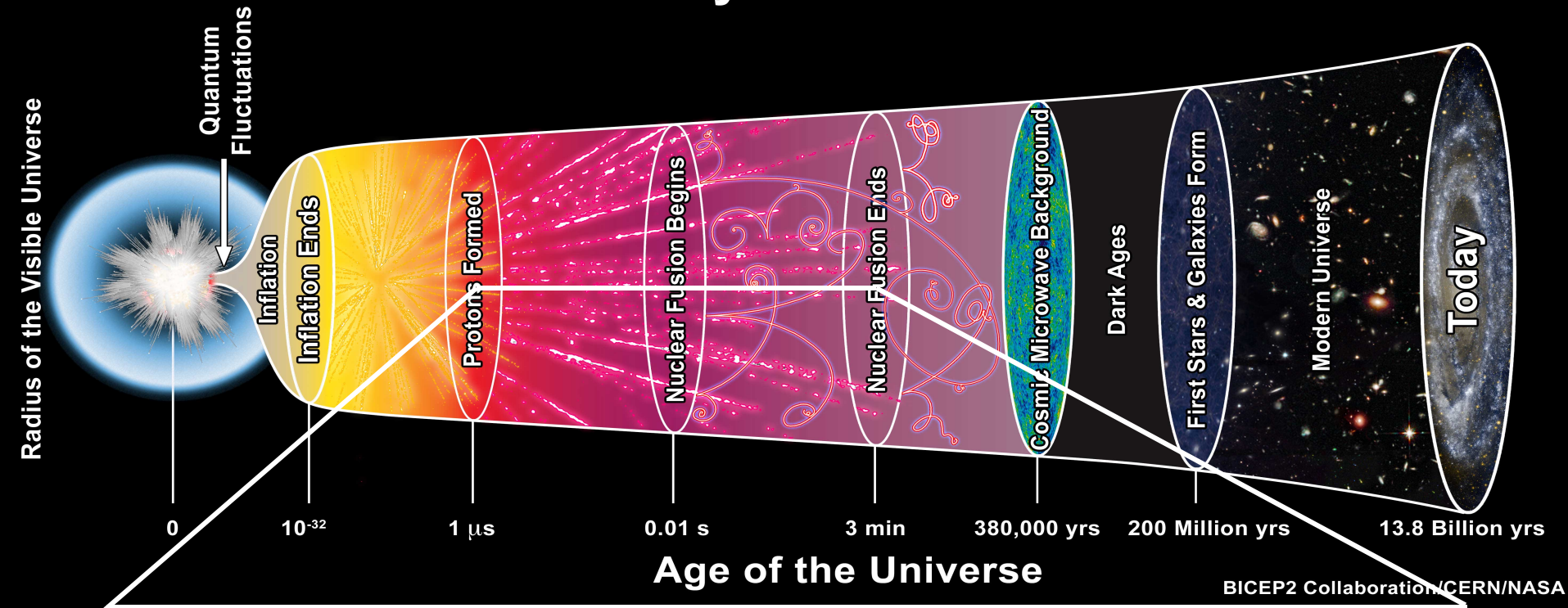
Sakharov conditions:

CP, C, B violation

Out of equilibrium



# History of the Universe



JGB  
(2019)

PBH=DM  
collapse

Baryogenesis

Nucleosynthesis

quark-hadron  
transition

hot-spot  
EWB

baryon  
dilution

light  
elements

200 MeV

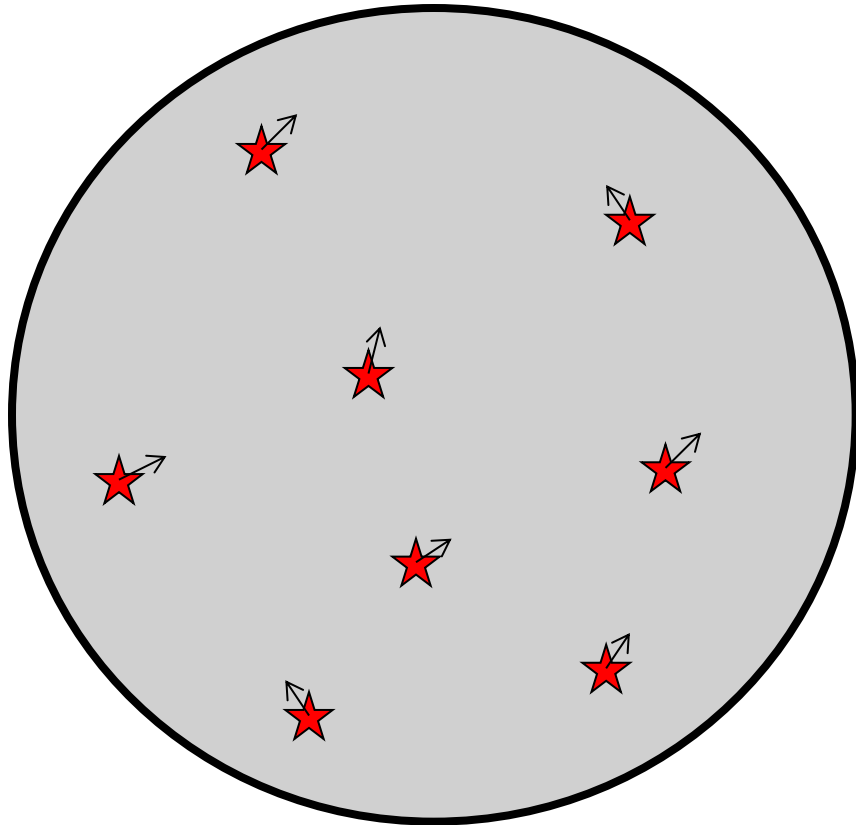
100 MeV

10 MeV

1 MeV

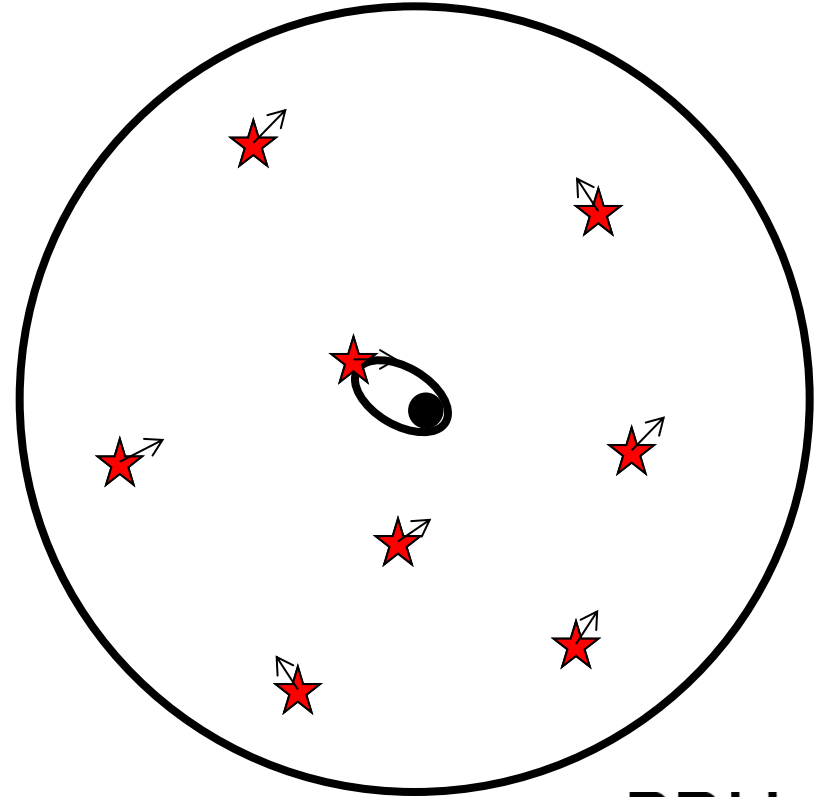
# Spatial distribution of DM

Thomson model



PDM

Rutherford model



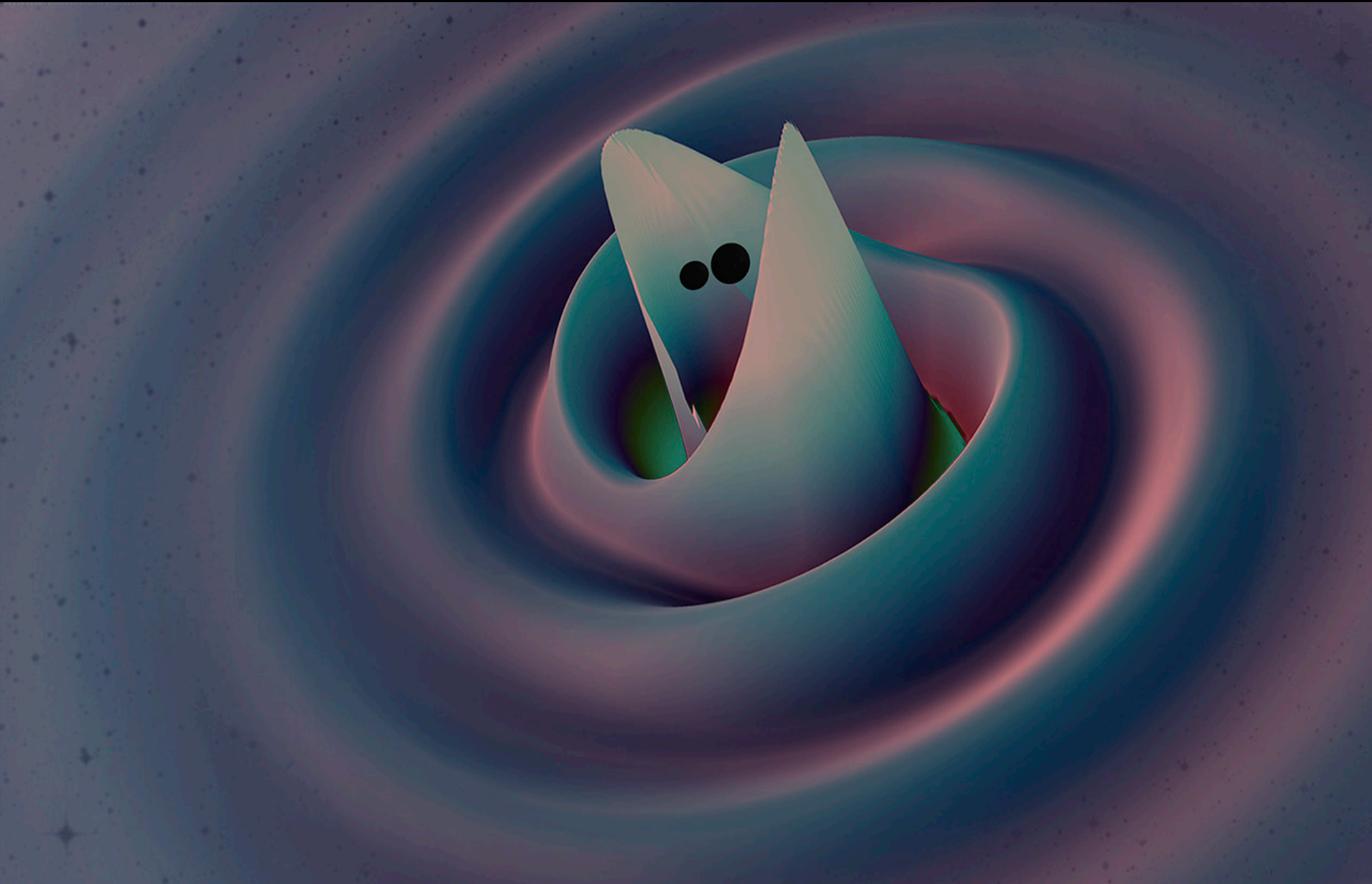
PBH

JGB (2017)

# Massive PBH = seeds of structure

- Massive primordial black holes with  $10^{-5} M_{\odot} < M_{\text{PBH}} < 10^6 M_{\odot}$ , which **cluster and merge** and could resolve some of the most acute problems of  $\Lambda$ CDM paradigm.
- $\Lambda$ CDM N-body simulations never reach the  $10^3 M_{\odot}$  particle resolution, so for them **PBH DM is as good as Particle DM.**
- PBH DM paradigm naturally incorporates all properties of collisionless CDM scenario on large scales but **differs on small scales.**

# GW emission



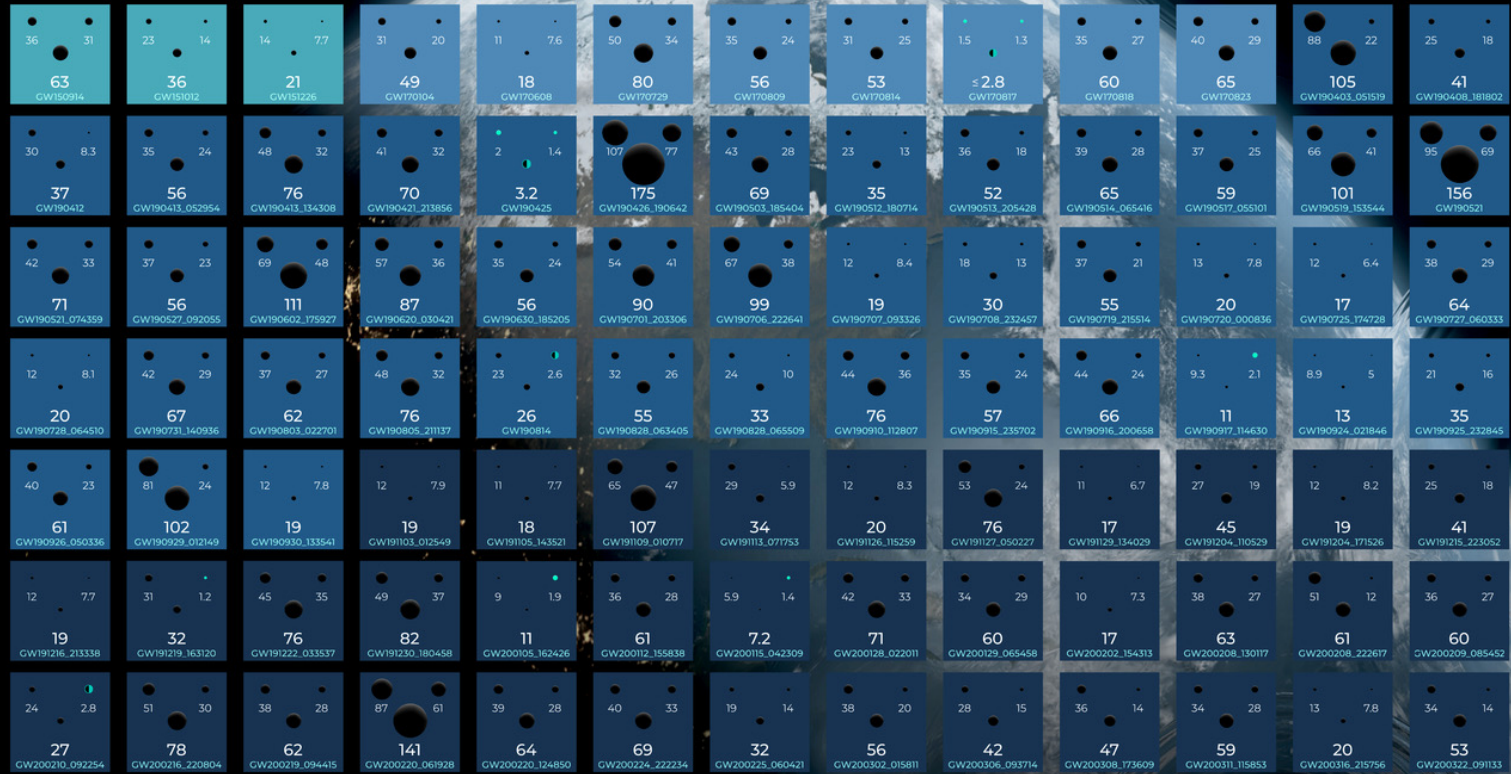


# GWTC-3

OBSERVING RUN  
01  
2015 - 2016

02  
2016 - 2017

03a+b  
2019 - 2020



**KEY**

- BLACK HOLE
- NEUTRON STAR (SHOWN AT 1/10 SCALE)
- UNCERTAIN OBJECT
- PRIMARY MASS
- SECONDARY MASS
- FINAL MASS
- DATE (TIME)

UNITS ARE SOLAR MASSES  
1 SOLAR MASS =  $1.989 \times 10^{30}$  kg

Note that the mass estimates shown here do not include uncertainties, which is why the final mass is typically higher than the sum of the primary and secondary masses. In actuality the final mass is lower.

The events listed here pass one of two thresholds for detection. They either have a probability of being astrophysical of at least 50% or they pass a false alarm rate threshold of less than 1 per 3 years.

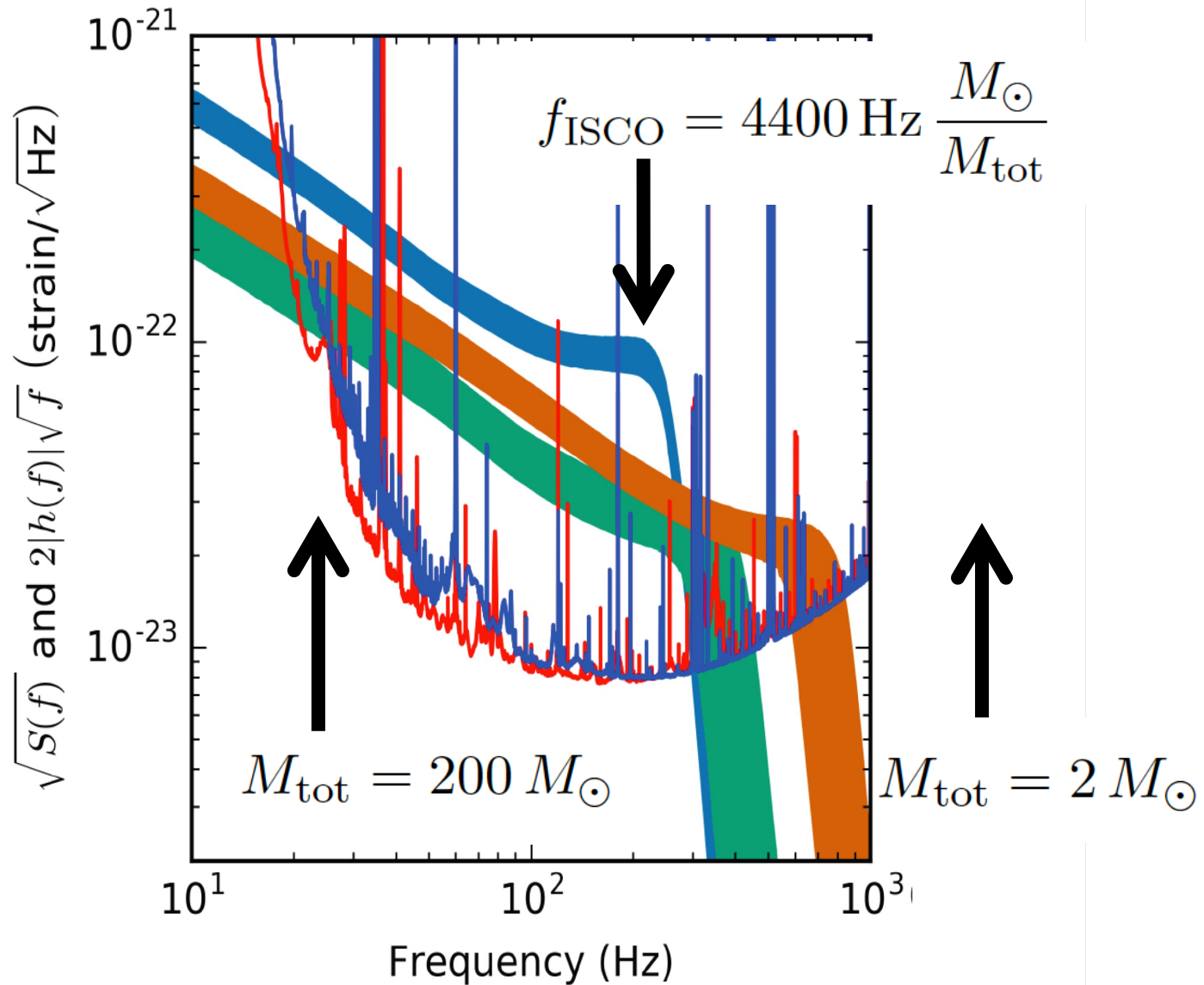
## GRAVITATIONAL WAVE MERGER DETECTIONS SINCE 2015



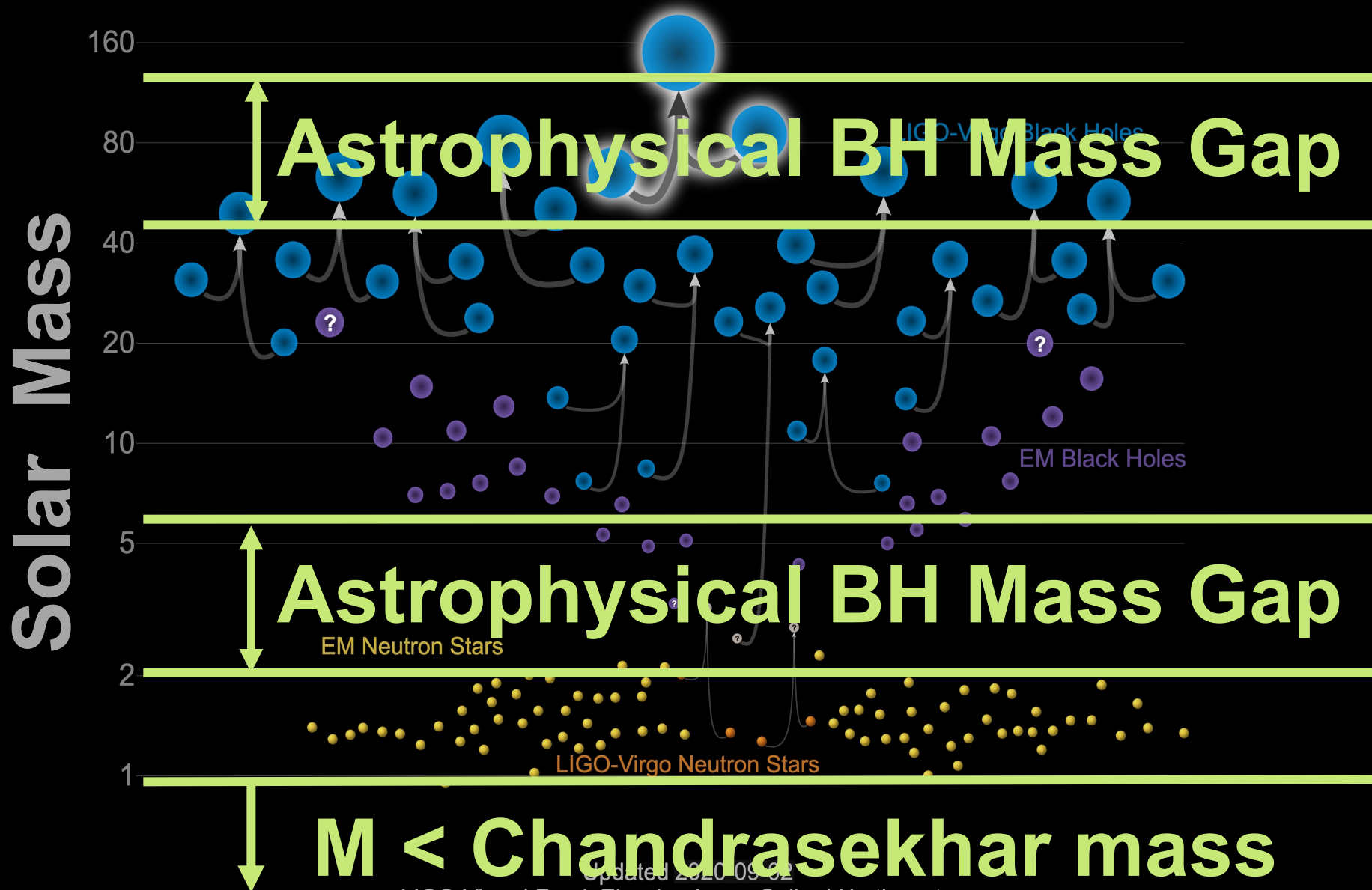
LIGO Center of Excellence for Gravitational Wave Discovery



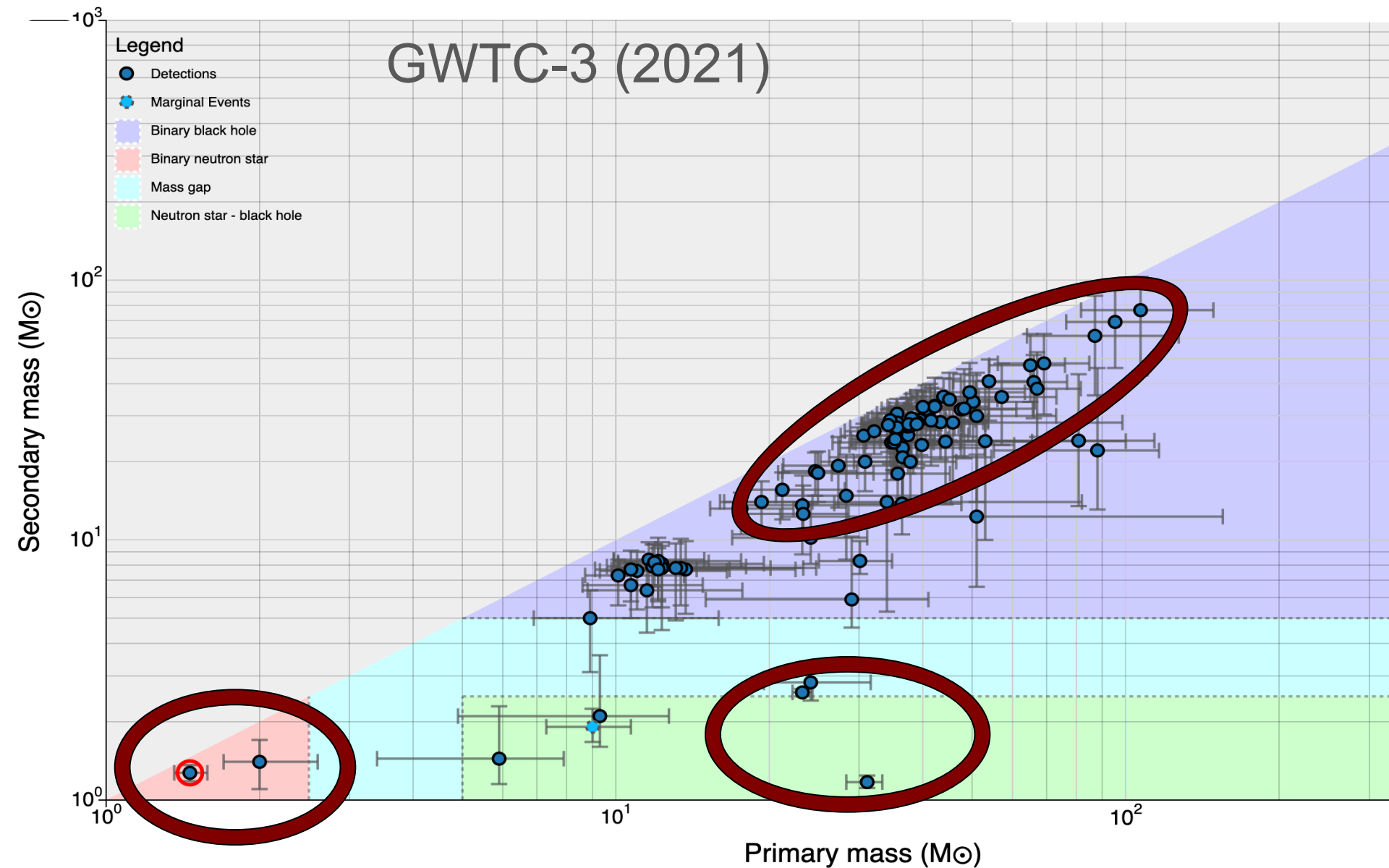
# LVC BBH events



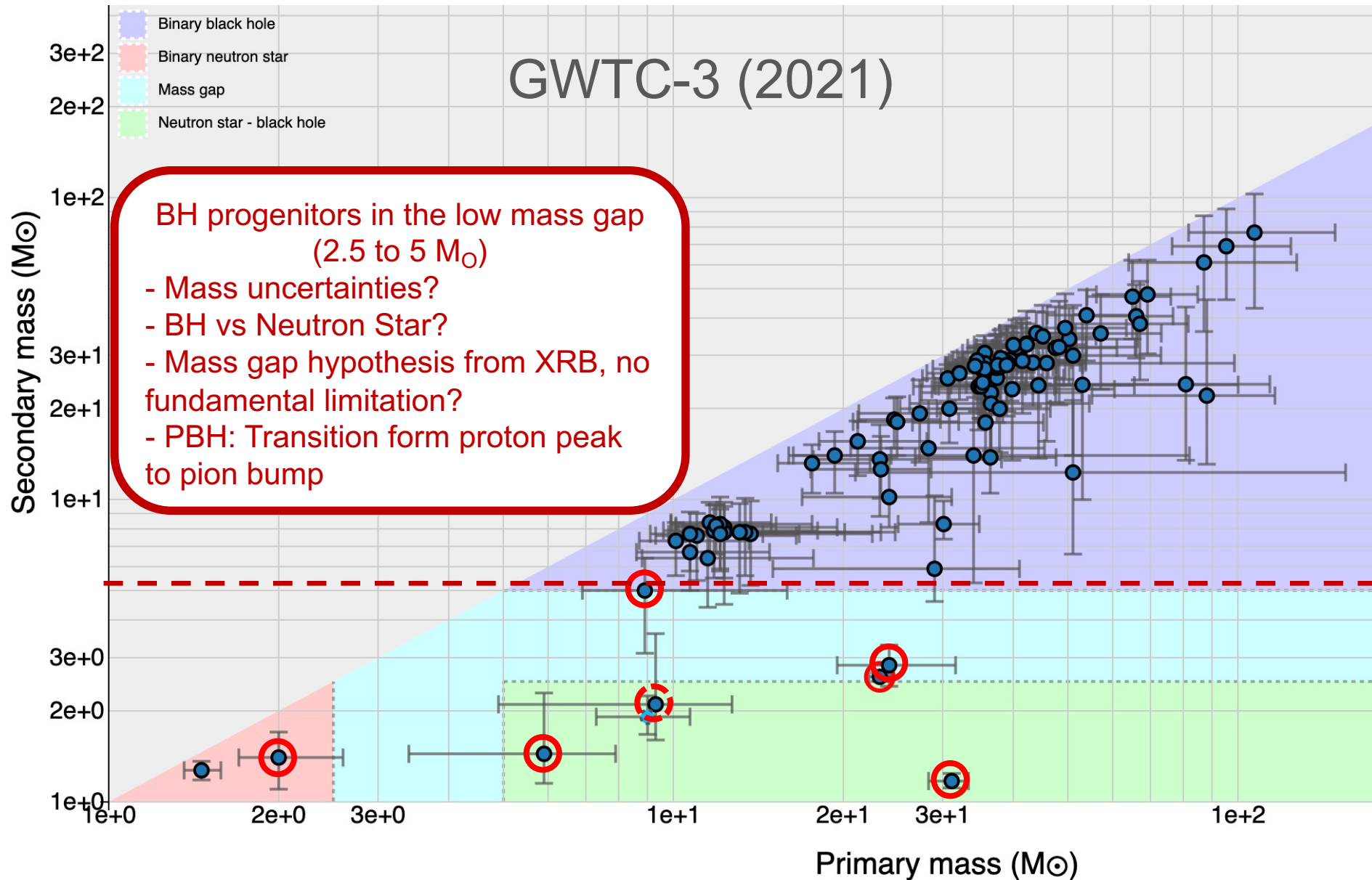
# Black Holes and Neutron Stars



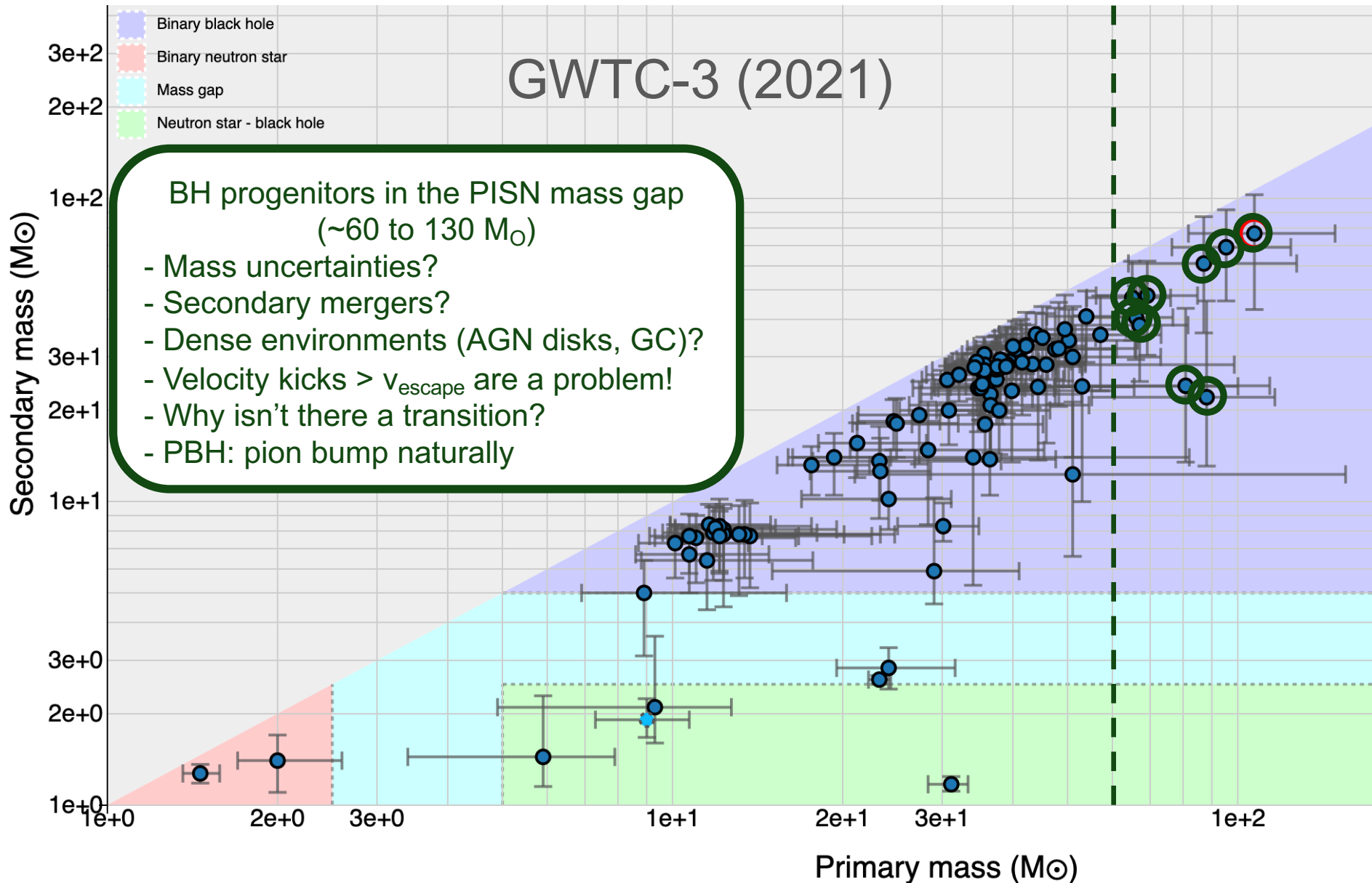
# Primary and secondary masses



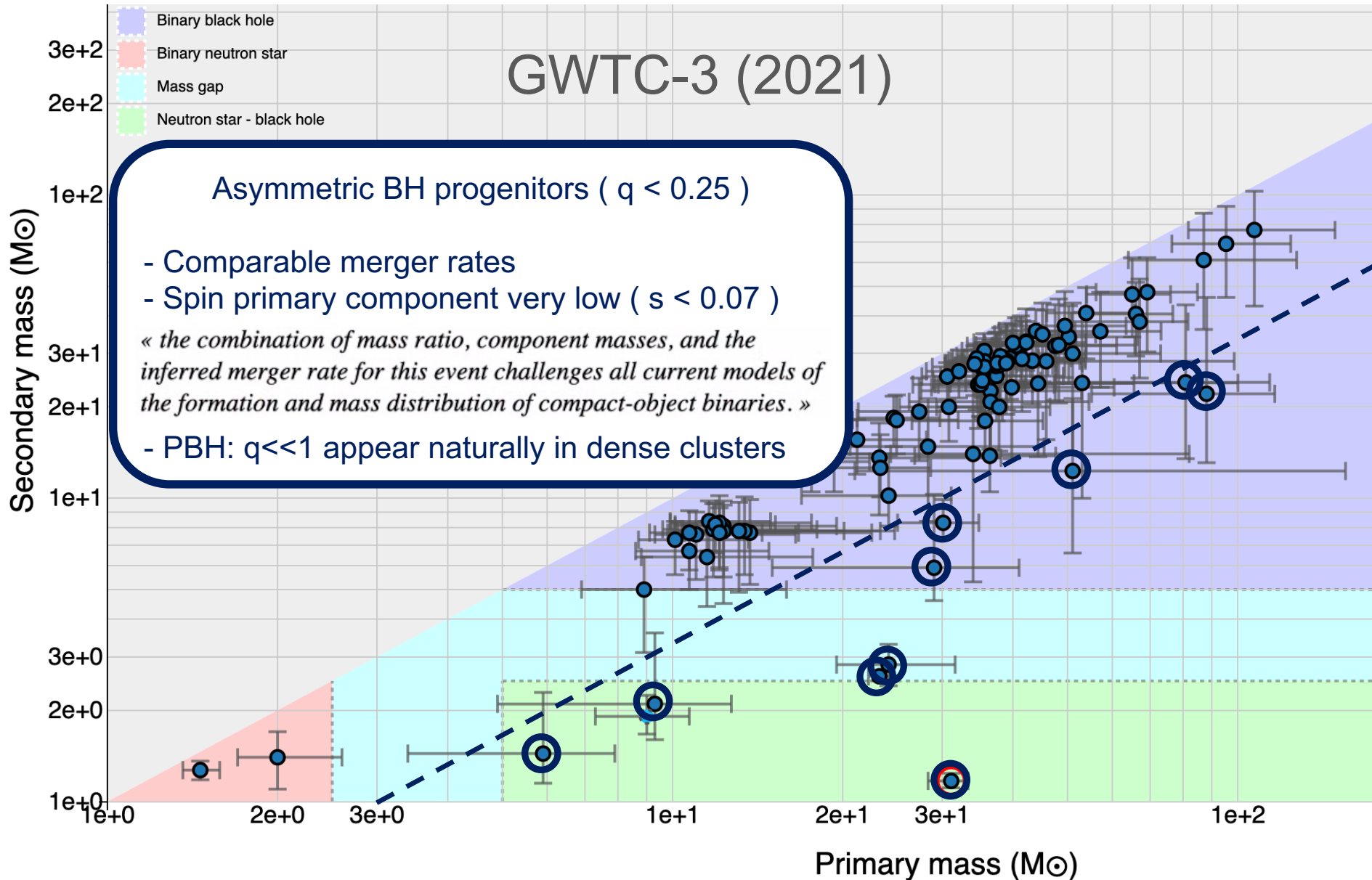
# Are LIGO/Virgo BH Primordial?



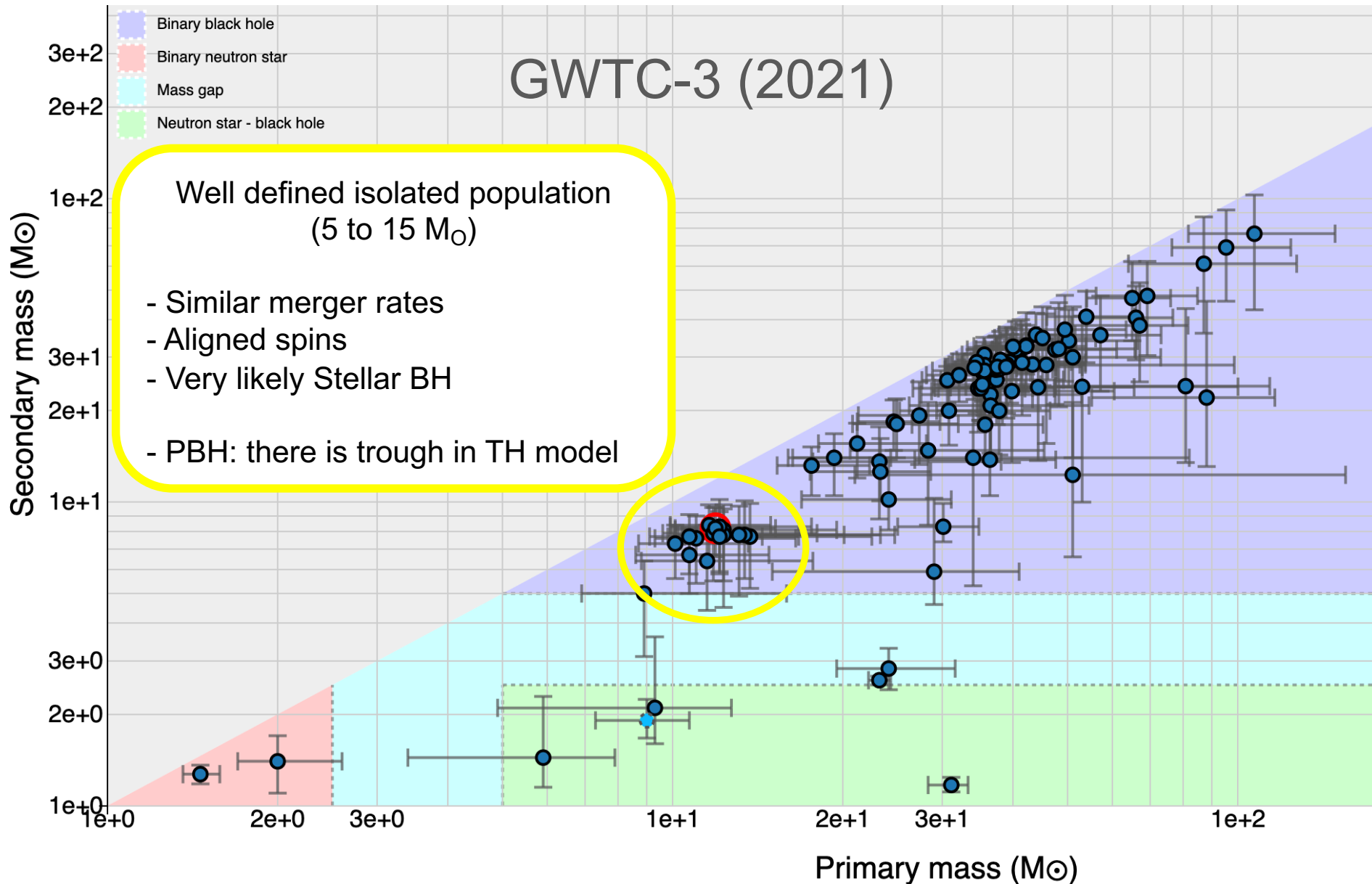
# Are LIGO/Virgo BH Primordial?



# Are LIGO/Virgo BH Primordial?



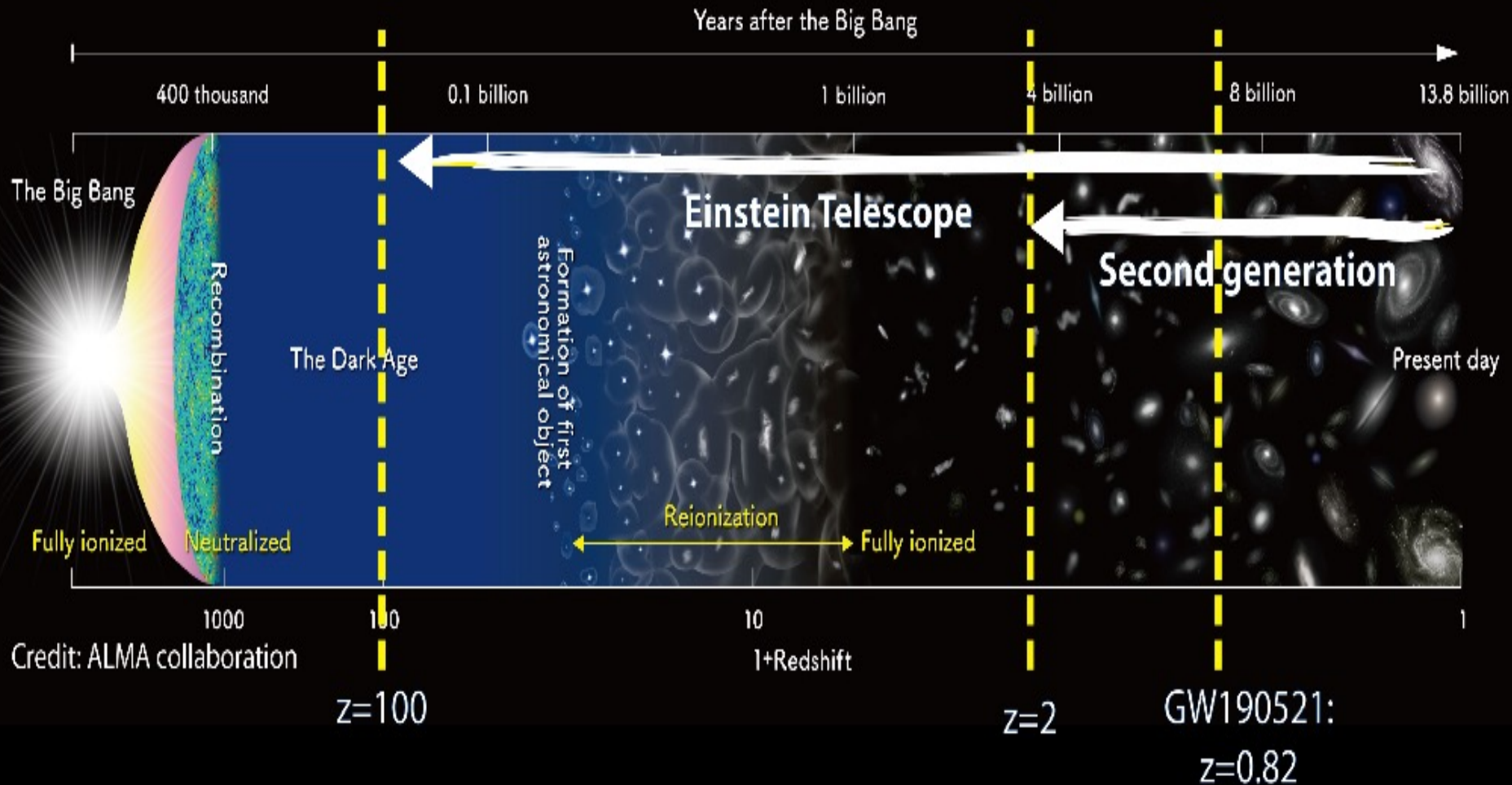
# Are LIGO/Virgo BH Primordial?



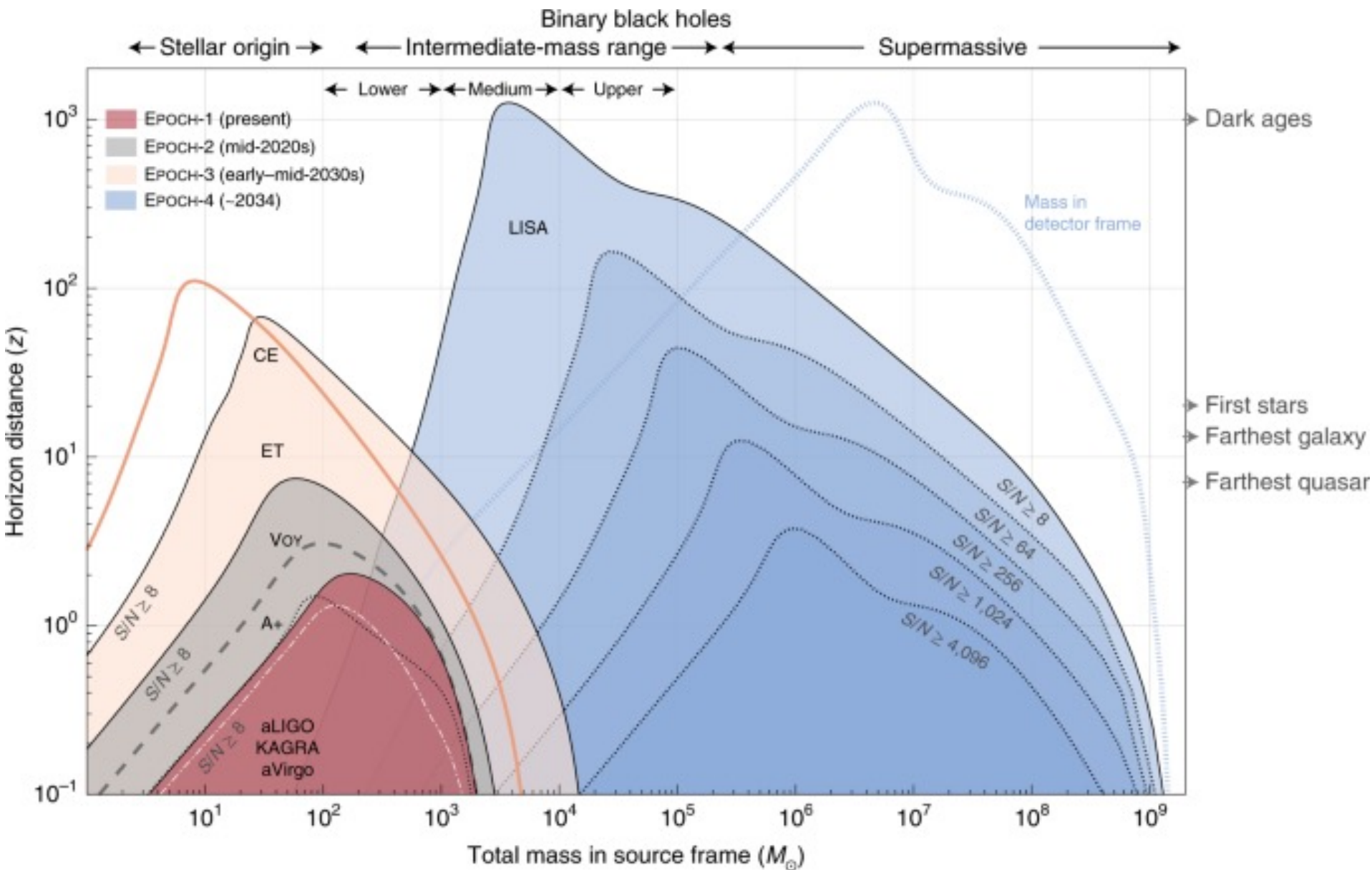


# The future of GW (G3)

## Detection horizon for black-hole binaries



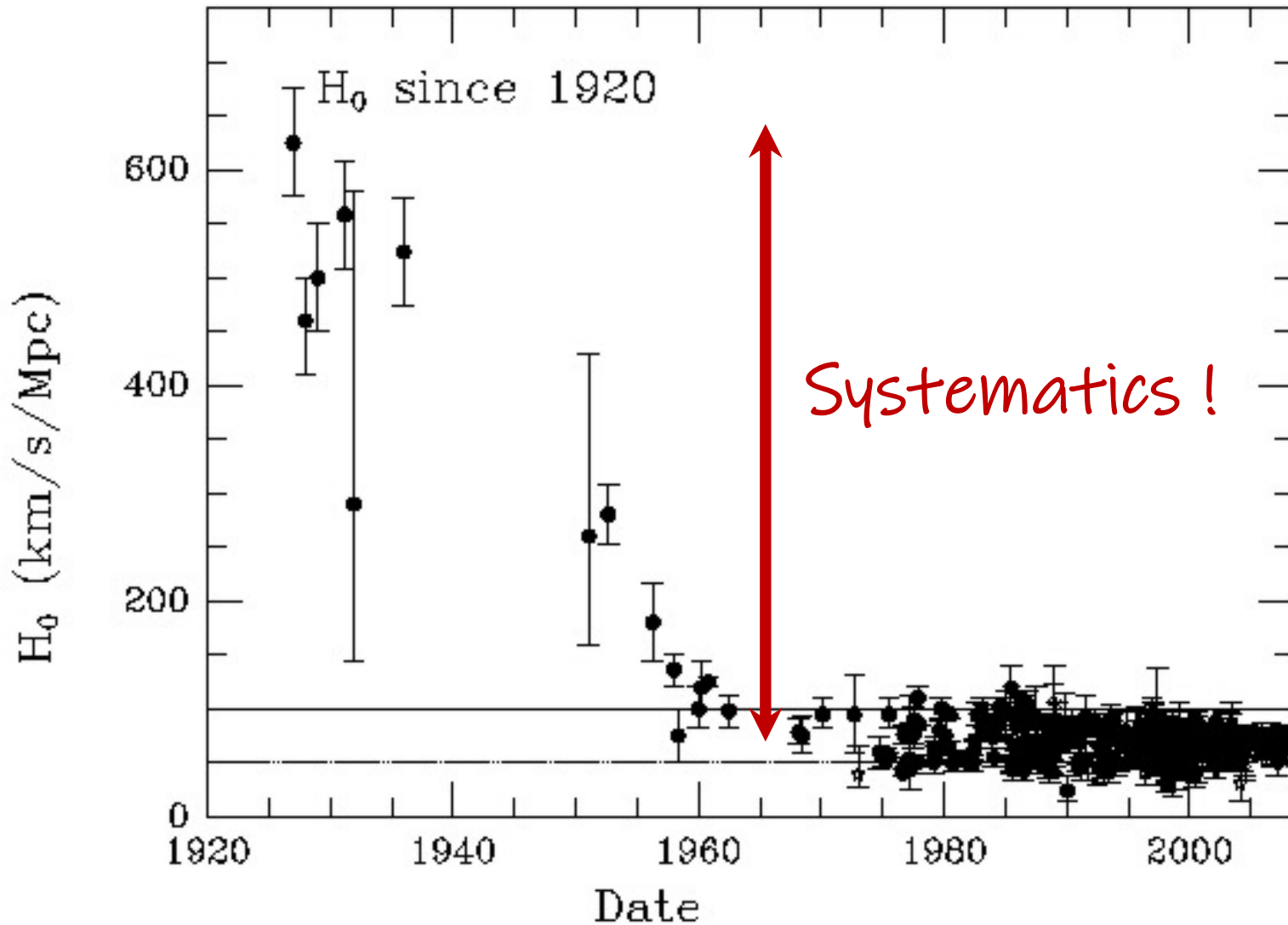
# BBH sensitivity in future G3 GW



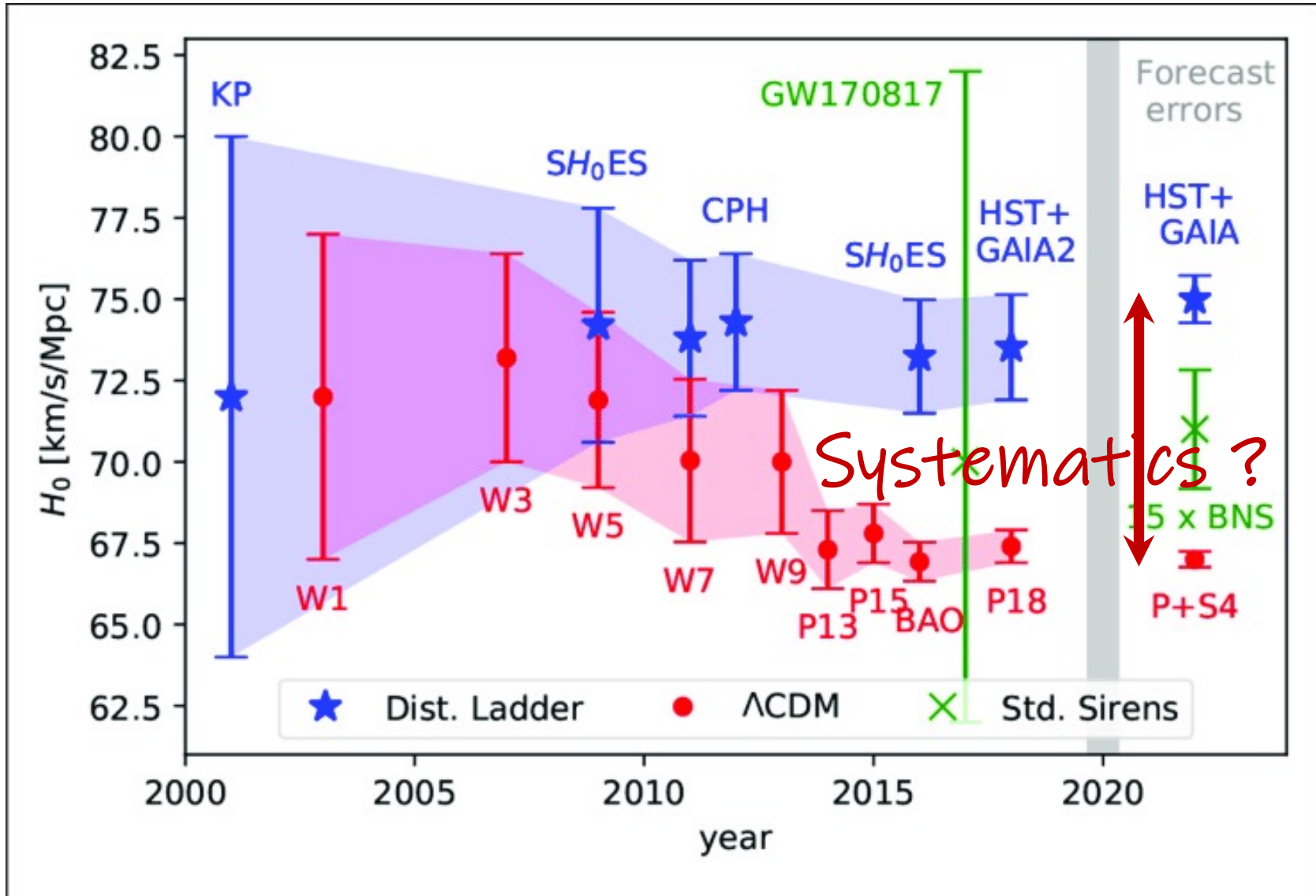
# Partial Summary

- Quantum diffusion inevitably generates PBH
- Thermal history predicts PBH have multimodal mass distribution  $\sim 10^{-5}, 1, 100, 10^5 M_{\odot}$  ( $10^{-10} M_{\odot}$  also?)
- The predicted PBH spin and mass distribution has been measured by LIGO/Virgo + OGLE around 1-100  $M_{\odot}$  (features: peak & plateau)
- Other peaks could be explored with microlensing
- PBH scenario can explain various cosmic conundra
- Paradigm shift in Structure Formation of Universe
- Very rich phenomenology: multiscale, multiepoch, multiprobe => Future GB detectors (ET, LISA, GAIA)

# Hubble Tension ca. 2000



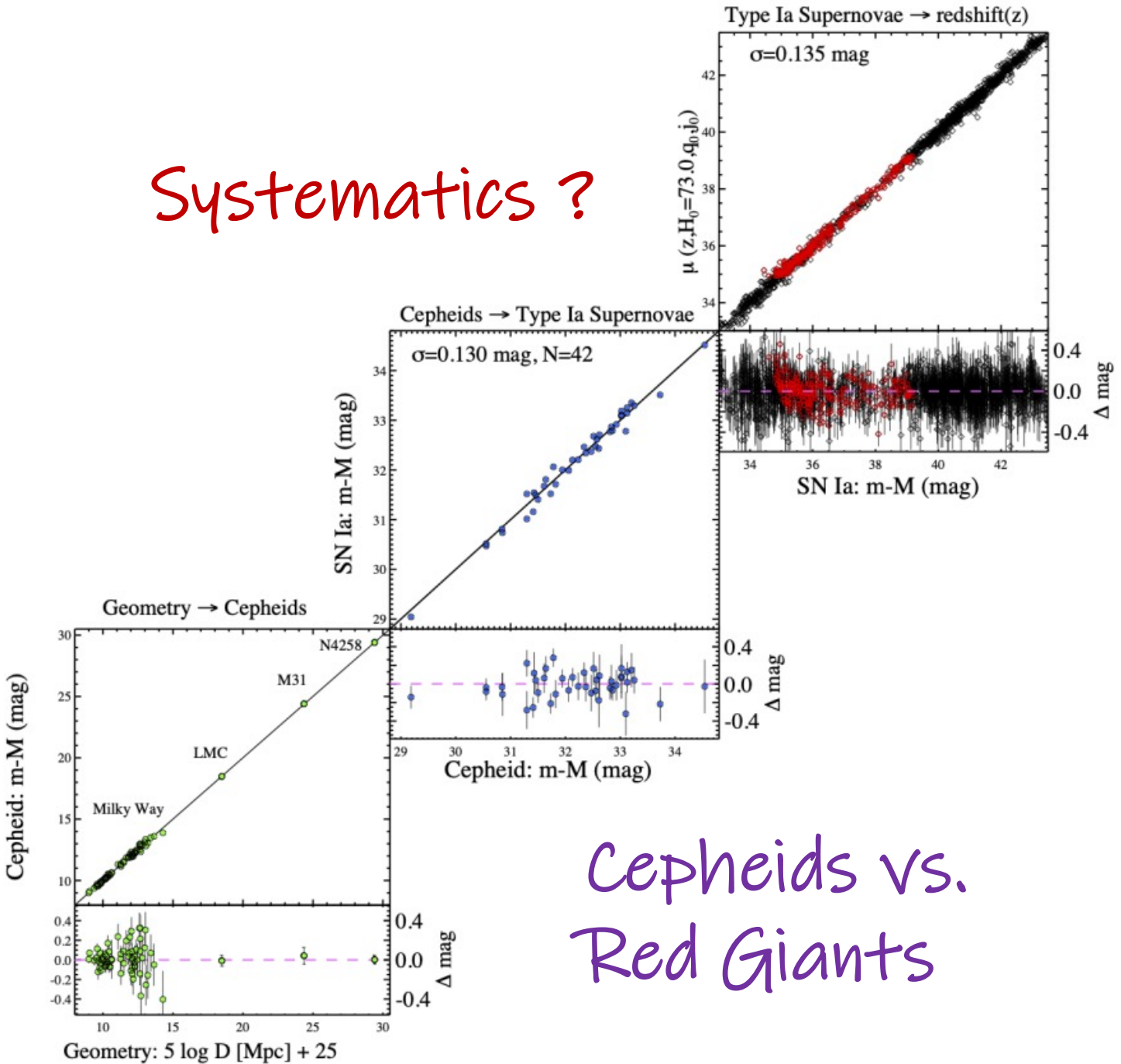
# Hubble Tension ca. 2020



Benson et al. (2017)

# Hubble Tension ca. 2021

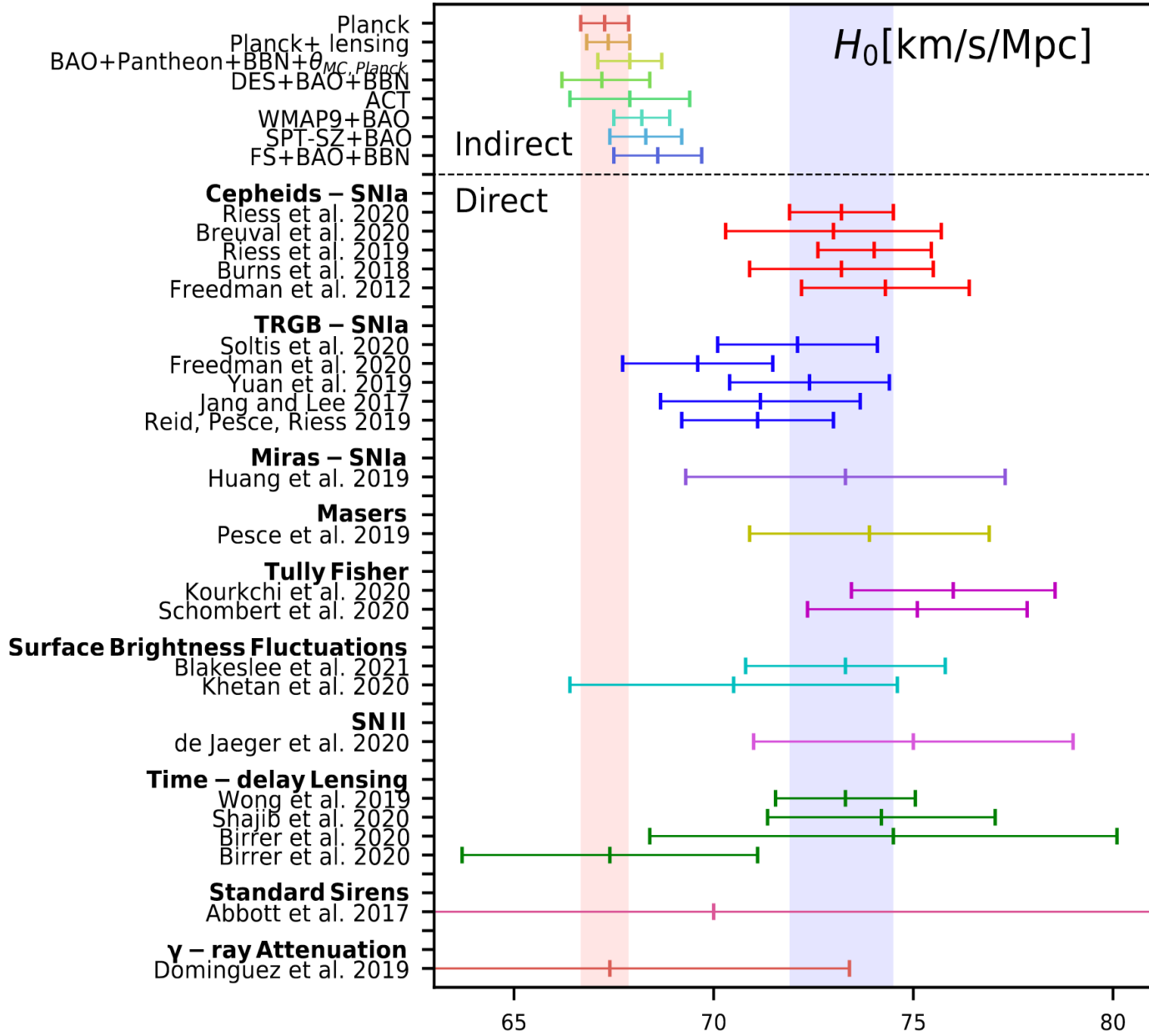
Systematics ?



Cepheids vs.  
Red Giants

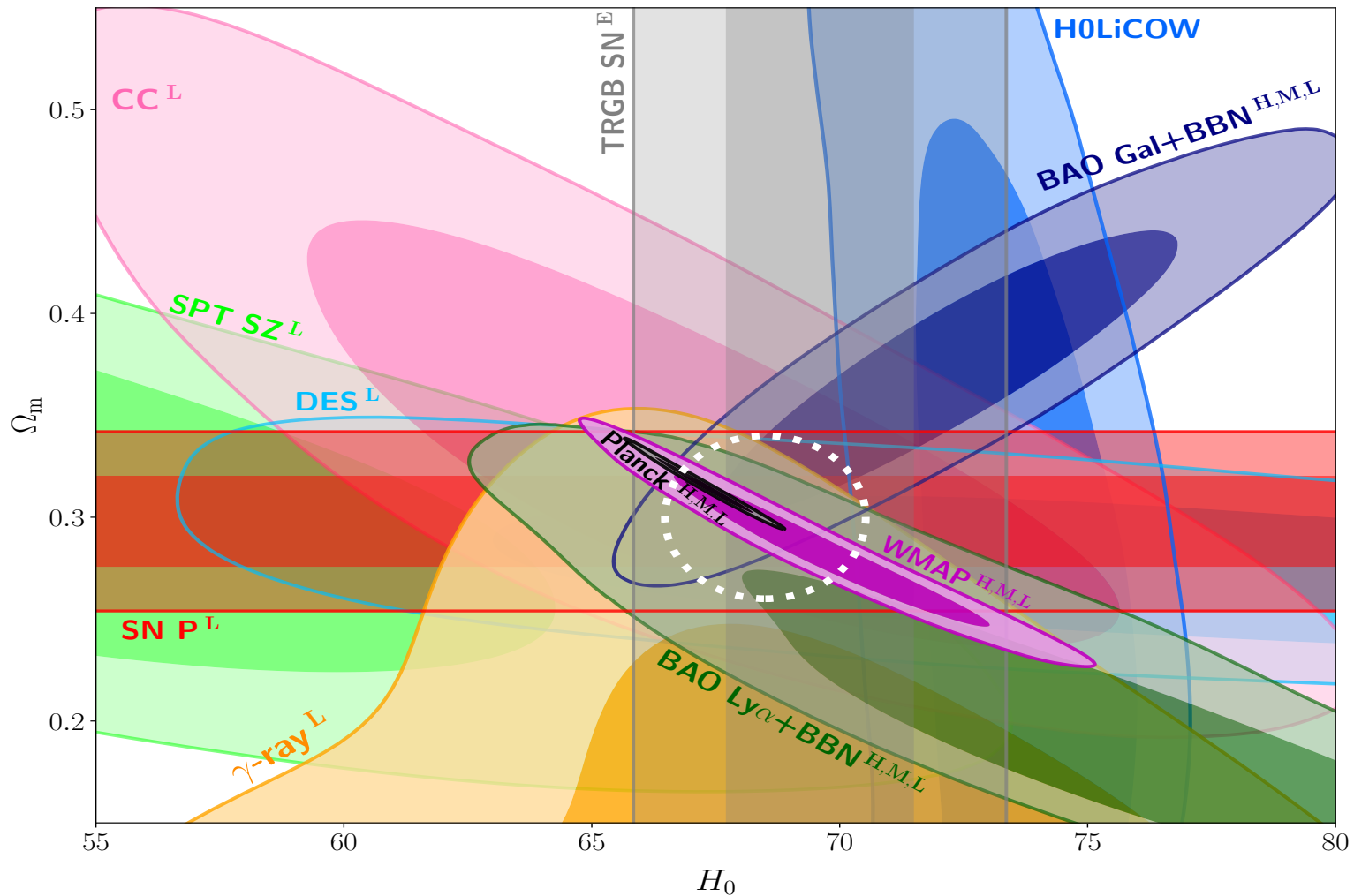
Di Valentino et al. (2021)

# Hubble Tension ca. 2021



Di Valentino (2021)

# Hubble Tension ca. 2022

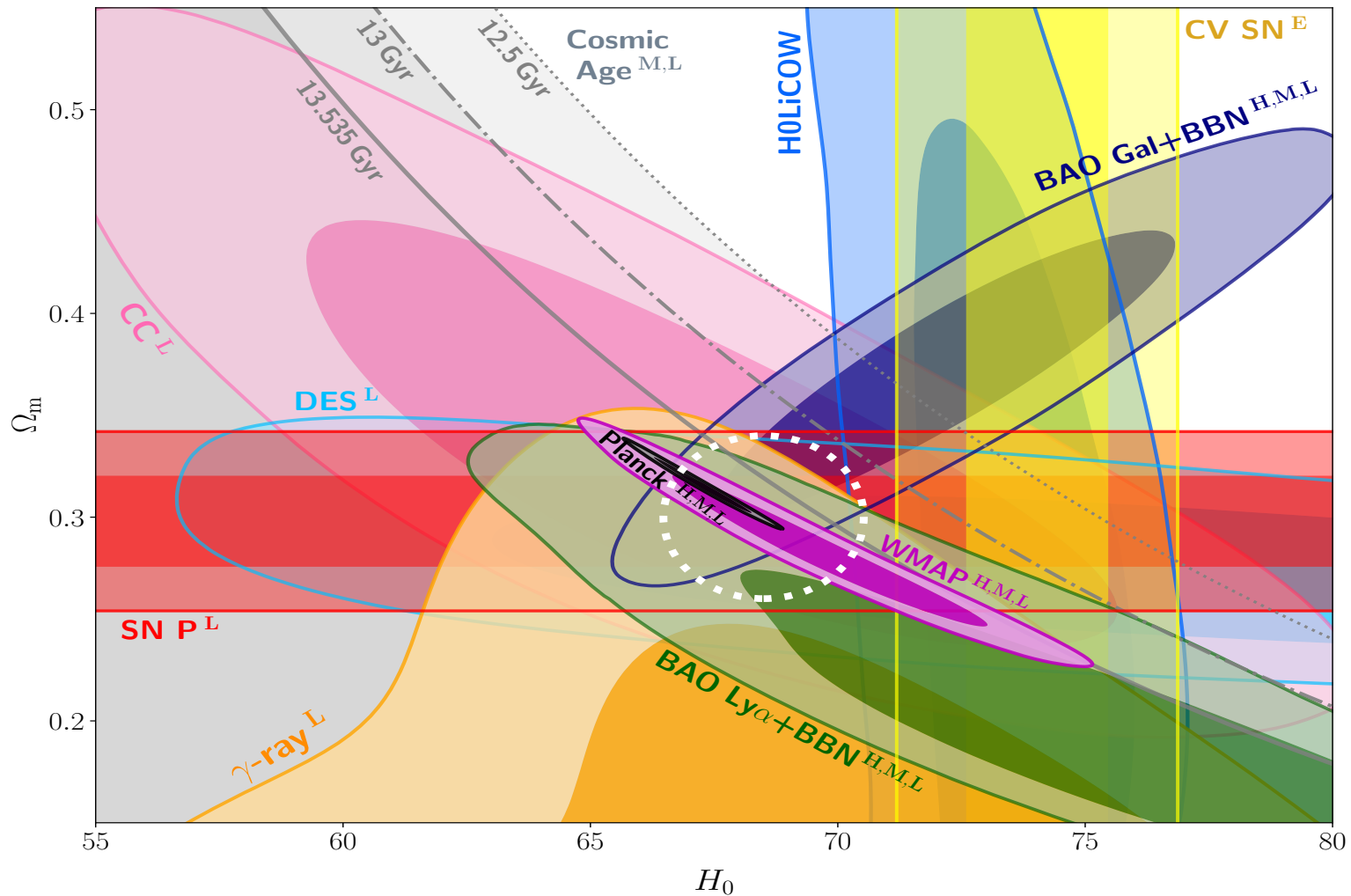


Mack et al. (2022)

Contours correspond to SN P (red), DES (light blue), CC (pink), H0LiCOW (blue), BAO Gal (navy), BAO Ly $\alpha$  (green),  $\gamma$ -ray (orange), WMAP (magenta), Planck (black), CV SN (yellow) and some guiding cosmic-age constraints ( $t_* = 13.535, 13$  and  $12.5$  Gyr; orange). See the text for descriptions and sources of those constraints. Each constraint in the figure is labeled according to whether it can be changed by nonstandard high- $z$  models (H), mid- $z$  models (M), low- $z$  models (L), or local environmental factors (E). See the text for the definition of those model categories. We leave the H0LiCOW technique without a label because it is relatively insensitive to the underlying cosmological model.



# Hubble Tension ca. 2022



Mack et al. (2022)

Contours correspond to SN P (red), DES (light blue), CC (pink), H0LiCOW (blue), BAO Gal (navy), BAO Ly $\alpha$  (green),  $\gamma$ -ray (orange), WMAP (magenta), Planck (black), CV SN (yellow) and some guiding cosmic-age constraints ( $t_* = 13.535, 13$  and  $12.5$  Gyr; orange). See the text for descriptions and sources of those constraints. Each constraint in the figure is labeled according to whether it can be changed by nonstandard high- $z$  models (H), mid- $z$  models (M), low- $z$  models (L), or local environmental factors (E). See the text for the definition of those model categories. We leave the H0LiCOW technique without a label because it is relatively insensitive to the underlying cosmological model.

# Forces in Physics

- **Fundamental Forces**

Gravitation, Strong, Weak, E.M.

- **Residual Forces**

Molecular, Nuclear, Surface Tension

- **Collective Forces**

Brownian motion,

Entropic Forces

$$F dx = dW = -dU + TdS \Rightarrow F = -\frac{dU}{dx} + T\frac{dS}{dx}$$

# Entropic forces in GR

$$\mathcal{S} = \frac{1}{2\kappa} \int d^4x \sqrt{-g} R + \int d^4x \mathcal{L}_m(g_{\mu\nu}, S)$$

Thermodynamical constraint

$$\delta\mathcal{S} = \int d^4x \left( \frac{1}{2\kappa} \frac{\delta(\sqrt{-g}R)}{\delta g^{\mu\nu}} + \frac{\delta\mathcal{L}_m}{\delta g^{\mu\nu}} \right) \delta g^{\mu\nu} + \int d^4x \frac{\partial\mathcal{L}_m}{\partial S} \delta S = 0$$

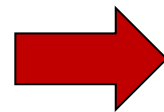
Variational constraint: 2<sup>nd</sup> law thermodynamics

$$\frac{\partial\mathcal{L}_m}{\partial S} \delta S = \frac{1}{2} \sqrt{-g} f_{\mu\nu} \delta g^{\mu\nu}$$

Non-equilibrium Einstein field equations

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \kappa (T_{\mu\nu} - f_{\mu\nu})$$

entropic force



$$D^\mu T_{\mu\nu} = D^\mu f_{\mu\nu}$$

Bianchi identities

# Entropic forces in GR

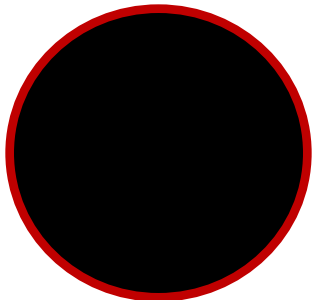
Temperature and Entropy from the gravity sector

- Horizon  $H$  with induced metric  $h$

$$\mathcal{S}_{\text{GHY}} = \frac{1}{8\pi G} \int_H d^3y \sqrt{h} K = \frac{1}{8\pi G} \int_H dt \sin\theta d\theta d\phi \sqrt{h} K$$

- Schwarzschild black hole

$$ds^2 = - \left(1 - \frac{2GM}{r}\right) dt^2 + \left(1 - \frac{2GM}{r}\right)^{-1} dr^2 + r^2 d\Omega_2^2$$



$$n = -\sqrt{1 - \frac{2GM}{r}} \partial_r$$

normal vector to  
 $S_2$  of radius  $r$

# Entropic forces in GR

$$S_{\text{GHY}} = \frac{1}{8\pi G} \int_H d^3y \sqrt{h} K = \frac{1}{8\pi G} \int_H dt \sin\theta d\theta d\phi \sqrt{h} K$$

$$\sqrt{h}K = (3GM - 2r) \sin\theta \quad \text{at event horizon } r = 2GM$$

$$S_{\text{GHY}} = -\frac{1}{2} \int dt M c^2 = - \int dt T_{\text{BH}} S_{\text{BH}}$$

$$T_{\text{BH}} = \frac{\hbar c^3}{8\pi G M}$$

Classical (emergent)

quantum origin

$$S_{\text{BH}} = \frac{A c^3}{4G\hbar} = \frac{4\pi G M^2}{\hbar c}$$

# Entropic forces in GR

- Contribution to bulk entropy of the inevitable Schwarzschild black hole component of Dark Matter

Assuming BH total comoving number is conserved, their Total energy density and entropy density ( $\hbar = c = 1$ ) is

$$\rho_{BH} = n_{BH} M, \quad s_{BH} = n_{BH} 4\pi G M^2$$

Therefore 
$$a^3 \frac{d}{dt}(\rho_{BH} a^3) = T_{BH} \frac{d}{dt}(s_{BH} a^3) = 0.$$

No contribution to entropic force of the universe unless multiple black hole mergers or significant mass accretion, which may change mass or number density of black holes.

# Entropic forces in FLRW

Non-equilibrium thermodynamics in expanding universe

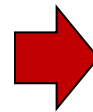
$$ds^2 = -N(t)^2 dt^2 + a^2(t) \left( \frac{dr^2}{1 - kr^2} + r^2 d\Omega_2^2 \right)$$

$$T^{\mu\nu} = (\rho + p)u^\mu u^\nu + pg^{\mu\nu}$$

$$D^\mu T_{\mu\nu} = D^\mu f_{\mu\nu}$$

2<sup>nd</sup> law thermodynamics

$$TdS = d(\rho a^3) + p d(a^3)$$



$$\dot{\rho} + 3H(\rho + p) = \frac{T\dot{S}}{a^3}$$



Hamiltonian constraint

$$\dot{a}^2 + k = \frac{8\pi G}{3} \rho a^2$$

Friedmann/Raychaudhuri equation

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3}(\rho + 3p) + \frac{4\pi G}{3} \frac{T\dot{S}}{a^3 H}$$

# Entropic forces in FLRW

- Causal Cosmological Horizon  $H$


$$\sqrt{h}K = 2N(t) r a \sqrt{1 - kr^2} \sin \theta \quad \text{Trace extrinsic curvature}$$

$$d_H = a \eta \quad \text{Causal horizon distance}$$

$$r_H = \sinh(\eta\sqrt{-k}) / \sqrt{-k} \quad \text{Conformal time } \eta$$

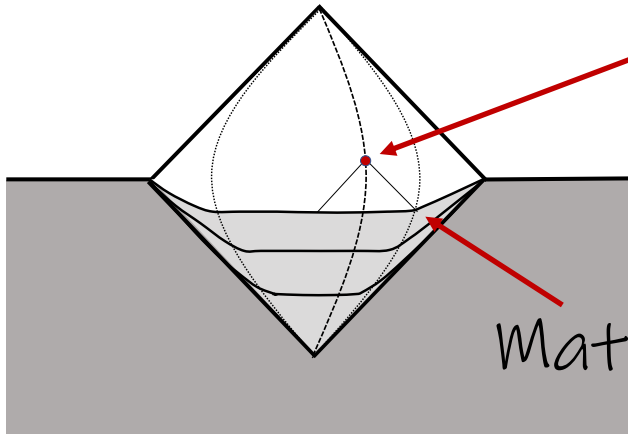
$$S_{GHY} = -\frac{1}{2G} \int dt N(t) \frac{a}{\sqrt{-k}} \sinh(2\eta\sqrt{-k})$$

$$= - \int dt N(t) T_H S_H = - \int dt N a^3 \rho_H$$

$$T_H = \frac{\hbar c \sinh(2\eta\sqrt{-k})}{2\pi a r_H^2 \sqrt{-k}}, \quad S_H = \frac{c^3 \pi a^2 r_H^2}{\hbar G} \quad \text{Emergent}$$




# Cosmic Acceleration



Observer's causal horizon

$$\rho_H a^2 = \frac{T_H S_H}{a} = \frac{1}{2G} \frac{\sinh(2a_0 H_0 \eta)}{a_0 H_0}$$

Matching:  $H^2 = \frac{8\pi G}{3} \rho - \frac{k}{a^2} \implies \sqrt{-k} = a_0 H_0$

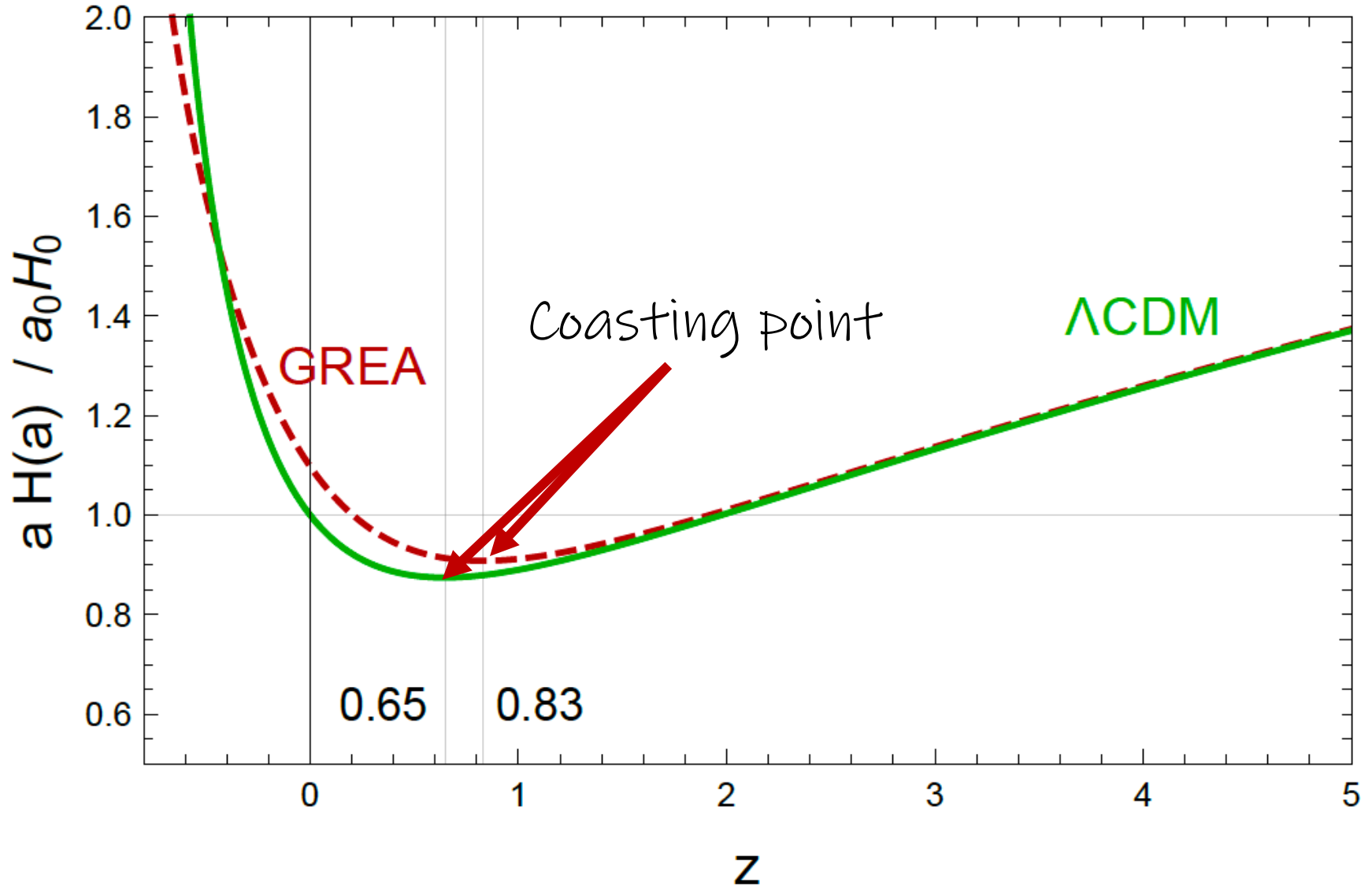
Hamiltonian constraint in conformal time  
(primes denote derivatives w.r.t.  $\tau = a_0 H_0 \eta$ )

$$\left(\frac{a'}{a}\right)^2 = \Omega_M \left(\frac{a_0}{a}\right) + \Omega_K + \frac{4\pi}{3} \Omega_K^{3/2} \sinh(2\tau)$$

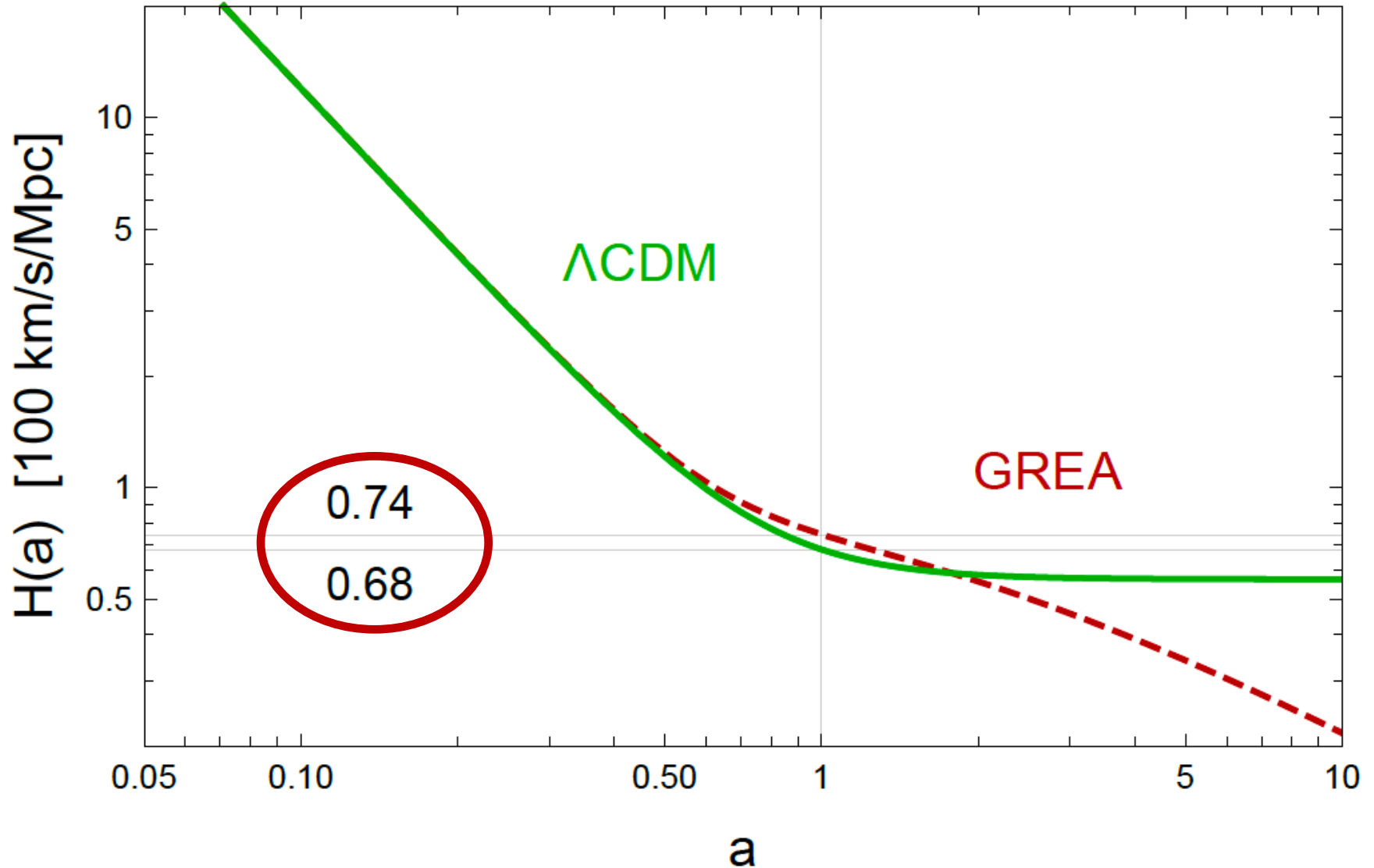
Entropic force term

Note:  $\Lambda = 0$

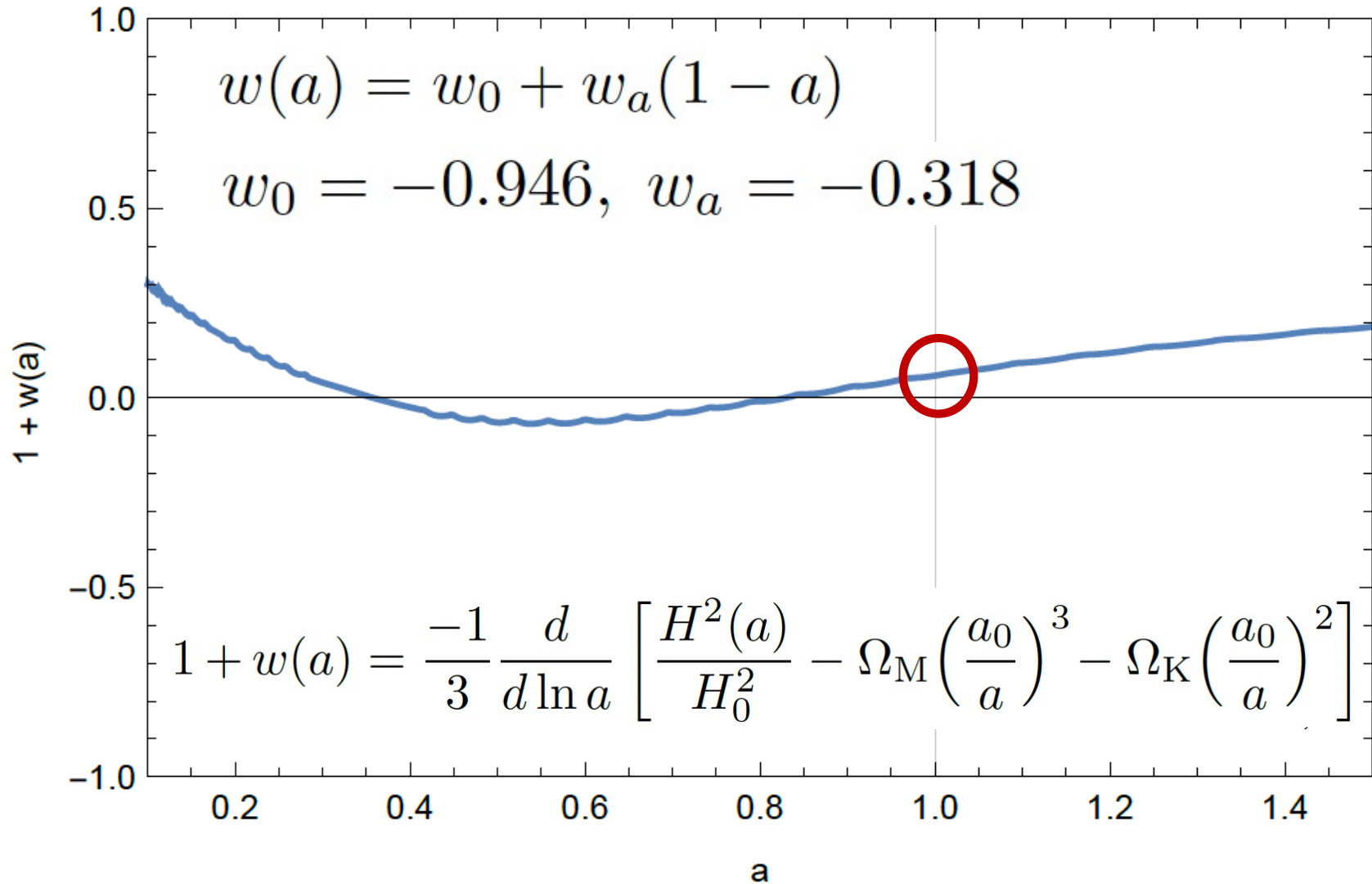
# Cosmic Acceleration



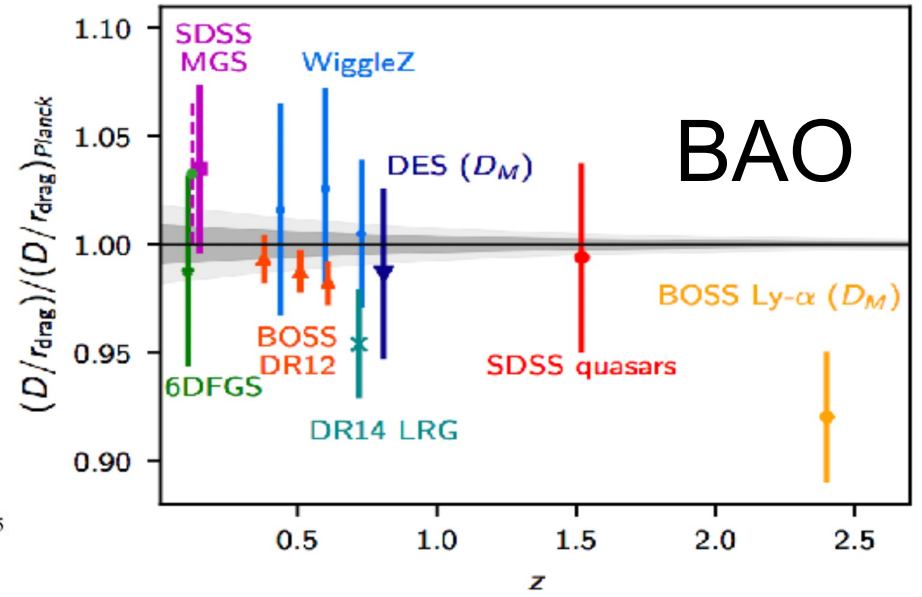
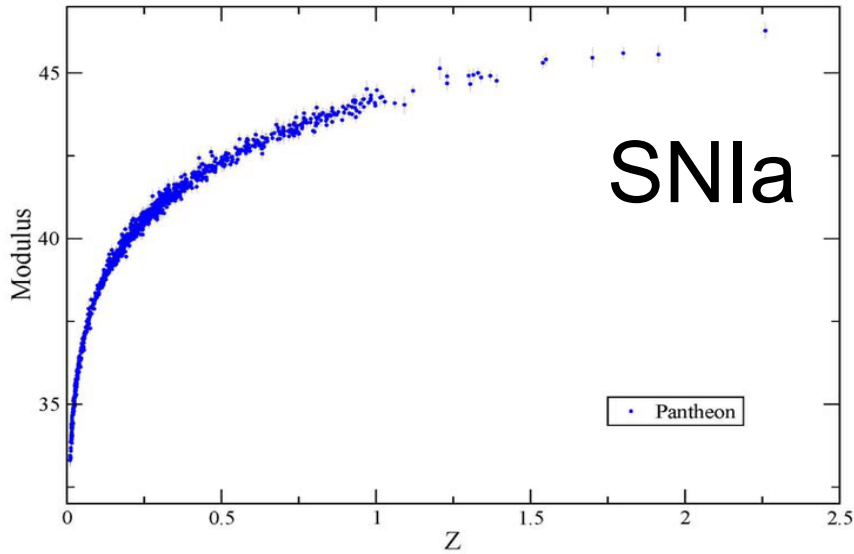
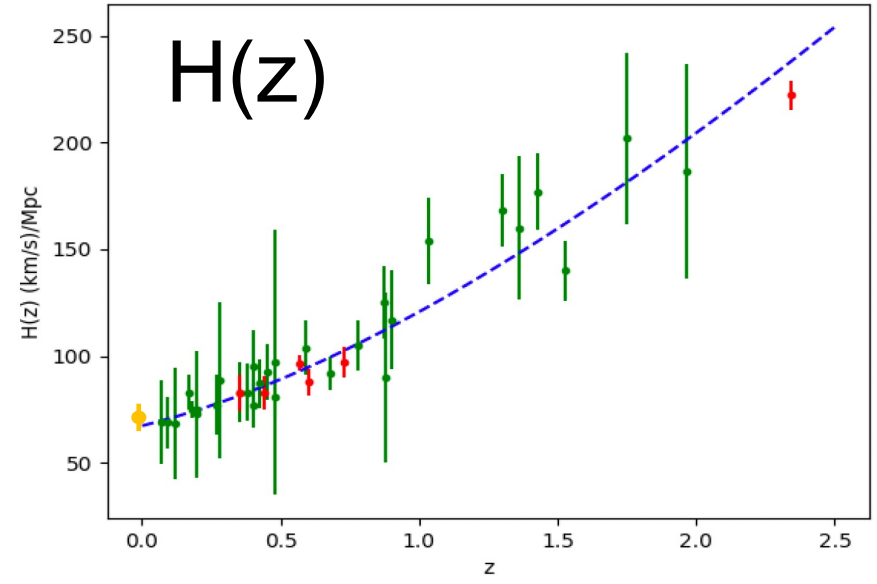
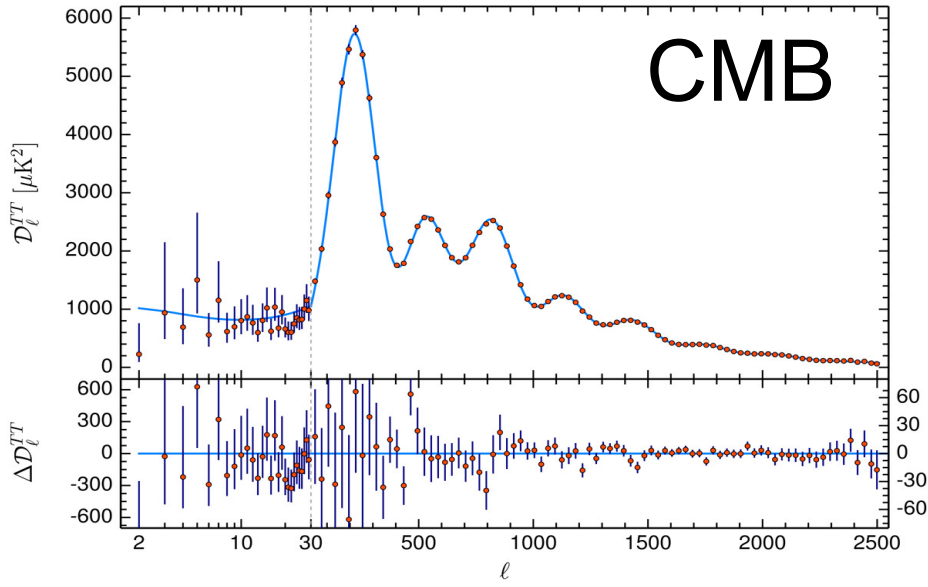
# Cosmic Acceleration



# Cosmic Acceleration

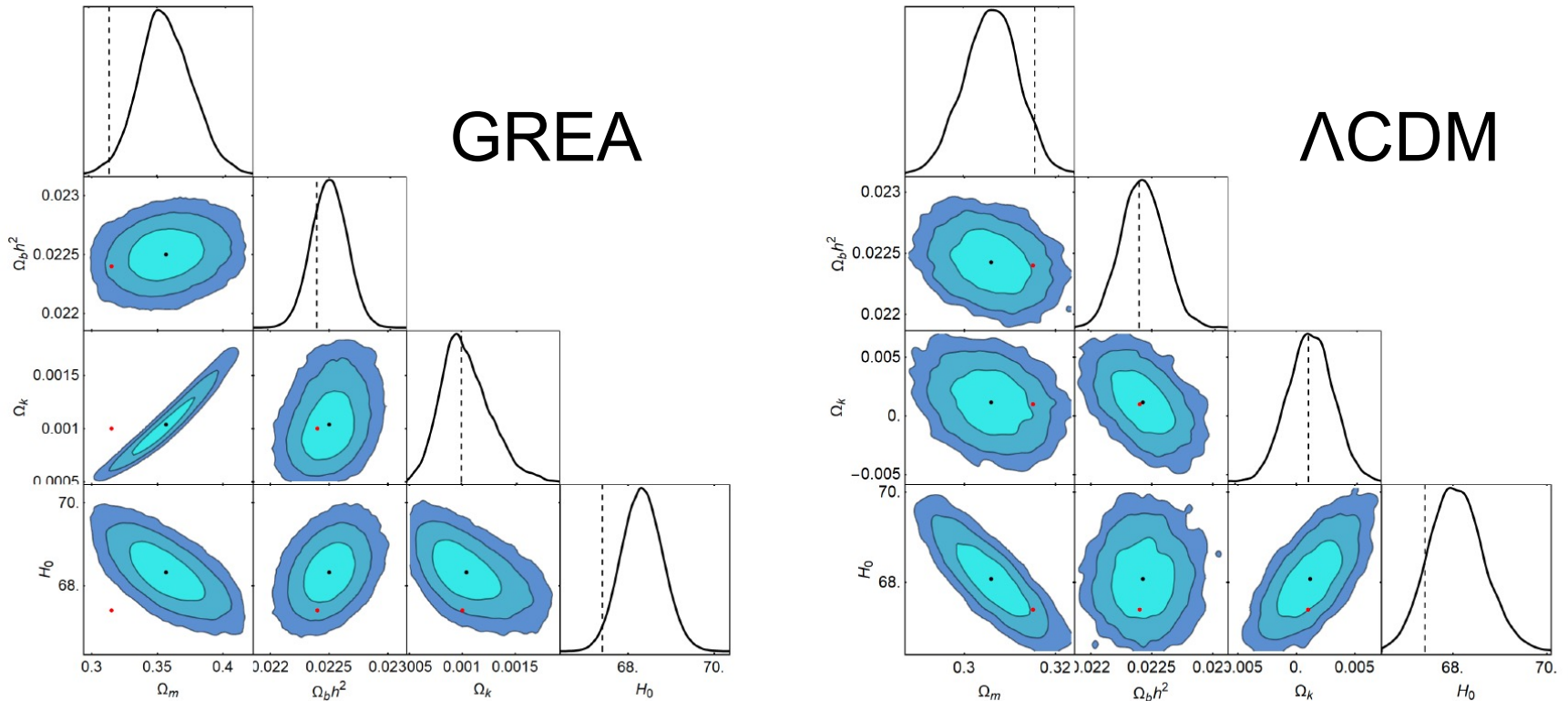


# Cosmo Observations



# Cosmic Constraints

Arjona, Espinosa, JGB & Nesseris (2021)



Model	$\Omega_{m,0}$	$\Omega_{b,0}h^2$	$\Omega_{k,0}$	$H_0$	$\chi^2_{min}$	$\log Z(1)$
$\Lambda$ CDM	$0.3057 \pm 0.0056$	$0.0224 \pm 0.0002$	$0.0012 \pm 0.0018$	$68.08 \pm 0.58$	1075.63	-557.515
GREY	$0.3522 \pm 0.0190$	$0.0225 \pm 0.0001$	$0.0010 \pm 0.0002$	$68.38 \pm 0.48$	1071.35	-548.509

# Cosmic Constraints

Arjona, Espinosa, JGB & Nesseris (2021)

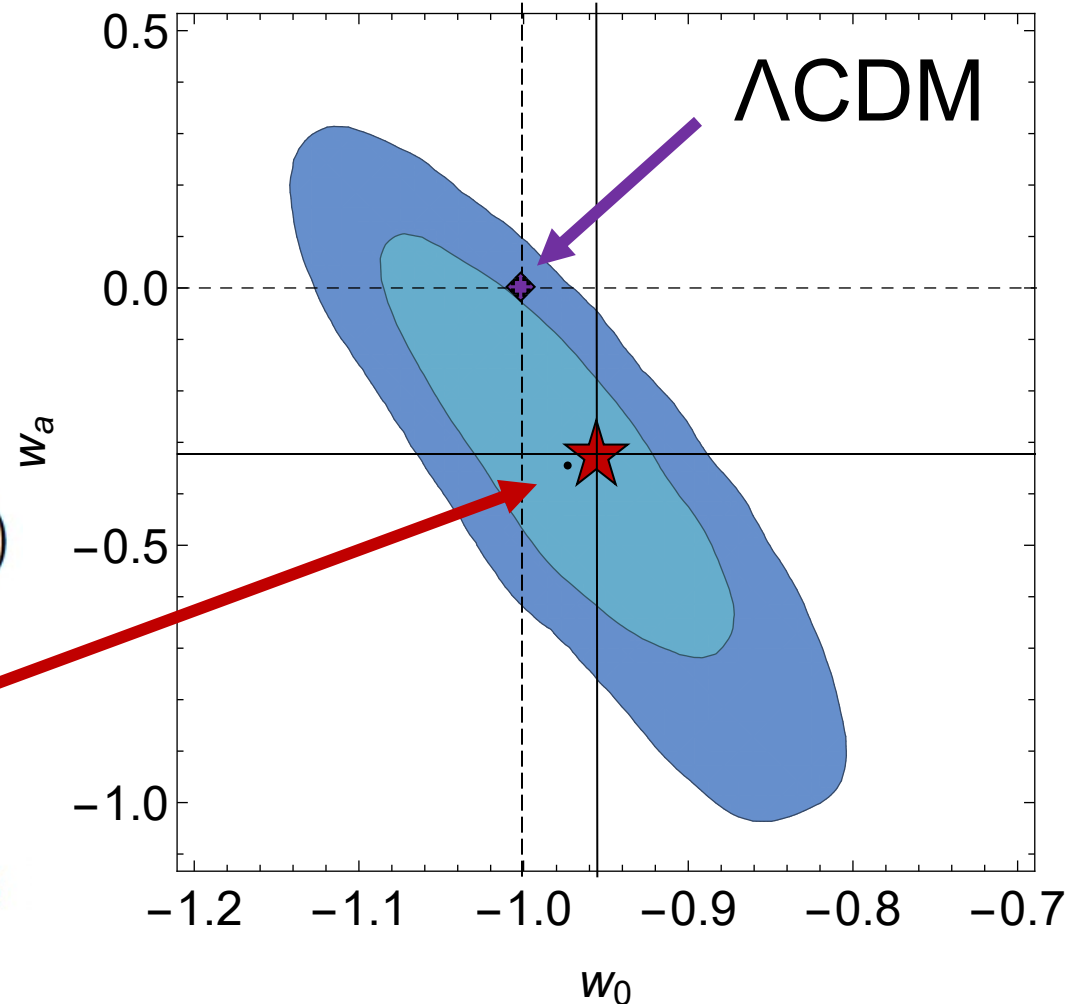
Same data  
but with

$(w_0, w_a)$  free:

$$w(a) = w_0 + w_a(1 - a)$$

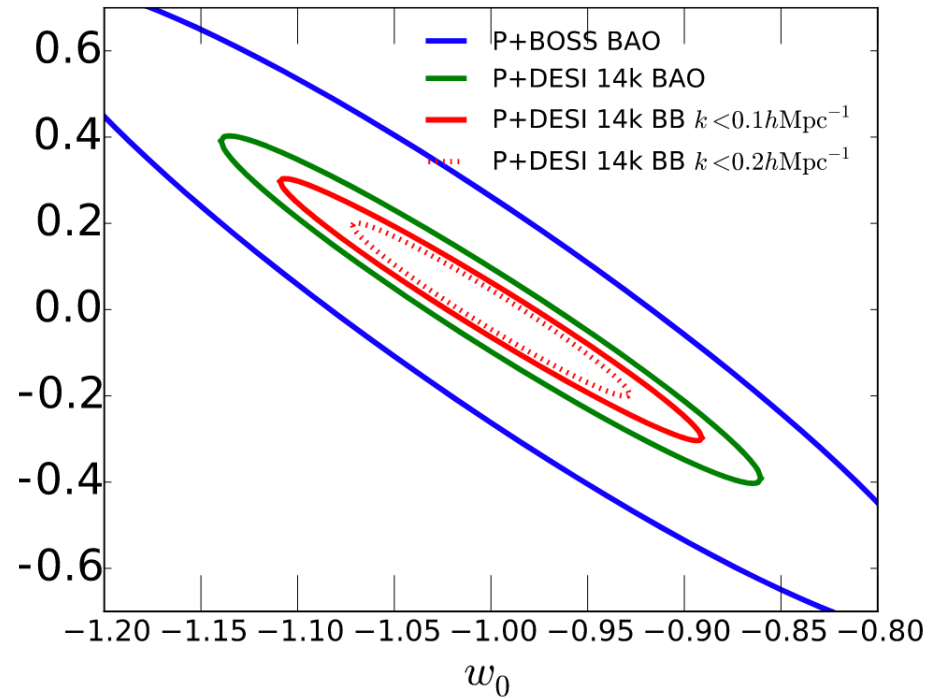
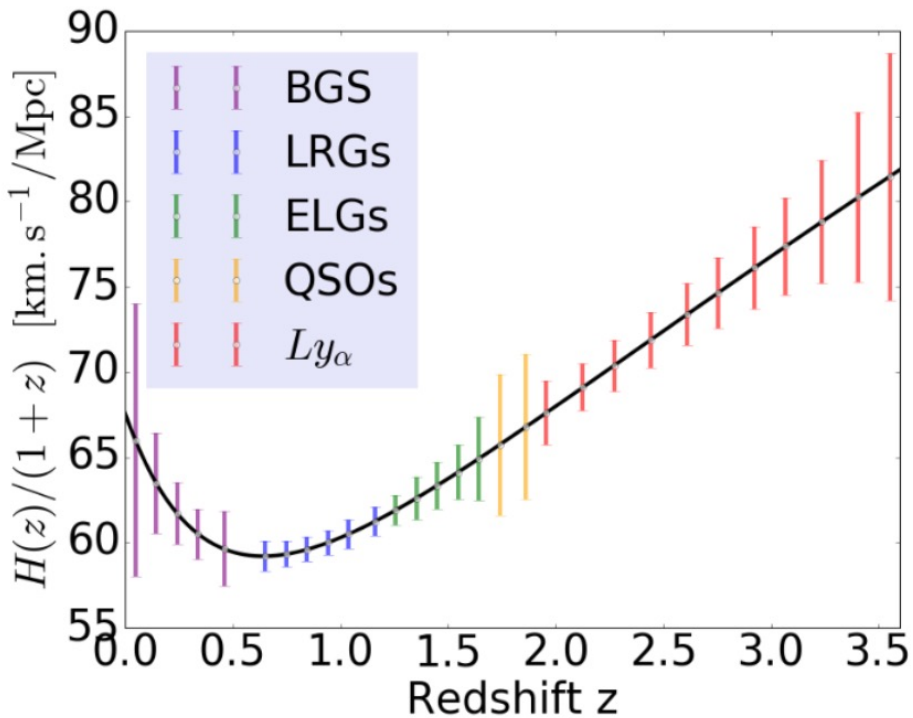
**GREA**

$$w_0 = -0.946, w_a = -0.318$$



# Future Constraints

DESI Coll. (2016)



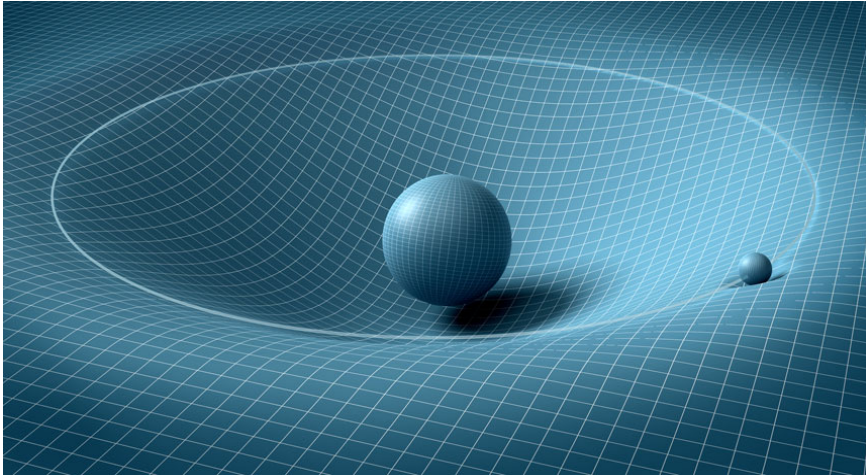


# Conclusions

- Non-equilibrium phenomena in GR: entropic forces
- ADM (3+1) slicing: Raychaudhuri eq. grav. collapse
- Cosmic acceleration from first principles
- No need for a Cosmological Constant
- Future Infinity is Flat Space Minkowsky
- Just QFT, GR and Non eq. Thermodynamics
- Multiple consequences for Large Scale Structure
- Possible solution to the  $H_0$  tension
- Future: Preheating after inflation (Big Bang)
- Future: Cosmic Web Entropy: Cosmic Voids

# On the shoulders of giants

General Relativity



Thermodynamics



Quantum Mechanics

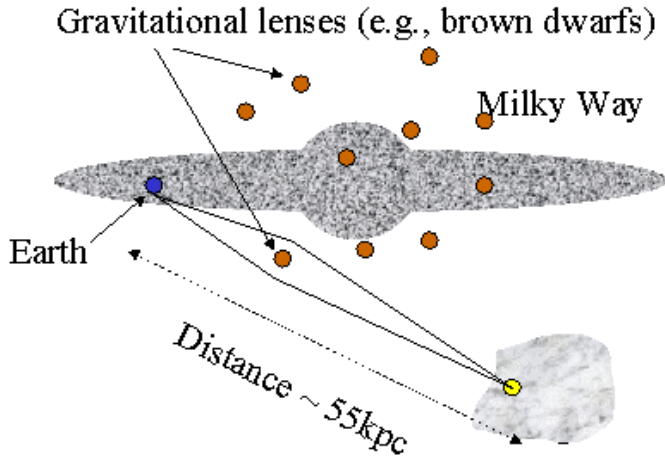


	1 <sup>st</sup>	2 <sup>nd</sup>	3 <sup>rd</sup>				
QUARKS	2.3 M <b>u</b> 2/3 1/2 up	1.27 G <b>c</b> 2/3 1/2 charm	173.1 G <b>t</b> 2/3 1/2 top	strong nuclear force	electromagnetic force	126 G <b>H</b> higgs 0 0	
	4.8 M -1/3 1/2 <b>d</b> down	95 M -1/3 1/2 <b>s</b> strange	4.2 G -1/3 1/2 <b>b</b> bottom				0 0 1 <b>g</b> gluon
	0.511 M -1 1/2 <b>e</b> electron	105.7 M -1 1/2 <b>μ</b> muon	1.78 G -1 1/2 <b>τ</b> tau				0 0 1 <b>γ</b> photon
LEPTONS	<2.2 0 1/2 <b>ν<sub>e</sub></b> e neutrino	0.17 M 0 1/2 <b>ν<sub>μ</sub></b> μ neutrino	<15.5 M 0 1/2 <b>ν<sub>τ</sub></b> τ neutrino	weak nuclear force	80.4 G ±1 1 <b>W</b> W boson	91.2 G 0 1 <b>Z</b> Z boson	
	FERMIONS				GAUGE BOSONS		

**Backup slides**

**Long duration  
microlensing  
events  
OGLE-GAIA**

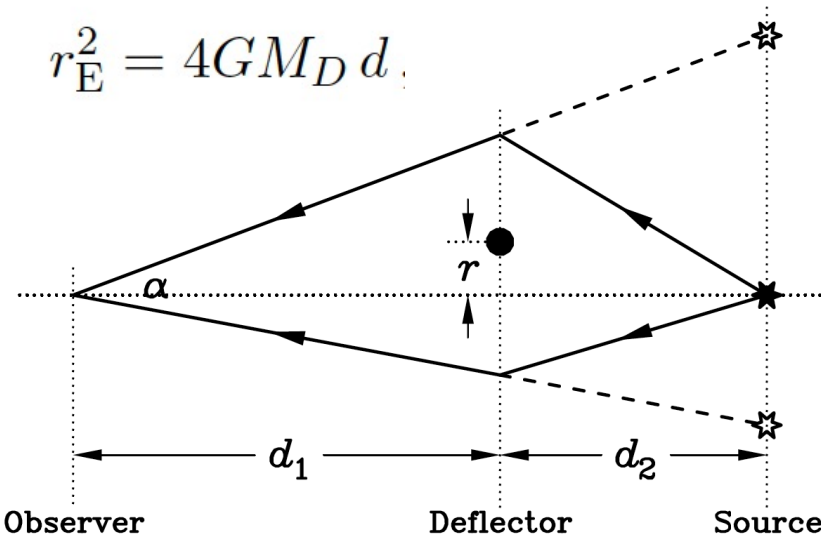
# Microlensing



$$A = \frac{2 + u^2}{u\sqrt{4 + u^2}} \quad u = \frac{r}{r_E} \quad \text{amplification}$$

$$\overline{Dt} = \frac{r_E}{v} = \frac{\sqrt{4GM_D d}}{v} \quad \text{average } \frac{1}{2} \text{ crossing}$$

$$r_E^2 = 4GM_D d$$



$$d = \frac{d_1 d_2}{d_1 + d_2}$$

$$M_D = 100 M_\odot \Rightarrow \overline{Dt} = 4 \text{ years}$$

$$M_D = 10 M_\odot \Rightarrow \overline{Dt} = 1.23 \text{ years}$$

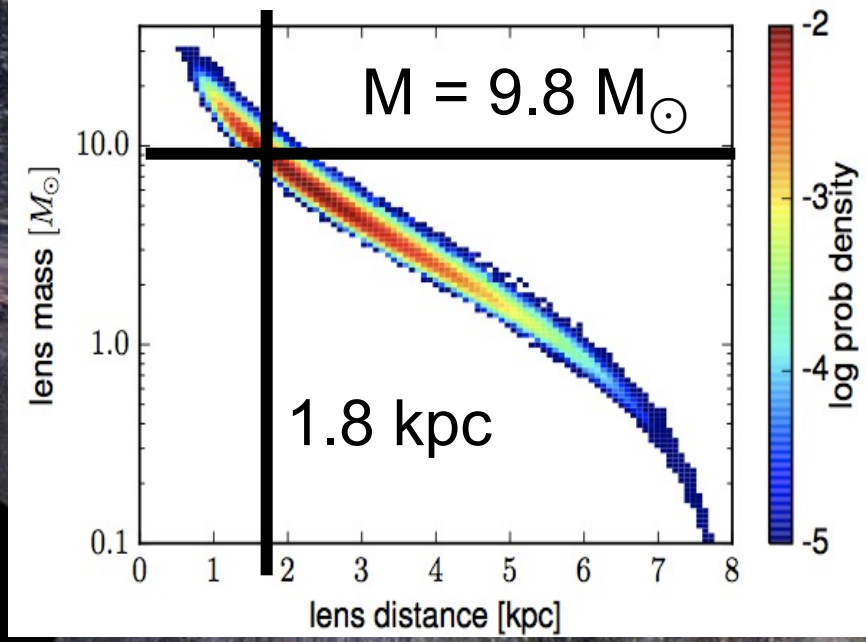
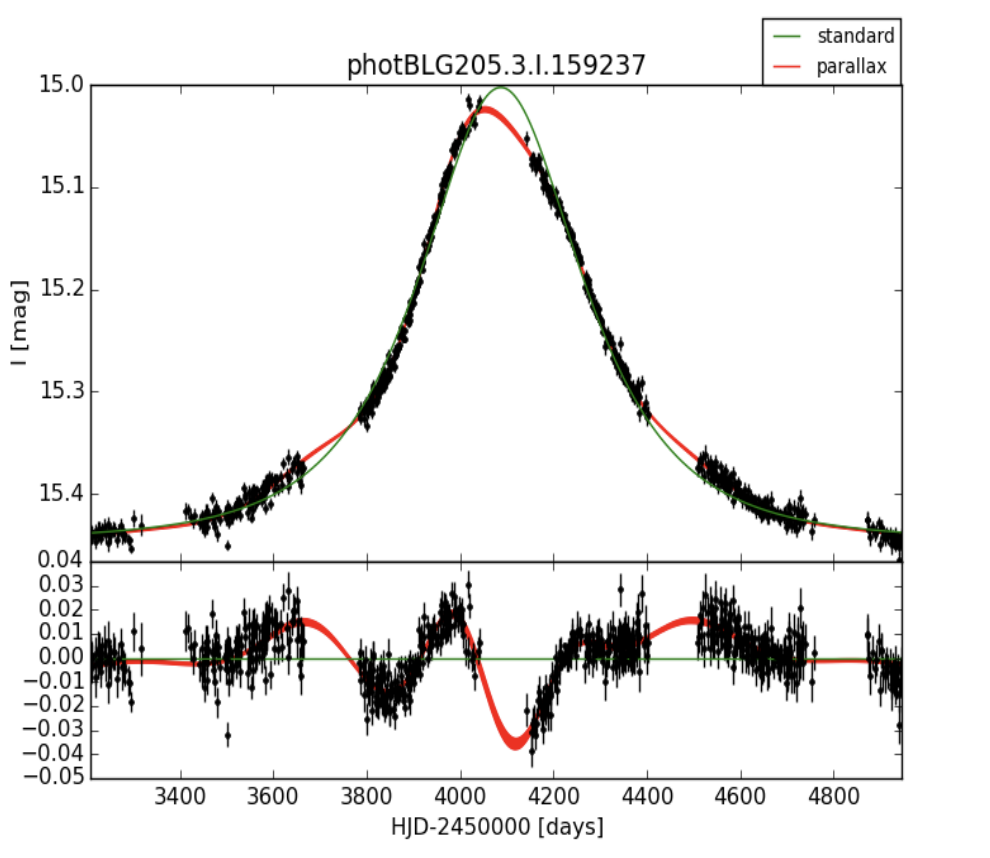
$$M_D = 1 M_\odot \Rightarrow \overline{Dt} = 5 \text{ months}$$

$$M_D = 0.1 M_\odot \Rightarrow \overline{Dt} = 1.5 \text{ months}$$

$$M_D = 0.01 M_\odot \Rightarrow \overline{Dt} = 2 \text{ weeks}$$

# OGLE3-UL-PAR-02 - candidate BH

Wyrzykowski (2016)



OGLE photometry  
from 2001-2008  
and microlensing model



$$\frac{r_E}{v} = \frac{\sqrt{4GM_D d}}{v}$$

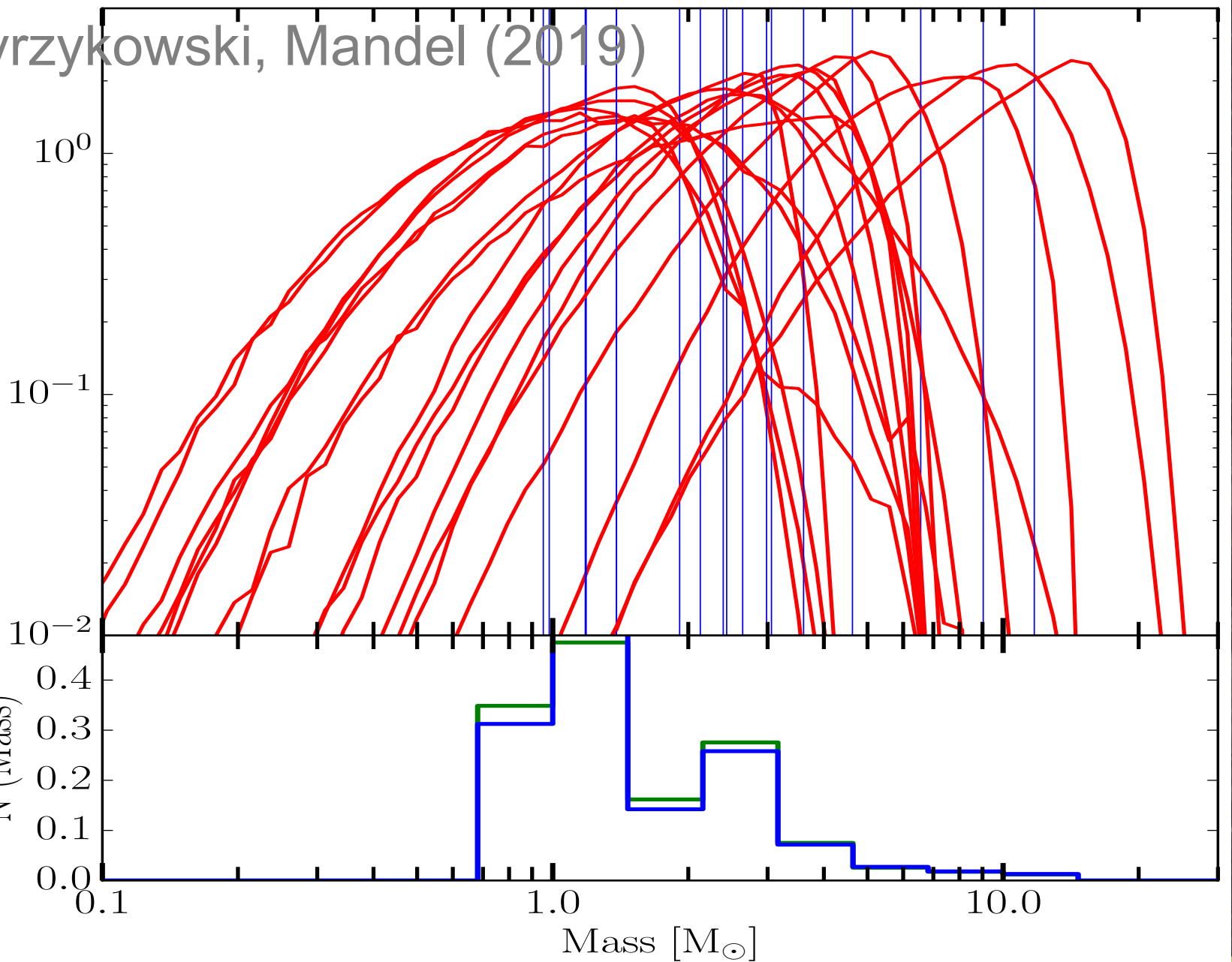
**Mass, Distance** (degenerated estimate)

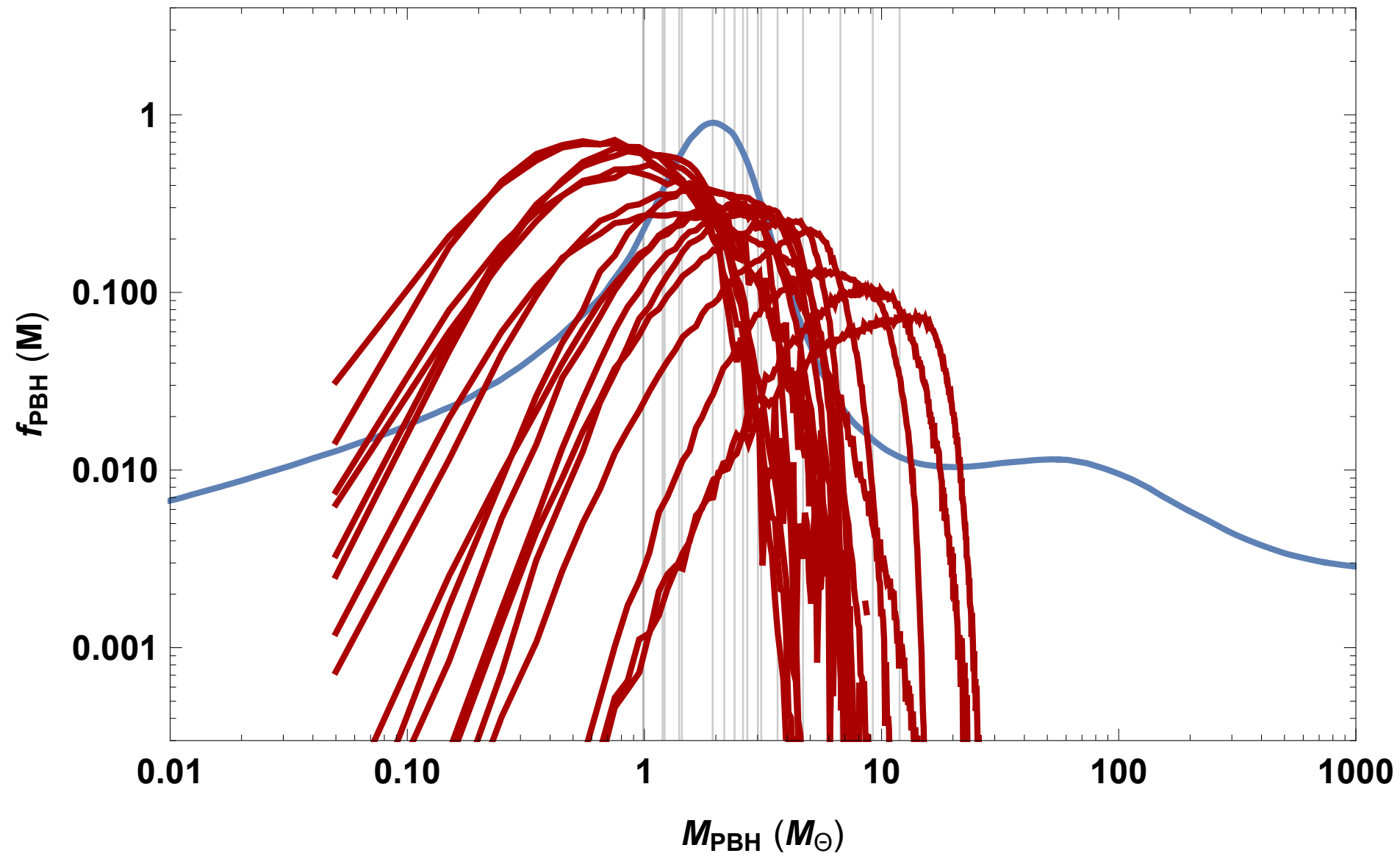
# MASS FUNCTION

Wyrzykowski, Mandel (2019)

Probability density

$N(\text{Mass})$



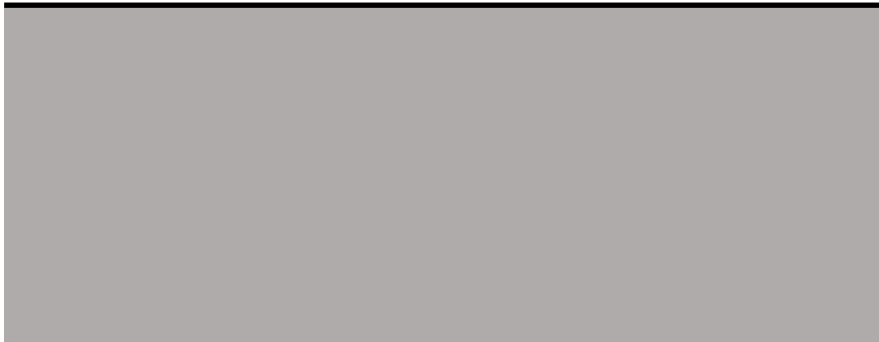




# Open Universe model for GREA

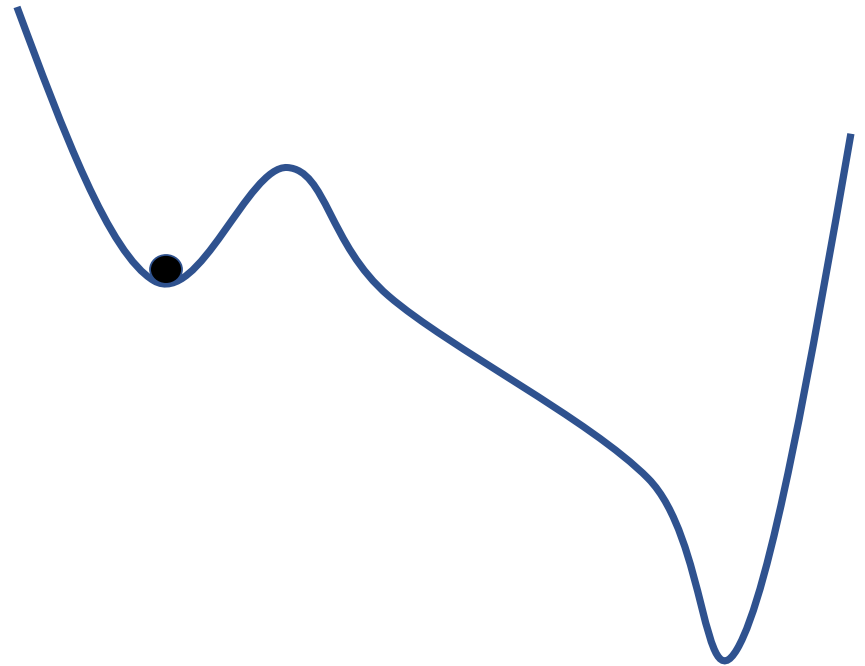
# Cosmic Acceleration

Penrose diagram



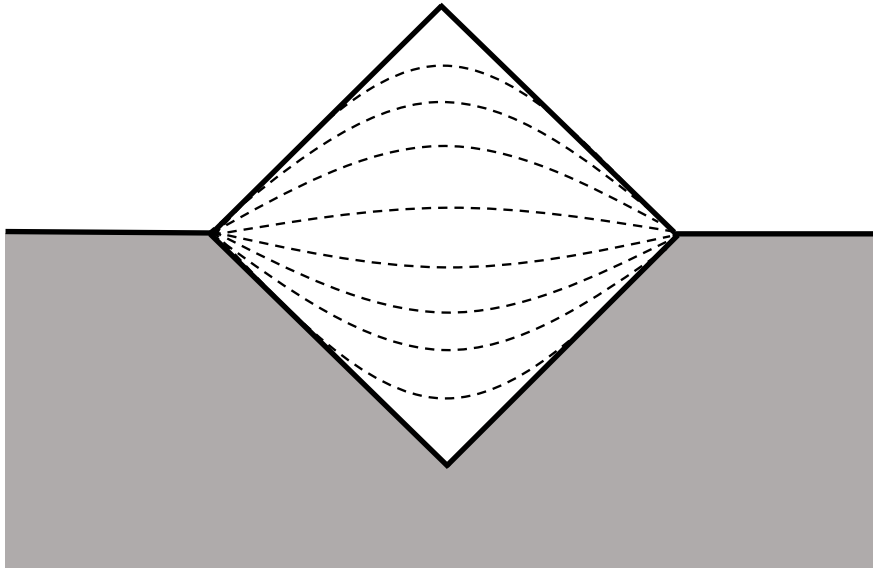
de Sitter

Potential



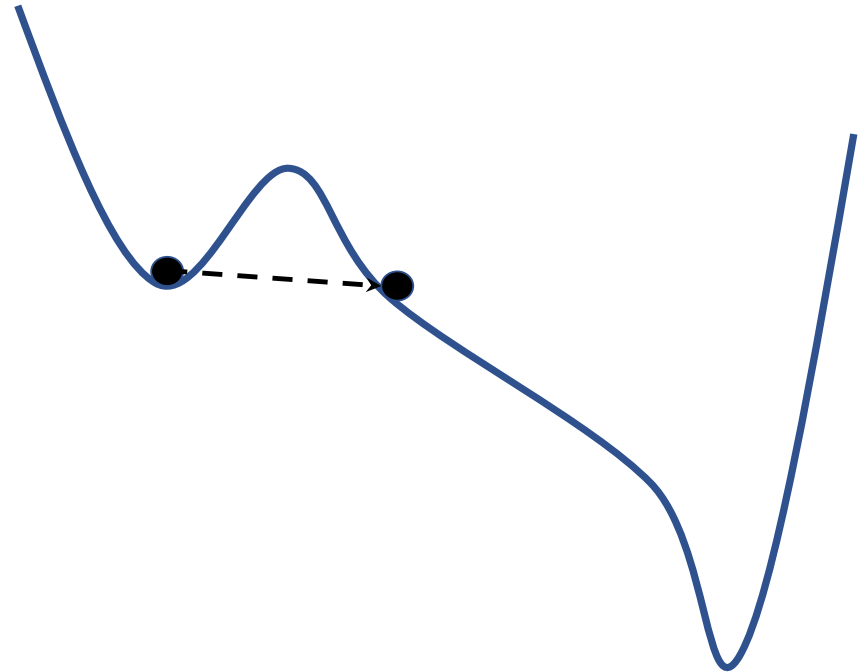
# Cosmic Acceleration

Penrose diagram



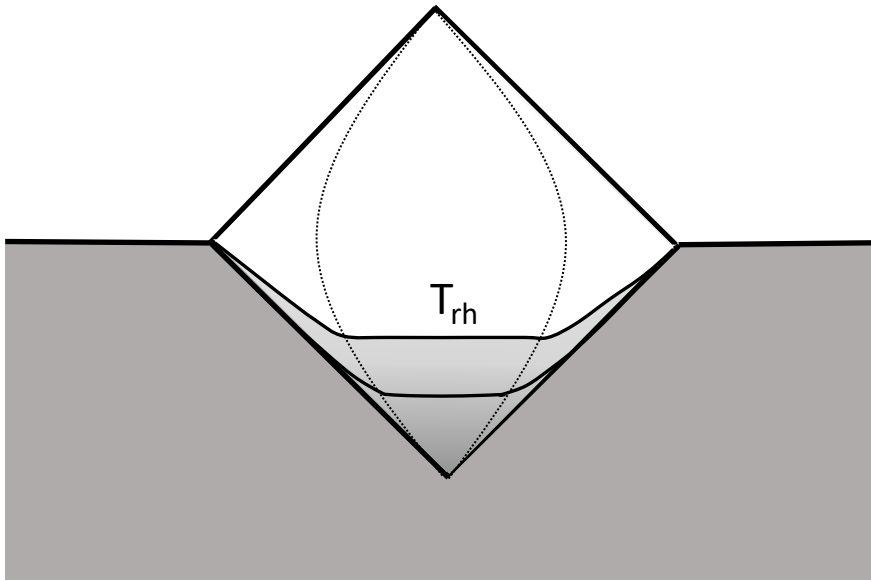
Open empty universe

Potential



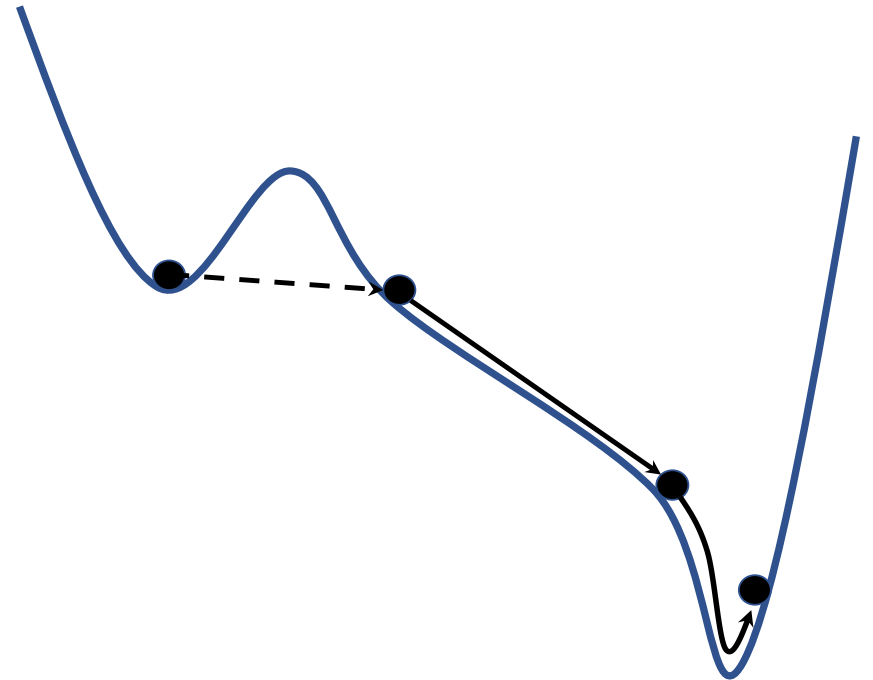
# Cosmic Acceleration

Penrose diagram



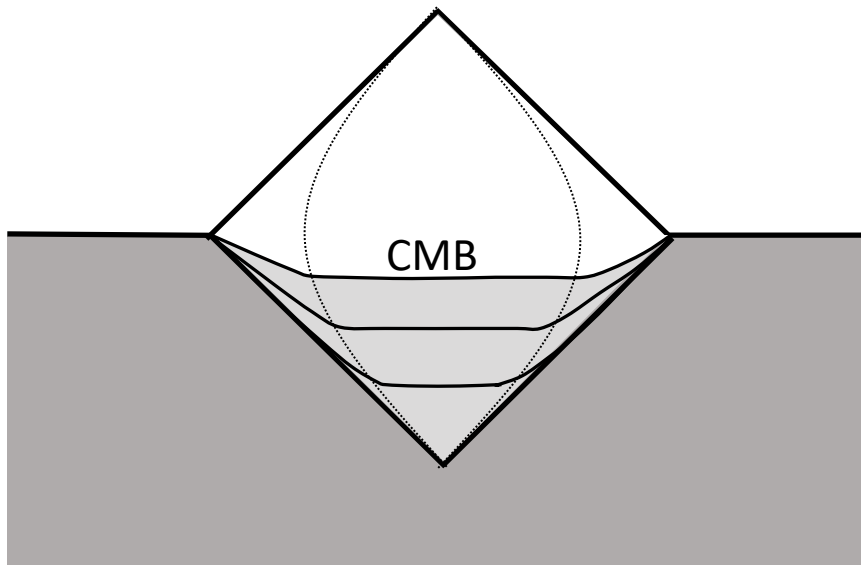
Flat reheated universe

Potential



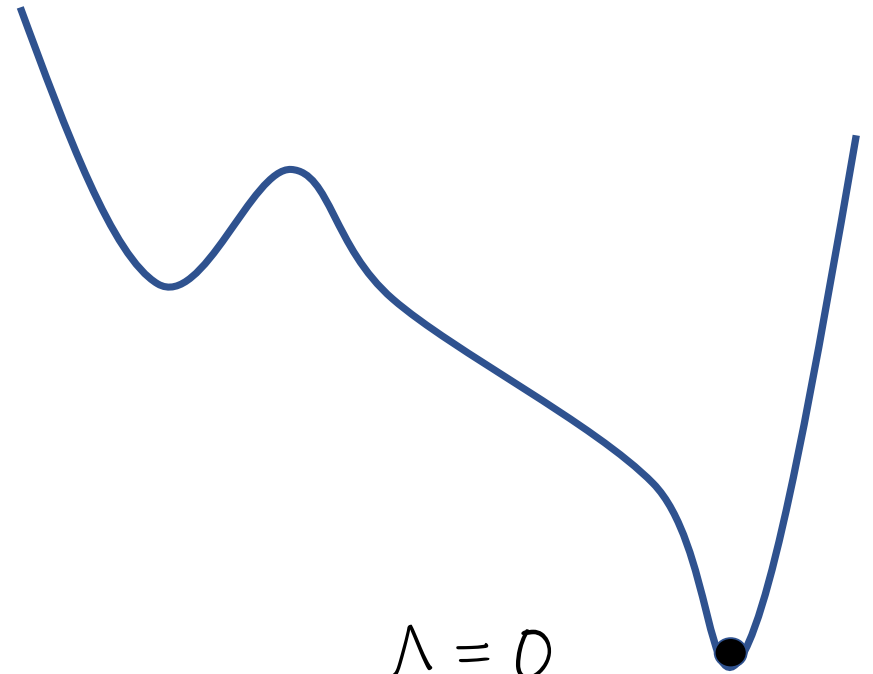
# Cosmic Acceleration

Penrose diagram



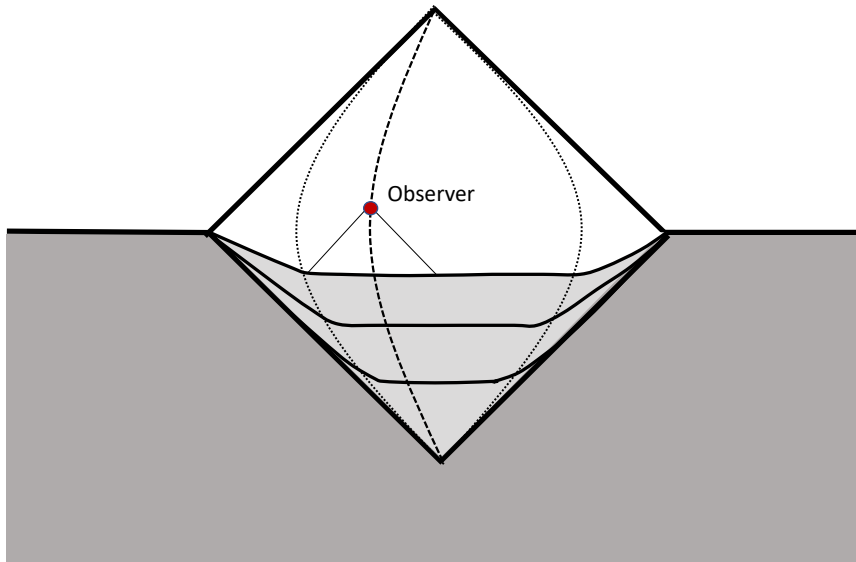
Flat late universe

Potential



# Cosmic Acceleration

Penrose diagram



Flat late universe

Potential

