Introduction to Octopus: Optical properties of finite systems

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The dynamical polarizability is the ratio of the induced dipole moment to the perturbing electric field:

$$\alpha(\omega) = \frac{\delta \boldsymbol{p}(\omega)}{\boldsymbol{E}(\omega)} = \frac{1}{\boldsymbol{E}(\omega)} \int d\boldsymbol{r} \, \boldsymbol{r} \, \delta n(\boldsymbol{r}, \omega)$$

 $\alpha(\omega)$ is related to the optical absorption cross-section:

$$\sigma(\omega) = \frac{4\pi\omega}{c} \Im \mathfrak{m} \left\{ \operatorname{Tr} \left[\alpha(\omega) \right] \right\}$$

The dynamical polarizability can be computed by solving directly the time-dependent Kohn-Sham equations:

- Take the DFT ground state wavefunctions $\varphi_i(r)$.
- Excite all the frequencies of the system by applying the appropriate instantaneous perturbation $\delta v(\mathbf{r}, t) = -Ex_j \delta(t)$.
- Use TDDFT to propagate the wavefunctions in time:

$$\varphi_i(\boldsymbol{r}, t + \Delta t) = \hat{T} \exp\left\{-\mathrm{i} \int_t^{t + \Delta t} \mathrm{d}t \, \hat{H}_{\mathrm{KS}} \varphi_i(\boldsymbol{r}, t)\right\}$$

and keep track of the density n(r, t).

• Compute the polarizability $\alpha_{ij}(\omega) = \frac{1}{E(\omega)} \int d\mathbf{r} \, x_i \delta n(\mathbf{r}, \omega).$

Casida equations:

• Pseudo-eigenvalue equation of the form:

$$\hat{R}F_q = \Omega_q^2 F_q$$

- Eigenvalues Ω_q^2 are the square of the excitation energies
- Eigenvectors are related to the oscilator strenghts
- \hat{R} is a matrix that involves pairs of occupied and unoccupied KS states

Other methods implemented in Octopus

Sternheimer equations:

- Relies on the calculation of the first order variations of the KS wavefunctions $\psi_m'({\bf r},\pm\omega)$
- Equations have the following form

$$\left[\hat{H}_{\rm KS} - \epsilon_m \pm \omega + i\eta\right] \psi'_{m,i}(\boldsymbol{r}, \pm \omega) = -\hat{P}_c \hat{H}'(\pm \omega) \psi_m(\boldsymbol{r}) \,,$$

• \hat{H}' is the first order variation of the Kohn-Sham Hamiltonian:

$$\hat{H}'(\omega) = x_i + \int \mathrm{d}\mathbf{r}' \, \frac{\delta n_i(\mathbf{r}',\omega)}{|\mathbf{r}-\mathbf{r}'|} + \int \mathrm{d}\mathbf{r}' \, f_{\mathrm{xc}}(\mathbf{r},\mathbf{r}',\omega) \delta n_i(\mathbf{r}',\omega) \,.$$

Real time propagation

Pros

- Favourable scaling with system size
- Does not require the calculation of empty states
- Easy to extend to other perturbations/responses
- Allows to go beyond linear-response
- Only requires knowledge of $v_{\rm xc}$

Cons

Slow for small systems

Casida equations

Pros

• Fast for small systems

Cons

- Requires calculation of empty states
- Requires computation of large matrices
- Unfavourable scaling with system size

Sternheimer equations

Pros

- Favourable scaling with system size
- Does not require the calculation of empty states
- Allows to go beyond linear-response

Cons

• Equations needs to be solved one frequency at a time

You can find the tutorials under this link: https://octopus-code.org/documentation/main/tutorial/

Optical response series:

- Lesson 1: Optical spectra from time-propagation
- Lesson 2: Convergence of the optical spectra
- Lesson 3: Optical spectra from Casida
- Lesson 4: Optical spectra from Sternheimer
- Lesson 5: Triplet excitations
- Lesson 6: Use of symmetries in optical spectra from time-propagation

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Have Fun !

