## Tensor network states for relativistic quantum field theory

Entangle This V





#### Goal: strongly coupled relativistic field theories

 $QCD \equiv High T_c$  supra of HEP

#### Goal: strongly coupled relativistic field theories

 $QCD \equiv High T_c$  supra of HEP

Monte Carlo on Wick-rotated lattice-discretized = only game in town



Science, 2008, BMW collaboration

#### With tensor network states



- ▶ 3+1 dimensions
- Relativistic fermions
- Gauge fields
- ► Taking the continuum limit for relativistic models ← today

**Objective**: understand the continuum on the simplest non-trivial model:  $\phi_2^4$ 

# Relativistic field theory as a condensed matter system

## Casual definition of a relativistic scalar field $\phi_2^4$



#### Hamiltonian

A continuum of nearest neighbor coupled anharmonic oscillators

$$\hat{H} = \int_{\mathbb{R}} \mathrm{d}x \quad \frac{\hat{\pi}(x)^2}{2}_{\text{on-site inertia}} + \frac{[\nabla \hat{\phi}(x)]^2}{2}_{\text{spatial stiffness}} + \frac{m^2 \hat{\phi}^2(x)}{2}_{\text{on-site potential } \hat{V}} + g \hat{\phi}^4(x)$$

with  $[\hat{\phi}(x), \hat{\pi}(y)] = i\delta(x - y)\mathbb{1} - i.e.$  bosons / harmonic oscillators

## Better definition of $\phi_2^4$

#### Renormalized $\varphi_2^4$ theory

$$H = \int \mathrm{d}x \; \frac{:\pi^2:_m}{2} + \frac{:(\nabla \phi)^2:_m}{2} + \frac{m^2}{2}:\phi^2:_m + g:\phi^4:_m$$

## Better definition of $\phi_2^4$

#### Renormalized $\phi_2^4$ theory

$$H = \int dx \, \frac{:\pi^2:_m}{2} + \frac{:(\nabla \Phi)^2:_m}{2} + \frac{m^2}{2}: \Phi^2:_m + g: \Phi^4:_m$$

- 1. Rigorously defined relativistic QFT without cutoff (Wightman QFT)
- **2.** Vacuum energy density  $\varepsilon_0$  finite for all g
- **3.** Difficult to solve unless  $g \ll m^2$  not integrable
- 4. Phase transition around  $f_c = \frac{g}{4m^2} = 11$  i.e.  $g \simeq 2.7$  in mass units

## Two (main) games in town

#### **Perturbation theory**

+ resummation

$$\Lambda = -12 \bigoplus g^2 + 288 \bigoplus g^3 + \\ - \left( 2304 \bigoplus + 2592 \bigoplus + 10368 \bigoplus \right) g^4 + \mathcal{O}(g^5)$$

$$\Gamma_2 = -96 - g^2 + \left[1152 + 3456 \right] g^3 - \left[41472 + 13824 \right] g^4 + 13824 + 1$$

state of the art is  $O(g^8)$ 

arXiv:1805.05882 Serone, Spada, Villadoro

#### Lattice Monte-Carlo



arXiv:1807.03381 Bronzin, De Palma, Guagnelli

## Short distance troubles

### Similarity between relativistic and critical models

► A critical model is scale invariant in the IR

$$\langle \mathfrak{O}(x)\mathfrak{O}(y)\rangle \stackrel{\sim}{\underset{|x-y|\to+\infty}{\sim}} \frac{1}{|x-y|^{2\Delta_0}}$$

### Similarity between relativistic and critical models

A critical model is scale invariant in the IR

► A relativistic QFT is scale invariant in the UV

$$\langle \mathfrak{O}(x)\mathfrak{O}(y)\rangle \underset{|x-y|\to 0}{\sim} \quad \frac{1}{|x-y|^{2\Delta_{\mathfrak{O}}}}$$

## Similarity between relativistic and critical models

A critical model is scale invariant in the IR

$$\langle \mathfrak{O}(x)\mathfrak{O}(y)\rangle \stackrel{\sim}{\underset{|x-y| \to +\infty}{\sim}} \frac{1}{|x-y|^{2\Delta_0}}$$

► A relativistic QFT is scale invariant in the UV

$$\langle \mathfrak{O}(x)\mathfrak{O}(y)\rangle \underset{|x-y|\to 0}{\sim} \quad \frac{1}{|x-y|^{2\Delta_{\mathfrak{O}}}}$$

#### **Consequence on entanglement**

With a UV cutoff  $\Lambda = 1/a$  in 1 + 1 dimensions:

 $S\propto \log(\Lambda)$ 

 $\implies$  infinite amount of information in high frequency modes

#### **Consequence for lattice discretizations**

1. easy: taking thermodynamic limit



#### **Consequence for lattice discretizations**

1. easy: taking thermodynamic limit



2. hard: taking small lattice spacing



## **Consequence for lattice discretizations**



2. hard: taking small lattice spacing



A finely discretized relativistic QFT, seen as a lattice model, is almost critical.



 $f_c$  estimate continuum extrapolation with GILT-TNR Clément Delcamp, AT, 2020

## UV "criticality" is usually milder than IR criticality

UV CFT tend to be kind

For QFT that are either

- 1. super renormalizable or
- 2. asymptotically free

the critical behavior at short distance is free

## UV "criticality" is usually milder than IR criticality

UV CFT tend to be kind

For QFT that are either

- 1. super renormalizable or
- 2. asymptotically free

the critical behavior at short distance is free

E.g. for  $\varphi_2^4$  at short distances

$$H \longrightarrow H_0 = \int \mathrm{d}x \; \frac{:\pi^2:_m}{2} + \frac{:(\nabla \Phi)^2:_m}{2} + \frac{m^2}{2}: \Phi^2:_m$$

which is exactly solvable

#### Objective

Stop wasting parameters on short distance criticality

- 1. Disentangle the trivial UV behavior
- $\ensuremath{\mathbf{2.}}$  Put some tensor network on top to deal with the IR

## Gaussian disentangling

### Disentangle short distance criticality

1 – Bogoliubov transform

Define modes  $a(p), a^{\dagger}(p)$  as

$$a(p) = \frac{1}{\sqrt{2}} \left( \sqrt{\omega_p} \, \phi(p) + i \frac{\pi(p)}{\sqrt{\omega_p}} \right) \quad \text{with} \quad \omega_p = \sqrt{p^2 + m^2}$$

which verify  $[a(p),a^{\dagger}(q)]=2\pi\,\delta(p-q)$  and yield

$$H_0 = \int_{\mathbb{R}} \mathrm{d}p \, \omega_p \, a_p^{\dagger} a_p$$

The ground state of  $H_0$  is the Fock vacuum, *i.e.*  $|\text{GS}\rangle = |0\rangle$  with  $\forall p, a_p|0\rangle = 0$ 

### Disentangle short distance criticality

2 – Go back to real space

Fourier transform the modes  $a_p$ 

$$\mathsf{a}(x) = rac{1}{2\pi} \int_{\mathbb{R}} \mathsf{d}p \; e^{ipx} \: a_p$$

which enforces  $[a(x), a^{\dagger}(y)] = \delta(x - y)$ 

### Disentangle short distance criticality

#### 2 – Go back to real space

Fourier transform the modes  $a_p$ 

$$\mathsf{a}(x) = rac{1}{2\pi} \int_{\mathbb{R}} \mathsf{d}p \; e^{i p x} \; \mathsf{a}_p$$

which enforces  $[a(x), a^{\dagger}(y)] = \delta(x - y)$ 

#### Note

- **1.** We integrate with dp not  $\omega_p^{-1/2} dp$
- **2.**  $\phi$  is *not* a local function of *a*, *a*<sup>†</sup>

$$\phi(x) = \int_{\mathbb{R}} dy J(x-y) \left[ a(y) + a^{\dagger}(y) \right] \quad \text{with} \quad J(x) = \int_{\mathbb{R}} \frac{dp}{\sqrt{2\omega_p}} e^{ipx}$$

#### **Tensor network intuition**



We now have two possible ways to split  $\mathscr{H} = \mathscr{H}_{-} \otimes \mathscr{H}_{+}$ 

1. Standard one, yielding  $S\propto\log\Lambda$ 

 $\mathscr{H}_{+} = \mathsf{span} \big\{ \varphi(x_1) \cdots \varphi(x_n) | \Omega_{+} \rangle \big\} \ \text{ for } \ x \geqslant 0 \big\}$ 



We now have two possible ways to split  $\mathscr{H} = \mathscr{H}_{-} \otimes \mathscr{H}_{+}$ 

1. Standard one, yielding  $S\propto\log\Lambda$ 

 $\mathscr{H}_{+} = \operatorname{span}\left\{ \varphi(x_1) \cdots \varphi(x_n) | \Omega_{+} \rangle \right\} \text{ for } x \ge 0 \Big\}$ 



**2.** The free particle one  $S_{\text{free}}$ 

 $\mathscr{H}_{+} = \operatorname{span}\left\{a^{\dagger}(x_{1})\cdots a^{\dagger}(x_{n})|0
ight\} \text{ for } x \ge 0$ 



Super-renormalizability  $\implies$  Gaussian disentangling kills the divergent part of S:

#### Conjecture

For any bosonic QFT with strongly relevant interaction  $V(\phi)$  in 1 + 1d, the free particle entanglement entropy  $S_{\text{free}}$  is **finite** in the ground state

Super-renormalizability  $\implies$  Gaussian disentangling kills the divergent part of S:

#### Conjecture

For any bosonic QFT with strongly relevant interaction  $V(\phi)$  in 1 + 1d, the free particle entanglement entropy  $S_{\text{free}}$  is **finite** in the ground state

Hence the ground state has an efficient (continuous) MPS representation:



### Trading entanglement for (mild) non-locality



#### Trading entanglement for (mild) non-locality



*H* local in  $\phi(x)$  hence mildly non-local in a(x), e.g.

$$\int dx \phi(x)^2 = \int dx \ \int dx_1 \, dx_2 \, J(x_1 - x) J(x_2 - x) (a(x_1) + a^{\dagger}(x_1)) (a(x_2) + a^{\dagger}(x_2))$$





#### **Remarks on Gaussian disentanglement**

Idea used the lattice, in Quantum chemistry, for impurity models e.g.

- ► Krumnow, Veis, Legeza, and Eisert 2016
- ▶ Wu, Fishman, Pixley, Stoudenmire 2022

Here minor differences

- 1. The disentangler is not optimized (not needed)
- 2. The disentangler does not have a simple local representation
- 3. The disentangler makes the optimization well defined  $\rightarrow$  kills divergence

## Relativistic continuous matrix product states

### Relativistic continuous matrix product states

aka continuous matrix product states (CMPS) [Verstraete and Cirac 2010] on Gaussian disentanglement steroids

#### Definition

RCMPSs are a manifold of states parameterized by 2  $(D \times D)$  matrices Q, R

$$|Q,R
angle = {
m tr} \left\{ {\mathcal P} \exp \left[ \int {
m d}x \; Q \otimes {\mathbb 1} + R \otimes a^{\dagger}(x) 
ight] 
ight\} |0
angle$$

#### with

- $|0\rangle$  is the Fock vacuum of the free model  $H_0$
- ► trace taken over  $\mathbb{C}^{D}$
- ▶ P path-ordering exponential

### **Basic properties of RCMPS**

$$|Q,R\rangle = \mathrm{tr}\left\{ \mathcal{P}\exp\left[\int\mathrm{d}x\;Q\otimes\mathbb{1}+R\otimes a^{\dagger}(x)
ight]
ight\}|0
angle_{a}$$

Checklist:

- **1.** Extensive because of  $\mathcal{P} \exp \int$
- 2. Observables **computable** at cost  $D^3$  (non trivial!) requires  $[a(x), a^{\dagger}(y)] = \delta(x y)$
- 3. No UV problems

 $|0,0\rangle = |0\rangle$  is the ground state of  $H_0$  hence exact CFT UV fixed point  $\langle Q, R| : P(\Phi) : |Q, R\rangle$  is finite for all Q, R (not trivial!)

#### Tensor network intuition



In the **continuum limit** contracting a non-uniform ladder is numerically exact with high order Runge-Kutta.

## The variational algorithm

Optimization

Compute  $e_0 = \langle Q, R | h | Q, R \rangle$  and  $\nabla_{Q,R} e_0$ Minimize  $e_0$  with (geometric improvements of) gradient descent

## The variational algorithm

#### Optimization

Compute  $e_0 = \langle Q, R | h | Q, R \rangle$  and  $\nabla_{Q,R} e_0$ Minimize  $e_0$  with (geometric improvements of) gradient descent

Computations of  $e_0$  and  $\nabla e_0$  in a nutshell:

- 1.  $V_b = \langle : e^{b \Phi(x)} : \rangle_{QR}$  computable by solving an ODE with cost  $\propto D^3$
- **2.**  $\langle : \Phi^n : \rangle_{QR}$  computable doing  $\partial_b^n V_b \Big|_{b=0} \to \infty D^3$
- **3.**  $e_0 = \langle h \rangle_{QR}$  computable by summing such terms at cost  $D^3 \to \propto D^3$
- 4.  $abla e_0$  computable by solving the adjoint ODE (backpropagation)  $\rightarrow \propto D^3$

## The variational algorithm

#### Optimization

Compute  $e_0 = \langle Q, R | h | Q, R \rangle$  and  $\nabla_{Q,R} e_0$ Minimize  $e_0$  with (geometric improvements of) gradient descent

Computations of  $e_0$  and  $\nabla e_0$  in a nutshell:

- 1.  $V_b = \langle : e^{b \Phi(x)} : \rangle_{QR}$  computable by solving an ODE with cost  $\propto D^3$
- **2.**  $\langle : \Phi^n : \rangle_{QR}$  computable doing  $\partial_b^n V_b \Big|_{b=0} \to \infty D^3$

**3.**  $e_0 = \langle h \rangle_{QR}$  computable by summing such terms at cost  $D^3 \to \infty D^3$ 

4.  $abla e_0$  computable by solving the adjoint ODE (backpropagation)  $\rightarrow \propto D^3$ 

Functioning Julia implementation. OptimKit.jl to solve the Riemannian minimization, KrylovKit.jl to solve fixed point equations, DifferentialEquations.jl (Vern7 solver) to solve ODE. Soon Rcmps.jl?

#### Using the optimized state

After optimization:  $|Q,R\rangle \simeq |0\rangle_{\text{int.}}$  with  $\langle Q,R| \ \hat{h} |Q,R\rangle = e_0 + \varepsilon$ 

#### This gives:

► All equal-time *N*-point functions

 $\langle \phi(x_1) \phi(x_2) \cdots \phi(x_n) \rangle \simeq \langle Q, R | \phi(x_1) \phi(x_2) \cdots \phi(x_n) | Q, R \rangle$ 

at cost  $D^3$  by solving coupled linear ODEs

▶ In particular all Euclidean 2-point functions ⇒ spectral function

$$\langle \Phi(x)\Phi(0)\rangle = \int_0^{+\infty} \mathrm{d}\mu\,\mu\rho(\mu)K_0(\mu x)$$

## **Results:** $\phi_2^4$ energy density



**New**: *D* can now be pushed to 32 or even 64 with some effort

**Results:**  $\phi_2^4$  – field expectation value  $\langle \phi \rangle$ 



**New:** the mass can be fitted from 2-point function and agrees with RHT to  $10^{-3}$ 

#### Todo-list for continuous tensor networks

#### In 1+1 dimensions

- Solve Fermion / Gauge theories
- Go beyond strongly renormalizable interactions
- Do general CFT perturbations
- ► Compute more observables (masses, spectra, *c*-function...)

And of course the grand goal: do higher dimensions!

Come work on it in Paris with Edo and Karan!

## Summary

Problem

Relativistic QFT have infinite entanglement at short distance

**Solution** in 1 + 1d

$$|Q,R
angle = {
m tr} \left\{ {\mathcal P} \exp \left[ \int {
m d}x \ Q \otimes {\mathbb 1} + R \otimes a^{\dagger}(x) 
ight] 
ight\} |0
angle$$



- 1. Ansatz for 1+1 relativistic QFT  $% \left( {{\left( {T_{{\rm{T}}} \right)} \right)} \right)$
- **2.** The  $\phi(x) \rightarrow a(x)$  trick disentangles the divergent UV
- 3. The CMPS on top solves the rest
- 4. Efficient (cost poly D, error plausibly  $1/{\rm superpoly}\ D$  )