Quarkyonic effective field theory

Dyana C. Duarte

In collaboration with K.S. Jeong, S.Hernandez-Ortiz and L. McLerran Based on PRD 104, L091901, PRC 107, 065201

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Outline

• Brief motivation;

• Quarkyonic Matter for the EoS of neutron stars;

• Quarkyonic EFT;

• Final Remarks.



Many different approaches and possibilities!

- ► Extension of low-energy nuclear physics to higher densities: RMF and many body calculations with dependence of couplings and masses with n_B [Oertel et al., Rev.Mod.Phys. 89, 015007 (2017), Kaiser et al., Nucl. Phys. A. 697, 255 (2002), Drischler et al., Phys. Rev. Lett.122, 042501 (2019), Lonardoni et al., Phys. Rev. Res. 2, 022033 (2020)];
- Other interactions and degrees of freedom [Glendenning, Astrophys. J. 293, 470 (1985), Knorren, et al., Phys. Rev. C 52, 3470 (1995), Cai et Al., Phys. Rev. C 92, 015802 (2015)];
- Phase transition from nuclear to quark matter. [Annala et al. Nat. Phys. 16, 907 (2020), Nature Communications 14, 8451 (2023)].

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- Phase transition from nuclear to quark matter. [Annala et al. Nat. Phys. 16, 907 (2020), Nature Communications 14, 8451 (2023)].

Resultant model/theory must:

 $\sqrt{}$ Satisfy the well-known nuclear matter properties at saturation density; $n_B \sim (1-2)\rho_0$ $\sqrt{}$ Evolve to a phase of deconfined quarks at high densities; $n_B \gtrsim (2-5)\rho_0$??? $\sqrt{}$ Recover the pQCD results at asymptotically high densities and temperatures. $n_B \gtrsim 40\rho_0$

Bayesian Inference (gaussian regression)

Neutron star EoS is characterized by:

- Nucleon mass scale in the hadronic part;
- Near-conformal properties (no mass scale).

Table 1 | Characteristic features of dense strongly interacting matter

| | CEFT | Dense NM | Pert. QM | CFTs | FOPT |
|-----------------------------|------------|--------------|-----------|------|--|
| c _s ² | ≪1 | [0.25, 0.6] | ≲1/3 | 1/3 | $0 = dp/d\varepsilon$ |
| Δ | ≈1/3 | [0.05, 0.25] | [0, 0.15] | 0 | 1/3 – р _{РТ} /є = $(\varepsilon - 3p)/(3\varepsilon)$ |
| Δ' | ≈ 0 | [-0.4,-0.1] | [-0.15,0] | 0 | 1/3- Δ = $d\Delta/d\ln(\varepsilon)$ |
| d _c | ≈1/3 | [0.25, 0.4] | ≲0.2 | 0 | $\geq 1/(3\sqrt{2}) = \sqrt{\Delta^2 + (\Delta')^2}$ |
| Y | ≈2.5 | [1.95, 3.0] | [1, 1.7] | 1 | 0 $= d\ln(p)/d\ln(\varepsilon)$ |
| p/p _{free} | ≪1 | [0.25, 0.35] | [0.5, 1] | _ | $p_{	extsf{PT}}/p_{	extsf{free}} = p/p_{	extsf{free}}$ |

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Strongly interacting matter exhibits deconfined behavior in massive neutron stars

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| | | | | | |

Article

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Quarkyonic Matter



Drischler et al (2021).



Pressure when $\mu_Q > M_q$ and $n_B \neq 0$:

 $\mathcal{O}(N_c^0) \to \mathcal{O}(N_c)$ (hadron) \to (quarkyonic)

Weakly interacting quark system? Baryonic system?

Quarkyonic Matter can satisfy constraints to the EoS of NS





Good agreement with sound velocity obtained from an equation of state extracted from neutron stars properties using deep neural network.

K.S. Jeong, L. McLerran, S. Sen, Phys. Rev. C 101 035201 (2920)

Quarkyonic Matter can satisfy constraints to the EoS of NS



Model for the dependence of Δ with the baryon density:

$$\Delta = \frac{\Lambda^3}{k_{FB}^2} + \kappa \frac{\Lambda}{N_c^2}$$

McLerran and Reddy, PRL 122,122701 (2019) Three flavor excluded volume model, shell appearing dynamically with the increase of baryon density.

> DD, Jeong, Hernandez-Ortiz, PRC 102 065202 (2020)



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Our purpose: Inclusion of a negative metric ghost field, that fill precisely the same state as the quarks, and cancel away the degrees of freedom of unconstrained nucleons.

Nucleons can be thought of as an ensemble of quarks: If there is a quark Fermi sea with a chemical potential μ_Q in the low momentum states, the quarks composing the nucleon cannot occupy the same states.

$$\rho^G = \frac{1}{1 + e^{\beta(N_C E_Q - \mu_G)}} - \frac{1}{1 + e^{\beta(N_C E_Q + \mu_G)}}$$

Density of the states in which nucleons cannot exist: Region of quark Fermi sea up to $\mu_G = N_c \mu_Q$.

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of as nucleon energy:

$$\rho^{Q} = \frac{1}{1 + e^{\beta(E_{Q} - \mu_{Q})}} - \frac{1}{1 + e^{\beta(E_{Q} + \mu_{Q})}}$$
$$\rho^{N}_{\text{const.}} = \rho^{n} = \frac{1}{1 + e^{\beta(E_{N} - \mu_{N})}} - \frac{1}{1 + e^{\beta(E_{N} - \mu_{G})}}$$

Density of quarks in a noninteracting gas of quarks and nucleons;

Density of nucleons constrained to not propagate in the quark Fermi sea:

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Density of quarks in a noninteracting gas of quarks and nucleons;

Density of nucleons constrained to not propagate in the quark Fermi sea:

Nucleons are placed in a shell of momenta above the momenta of quark Fermi sea, where they are not Pauli blocked!

$$\mathcal{L}_{\text{eff}} = \overline{N} \left(i\partial - M_N + \gamma^0 \mu_N \right) N + \overline{G} \left(i\partial - M_N + \gamma^0 \mu_G \right) G + \overline{Q} \left(i\partial - M_Q + \gamma^0 \mu_Q \right) Q$$

<u>**Ghosts:**</u> Nucleon fields with M_N and $\mu_G = N_c \mu_Q$, with the same Lorentz structure of nucleonic field, satisfying anti-periodic boundary conditions in immaginary time. It is, however, a commuting field!

But what does it mean? Let $S(\mu, M)$ be the propagator for a given field.

 $Z_{\text{Grassman}} = \det^{-1} S(\mu, M)$ $Z_{\text{c-number}} = \det S(\mu, M)$

DD, K.S. Jeong, S. Hernandez-Ortiz, L.McLerran, PRD 104, L091901(2021)

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But what does it mean? Let $S(\mu, M)$ be the propagator for a given field.

- Independent parameters: nucleon density n_N and quark baryon density $n_Q^B = n_Q/N_c$
- At fixed total baryon density...

$$n_B = n_N - n_G(n_Q^B) + n_Q^B$$
$$dn_B = dn_N + \left(1 - \frac{dn_G(n_Q^B)}{dn_Q^B}\right) dn_Q^B = 0$$

... the configuration that minimized the energy density is....

$$d\varepsilon = \frac{\partial \varepsilon}{\partial n_N} dn_N - \frac{\partial \varepsilon}{\partial n_G(n_Q^B)} dn_G(n_Q^B) + \frac{\partial \varepsilon}{\partial n_Q^B} dn_Q^B = 0$$

... and this leads to the relation

$$\mu_B = \frac{d\varepsilon}{dn_B} = \mu_N$$

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- At fixed total baryon density...

$$n_B = n_N - n_G(n_Q^B) + n_Q^B \qquad \qquad \begin{array}{l} \mu_Q^B \simeq N_c \mu_Q \\ \text{and} \\ dn_B = dn_N + \left(1 - \frac{dn_G(n_Q^B)}{dn_Q^B}\right) dn_Q^B = 0 \qquad \qquad \begin{array}{l} \frac{dn_G(n_Q^B)}{dn_Q^B} \simeq N_c^3 \\ \frac{dn_Q^B}{dn_Q^B} \simeq N_c^3 \end{array}$$

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$$S = \int d^4x \left\{ \overline{\psi} \left(\frac{1}{i} \partial - g_v V - \gamma^0 \mu_N^* + M_N - g_\sigma \sigma \right) \psi \right. \\ \left. + \overline{G} \left(\frac{1}{i} \partial - g_v V - \gamma^0 \mu_G^* + M_N - g_\sigma \sigma \right) G \right. \\ \left. + \frac{M_\sigma^2}{2} \sigma^2 + \frac{M_v^2}{2} V^2 \right\} \\ \varepsilon = \varepsilon_{\rm kin}^N(\mu_N) - \left(1 - \frac{1}{N_c^3} \right) \varepsilon_{\rm kin}^N(N_c \mu_Q) \\ \left. + \frac{g_v^2}{2M_v^2} \left[n_v(\mu_N) - n_v(N_c \mu_Q) \right]^2 \\ \left. - \frac{g_\sigma^2}{2M_\sigma^2} \left[n_s(\mu_N) - \left(1 - \frac{1}{N_c^3} \right) n_s(N_c \mu_Q) \right]^2 \right]$$

DD, K.S. Jeong, S. Hernandez-Ortiz, L.McLerran, PRC 107, 065201.



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To work at finite temperature: Extremize free energy density with the same constraints.

$$f = \varepsilon - Ts = p + \mu_B n_B \qquad \qquad \mu_G = N_c (\mu_Q - \Delta M) + g_v V_0$$

DD, K.S. Jeong, S. Hernandez-Ortiz, L.McLerran, PRC 107, 065201 (2023).

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EFT with ghosts: Possibilities/Questions

- Inclusion of different nucleon and quark potentials directly in the Lagrangian density;
- Isospin asymmetry?
- Dynamic generation of the shell with different nucleonic interactions: Is it possible to construct a model that respect nuclear properties at saturation density and, at the same time respect the contraints from observational data?

 $k_G = N_c k_Q$

$$\bar{N}\Gamma_i N + \bar{G}\Gamma_i G$$

• Quarkyonic sigma model: How chiral symmetry restoration is related to quarkyonic matter?

$$M_N = g_N(\phi + i\pi\gamma^5)$$
$$M_Q = g_Q(\phi + i\pi\gamma^5)$$

• Future FAIR/NICA simulations: Possible clues of quarkyonic matter: How it would appear in the QCD phase diagram?

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Thanks to Sanjay Reddy, Toru Kojo, Rob Pisarski and Jinfeng Liao for all the useful discussions during the development of this worts./ 20

Thanks for your attention!

Dynamically generated shell of nucleons: Excluded Volume Model

$$n_{ex} = \frac{n_N}{1 - n_N/n_0} = \frac{N_f}{\pi^2} \int_{k_F}^{k_F + \Delta} dk \ k^2$$

$$\varepsilon = \frac{N_f}{\pi^2} \left(1 - \frac{n_N}{n_0} \right) \int_{k_F}^{k_F + \Delta} dk \ k^2 \sqrt{k^2 + M^2} + \varepsilon_Q$$
Free gas of quarks contribution
$$\begin{cases} n_Q = \frac{N_f}{\pi^2} \int_0^{k_Q} dk \ k^2 = \frac{N_f}{3\pi^2} k_Q^3 \\ \varepsilon_Q = \frac{N_c N_f}{\pi^2} \int_0^{k_Q} dk \ k^2 \sqrt{k^2 + m^2} \end{cases}$$

$$k_Q = k_F / N_c \qquad m = M / N_c$$

K.S. Jeong, L. McLerran, S. Sen, Phys. Rev. C 101 035201 (2020)

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At small k_Q quark density increase very fast to generate acceptable sound velocity \rightarrow Modification in the low density Fermi distribution in a way that does

not affect its behavior for large Fermi momenta:

 $1 \to \frac{\sqrt{k_Q^2 + \Lambda^2}}{k_Q}$

$$k_Q = k_F / N_c \qquad m = M / N_c$$



$$n_Q = \frac{N_f}{3\pi^2} \left[\left(k_Q^2 + \Lambda^2 \right)^{3/2} - \Lambda^3 \right]$$

$$\varepsilon_Q = \frac{N_c N_f}{\pi^2} \int_0^{k_Q} dk \; k \sqrt{k^2 + \Lambda^2} \sqrt{k^2 + m^2}$$

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K.S. Jeong, L. McLerran, S. Sen, Phys. Rev. C 101 035201 (2020)

Quarkyonic EFT: Excluded volume model

• By assuming $\mu_G = N_c \mu_Q$ one may derive the excluded volume model and verify the shell-like distribution on nucleons in the ideal gas limit.

$$\Omega = -p = \varepsilon_n - \mu_N n_n - (\mu_n - \mu_G) n_G + \varepsilon_{\tilde{Q}} - \mu_{\tilde{Q}} n_{\tilde{Q}}$$

$$m = \varepsilon_N - \varepsilon_G$$
$$= 4 \int_{k_G}^{k_N} dk \frac{k^2}{2\pi^2} \sqrt{k^2 + M_N^2}$$

 ${\mathcal E}$

$$n_n = n_N - n_G = 4 \int_{k_G}^{k_N} dk \frac{k^2}{2\pi^2}$$

$$k_{N,G} = \sqrt{\mu_{N,G}^2 - M_N^2}$$

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$$\begin{split} \varepsilon_n^{\text{ex}} &= \varepsilon_N^{\text{ex}} - \varepsilon_G^{\text{ex}} \\ &= 4 \int_{k_G^{\text{ex}}}^{k_N^{\text{ex}}} dk \frac{k^2}{2\pi^2} \sqrt{k^2 + M_N^2} \\ n_n^{\text{ex}} &= \frac{n_N - n_G}{1 - \frac{n_N - n_G}{n_0}} = 4 \int_{k_G^{\text{ex}}}^{k_N^{\text{ex}}} dk \frac{k^2}{2\pi^2} \\ &\quad k_{N,G}^{\text{ex}} = \sqrt{(\mu_{N,G}^{\text{ex}})^2 - M_N^2} \end{split}$$

Quarkyonic EFT: Excluded volume model

• The other quantities in the reduced volume are related to the ones of the total system from:

$$\varepsilon^{\text{ex}} = \frac{\varepsilon}{1 - \frac{n}{n_0}}$$
$$\mu^{\text{ex}} = \frac{\partial \varepsilon^{\text{ex}}}{\partial n^{\text{ex}}}$$

• At fixed total baryon density (and zero temperature) the system configuration is obtained by extremizing the energy density, and the chemical potentials are related as $\mu_B = \mu_{N.}$.